Calorimetry in HEP

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ESIPAP 2021 - 01-02/02/2021





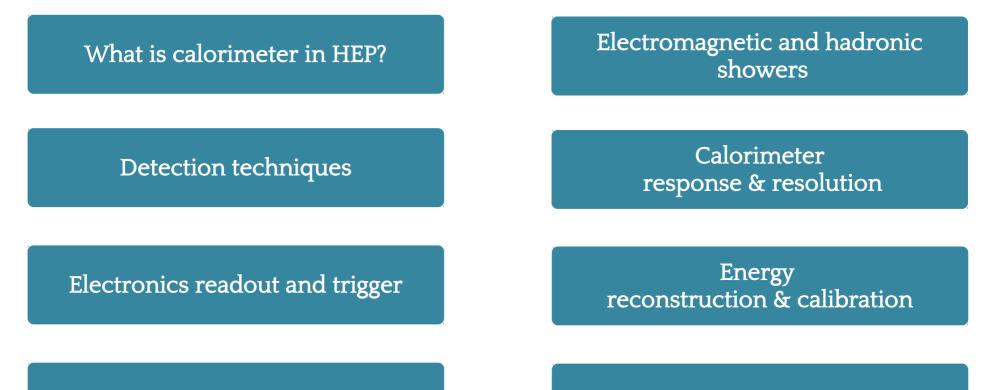








Lecture plan



Beyond calorimetry: Particle Flow

An example: CMS HGCAL

Reference

The following lecture was built upon several lectures and books

Lectures

- OC. Ochando, *Lectures on Calorimetry*, ESIPAP 2019
- OI. Wingerter-Seez, Calorimetry: Concepts and Examples, ESIPAP 2016
- E. Garutti, *The art of calorimetry*
- **O R. Wigmans**, *Calorimetry*, EDIT 2011
- OV. Boudry, *La Calorimetrie*, Ecole du detecteur a la mesure 2013
- OD. Cockerill, Introduction to Calorimeters, Southampton Lecture 2016
- OA. Zabi, Instrumentation for High Energy Physics, TES-HEP 2016
- OP. Janot, Particle-Flow event reconstruction from LEP to LHC, EDIT 2011

Books

- **O R. Wigmans**, *Calorimetry, Energy measurement in Particle Physics*, Oxford science publication
- **O C. Gruppen & B. Shwartz**, *Particle detectors*, Cambridge monographs on particle physics, nuclear physics and cosmology

A few words about myself

Physicist at the Laboratoire Leprince-Ringuet

O Located at the Ecole polytechnique, in the south of Paris

■ PhD thesis in ATLAS during the LHC Run 1 data taking period

O Jet energy calibration

○Z+jets cross section measurements

■ In CMS since 2012

 \bigcirc Higgs measurements (H \rightarrow ZZ^{*} \rightarrow 4 leptons, H $\rightarrow \tau\tau$)

O Electron energy reconstruction

O Phase 1 calorimeter trigger upgrade

O High Granularity Calorimeter (CMS endcap calorimeter Phase 2 upgrade)

Lecture plan



Electronics readout and trigger

Electromagnetic and hadronic showers

Calorimeter response & resolution

Energy reconstruction & calibration

Beyond calorimetry: Particle Flow

An example: CMS HGCAL

What is a calorimeter (in HEP)?

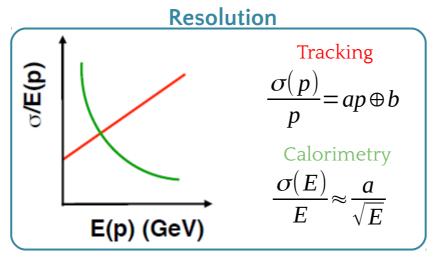
Originally a "calorimeter" is an instrument measuring heat ("calor" in latin) produced by some reactions (chemical or physical)

O In HEP it is quite different

- Detection of particles through total absorption in a block of matter
- Complementary to tracking detectors

O Trackers measures charged particle bending

O Calorimeters measure absorbed energy



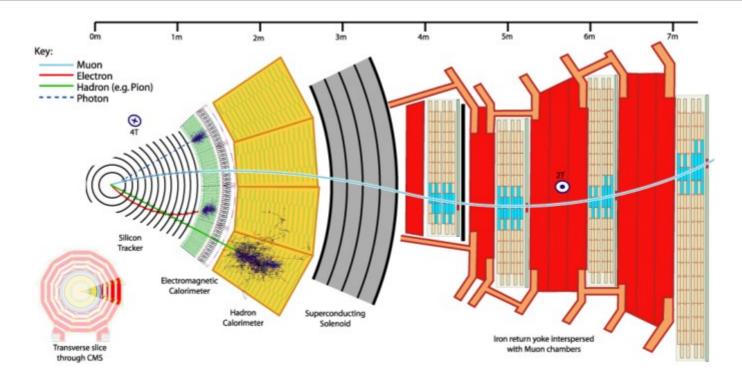
Calorimeters can measure both charged and neutrals



Typical HEP detectors

Muon Muon Spectrometer spectrometer Hadronic Calorimeter The dashed tracks Calorimeters are invisible to Neutrino the detector Proton Neutron Muon Electromagnetic Calorimeter Electron Solenoid magnet Photon Transition Radiation Tracking Tracker Trackers Pixel/SCT detector

CMS ECAL



CMS: typical Onion-like detector structure

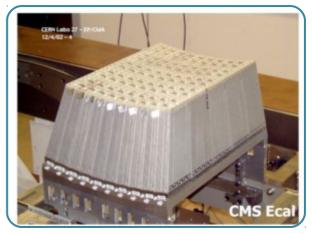
O Tracking + magnet (curvature of charged particles)

O EM and hadronic calorimeters

CMS ECAL

O Scintillating crystals + photodiodes

CMS ECAL module



Back to the origins: Nuclear radiation detectors

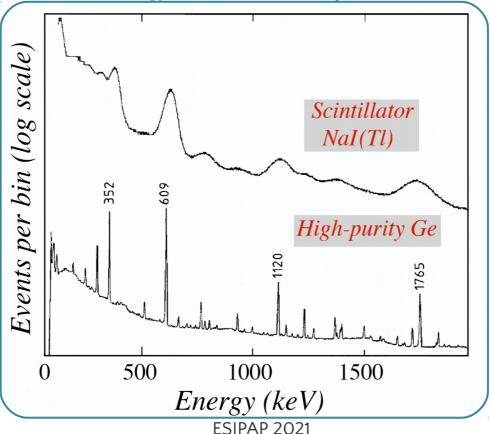
■ Late 40's

○ **Scintillating crystals** + Photomultiplier tubes (PMT)

 \bigcirc First calorimeters used in the detection of nuclear decays ($\alpha,\,\beta,\,\gamma)$

■ In the 60's

O First **semiconductor detectors** (silicon and germanium)



Energy resolution is important

Neutrino experiments at SPS (CERN)

■ In the 70's – 80's

O Deep inelastic **neutrino interactions**

O Weakly interacting

O Instrumented mass of -1 kTon

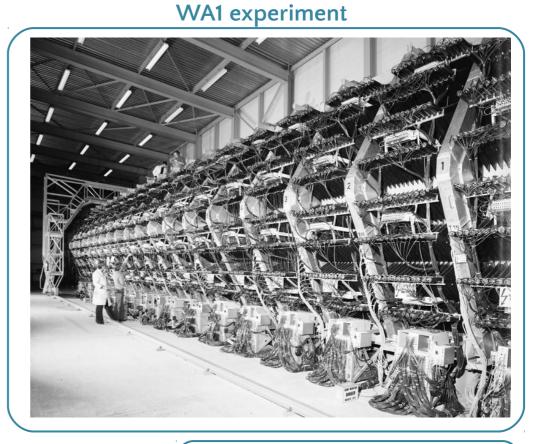
O Intense beams

WA1

○ Slabs of (magnetized) **Iron**

O Interleaved with **plastic scintillators**

O + wire chambers in the rear to track muons

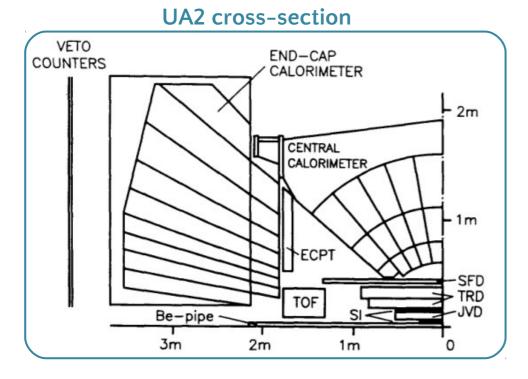


Scintillators + PMT



UA1 and UA2

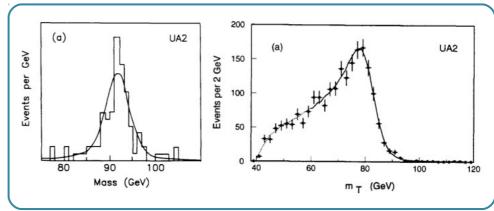
- UA1 & UA2 = Experiments at the SppS (CERN)
- UA2 (1981-1990)
 - O More focused on calorimetry
 - Lead (or Iron) + scintillator
- Discovery of the W and Z bosons



UA2 detector



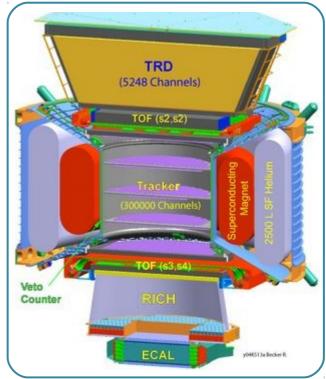
Z mass & W transverse mass



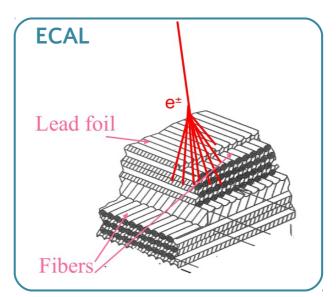
AMS



AMS cross-section



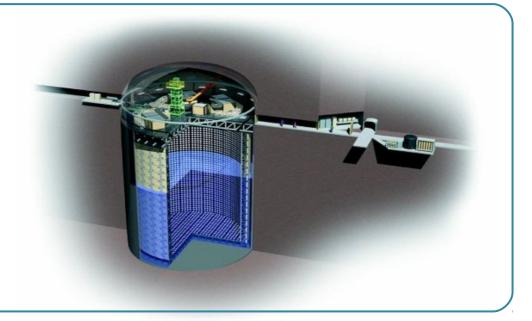
- AMS: experiment on the International Space Station
- Search presence of antimatter and dark matter
- Electromagnetic calorimeter
 - O Measure high energy electrons/positrons
 - O Discriminate against protons



(Super-)Kamiokande

Super-Kamiokande





- Water tank (> 2140 t) placed in underground mine
- Surrounded by 1k photomultipliers
- Scattering of neutrinos with electron or nuclei of water → Cerenkov light

■ 1990's:

O Measurement of solar neutrinos flux deficit

■ Followed by Super-Kamiokande

O Discovery of **neutrino oscillation**

Large PMT

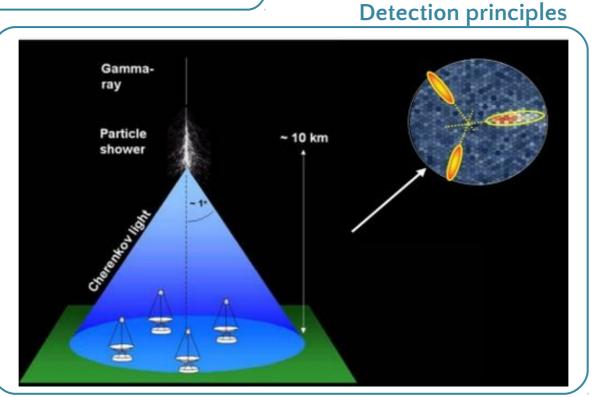


HESS

The 5 HESS telescopes



- Explore cosmic gamma rays
 O Interaction with the atmosphere
 O Emission of Cerenkov light
- Telescopes record this Cerenkov light on the ground



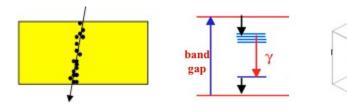
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Calorimeter: principles

Particles interact with matter Depends on particle and material P



Energy lost transferred to detectable signal *Light, electric signal, etc.*

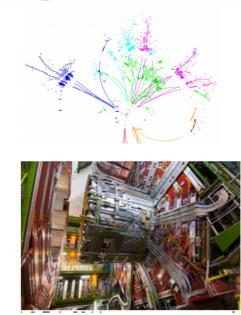


Signal collected and acquired *In the end: digitized signal*

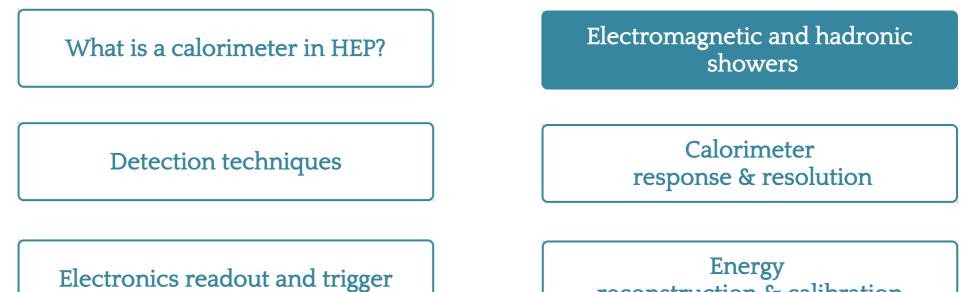
 $d = N \cdot e$

Calibration and reconstruction Infer initial particle energy, position and type

Everything together Build an experimental setup Many constraints to be satisfied



Lecture plan



reconstruction & calibration

Beyond calorimetry: Particle Flow

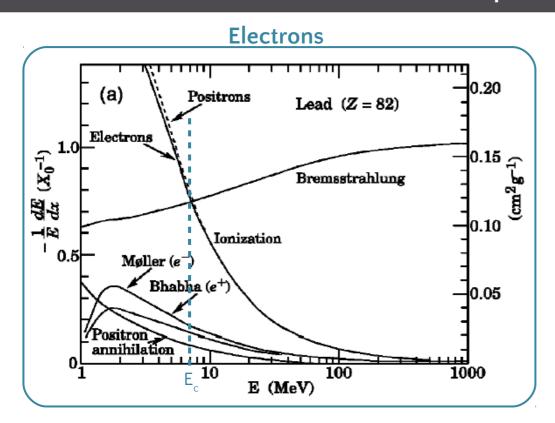
An example: CMS HGCAL

Electromagnetic interactions with matter

Photons Electrons Photoelectric effect Ionisation e **Compton scattering** m mm e e Bremsstrahlung **Pair production** High energy $\gamma + nucleus \rightarrow e^+e^- + nucleus$

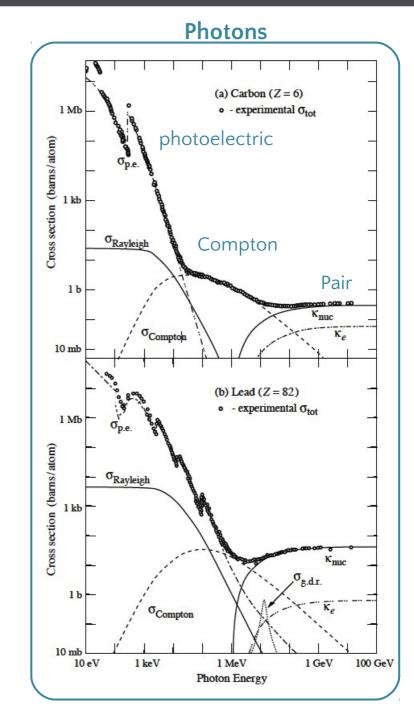
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Dominant processes



High energy electrons: Bremsstrahlung

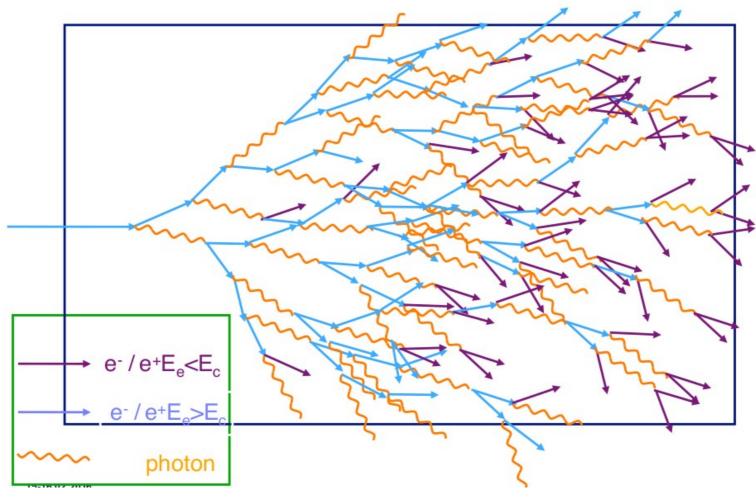
- High energy photons: Pair creation
- Below critical energy
 - O Energy loss through **ionisation / excitation** of the medium



Shower development

High energy particle creates a cascade of lower energy electrons and photons OThrough bremsstrahlung and pair production

When the critical energy is reached, secondary particles are slowly stopped (electrons) or absorbed (photons)

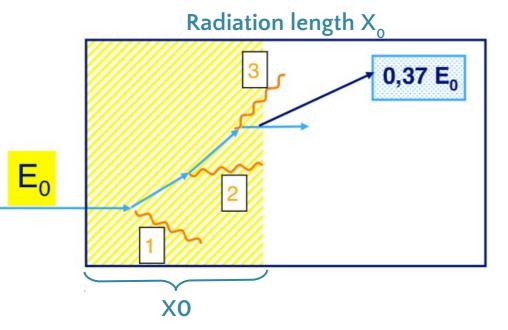


Energy loss and mean free path

Two dominant processes: Bremsstrahlung and pair production

Electron energy loss

$$-\left(\frac{dE}{dx}\right)_{rad} = \frac{E}{X_0}$$
$$E = E_0 e^{-x/X_0}$$
$$x(E_0/2) = X_0 \ln(2)$$



Pair prod. probability

$$\frac{dw}{dx} = \frac{1}{\lambda_{pair}} e^{-x/\lambda_{pair}}$$
$$\lambda_{pair} = \frac{9}{7} X_0$$

$$\blacksquare X_0 \approx \frac{180 A}{Z^2} g. cm^{-2}$$

O Conversion between the two with the material density

- Electrons loose half of their energy in about 2/3×X₀
- Photons convert in about 9/7×X₀

Simplistic shower model

• $x(E_0/2) = X_0 \ln(2)$ and $\lambda_{pair} = \frac{9}{7}X_0$, and the average roughly equals to X_0

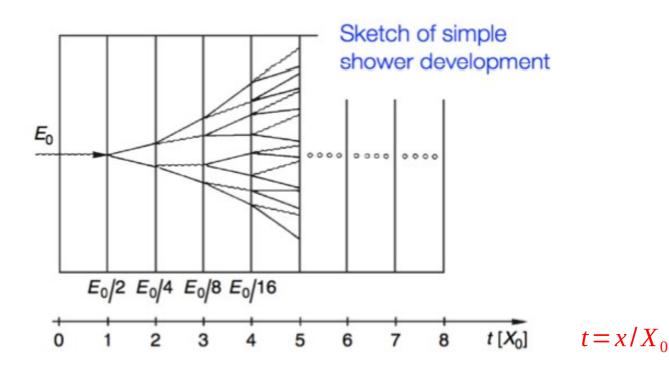
So we consider that, on average

 $\bigcirc \text{One particle duplication occurs every } X_0 \ (e \rightarrow e\gamma \ \text{or} \ \gamma \rightarrow ee)$

O With **equal sharing of energy** between the two produced particles

■ Stops at the critical energy **E** = **E**_c

O Reaches maximum number of particles = "shower maximum"



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Some EM shower properties

- Number of particles proportional to the initial energy
 - \bigcirc Energy per particle after depth t: $E = E_0 \cdot 2^{-t}$
 - O Shower maximum $t_{max} \propto \ln(E_0/E_c)$ (X₀ units)
- Shower lateral extent

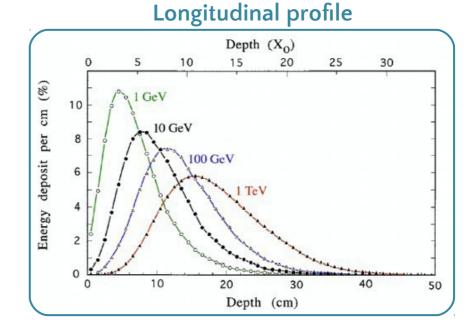
○ Narrow core

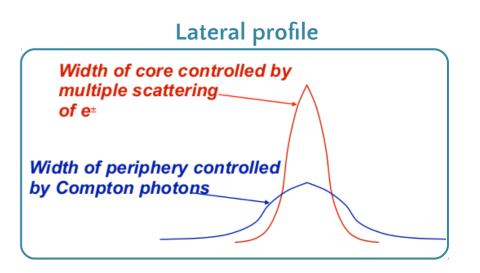
- Early stage of the shower
- 90% of shower contained in "Moliere" radius

$$R_{M} = \frac{21 \, MeV \times X_{0}}{E_{c}} \approx \frac{7 \, A}{Z} g \, cm^{-2}$$

⊖ Tails at larger angles

- Isotropic Compton scattering
- Beyond shower max





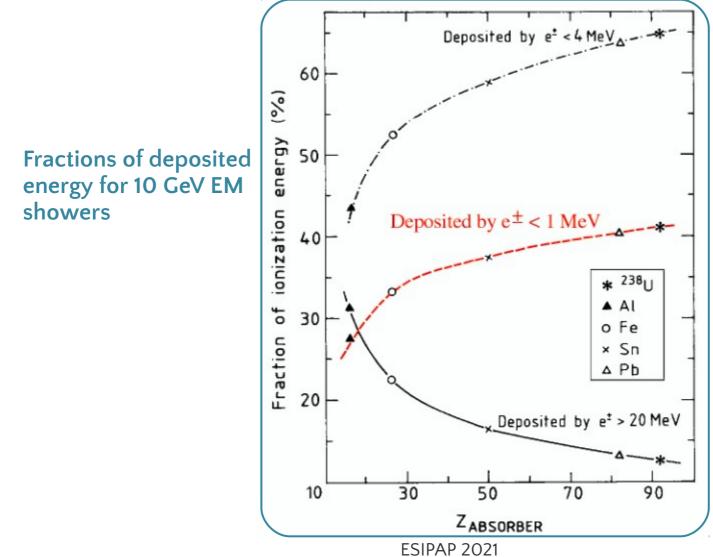
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Importance of low energy particles

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- Shower development driven by high-energy particles
- But phenomena at **E** < **E**_c **are important for calorimeter** properties

O In lead 40% of the energy deposited by electrons < 1MeV

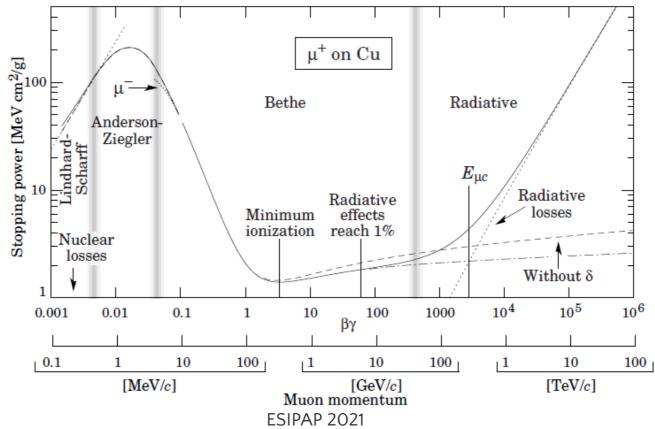


Useful quantities

Critical energy for solids & liquids $E_c = \frac{610 MeV}{Z+1.24}$	Critical energy $E_c = \frac{710 MeV}{Z + 0.92}$			
Radiation length (approximate formulas) $X_0 \approx \frac{180 A}{Z^2} g \cdot cm^{-2}$ $X_0 \approx \frac{716.4 A}{Z(Z+1) \ln(287/\sqrt{Z})} g \cdot cm^{-2}$	0			
Shower maximum (X ₀ units) $t_{max} = \ln (E_0/E_c) - \begin{pmatrix} 1 & \text{(electrons)} \\ 0.5 & \text{(photons)} \end{pmatrix}$	■ 95% longitudinal containment (X ₀ units) $L(95\%)/X_0 = t_{max} + 0.08Z + 9.6$			
Moliere radius (same as X ₀ for compound/mixture) $R_{M} = \frac{21 MeV \times X_{0}}{E_{c}} \approx \frac{7 A}{Z} g \cdot cm^{-2}$	■ 95% lateral containment $R(95\%)=2R_M$			
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The case of the muon

- Muons are charged leptons, like electrons, but much heavier
 - $O m_{\mu}/m_{e} \sim 200$
- Loss of energy by brem $-\left(\frac{dE}{dx}\right)_{brem} \propto \frac{E}{m^2}$
- Main mechanism for muons is ionization → **no shower**
 - O Very small energy deposits
 - O Radiation only above the TeV scale



Hadronic showers

Cascade of particles **initiated by a hadron**

O Strong interaction in addition to EM interaction

Many processes involved

○ Ionisation

O Hadron production (fragmentation, etc.)

○ Charge exchange

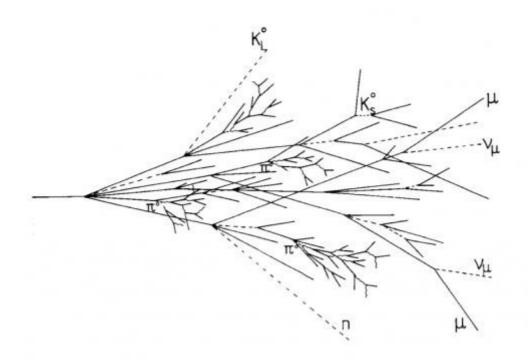
 $-\pi n \rightarrow \pi^{\circ} p$

O Spallation, fission

O Nuclear de-excitation

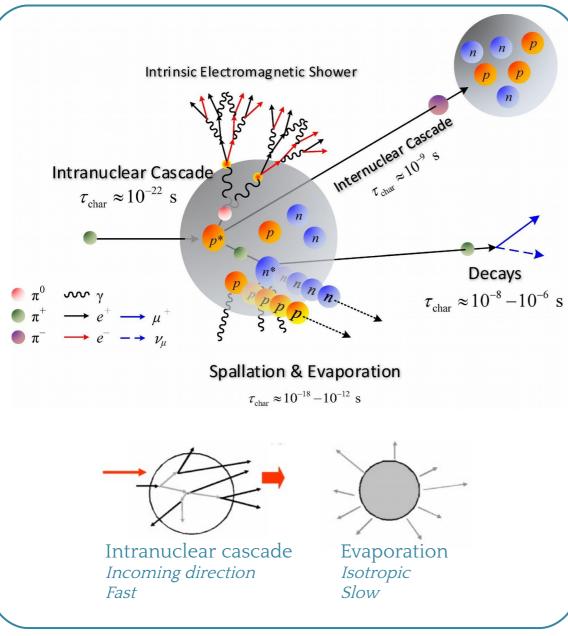
⊖ Pion decay

0...



Hadronic interactions

Hadronic shower evolution



■ 1) Hard collision

- O Can travel long distance before 1st interaction
- O Similar to a MIP

■ 2) Spallation

O Intra-nuclear cascade

O Frees protons and neutrons

O Nucleus excitation and deexcitation (evaporation)

"Typical" hadronic shower

Electromagnetic component

Contributions

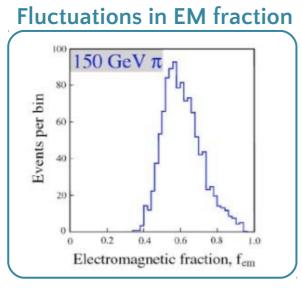
O Electrons & photons

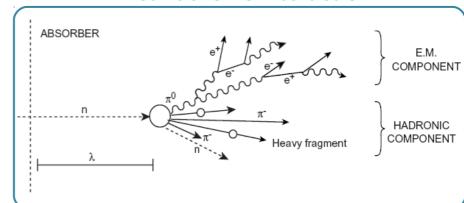
 \bigcirc Neutral pions (e.g. $\pi^{\circ} \rightarrow \gamma \gamma$)

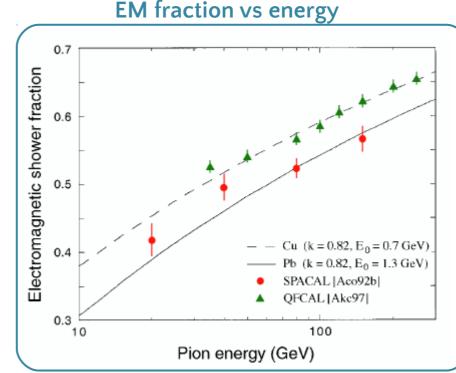
About 1/3 of π^o produced at each nuclear interaction

• On average, **EM fraction increases with** energy $\langle f_{em} \rangle = \langle E_{EM} / E_{tot} \rangle = 1 - (E / E_0)^{k-1}$ E_0 = average energy needed for π^0 production

k is related to the average multiplicity of π^{o} produced at each interaction







First hadronic interaction

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Non-EM components

Non-EM energy breakdown

Numbers for Lead

56% ionizing particles 2/3 are protons (from spallation) <E> - 50-100 MeV

34% invisible Break-up of nuclei

10% neutrons Very soft (typically a few MeV) On average 37n per deposited GeV

	Lead	Iron
Ionization by pions	19%	21%
Ionization by protons	37%	53%
Total ionization	56%	74%
Nuclear binding energy loss	32%	16%
Target recoil	2%	5%
Total invisible energy	34%	21%
Kinetic energy evaporation neutrons	10%	5%

A large part of energy losses is **invisible**

O Energy used to release protons and neutrons from nuclei

O Kinetic energy carried by recoil nuclei

Also significant fraction in evaporation neutrons

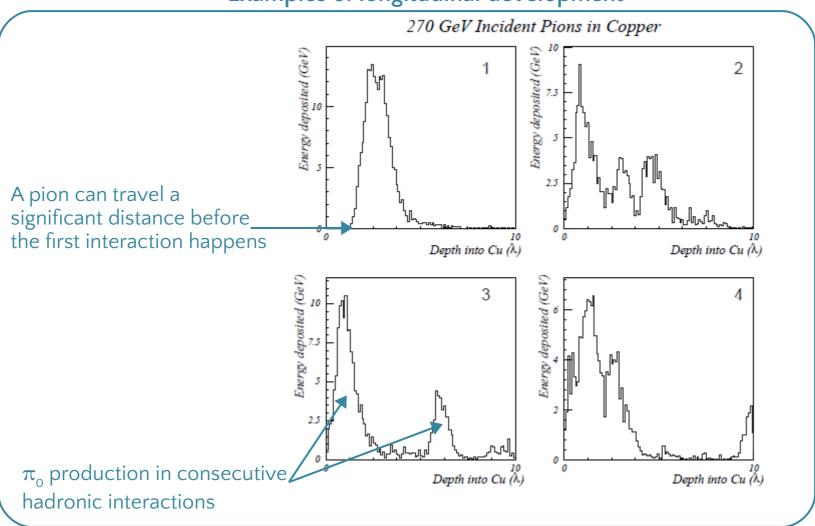
O Elastic scattering (large energy transfer for small nuclei, e.g. Hydrogen)

O Neutron capture (sizeable energy, but late w.r.t. main shower component)

Shower development

A hadronic shower doesn't have a profile which can be parameterized

■ The size of the 1st interaction will essentially determine the EM fraction



Examples of longitudinal development

Longitudinal profile

■ Hadronic shower governed by the interaction length $\lambda_{int} \propto A^{1/3}$

O Mean free path between inelastic interaction

Depth to contain shower increases with ln(E)

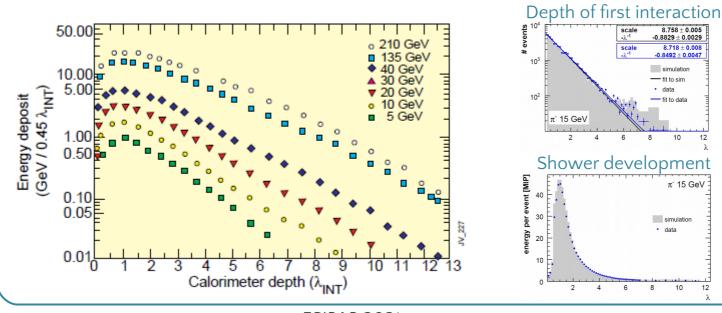
O Similarly to EM showers

But convolution of two components

O Depth of the first interaction

O Shower development

Longitudinal profile



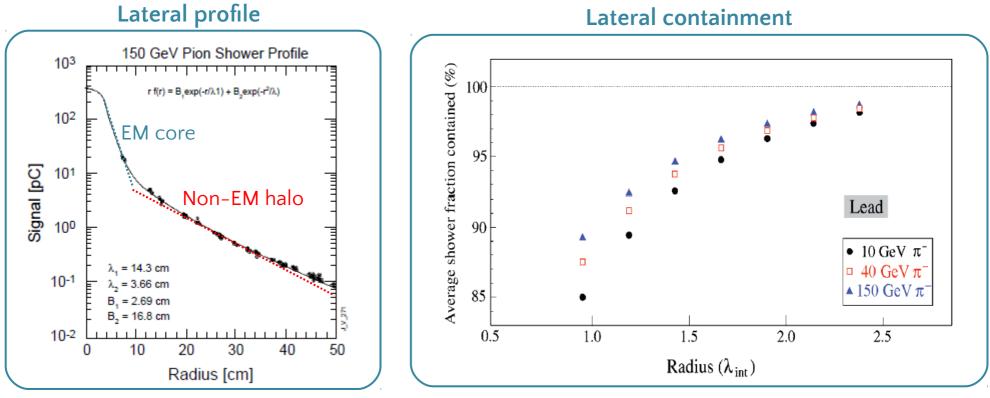
Examples of radiation and interaction length

	Z	ρ (g.cm -3)	E _c (MeV)	X ₀ (cm)	λ _{int} (cm)
Air				30 420	~70 000
Water				36	84
PbWO ₄		8.28		0.89	22.4
С	6	2.3	103	18.8	38.1
AI	13	2.7	47	8.9	39.4
L Ar	18	1.4		14	84
Fe	26	7.9	24	1.76	16.8
Cu	29	9	20	1.43	15.1
W	74	19.3	8.1	0.35	9.6
Pb	82	11.3	6.9	0.56	17.1
U	92	19	6.2	0.32	10.5

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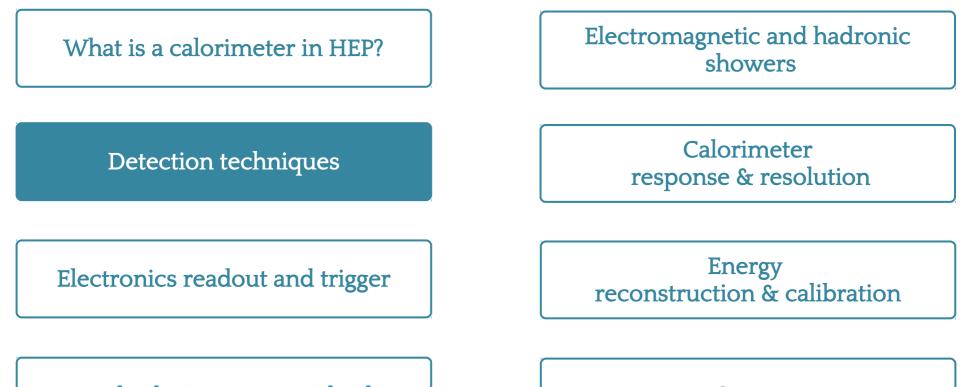
Lateral profile

- Lateral shower development has two components
 - ⊖ Electromagnetic core
 - Non-EM halo (mainly non-relativistic shower particles)
- EM shower core prominently present in the initial stages
- Energy dependence of the lateral containment is directly related to the EM fraction dependence



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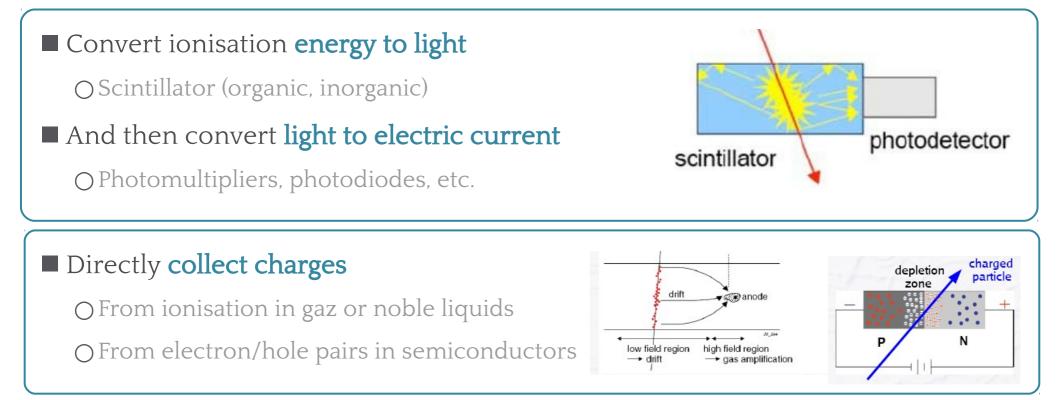
Lecture plan



Beyond calorimetry: Particle Flow

An example: CMS HGCAL

Measurement principles



One can also measure (very small) temperature increase

Oe.g. Bolometers

O Not presented here

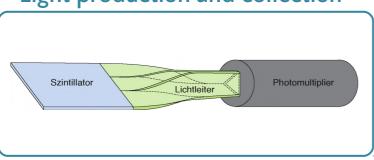
Scintillators

Excited atoms or molecules in scintillating medium

O De-excitation and emission of light (visible, UV, sometimes X-rays)

O Propagation of light (wavelength shifters can be used, light guides)

O Conversion into electric signal (photo-detector)



Light production and collection

2 types of scintillating material

Organic (plastics or liquids)

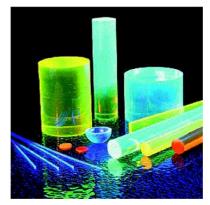
○ Inorganic (crystals)

2 types of light emission

 \bigcirc Fluorescence: prompt (ns $\rightarrow \mu$ s)

 \bigcirc Phosphorescence: emission over long period ($\mu s \rightarrow ms$)

Various scintillators



CMS PbWO4 crystal



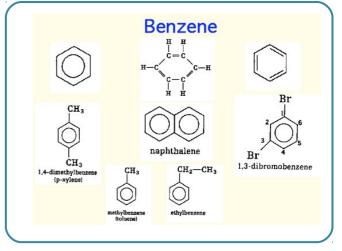
Organic scintillators

- Organic crystals, organic liquids, plastic scintillators
 - ○Aromatic hydrocarbon compounds (containing carbon rings)
 - OBase/solvent + fluor
- Transition of electrons between molecular levels
 scintillation
 - O Fast: few ns
 - ⊖ Fluorescent **UV light**

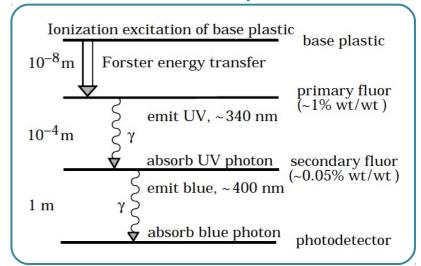
UV light absorbed in most organic material

- O Second fluorescent material for conversion in visible light
- ⊖a.k.a. wavelength shifter
- Usually made of **low Z / low density material**
 - $\bigcirc \rightarrow$ more volume
 - O But inexpensive

Aromatic hydrocarbons



De-excitation process



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Inorganic scintillator

Crystalline structure

- O Energy bands
- O Ionizing particles create free electrons and holes
- O Excites activation centres (impurities or doping)
- \bigcirc Decay \rightarrow light emission
- Slower than organic scintillator (> 100 ns)

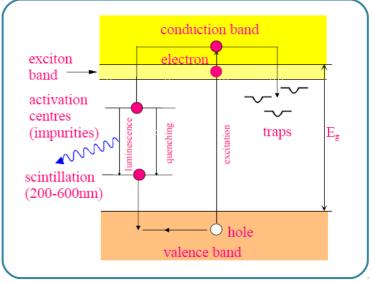
But high Z / high density

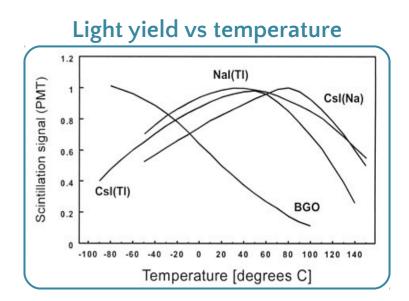
Material	Form	$\begin{array}{c} \lambda_{max} \\ (nm) \end{array}$	τ_f (ns)	$\underset{(g/cm^3)}{\rho}$	Photons per MeV
NaI(Tl) (20°C)	crystal	415	230	3.67	38,000
pure NaI (-196°C)	crystal	303	60	3.67	76,000
${\rm Bi}_4{\rm Ge}_3{\rm O}_{12}~(20^{\circ}{\rm C})$	crystal	480	300	7.13	8,200
Bi ₄ Ge ₃ O ₁₂ (-100°C)	crystal	480	2000	7.13	24,000
CsI(Na)	crystal	420	630	4.51	39,000
CsI(Tl)	crystal	540	800	4.51	60,000

Light output depends on temperature

O Needs good control and monitoring

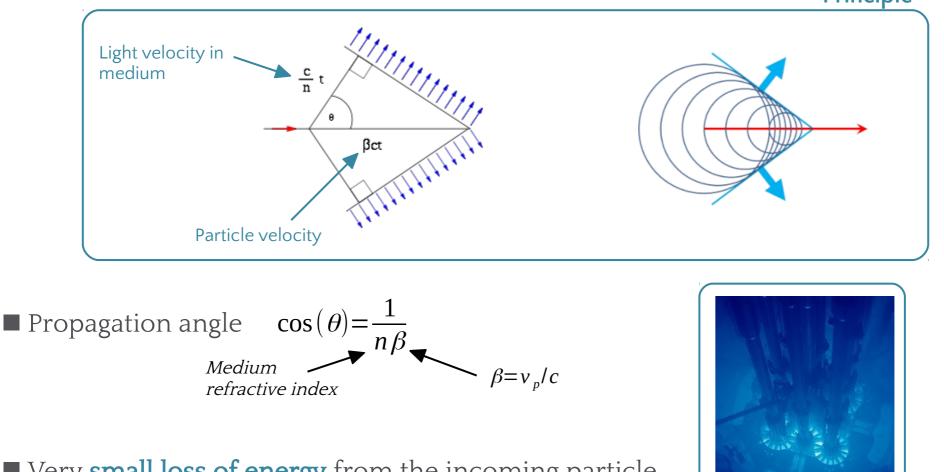
Light emission process





A bit on the Cerenkov effect

Collective effect when charged relativistic particle passes through matter at a speed higher than the speed of light (in the medium)
Principle



Very small loss of energy from the incoming particle

 \bigcirc e.g. -400 eV / cm for a particle with $\beta{\simeq}1$ in water

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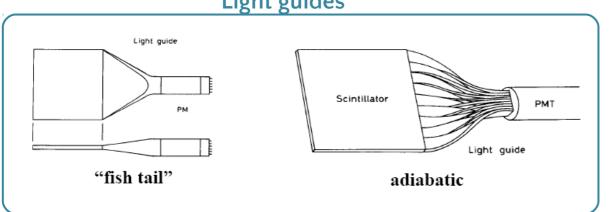
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Spectrum in $1/\lambda^2$

 \rightarrow appears blue/violet in the visible spectrum

Photodetectors

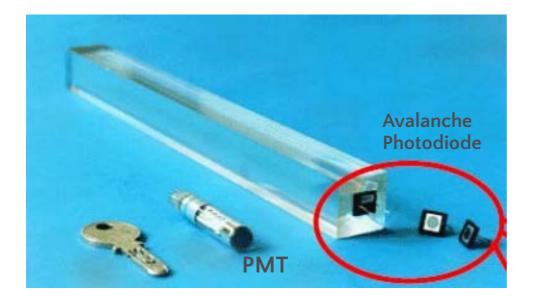
Conversion of scintillation (or Cerenkov) light to electric signals





Large PMT

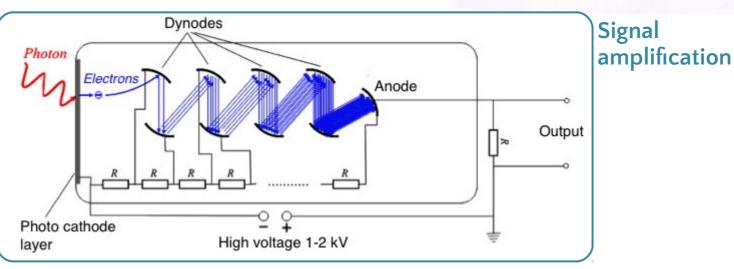




Light guides

Photomultipliers

- Photons hit a **photo cathode** → electrons (photoelectric effect)
- High voltage to accelerate the produced electrons
- Succession of dynodes to amplify the signal OProduces secondary electrons
- Can reach high amplification gains (10⁴ 10⁷)
- But several drawbacks
 - **O** Bulky
 - O Expensive,
 - O Sensitive to magnetic fields



Varieties of PMT



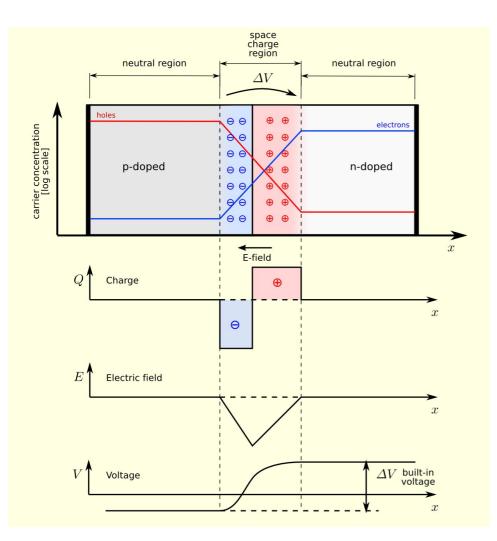
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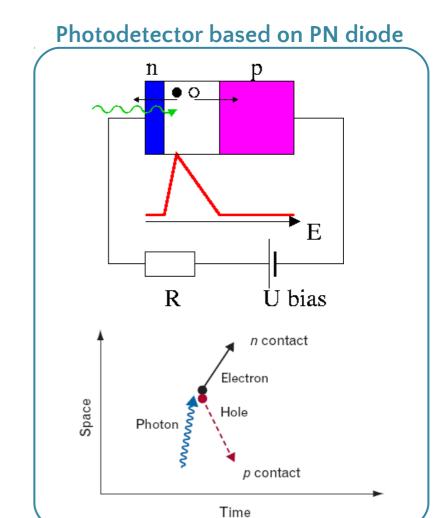
PN diodes

Semiconductor based photo detectors are more used nowadays

Based on PN diodes with reverse bias

O Photon creates electron-holes pair in the depletion region



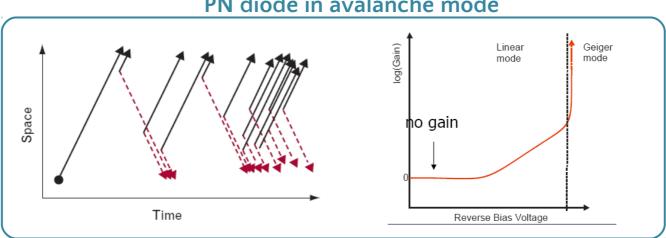


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Silicon PM, Avalanche Photo Diodes

Applying a high reverse voltage

O Creates electron-hole multiplication



PN diode in avalanche mode

Linear mode = Avalanche Photo Diode (APD)

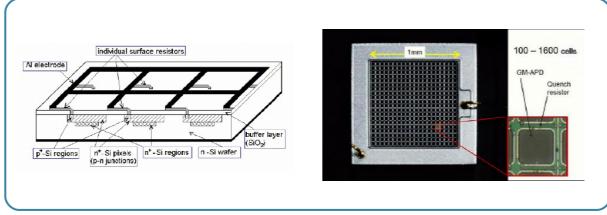
■ Geiger mode (or "Single Photon") APD

O Above breakdown voltage

O Binary mode

 $\bigcirc \rightarrow$ Arrays of G-APD = Silicon PM

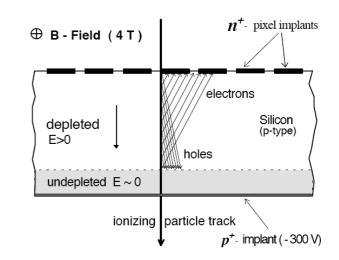
Array of Geiger mode APD

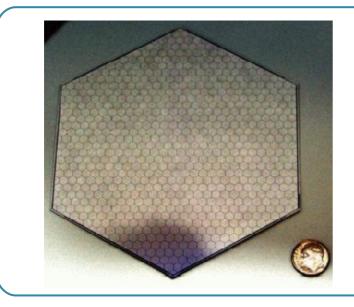


Direct charge collection with semiconductor detectors

Silicon detectors can also collect charges from ionization

- O Also used for tracking
- An electron can create many electron-hole pairs
 - O About 9000 e-h created / 100 microns
- Thin: few 100 microns thickness
- High bias voltage: few 100V





Silicon sensors with hexagonal cells



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Direct charge collection with noble liquids

Two processes in noble liquids

 \bigcirc Molecule **excitation** \rightarrow UV light emission

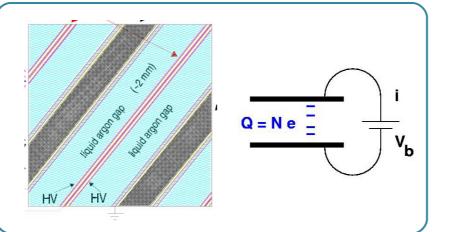
○ Molecule **ionization** → electron and ion drift

Dense material

○ Lots of charge → no charge amplification needed

O Good stability, good homogeneity

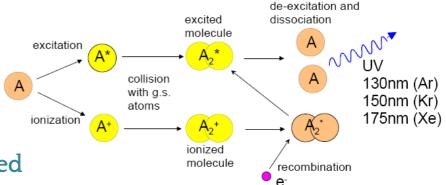
Charge collection



Similar principles can also be used in gas detectors

 \bigcirc But low density, so amplification needed \rightarrow less stable

O Larger detectors



Noble liquids characteristics

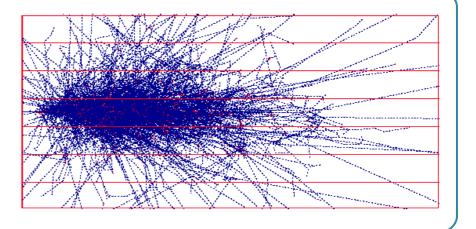
	Ar	Kr	Xe
Z	18	36	58
А	40	84	131
X ₀ (cm)	14	4.7	2.77
E _C (Mev)	41.7	21.5	14.5
R _M (cm)	7.2	4.7	4.2
W (eV/pair)	23.3	20.5	15.6
v drift (mm/µs)	10	5	3

Two calorimeter types

Homogeneous calorimeter

⊖ Single medium for

- Shower development (dense material)
- Signal collection
- O "All" energy deposited is collected



Sampling calorimeter

\bigcirc Two materials

- One for shower development (absorber)
- One for signal collection (detectors / active material)
- Only energy deposited in active material is collected
 - The shower is sampled

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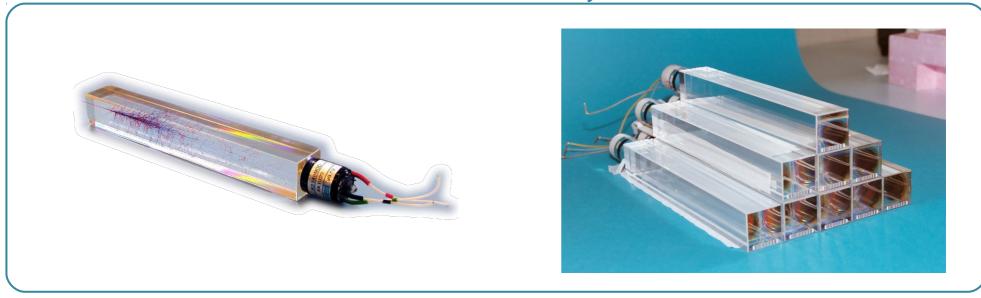
Homogeneous calorimeters

- Based on scintillating crystals with high density and high Z
- Very good energy resolution and linearity

But

- O Very expensive
- O Radiation damages can be a problem
- O No longitudinal (depth) segmentation

CMS ECAL PbWO4 crystals



Sampling calorimeter

- Absorber with high-density material
- Interleaved with active readout devices
- Lower cost than homogeneous calorimeters
- And can have longitudinal segmentation

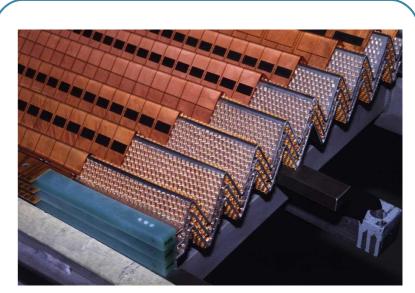
But

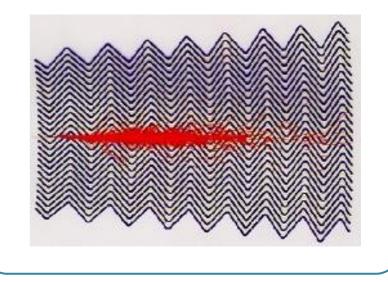
- O Only part of the shower energy is collected
- O Fluctuation of energy deposited in active layers
 - Proportional to number of charged particles

$$\frac{\sigma(E)}{E} \propto \frac{1}{\sqrt{n_{ch}}}$$

(And less charges deposited compared to homogeneous calorimeters)

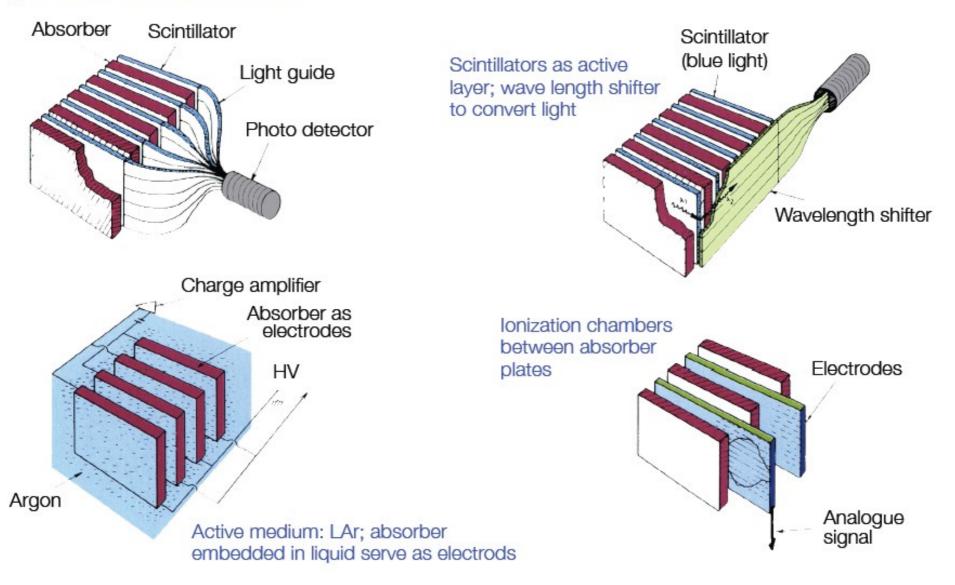
ATLAS Liquid Argon ECAL





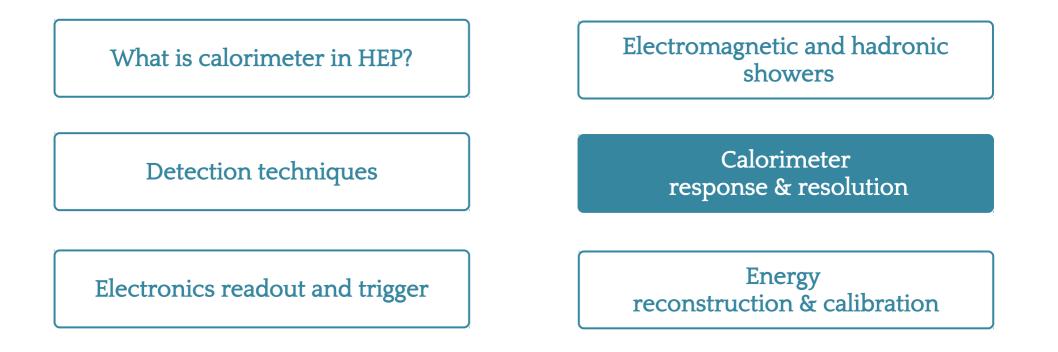
Some types of sampling calorimeters

Scintillators as active layer; signal readout via photo multipliers



Possible setups

Lecture plan

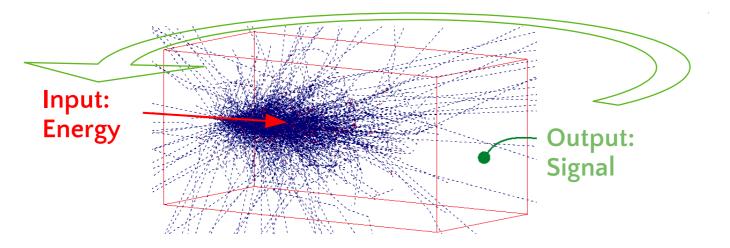


Beyond calorimetry: Particle Flow

An example: CMS HGCAL

Measurement of showers

From collected signal back to the energy of the particle



■ Average signal collection → **response** of the calorimeter to the **input energy**

- O Ideally proportional to the input energy (linearity)
- O Homogeneous and sampling calorimeters behave differently
- O The response differs between EM and hadronic showers

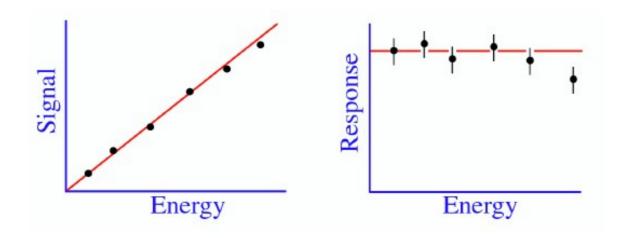
■ Fluctuations of the collected signal → calorimeter **resolution**

○ For a given energy there are shower to shower variations of the signal

O Contributions to these fluctuations differ with the energy

Calorimeter response

- Response: average signal per unit of deposited energy
- A linear calorimeter has a constant response
 - O The signal is proportional to the deposited energy



In general electromagnetic calorimeters are linear

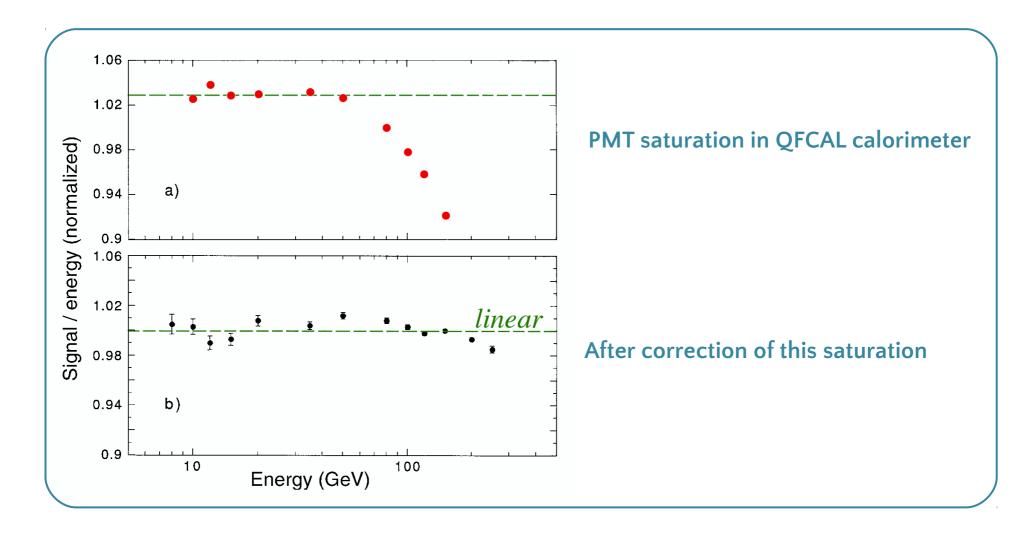
O All energy deposited through ionization / excitation of absorber

Hadronic calorimeters are not

Sources of non-linearity (1/2)

Instrumental effects

Oe.g. saturation of scintillators, photo-detectors, electronics



Sources of non-linearity (2/2)

■ Non-linearity appears if response depends on something that varies with energy

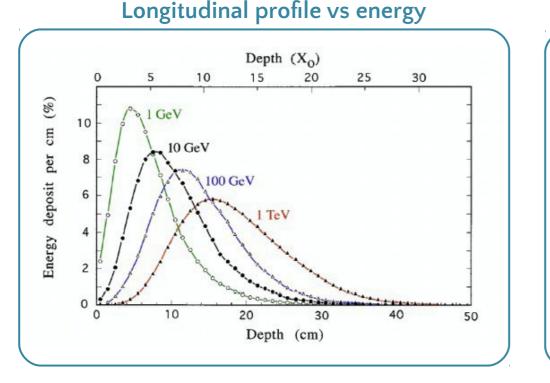
O e.g. if deposited energy counts differently depending on depth

- Since depth increases with energy

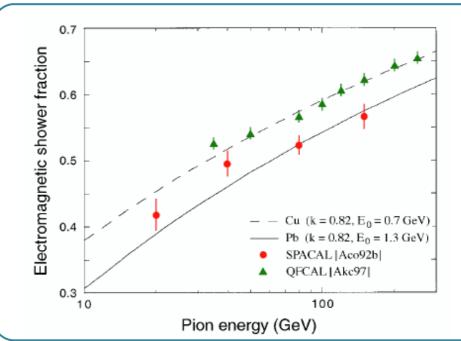
O Electromagnetic and hadronic energies count differently

- And **EM fraction increases with energy** → non linear hadronic calorimeter response

Energy leakage







Sampling: EM vs mip response

Homogeneous calorimeters

○ Same deposit mechanisms for EM showers and mips → same response (e/m=1)

Sampling fractions defined for a minimum ionizing particle (mip)

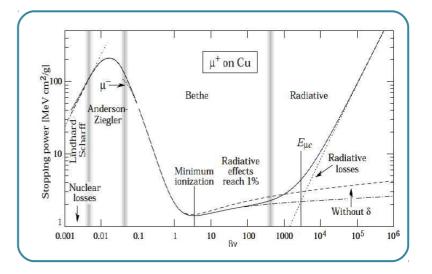
 $f_{samp} = \frac{E_{mip}^{loss \text{ in active}}}{E_{mip}^{total \, loss}}$

But EM showers are not sampled like mips

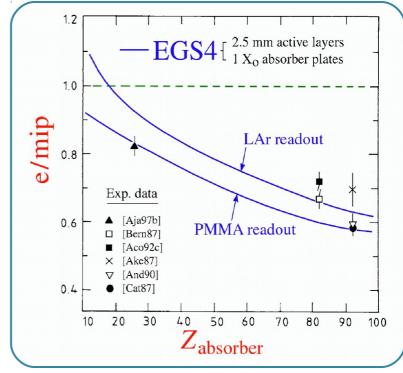
O Photoelectric effect $\sigma_{pe} \propto Z^5$

- O Soft photons are very inefficiently sampled due to the Z asymmetry between absorber and detector
- O Only photoelectrons produced near the boundary (<1mm) between active and passive material produce a signal

Sampling calorimeters: e/m < 1 (or << 1)



e/m vs absorber Z



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Sampling: e/m dependence with shower depth

e/m changes as the shower develop

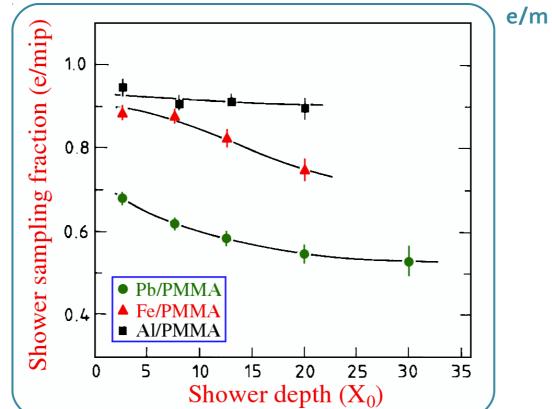
Because the shower composition evolves

O Early phase: relatively fast shower particles (pairs)

O Tails: dominated by Compton and photoelectric electrons

Longitudinally segmented calorimeter

O Must calibrate differently vs depth



e/m vs shower depth

Hadronic response and compensation

Response to the hadronic part usually smaller than to the EM part Oe/h > 1

O Invisible nuclear binding energy, escaping muons and neutrinos

○ Saturation effects, etc.

■ e = h → compensating calorimeter

O Can be obtained in non-homogeneous calorimeters

○ Homogeneous calorimeters are in general non-compensating

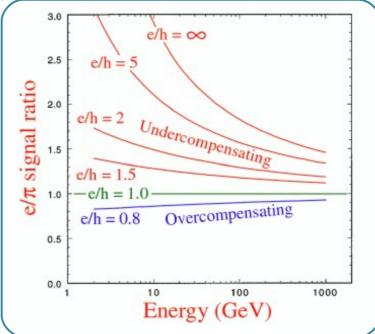
e/h not directly measurable

O Uses pion response

$$e/\pi = \frac{e}{f_{EM}e + (1 - f_{EM})h} = e/h \frac{1}{1 + f_{EM}(e/h - 1)}$$

• EM fraction increases with $E \rightarrow$ non linearity





Non-linearity and e/h

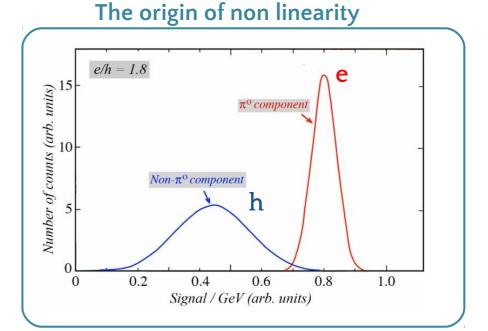
■ Non-linearity determined by e/h

$$\frac{\pi(E_1)}{\pi(E_2)} = \frac{f_{EM}(E_1) + [1 - f_{EM}(E_1)] \cdot e/h}{f_{EM}(E_2) + [1 - f_{EM}(E_2)] \cdot e/h}$$

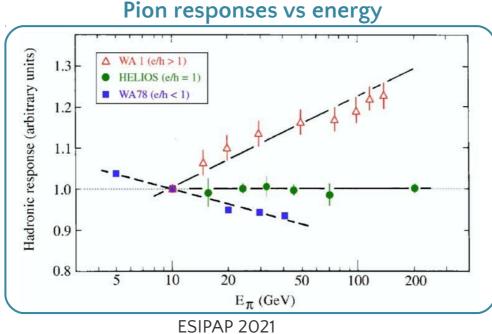
O Assuming linearity for EM showers

○ And e/h constant with energy

$$e/h=1 \Rightarrow \frac{\pi(E1)}{\pi(E2)}=1$$



■ Inversely: measurement of non-linearity is one method to determine e/h



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Compensation

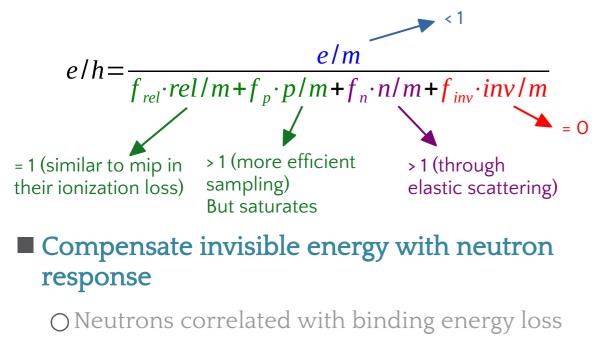
Non-electromagnetic shower energy components

 \bigcirc Ionization by charged pions (relativistic shower component) \rightarrow \mathbf{f}_{rel}

 \bigcirc Spallation protons (non-relativistic shower component) \rightarrow \mathbf{f}_{p}

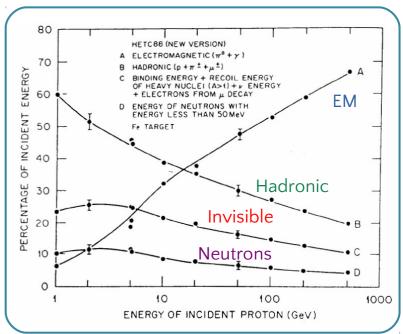
 \bigcirc Kinetic energy carried by **evaporation neutrons** \rightarrow \mathbf{f}_n

 \bigcirc The energy used to **release protons and neutrons from nuclei**, and the kinetic energy carried by **recoil nuclei** do not lead to a signal $\rightarrow \mathbf{f}_{inv}$



Reduce e/m further

Sharing of energy between components



Increasing neutron response

Elastic scattering higher with light nucleus: $f_{elastic} = 2 A / (A+1)^2$

○0.5 for Hydrogen

○0.005 for Lead

Recoil protons can be measured

■ Pb / H₂ calorimeter structure with 50/50 sharing

```
O 1 \text{ MeV neutron deposits 98\% in H}_2 
O mip deposits 2.2% in H<sub>2</sub> n/\text{mip} = 45
```

■ Pb / H₂ calorimeter structure with 90/10 sharing

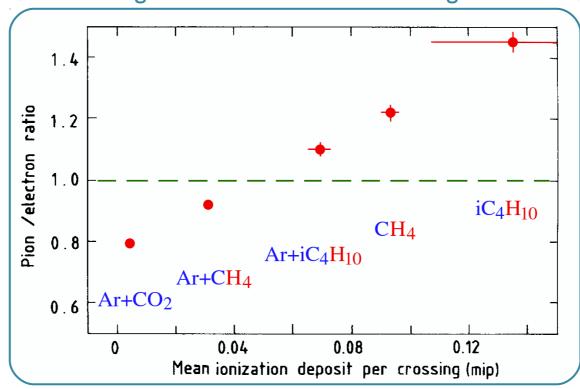
 $\bigcirc 1 \text{ MeV}$ neutron deposits 87% in H₂ $\land n/m$

 \bigcirc mip deposits 0.25% in $H_{\scriptscriptstyle 2}$

Tuning neutron response with active material

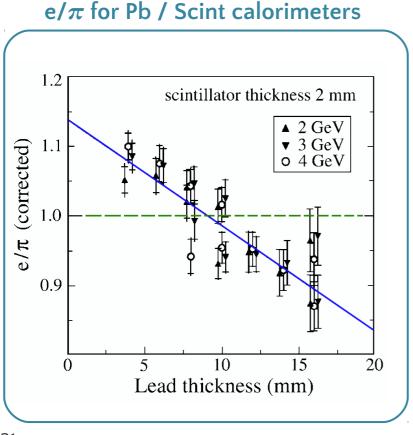
The key to boost hadronic response is to use hydrogenous active material
 The response can be tuned using more or less hydrogeneous material

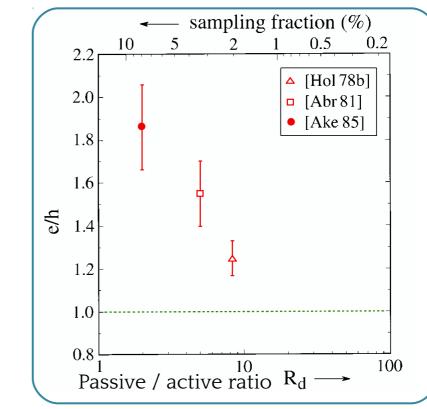
 π /e for several materials, with different energy deposits by slow neutrons (Uranium / gas calorimeter with different gas mixtures)



Tuning neutron response with sampling fraction

- Finer tuning can be obtained by **adjusting the sampling frequency**
- Works best with Lead and Uranium
 - O e.g. a ratio of 4:1 gives compensation for Pb/Scint
- In principle also possible with iron, but only a few neutrons generated ○ Ratio > 10:1 needed → deterioration of longitudinal segmentation





e/h for Fe / Scint calorimeters

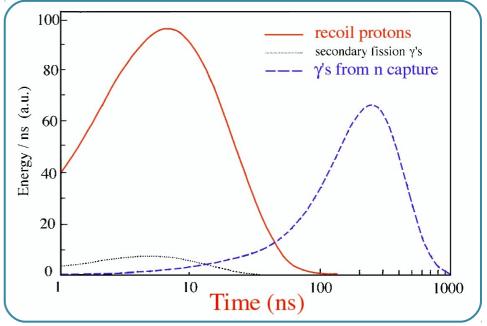
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Energy released by slow neutrons

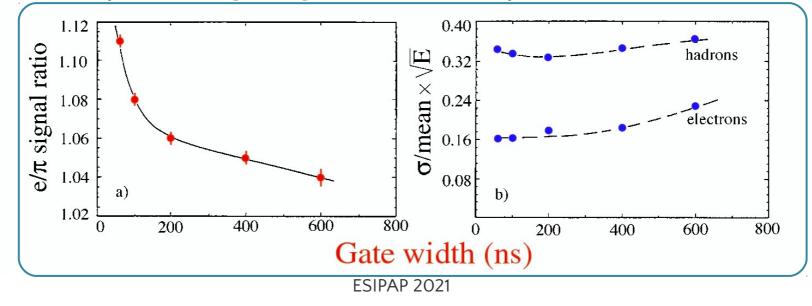
- Large fraction of neutron energy captured and released after 100 ns
- Needs long integration time to collect this energy
- Trade-off between compensation and noise integration (resolution degradation)

01/02/21

Time structure of neutron-induced processes in U / Scint calorimeter



Impact of charge integration time on compensation and resolution



Decreasing EM response

Electrons and photons are sampled less efficiently when using high-Z absorber

 \bigcirc Photoelectric effect cross section $\propto Z^5$

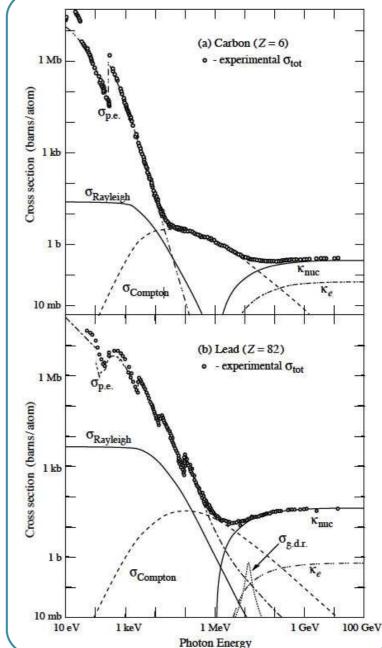
O Photons < 1MeV captured in absorber

Recipe for compensating hadron calorimeter

O High Z absorber

O Hydrogenous active medium

O Precisely tuned sampling fraction



Cross sections in Carbon and Lead

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Pros & cons of compensating calorimeters

Pros

O Same energy scale for electrons, hadrons and jets

O Just need to calibrate with electrons

O Excellent hadronic resolution

O Linearity, Gaussian response distribution

Cons

○ Small sampling fraction → EM energy resolution limited

O Compensation relies on detecting neutrons

- Large integration volume
- Long integration time (~50 ns) \rightarrow noise integration

Energy resolution

- Calorimeter's energy resolution is determined by **fluctuations**
- Input energy E x N number of secondary particles
 - O Poisson distribution of N $\rightarrow \sigma(E)/E \propto \sqrt{N}/N \propto 1/\sqrt{E}$
 - O Although in reality only a fraction can be detected (threshold effects)

Other types of fluctuations

- O Signal quantum fluctuations (e.g. photoelectron statistics)
- O Sampling fraction
- O Shower leakage
- O Instrumental effects (electronic noise, light attenuation, non-uniformity, etc.)
- O Hadronic-specific fluctuations (EM fraction, invisible energy)

Sampling fluctuations

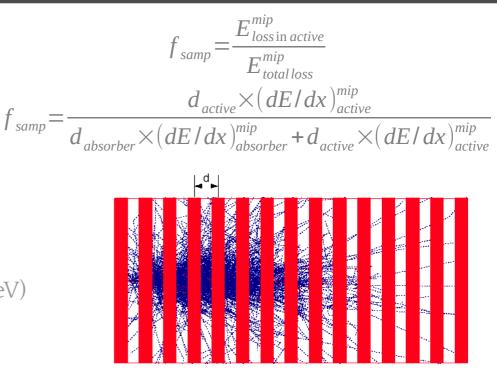
- Two aspects in sampling fluctuations
- **Sampling fraction**: fraction of energy deposited in active material

 \bigcirc Lower sampling \rightarrow less particles collected \rightarrow larger fluctuations

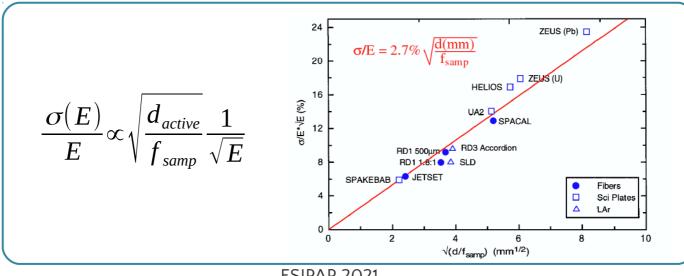
Active layer thickness

O Large fraction of low energy electrons (< 1MeV) produced in absorber

O Traveling a small distance in active material



Thicker active layer → worse resolution Lower sampling \rightarrow worse resolution



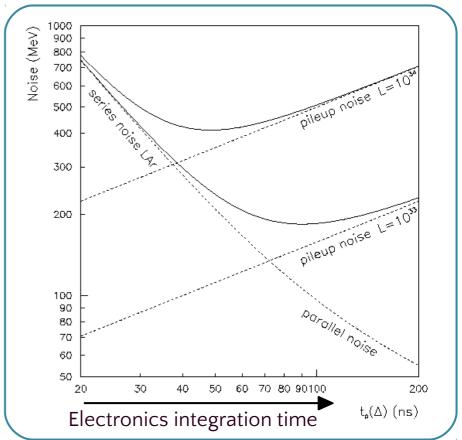
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Noise

Noise fluctuations are constant in energy

- O → Impact resolution in 1/E (mainly low energy)
- Usually comes from the electronics readout system
- But at hadron colliders
 - O Contributions from **pile-up interactions**
 - O = fluctuations due to multiple low energy collisions

Electronic noise vs pile-up noise (ATLAS LAr calorimeter)

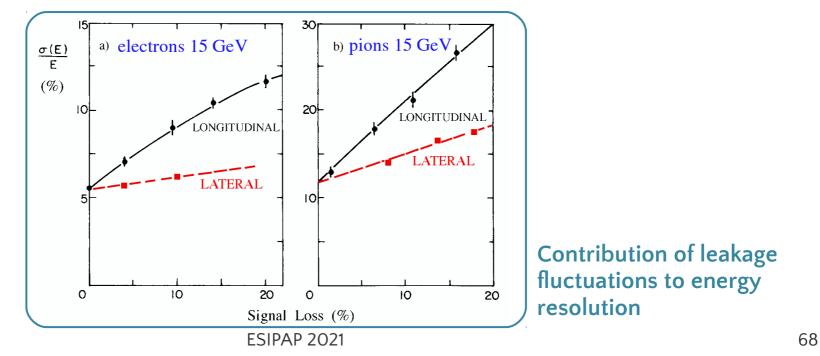


Leakage

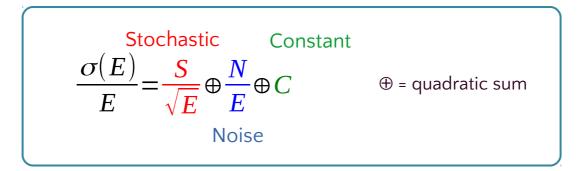
- Energy from secondary particles escaping measurement
 - O Non-Poissonian fluctuations
- Longitudinal leakage (rear of the detector)
 - OA detector is never infinitely deep
 - O Dangerous since increases as log(E)
 - O Alleviated if calorimeter "sufficiently" deep

Lateral leakage

- O One tends to limit the lateral size over which the signal is integrated
- O Need to limit integration of channels with low S/N
- O Need to limit integration of nearby showers



Energy resolution: parametrization



Stochastic term

O Everything with a **Poisson-like** statistics

O Intrinsic particle fluctuations, sampling, quantum fluctuations

Noise term

O Internal (e.g., electronics) and external (e.g. pileup) noise

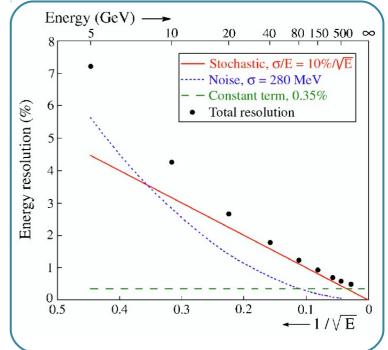
Constant term

O Fluctuations due to **leakage**

O Imperfections in construction, **non-uniformity**

- Local variations of temperature, light attenuation, material thicknesses, etc.

Energy resolution in EM ATLAS barrel calorimeter



Homogeneous vs sampling calorimeters

Technology (Experiment)	Depth	Energy resolution	Date
NaI(Tl) (Crystal Ball)	$20X_0$	$2.7\%/E^{1/4}$	1983
$Bi_4Ge_3O_{12}$ (BGO) (L3)	$22X_0$	$2\%/\sqrt{E}\oplus 0.7\%$	1993
CsI (KTeV)	$27X_0$	$2\%/\sqrt{E} \oplus 0.45\%$	1996
CsI(Tl) (BaBar)	$16 - 18X_0$	$2.3\%/E^{1/4} \oplus 1.4\%$	1999
CsI(Tl) (BELLE)	$16X_0$	1.7% for $E_{\gamma} > 3.5 \text{ GeV}$	1998
$PbWO_4 (PWO) (CMS)$	$25X_0$	$3\%/\sqrt{E}\oplus 0.5\%\oplus 0.2/E$	1997
Lead glass (OPAL)	$20.5X_0$	$5\%/\sqrt{E}$	1990
Liquid Kr (NA48)	$27X_0$	$3.2\%/\sqrt{E} \oplus \ 0.42\% \oplus 0.09/E$	1998
Scintillator/depleted U (ZEUS)	$20 - 30X_0$	$18\%/\sqrt{E}$	1988
Scintillator/Pb (CDF)	$18X_0$	$13.5\%/\sqrt{E}$	1988
Scintillator fiber/Pb spaghetti (KLOE)	$15X_{0}$	$5.7\%/\sqrt{E} \oplus 0.6\%$	1995
Liquid Ar/Pb (NA31)	$27X_0$	$7.5\%/\sqrt{E}\oplus 0.5\%\oplus 0.1/E$	1988
Liquid Ar/Pb (SLD)	$21X_0$	$8\%/\sqrt{E}$	1993
Liquid Ar/Pb (H1)	$20 - 30X_0$	$12\%/\sqrt{E}\oplus 1\%$	1998
Liquid Ar/depl. U (DØ)	$20.5X_{0}$	$16\%/\sqrt{E}\oplus 0.3\%\oplus 0.3/E$	1993
Liquid Ar/Pb accordion (ATLAS)	$25X_{0}$	$10\%/\sqrt{E} \oplus 0.4\% \oplus 0.3/E$	1996

Table 33.8: Resolution of typical electromagnetic calorimeters. E is in GeV.

Homogeneous

Sampling

Fluctuations in hadron showers

Same types of fluctuations as in EM showers +

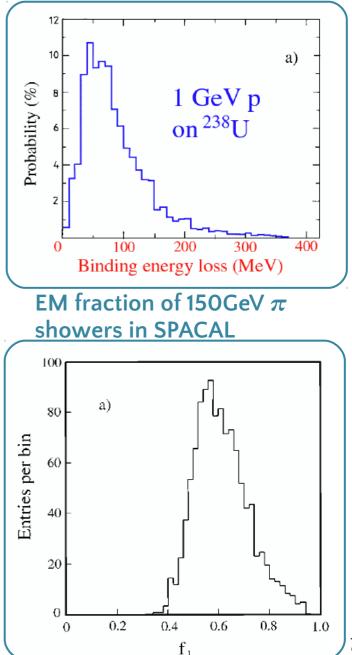
Fluctuations in visible energy

- O Fluctuations in losses due to nuclear binding energy
- O Note: Correlation with the number of neutrons produced in spallation reactions

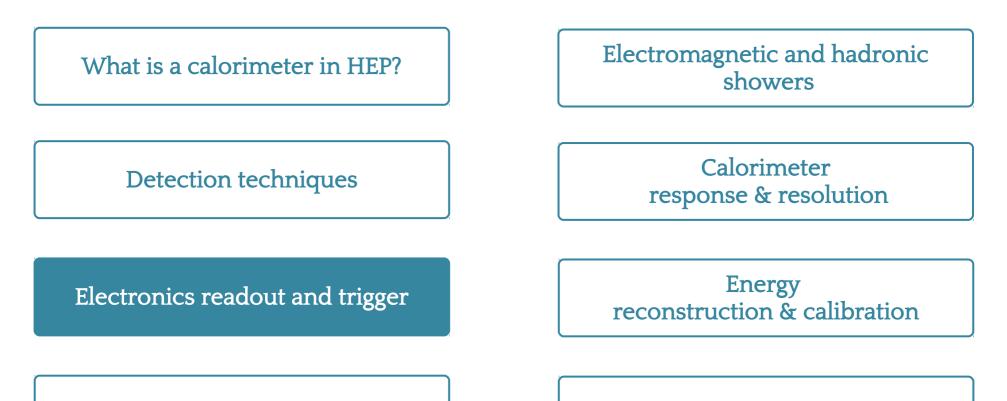
Fluctuation in the EM shower fraction

- \bigcirc Dominating effect in most hadron calorimeters, where e/h \neq 1
- \bigcirc Due to the irreversibility of π_0 production \rightarrow asymmetry in EM fraction distribution
- O Ideally need to measure the EM fraction for each shower

Binding energy loss for 1 GeV proton in Uranium



Lecture plan

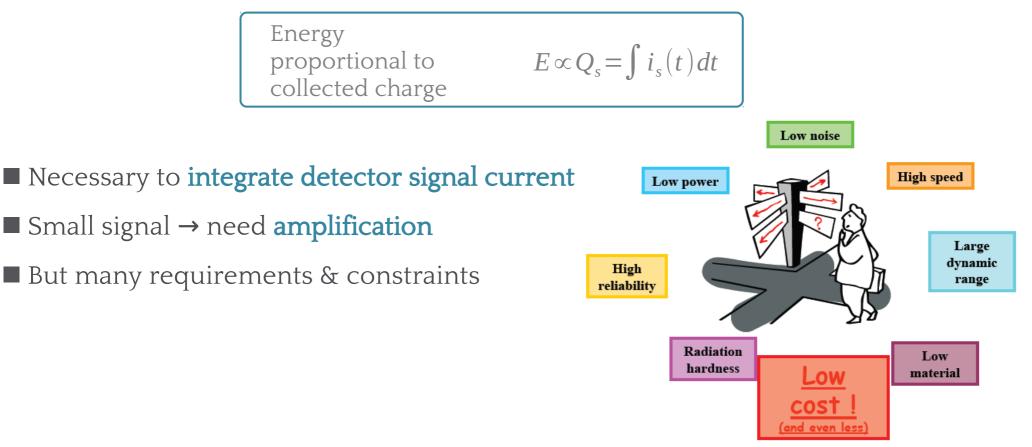


Beyond calorimetry: Particle Flow

An example: CMS HGCAL

Signal acquisition

- Determine energy deposited and event time in detector
- Detector signal = short current pulse
 - O Thin silicon detector (10-300 μm): 100ps-30ns
 - O Thick (-cm) Si or Ge detector: 1-10 μ s
 - O Scintillator + PMT/APD: 100ps 10μs



Data acquisition



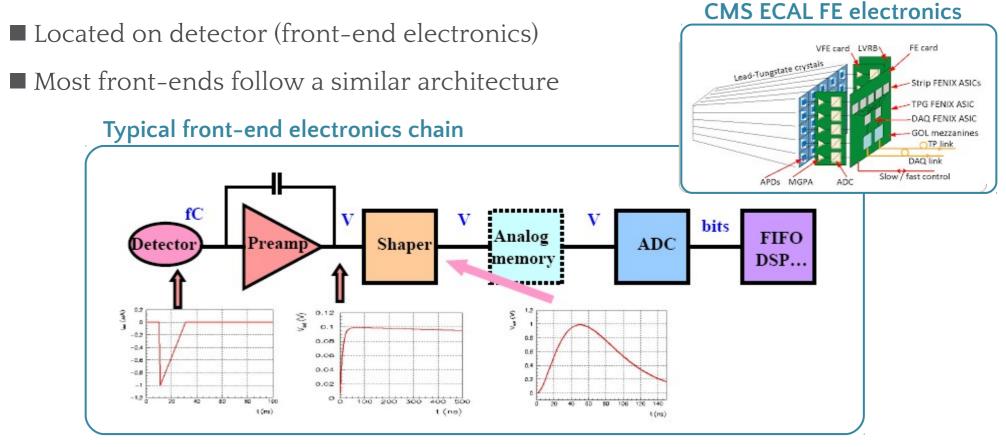
- Signal is **converted to digita**l values
- Put in **buffers** until it is read out by the DAQ (data acquisition) system and transferred to mass storage (disk, tape)
- In many cases, all the data cannot be stored or transferred

```
\bigcirc \rightarrow Trigger system
```

O Real-time processing of events on reduced data

OAccept / reject decision

Readout electronics



Very small signals (fC)

O Amplification needed

O Optimisation of S/N (shaper)

Need time to decide to keep the event or not \rightarrow memory

Conversion from analogue signal to digital values (ADC pulse sampling)

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Pulse shaping

Goal = increase S/N ratio

Cut low frequency and high frequency noise

O Limit bandwidth

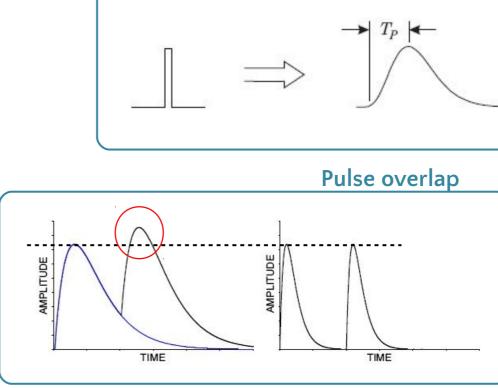
⊖ Shaping = filtering

- Limited frequency band with characteristic time consistent with the input pulse
- In case of successive signal pulses
 O Shaped pulse need to be short enough
 O Avoid signal overlap

Two conflicting effects

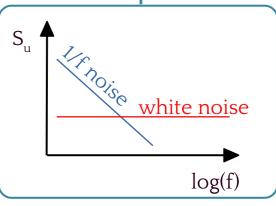
O Filter noise

O Avoid overlap



SENSOR PULSE

Noise spectrum



Shaper output

SHAPER OUTPUT



CR-RC shaper

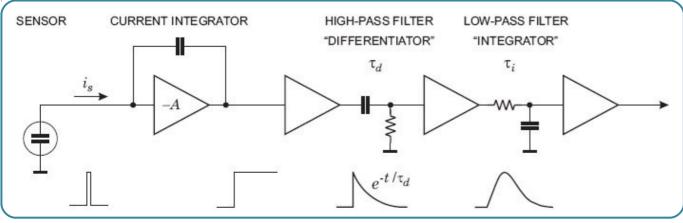
CR circuit = "differentiator" = high-pass filter

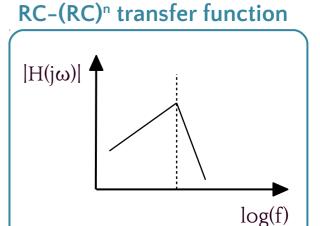
O Sets the duration of the pulse

RC circuit = "integrator" = low-pass filter

O Increases the rise time to limit high-frequency noise

Integrator + RC-RC chain

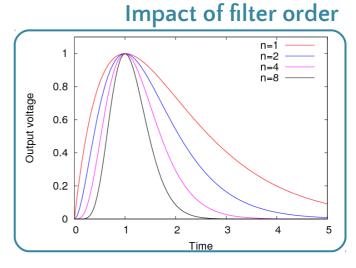




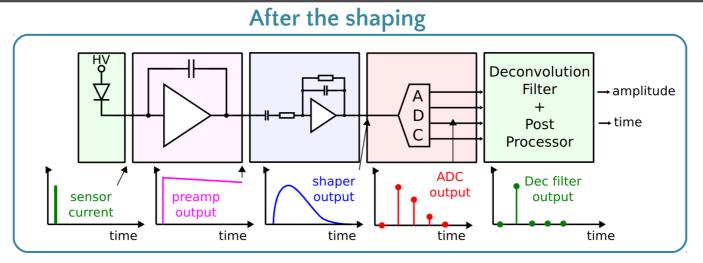
Can also have additional RC or CR steps

 \bigcirc e.g. CR-(RC)²

OCR-(RC)ⁿ approximates Gaussian shape

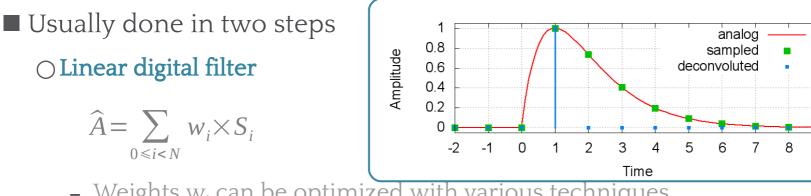


Amplitude measurement



Pulse at output shaper = convolution of input signal and readout chain response

Can perform **deconvolution** of the two, using the sampled pulse values



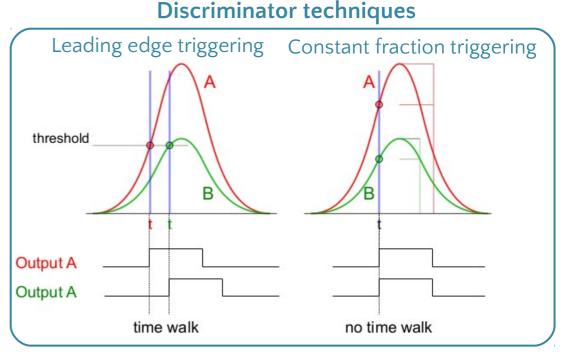
- Weights w_i can be optimized with various techniques
- Takes into account noise characteristics and known pulse shape

O Peak finder to find maximum

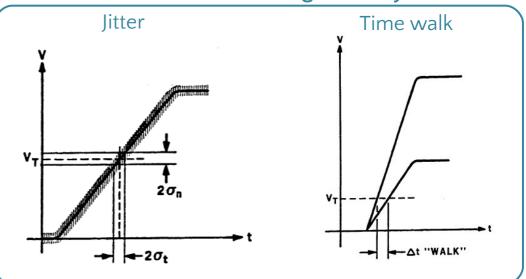
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Time measurement

- Time measurement based on discriminators
 - Compare voltage level of signal with a given level (threshold)
- Often discriminator output depends on the signal amplitude
 - O = Time walk
- Slope to noise ratio characterizes the time measurement
- Measurement affected by
 - **Jitter**: due to high-frequency noise
 - Time walk (can eventually be corrected)
- Ideally needs
 - O Fast rise time of the pulse
 - O Low noise



Measurement degraded by



01/02/21

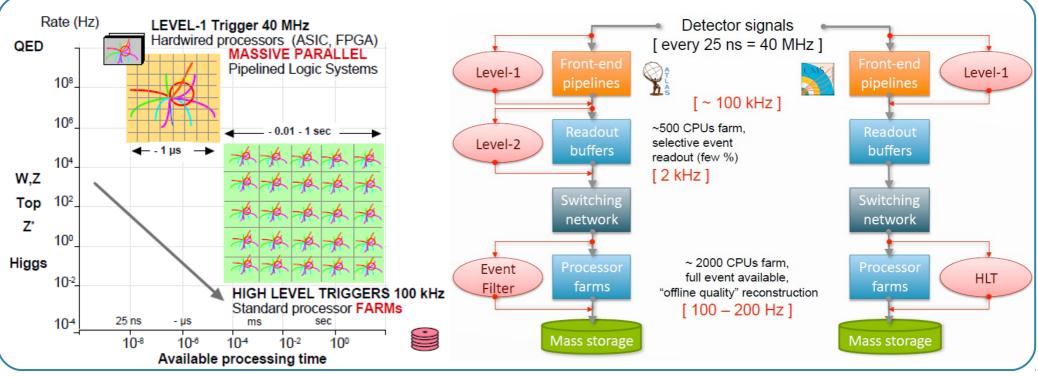
Trigger

In many cases it is impossible to store or transfer all the data from a calorimeter O e.g. at the LHC

Need fast trigger system to eliminate uninteresting events

O Trade-off between high performance (signal efficiency / background rejection)O And fast decision / low latency

One solution: multi-stage system



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Rate reduction in multi-stage systems

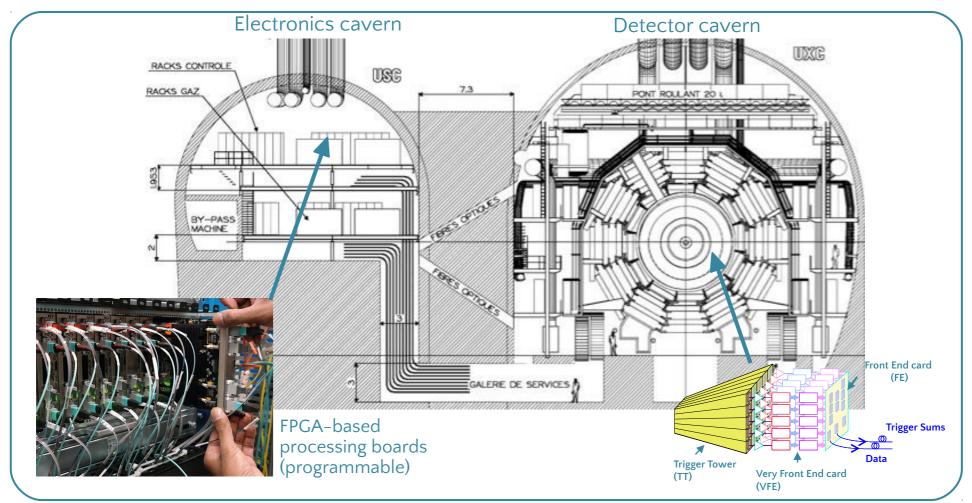
Off-detector electronics

■ First trigger level (L1) processing : hardware based (ASIC, FPGA)

O Data simplification / coarsification on-detector (in **front-end electronics**)

O Reconstruction and decision making off-detector (in **back-end electronics**)

O Linked with optical fibers



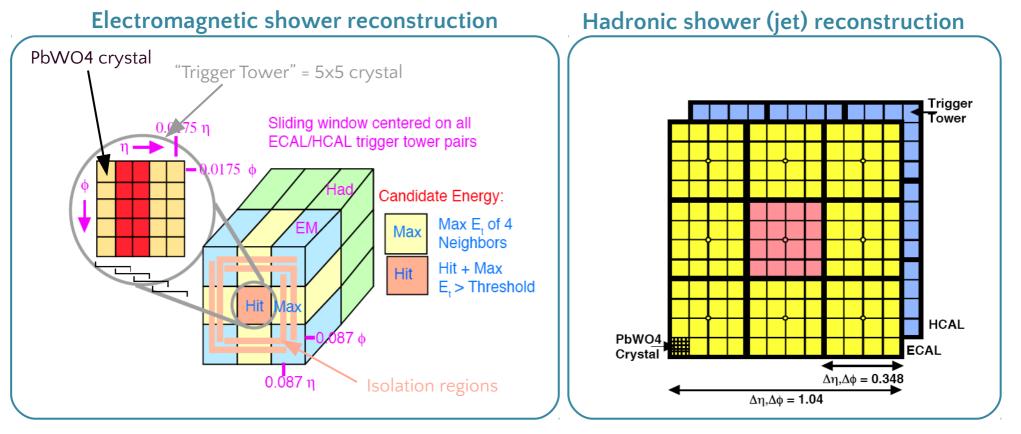
Calorimeter trigger objects, examples (CMS, Run 1)

Coarse granularity

O Cannot send the full granular data out of the calorimeter

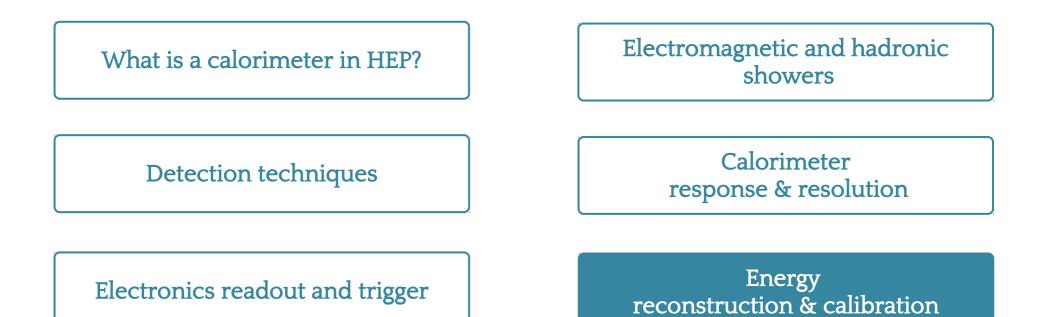
Simple reconstruction and identification of the showers
OFaster

Apply cut on the (transverse) energy of these objects



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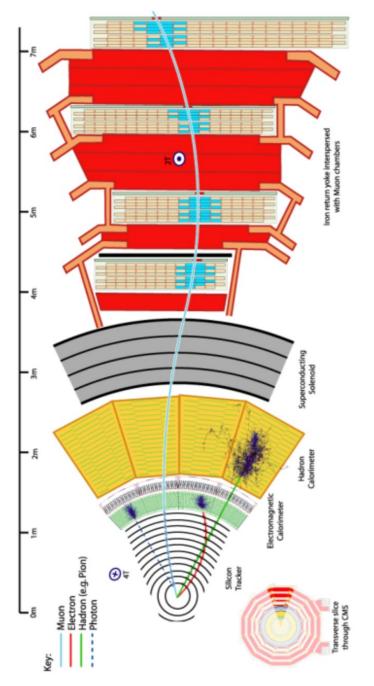
Lecture plan



Beyond calorimetry: Particle Flow

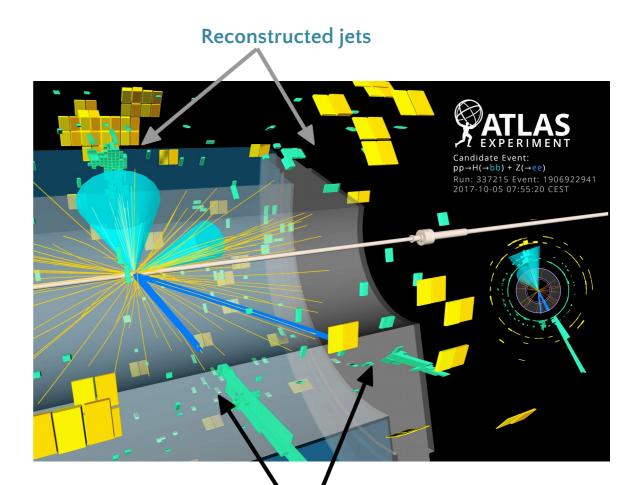
An example: CMS HGCAL

Reconstructing particles in calorimeters



Focus on objects at colliders O Electrons and photons

O Hadronic jets

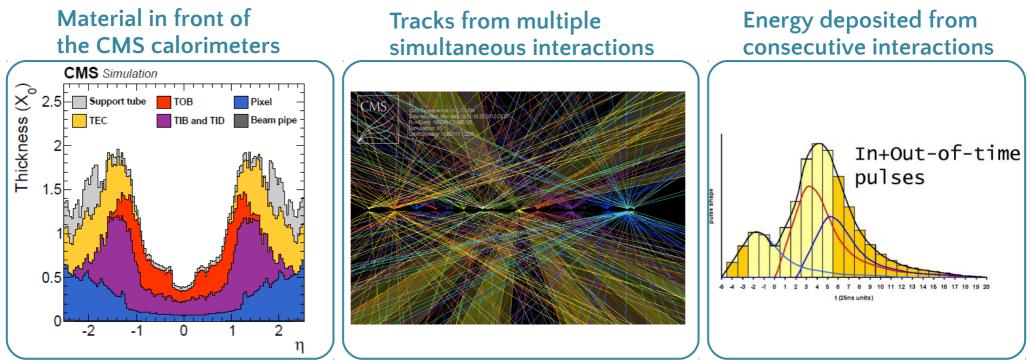


Real-life conditions

- Real conditions of a full detector in its environment are harsh
 - O Very high number of calorimeter channels (100k to several millions)
 - Magnetic field (impact on photodetectors, electronics, mechanics)
 - O Material in front of the calorimeter
 - **Radiations**, Pile-up (in-time and out-of-time)

0...

Degrades performance compared to standalone devices or test beams



Electrons and photons reconstruction

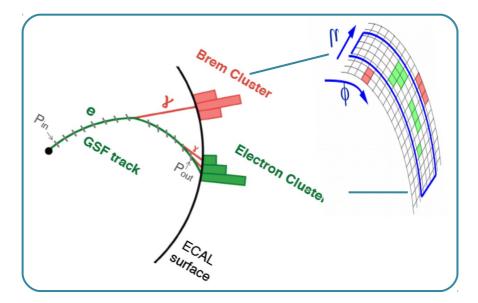
■ Material in front of the calorimeter (cables, cooling, mechanical support, etc.)

○ Electrons initiate showers before reaching the ECAL (e.g. 40-80%)

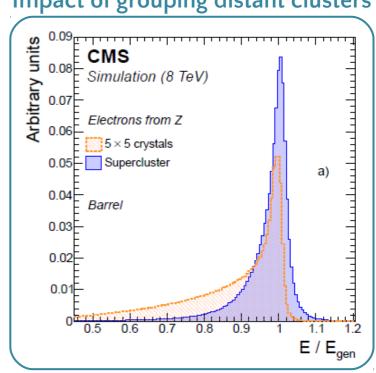
○ Photons convert (e.g. 20-40%) in e+e- pairs

■ Magnetic field → radiated energy spread along phi

OOne electron / photon can produce several distinct showers

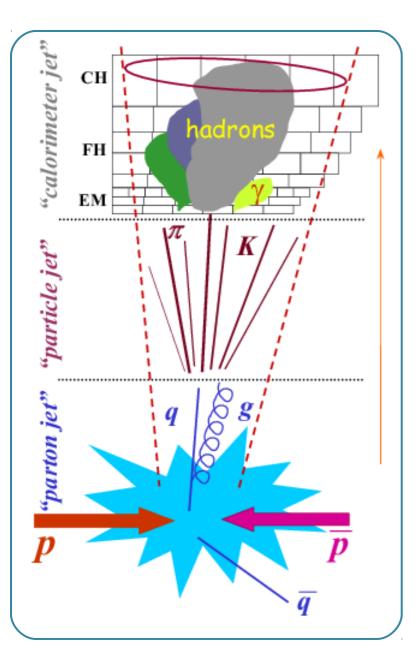


Reconstruction algorithms need to
 Cluster topologically connected deposits
 Associate distant clusters together



Impact of grouping distant clusters

From single hadrons to jets



- Quarks and gluons (a.k.a. partons) can be produced from collisions (e.g. protons)
- They produce secondary partons
 O Parton shower
- These partons turns into collimated hadrons
 OHadronization
- This set of collimated hadrons is called a "jet"
 O Each hadron will shower in a calorimeter
 O A jet will create a set of overlapping showers
- Jets are reconstructed in two steps
 - **O Clustering** of topologically connected deposits
 - **Grouping clusters** together with distancebased association (creates cone-like objects)

Jet algorithms

Usually inputs are clusters of energy deposits

O But can be any vector-like object

- Two main classes of algorithms
 - Cone algorithms (older)

○ Sequential clustering algorithms (used nowadays)

Sequential algorithms based on a distance

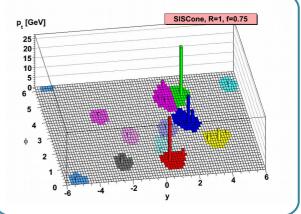
O Most commonly used: anti-kt

$$d_{ij} = min(\frac{1}{p_{Ti}^2}, \frac{1}{p_{Tj}^2}) \times \frac{R_i^2}{R}$$

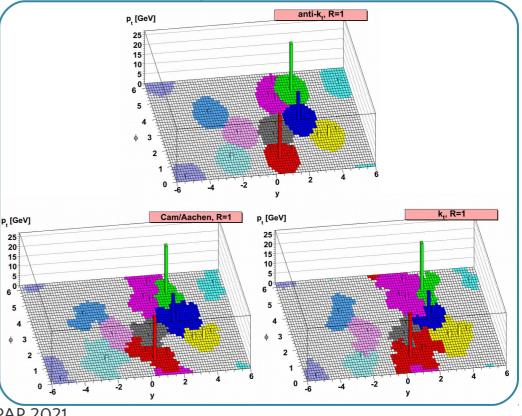
 \bigcirc Cluster first hard objects \rightarrow very stable

O Some others more sensitive to the jet substructure

Example of cone algorithm







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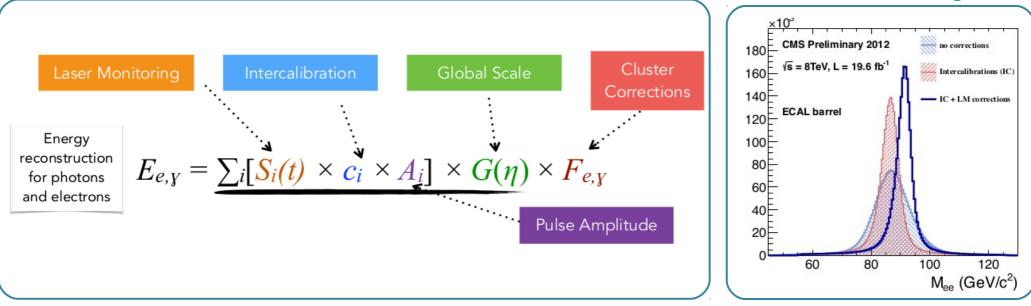
Calorimeter calibration / energy reconstruction

Typical calibration components

○ Pulse amplitude → energy (charge → energy deposited, can include sampling fraction)

O Response effects from monitored parameters

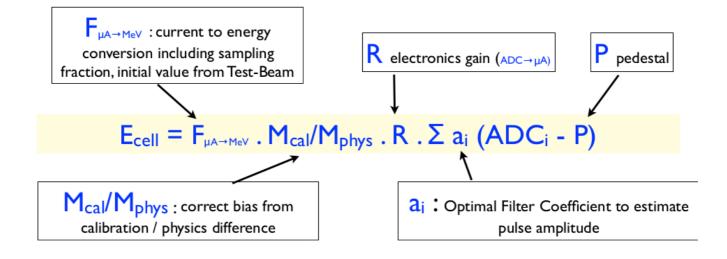
- O Intercalibration (looking at differences of response in different places)
- Leakage corrections (energy outside of cluster)
- ⊖Absolute energy scale correction



Example in CMS ECAL

Impact of intercalibration and laser monitoring

Pulse calibration

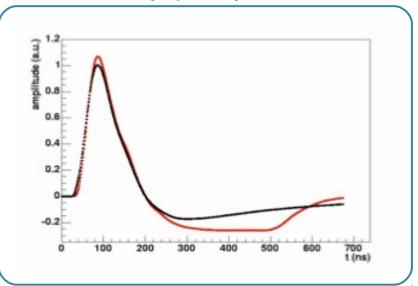


Correct the electronics chain response

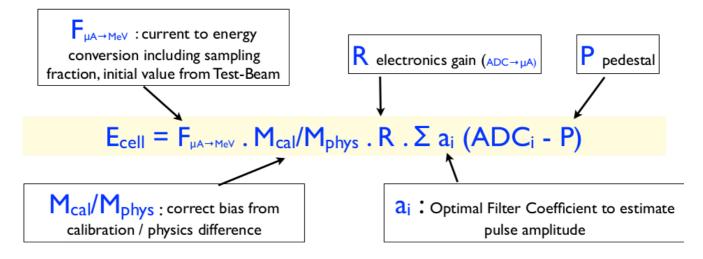
O Linear filtering of ADC samples (seen before)

- Can derive the response by injecting known signal (with dedicated calibration circuits)
 - Differences between calibration pulses and physics pulses need to be taken into account

Calibration and physics pulses differences



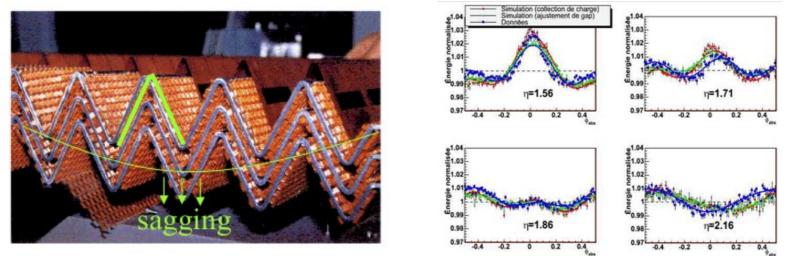
Pulse calibration



Take into account sampling fraction in case of sampling calorimeters

Local variations of the sampling fraction need to be understood as precisely as possible
Cravity effects impact sampling fraction

Gravity effects impact sampling fraction



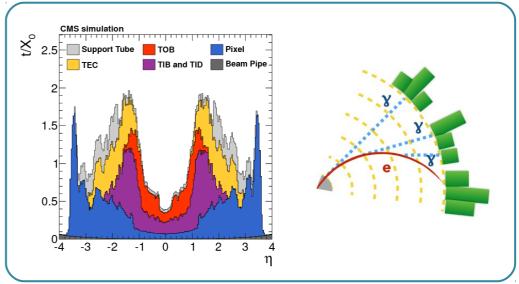
Cluster energy corrections

Several sources of energy reconstruction inefficiencies

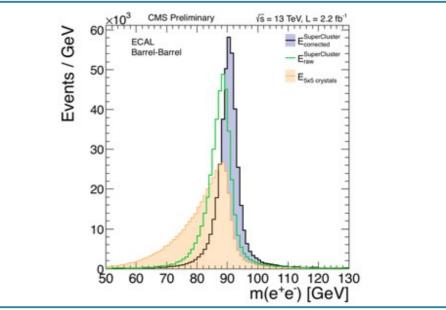
- O Threshold applied (e.g. to reduce noise impact)
- O Unclustered energy (linked to calorimeter depth, cluster size, interactions with material in front of the calorimeter)
- Remaining noise propagating to the final reconstructed cluster
- Requires reconstruction algorithms optimizations
- And corrections to take into account the remaining effects

O Derived from detailed detector simulations

Material in front of the CMS ECAL



Impact of clustering and corrections on reconstructed Z mass



Monitoring

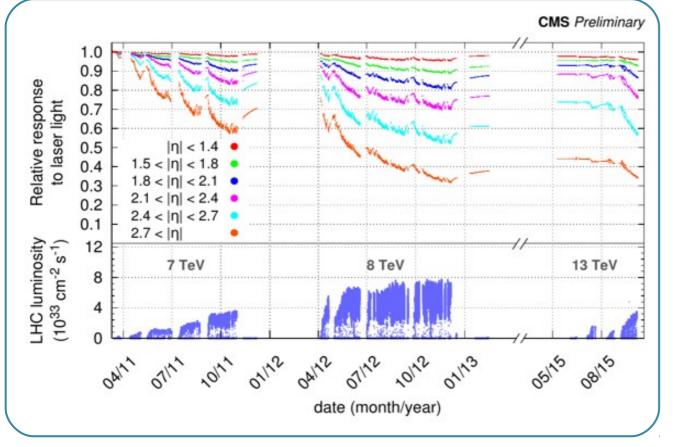
Some parameters affecting the detector response change with time

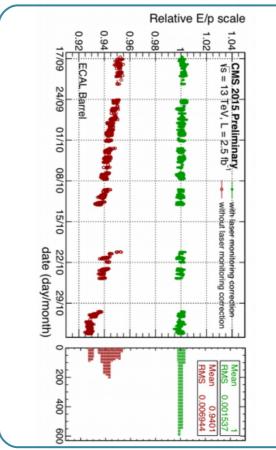
O Temperature, radiation effects (e.g. light attenuation), etc.

These parameters are **monitored** and the response is corrected accordingly

CMS ECAL PbWO4 crystal loss of transparency monitored with a laser system (light injection)

Effect of monitored crystal response on reconstructed electron energy





Intercalibration

There are always local effects which cannot be perfectly simulated nor measured

■ Need to equalize the response of each sensor one to an other

○ In situ (e.g. using collision data) intercalibration

In general, several methods are combined

O Using detector symmetries

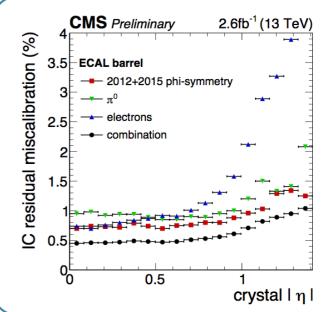
OUsing "standard candles" (e.g. known resonances)

O Comparing with other sub-detectors

Intercalibration methods used in CMS ECAL

Method	Description	Timescale
φ-symmetry	Energy flux around ϕ rings (constant η) should be uniform - IC corrects for non- uniformity	~days
⊓⁰∕η→γγ	In a $\pmb{\varphi}$ ring, use IC to improve M($\pmb{\gamma}\pmb{\gamma}$) resolution for π^0 and $\pmb{\eta}$ resonances	~months
E/p	Compare isolated electron energy from ECAL and Tracker, calculate IC to correct discrepancies	statistically limited

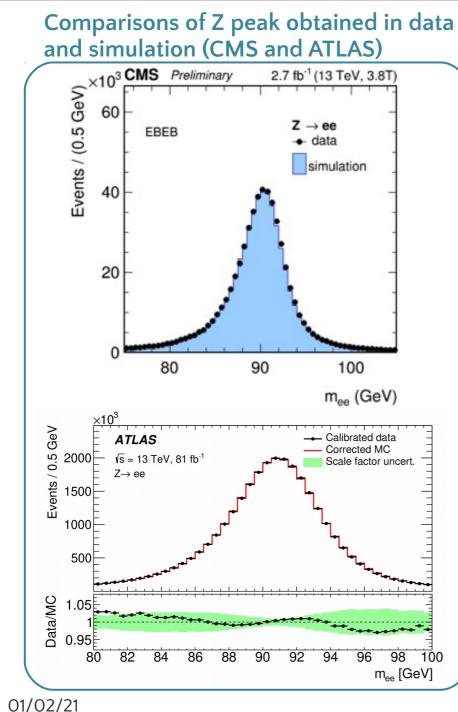
Precisions obtained with each intercalibration methods



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Absolute calibration

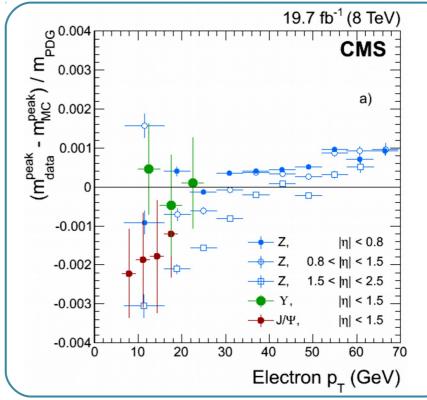
Comparisons of Z peak obtained in data and simulation (CMS and ATLAS)



Calibration factors to set the overall scale

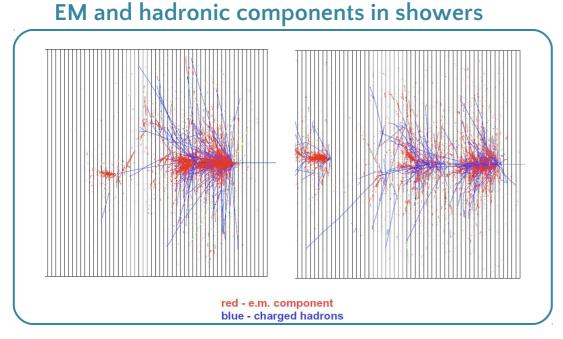
- In situ, based on standard candles (e.g. Z→ee)
- O Match data Z peak lineshape to simulation

Linearity check with different resonances

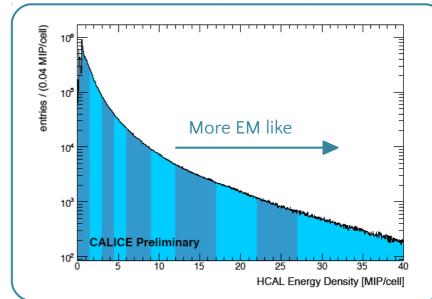


Jet energy measurement

- Similar techniques are used to calibrate jets made of hadron showers
- Additional techniques for non-compensating calorimeters
 - O Need to measure or estimate the EM fraction
 - O Apply different weights according to the EM fraction
 - O Called "software compensation"
- EM shower are narrow and dense and hadronic showers are more diffuse
 - $\bigcirc \rightarrow$ Apply weights according to energy density



Cell energy density

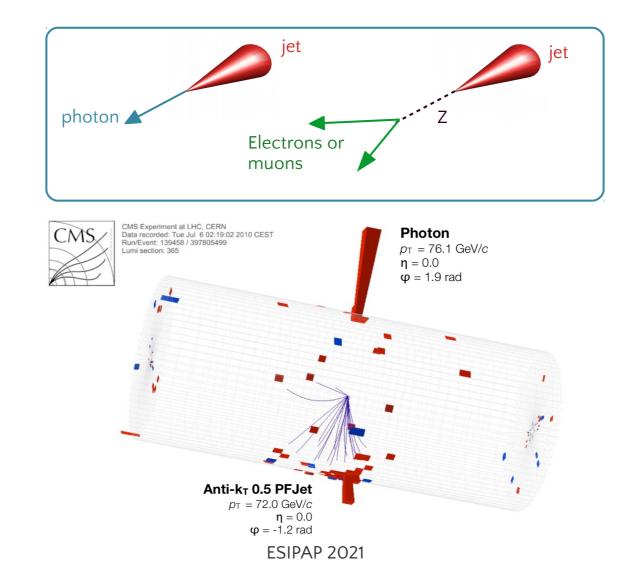


Jet in situ calibration

The jet energy resolution is poorer compared to other particles

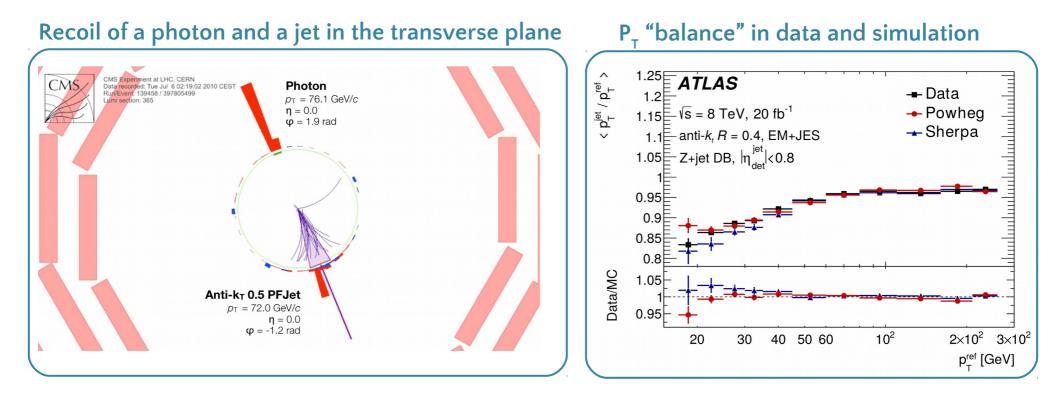
O Electrons, photons, muons

■ Use the **recoil between a jet and precisely measured objects**

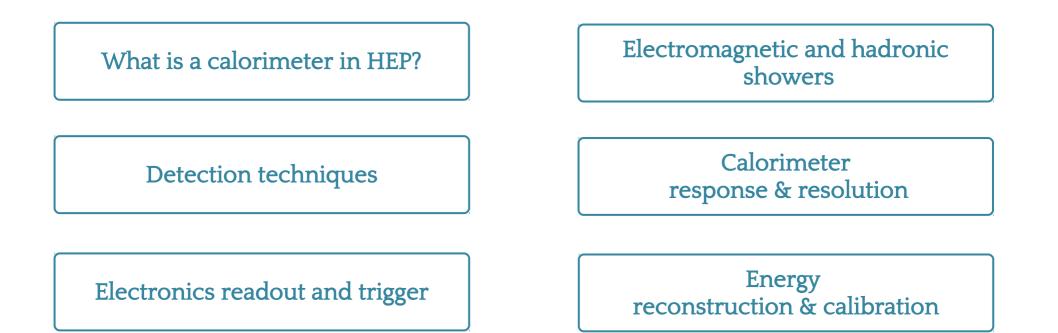


Jet in situ calibration

In the transverse plane, objects in an event are recoiling one against the othersOne corrects for discrepancies in data and simulation



Lecture plan



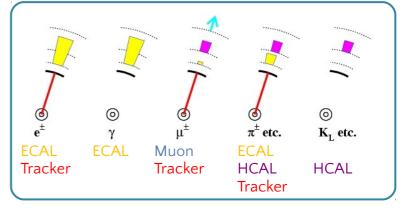
Beyond calorimetry: Particle Flow

An example: CMS HGCAL

Reconstruction beyond calorimeters

- Electrons, photons, hadrons, etc. produce different signatures in the different subdetectors
 - O Photon: mainly ECAL
 - O Electron: ECAL + tracker
 - O Charged hadrons: all calorimeters + tracker
 - O Neutral hadrons: all calorimeters
 - O Muons: Mainly muon chambers + tracker

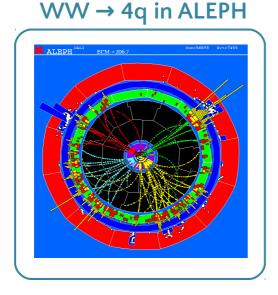




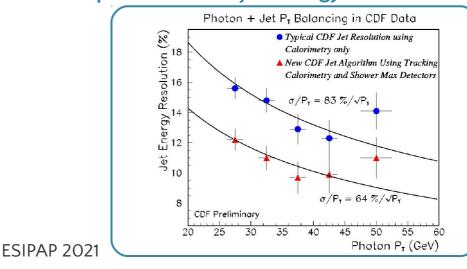
■ The idea is to **combine the information from all subdetectors**

O Can better identify objects and measure / calibrate their energy

■ Pioneered in ALEPH at LEP (90's) and used later in other detectors

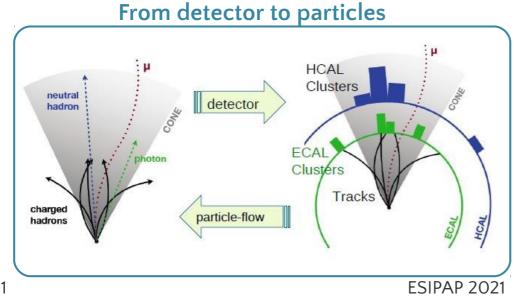


Improvement in jet energy resolution in CDF

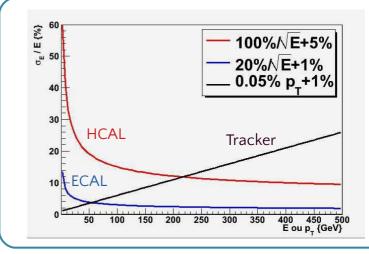


Particle Flow

- Sub-detectors are complementary, which is why combining them brings something more
- The ultimate goal is to **reconstruct each individual particle**
 - O In particular particles within a jet
 - \bigcirc Charged hadrons and low $p_{\scriptscriptstyle T}$ electrons better measured with the tracker
 - O Photons measured by the ECAL
 - O Neutral hadrons can only be measured by the HCAL
- Better identify EM and hadronic components (can apply software compensation)
- Can have a global description of collision events (pile-up, jet substructure, etc.)



Typical resolutions



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Particle Flow challenges

■ Jet energy resolution (forgetting correlations):

$$\sigma_{jet}^{2} = \sigma_{h^{+}}^{2} + \sigma_{\gamma}^{2} + \sigma_{h^{0}}^{2} + \sigma_{confusion}^{2} + \sigma_{threshold}^{2} + \sigma_{losses}^{2}$$
 Reconstruction contributions

○ <u>Threshold</u>: energy cuts applied

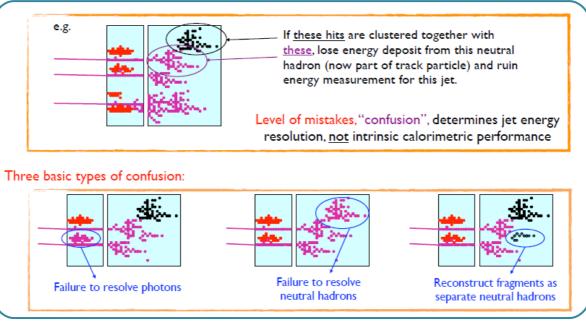
○ Losses: imperfect reconstruction

O <u>Confusion</u>: wrong identification of energy deposits (plays a major role)

Need an efficient linking procedure between sub-detectors

O Avoid **double counting** of energy

O Avoid to apply **wrong calibration** weights



Confusion effects

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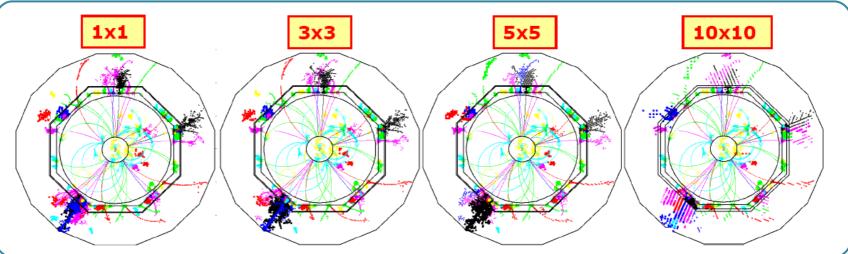
Particle Flow ingredients

- Good separation of particles
 - \bigcirc High magnetic field integral (B×R)
 - ⊖ High granularity
- **Low amount of material** before the calorimeters
 - O Light tracker, calorimeters inside the coil

Small Moliere radius (dense calorimeters)

O Minimize overlaps between showers

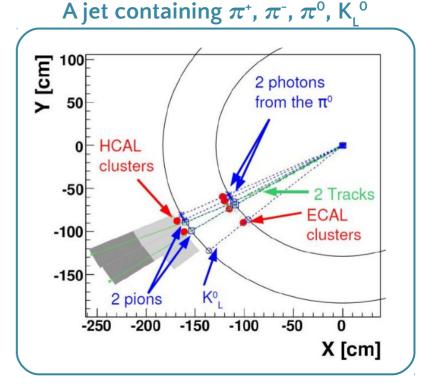
Efficient tracking

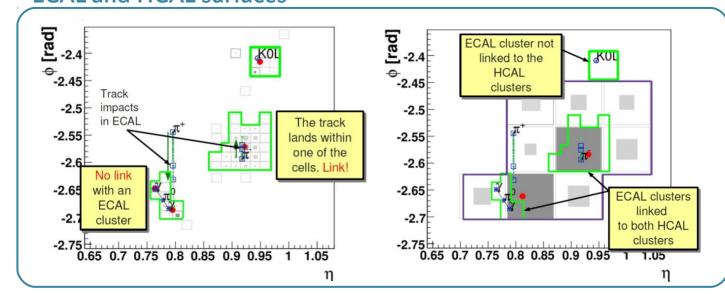


Impact of HCAL granularity

Particle Flow in CMS

- CMS not designed for Particle Flow
 - Though meets several of the criteria for a good PF
 - O Large field integral: B×R = 4.9 T.m
 - O Excellent ECAL resolution, granularity and small Moliere radius
 - O Excellent tracking





ECAL and HCAL surfaces

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Particle Flow in CMS

Particle Flow improved jet energy resolution significantly

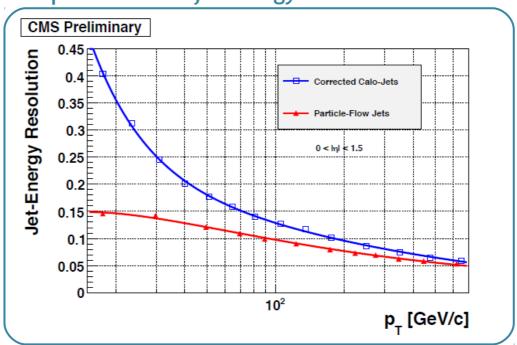
 \bigcirc In particular at low $p_{\scriptscriptstyle T}$

O Where the tracker contribution is important

But considerable challenges

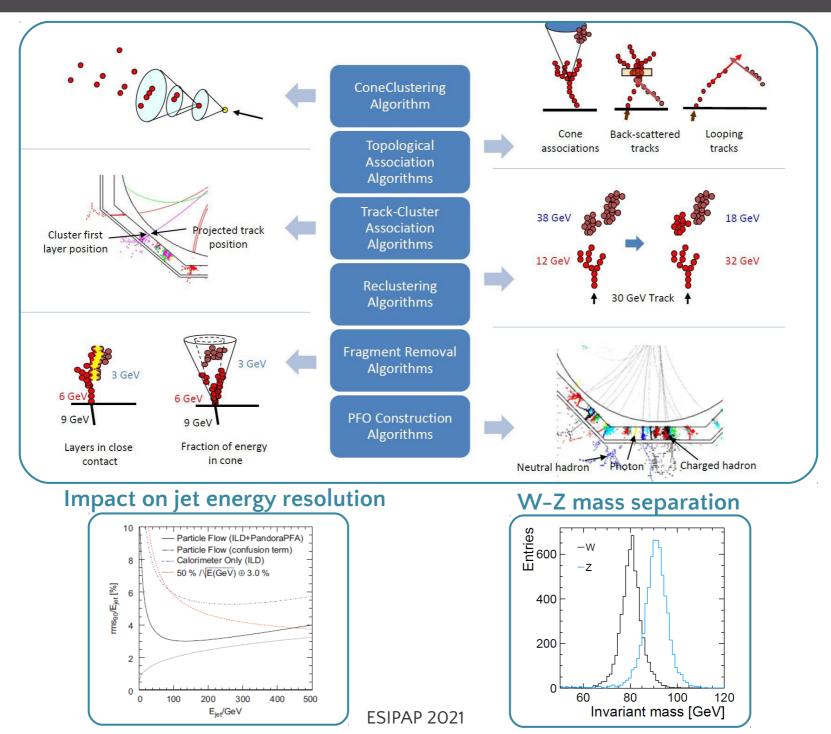
 \bigcirc Up to 2X₀ of tracker material

O Pile-up and very high density of particles



Improvement in jet energy resolution from PF

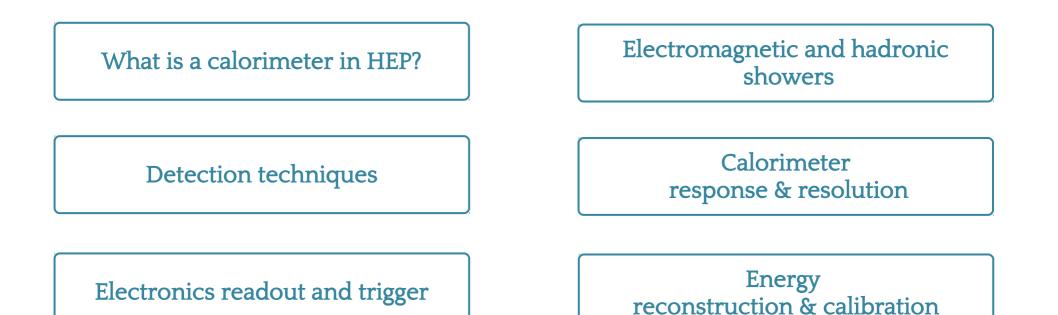
Particle Flow at the ILC



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Lecture plan



Beyond calorimetry: Particle Flow

An example: CMS HGCAL

High Luminosity LHC

Luminosity upgrade of the LHC

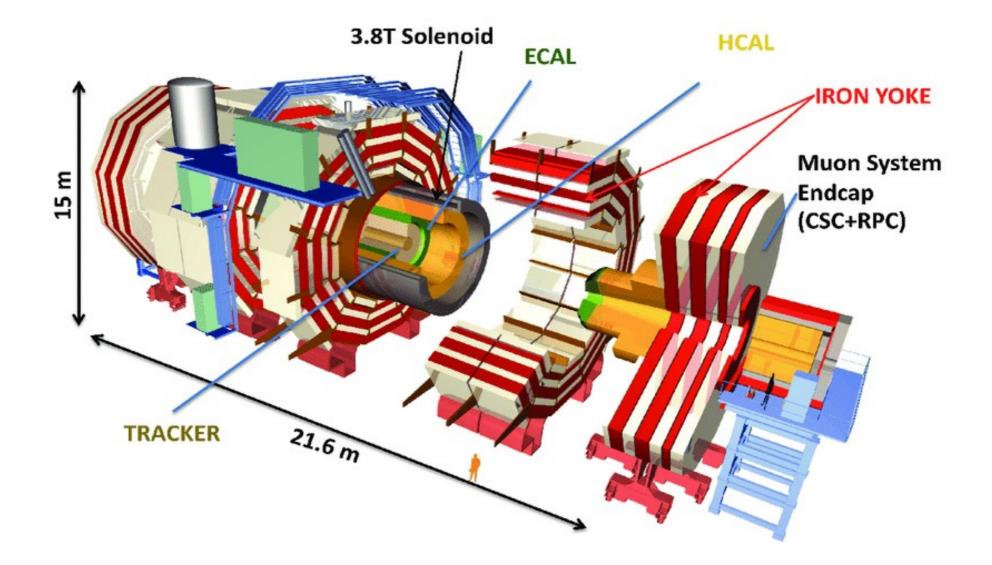
○ Starting in **2027**

O Installation in 2025-2027

■ Together with an **upgrade of the detectors**

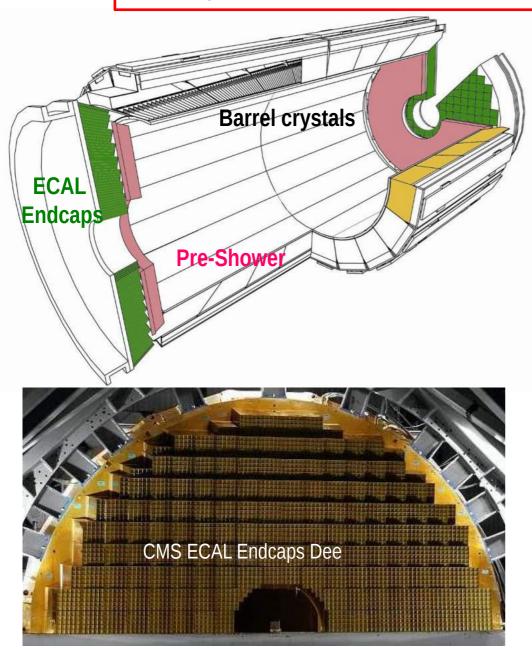


CMS

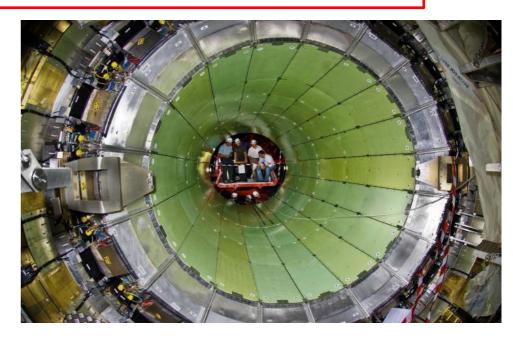


CMS ECAL

Homogenous calorimeter made from 75848 PbWO₄ scintillating crystals



01/02/21



- Barrel (|η|<1.48), ~67 t
- 61200 crystals over 36 super-modules

- Endcaps (1.48<|η|<3), ~23 t
- 14648 crystals over 4 Dees (2 per endcap)
- Preceded by Pb/Si Pre-Shower

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Material: inorganic scintillator

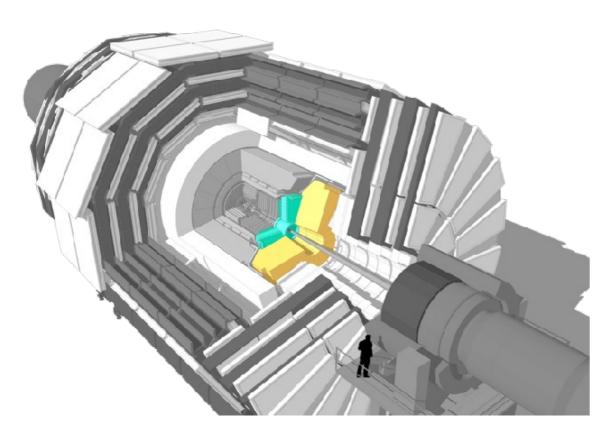
CMS crystals: PbWO₄

Excellent energy resolution $X_0 = 0.89$ cm \rightarrow compact calorimeter (28 cm for 26 X₀) $R_M = 2.2$ cm \rightarrow compact shower development Fast light emission (80% in less than 15 ns) Radiation hard (10⁵Gy) But

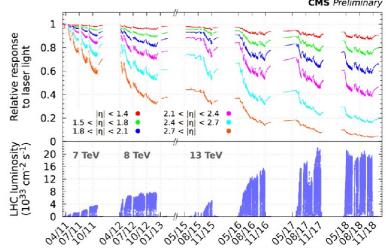
Low light yield (150 γ/MeV) Response varies with dose Response temperature dependance

CMS endcap calorimeters for HL-LHC

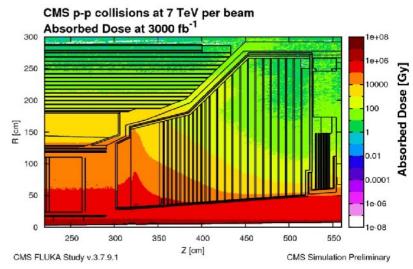
- CMS endcap calorimeters will need to be replaced
 - O ECAL crystals and HCAL scintillators suffer from **irreparable radiation damage** after 500 fb⁻¹



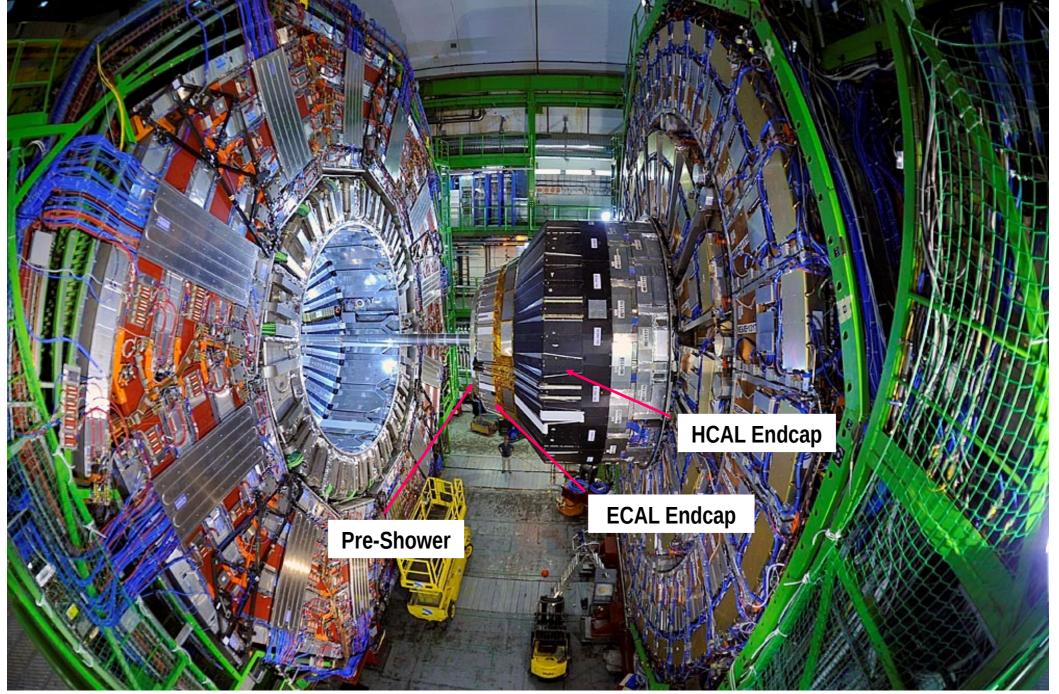
Loss of transparency of ECAL crystals



Absorbed dose at the end of HL-LHC



CMS endcap

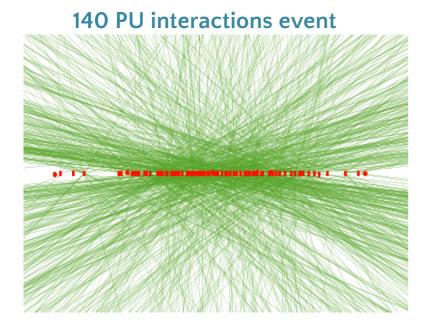


Challenges: pile-up (PU)

HL-LHC nominal parameters

O140-200 quasi simultaneous interaction each bunch crossing

- Needs detectors that can survive high radiation but can also disentangle all these simultaneous interactions
 - O Needs high granularity
 - O And precise timing information



CMS Simulation <u> = 200(su) Simulated Vertices 3D Reconstructed Vertices + 0.6 4D Reconstructed Vertices 4D Tracks 0.4 0.2 0 -0.2 -0.4-5 5 10 -100 z (cm)

Space-time view of interaction vertices

HGCAL overview

Radiation hard detector measuring energy and time **Silicon sensors** in ECAL and highest-radiation region of HCAL ○ Scintillating tiles + SiPM readout in low-radiation region ■ 620 m² of silicon sensors, **6M channels** ■ 400 m² of scintillator, 240k channels silicon sutron moderator CE-E ECAL (CE-E): <u>HCAL (CE-H)</u>: **28 layers**, 25 X_0 , 1.3 λ **22 layers**, 8.5 λ Steel, Cu absorbers Pb, Cu, CuW absorbers

Layer structure

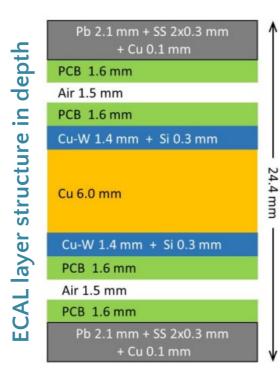
Silicon

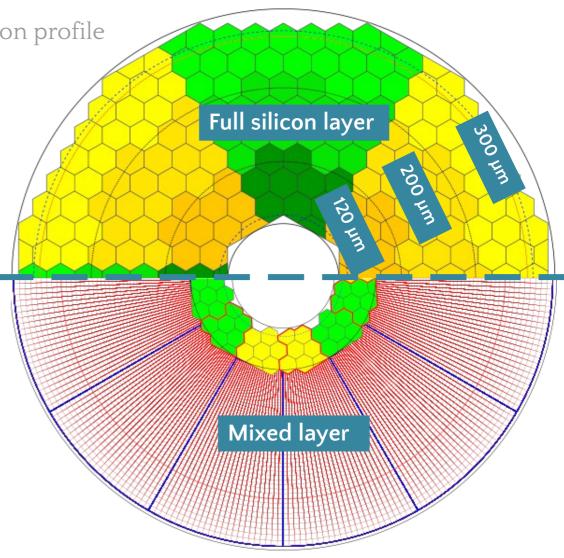
$\bigcirc 8$ hexagonal wafers

O Three thicknesses following radiation profile

- 120, 200 & 300μm
- Scintillator

 \bigcirc Tile size varies with eta (1.5°)

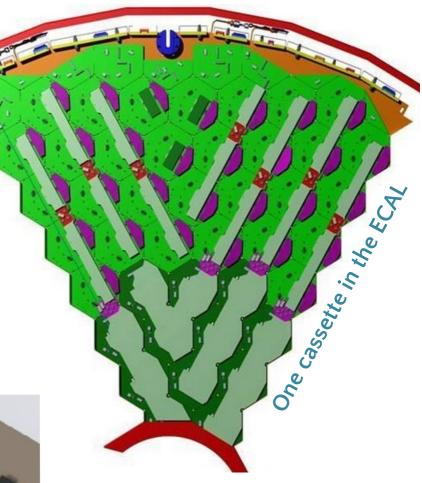




Layer components

- Detector is organized into cassettes made of a cooling plate with modules mounted on it
- Frontend electronics located on the modules
- Readout and control through a system of engine/wagon motherboards





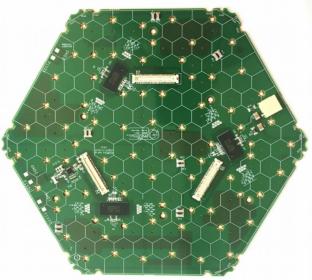
Silicon modules

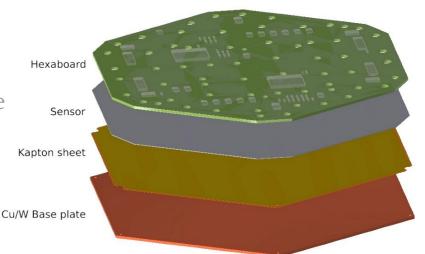
- Stack of sensor, readout PCB (Printed Circuit Board) and baseplate
 - Several HGCROC (HGC readout chip) ASICs on the board
- "High density" and "Low density" modules

○ 0.5 cm² cell area in the highest radiation region

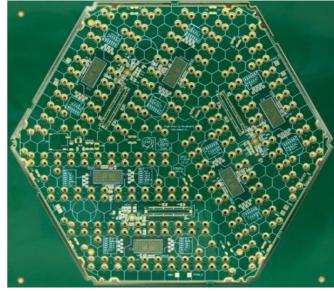
○1.1 cm² otherwise

Low density module 192 channels 3 readout chips





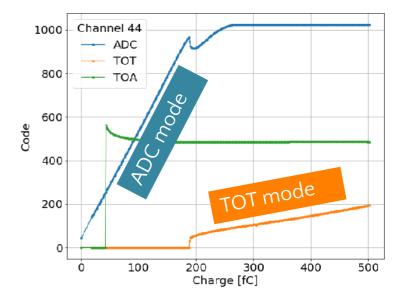
High density module 432 channels 6 readout chips

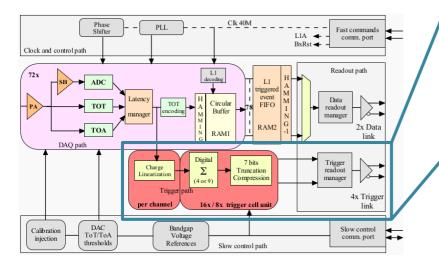


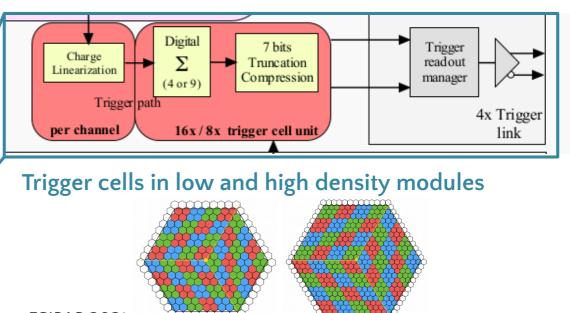
The readout chip (HGCROC)

ASIC amplifying, shaping, digitizing collected charges

- O Two modes: "**ADC**" (low charges) &"**TOT**" (high charges)
- O Data buffered before read out
- Trigger path
 - O ADC and TOT linearization
 - ⊖ Sums of channels → **Trigger cells**
 - O Energy compression







HGCAL prototype in beam test experiments

CE-E

- O Double sided mini-cassettes
- O Lead absorber
- ○26 X0, 1.4 L



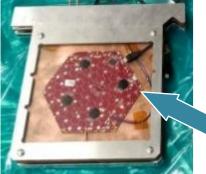
CE-H

Daisy arrangement of the modules
Steel absorber
3.4 L

CALICE AHCAL

○ Scintillator + SiPM○ Steel absorber○ 4.4 L

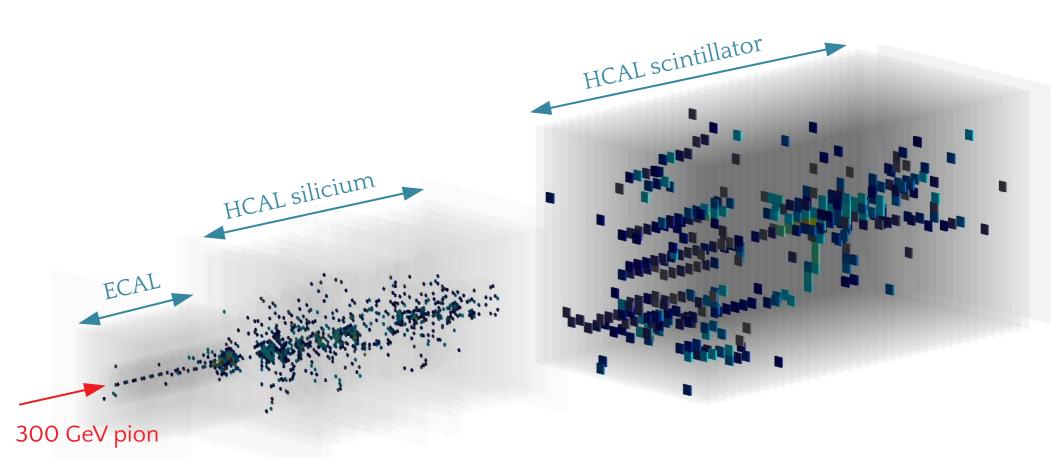






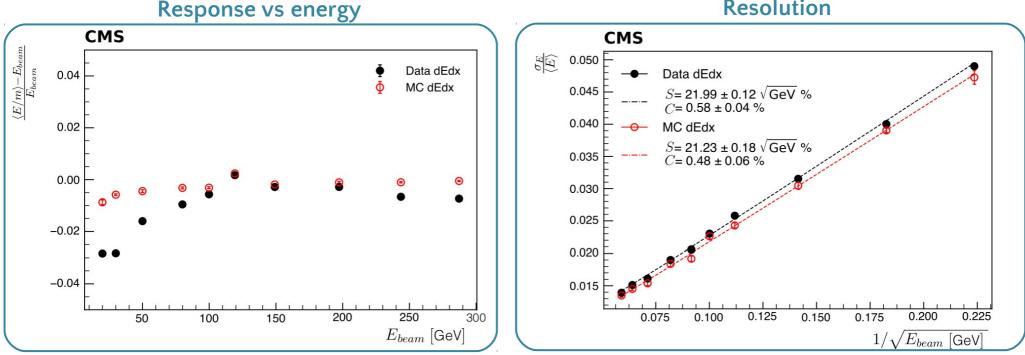


Display of pion shower



Electromagnetic showers reconstruction

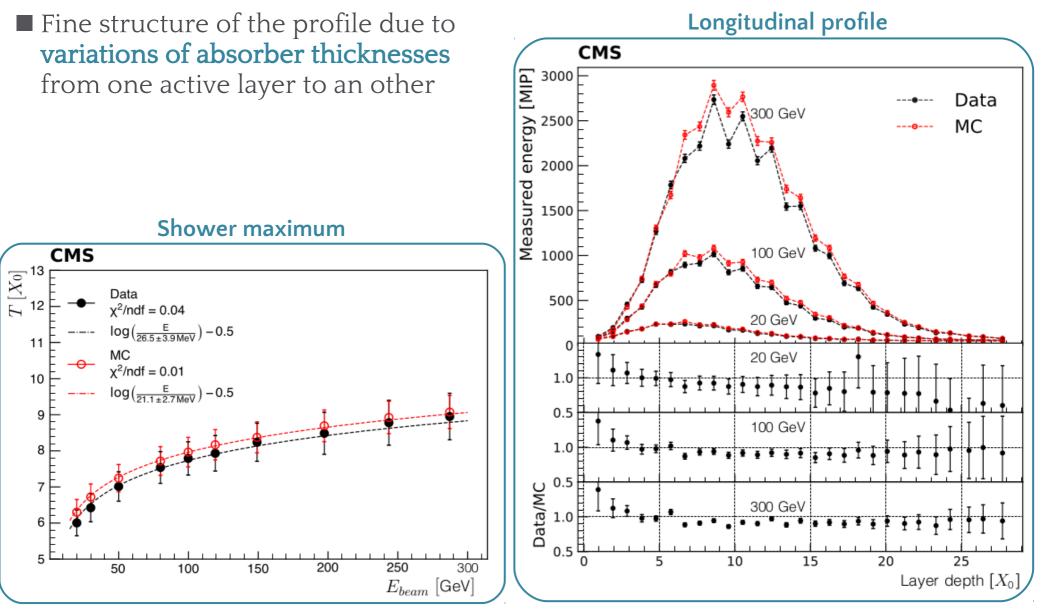
- Response to electromagnetic showers
 - O Good linearity
 - O Small drop of response for lower energy showers
- Resolution: stochastic term around 22%
- Simulation slightly optimistic



Resolution

Electromagnetic shower profile

Logarithmic dependency of the shower depth



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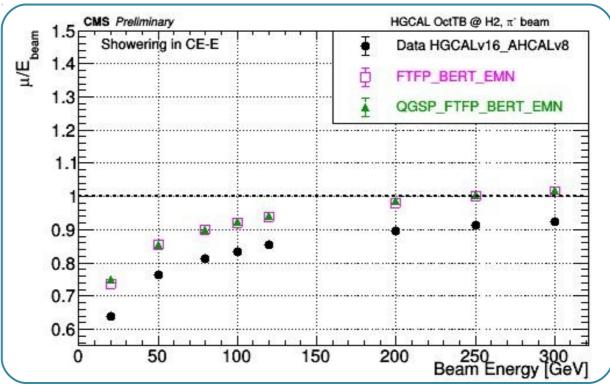
Hadronic shower reconstruction

First a calibration based on electrons in the ECAL part and pions in the HCAL part

As expected, non linear response for hadronic showers

O This is a **non compensating calorimeter**

O Need more sophisticated calibration techniques (e.g. software compensation)



Pion shower non linearity