

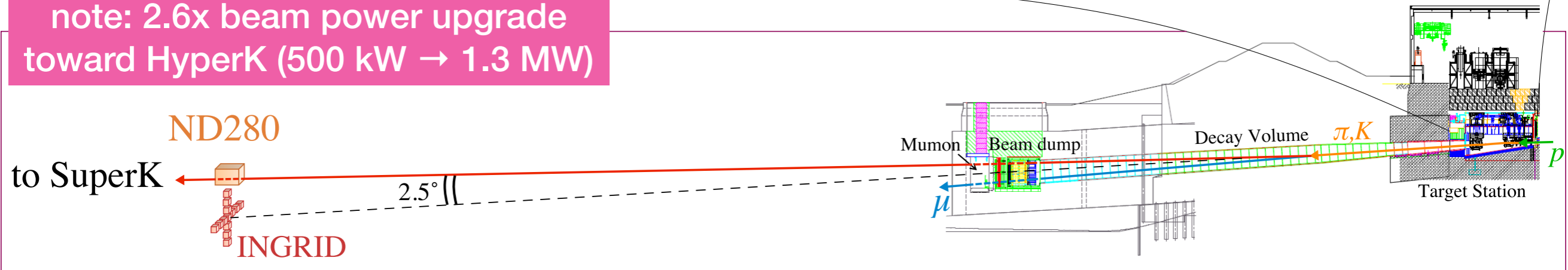
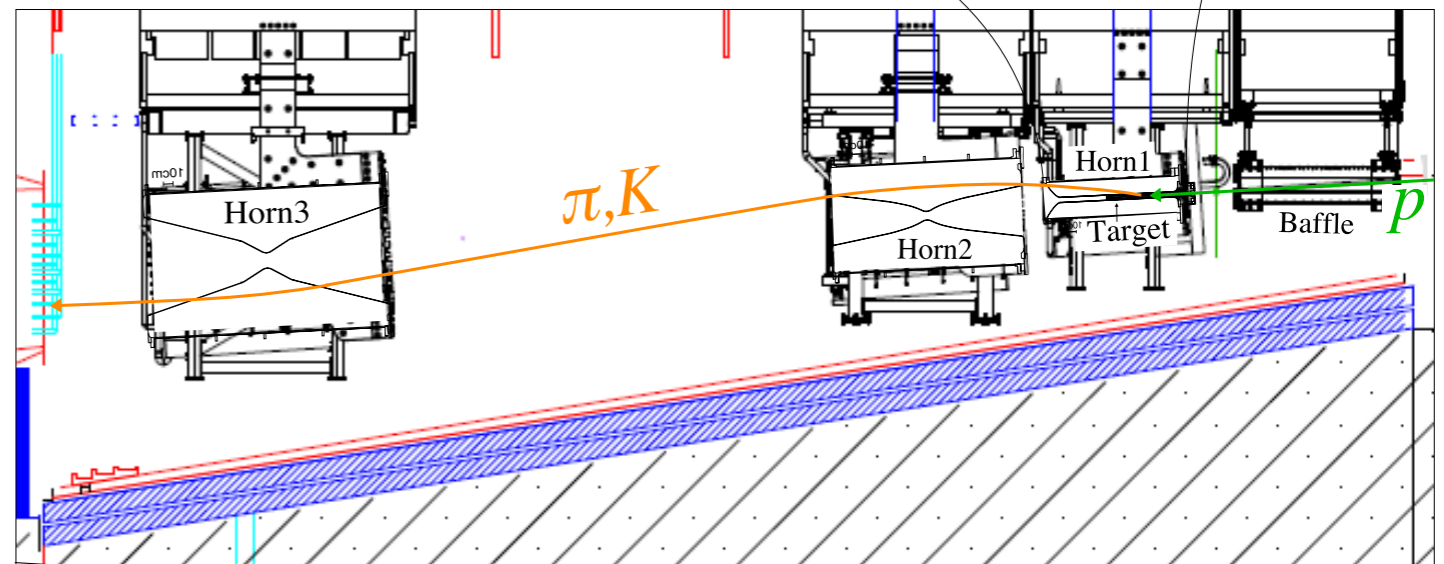
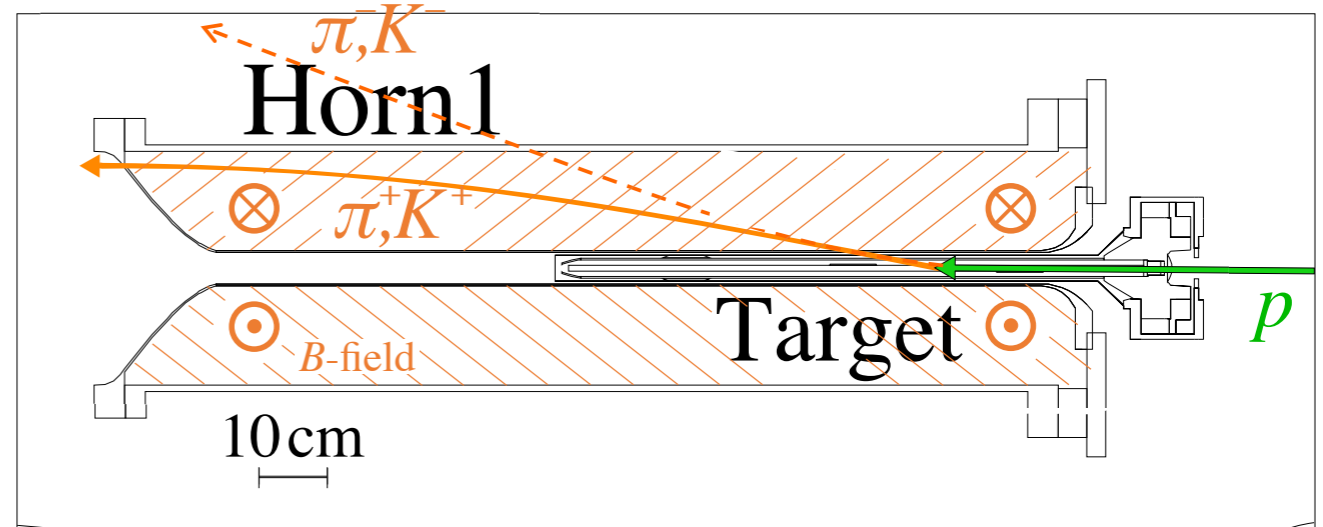
Low energy hadron production motivation for T2K/HK beam

NA61 low-E workshop, 2020-12-10
Lukas Berns (T2K/HK)

J-PARC ν -beamline

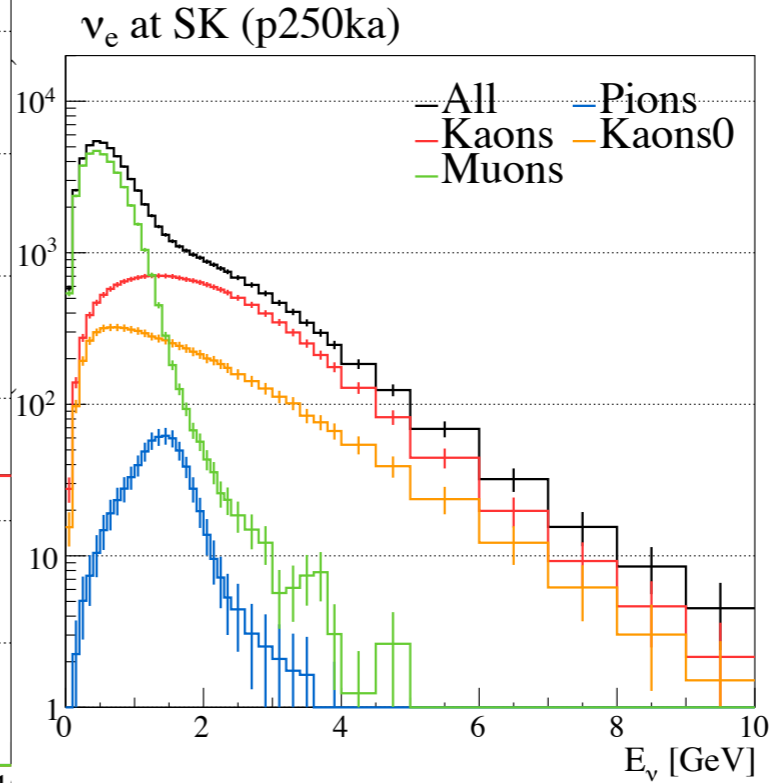
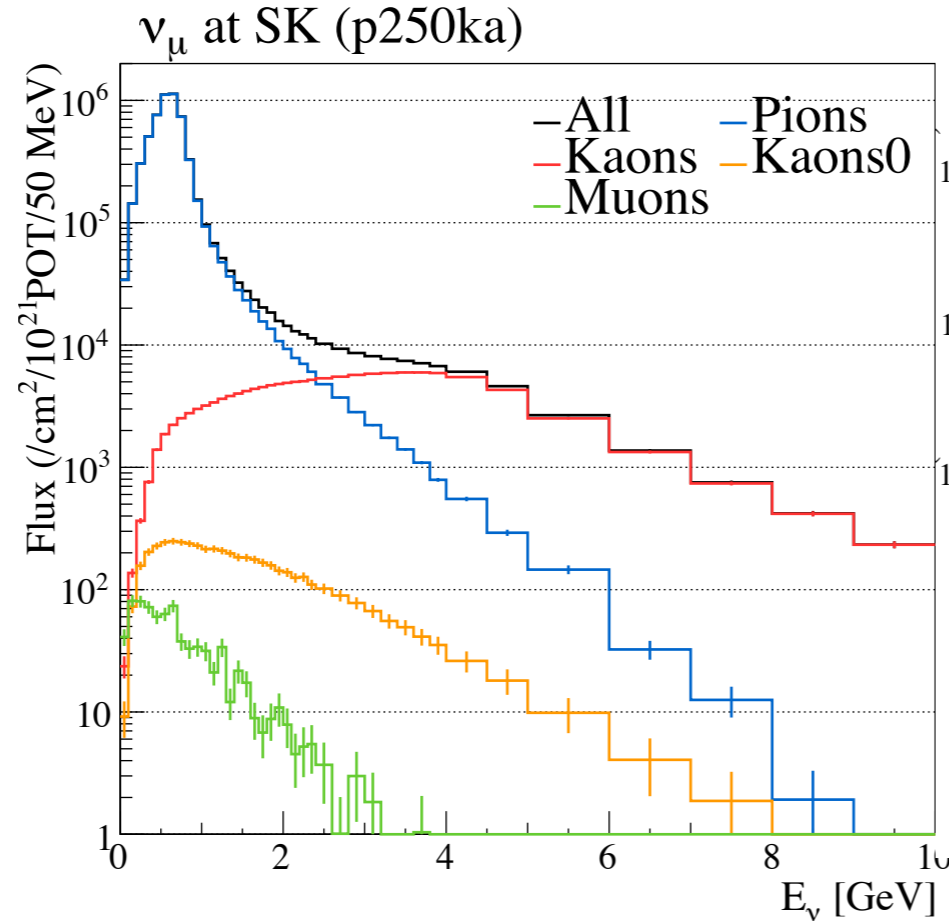
- 30 GeV protons produce π, K in 90 cm graphite target
- Three magnetic horns selectively focus π^+, K^+ or π^-, K^- to produce ν_μ or $\bar{\nu}_\mu$ beam (decay in-flight).
- Muon monitors and on-axis ν detector (INGRID) monitor beam stability and direction.

note: 2.6x beam power upgrade toward HyperK (500 kW \rightarrow 1.3 MW)



Right-sign ν_μ flux with 650 MeV peak from π^\pm decay. Above 3 GeV K^\pm decay.

Above each peak contribution from mis-focused and forward particles.

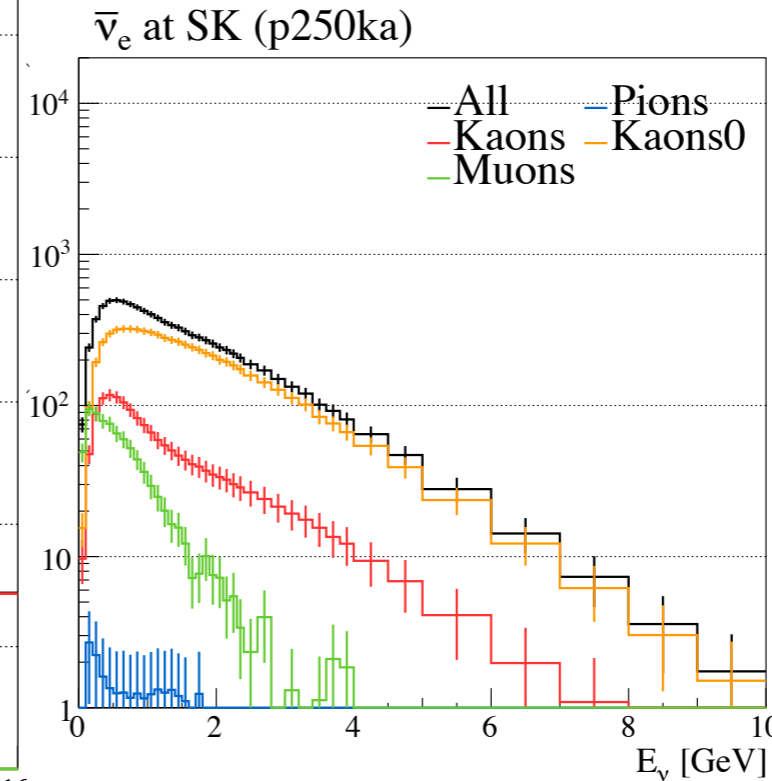
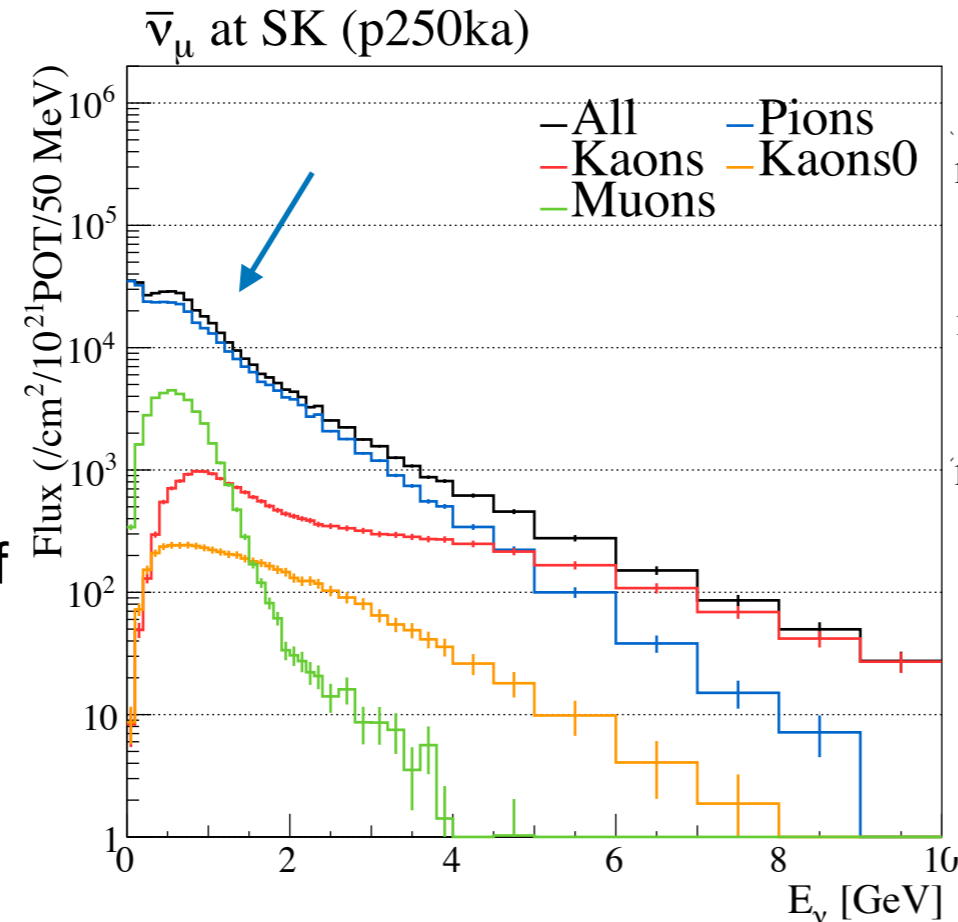


Small ν_e contamination from μ -decay in tunnel and K -decay.

For RHC the contribution from K_L^0 increases.

Flux composition is $\nu_\mu > \bar{\nu}_\mu > \nu_e > \bar{\nu}_e$ (in neutrino-mode) roughly differing by factor 10 each.

Wrong sign component from **rescattered pions (low-E)** and forward mesons from target downstream face of target that “miss” the horns (high-E).



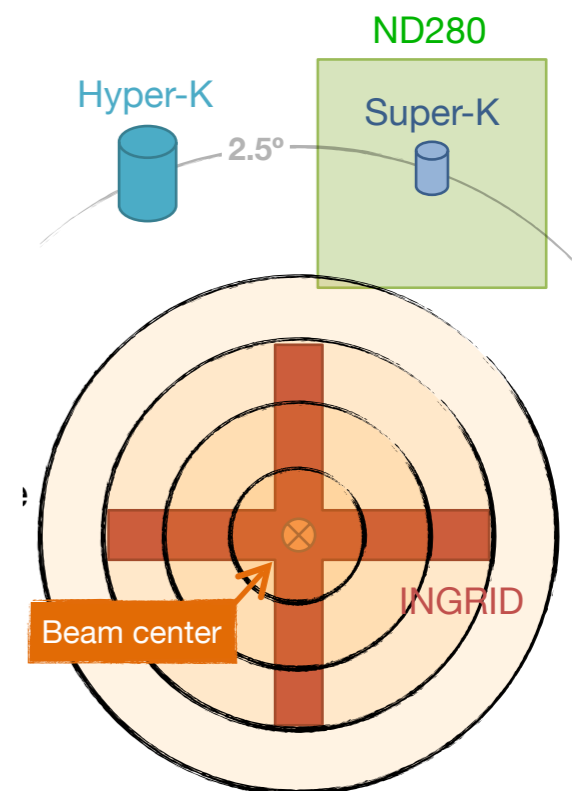
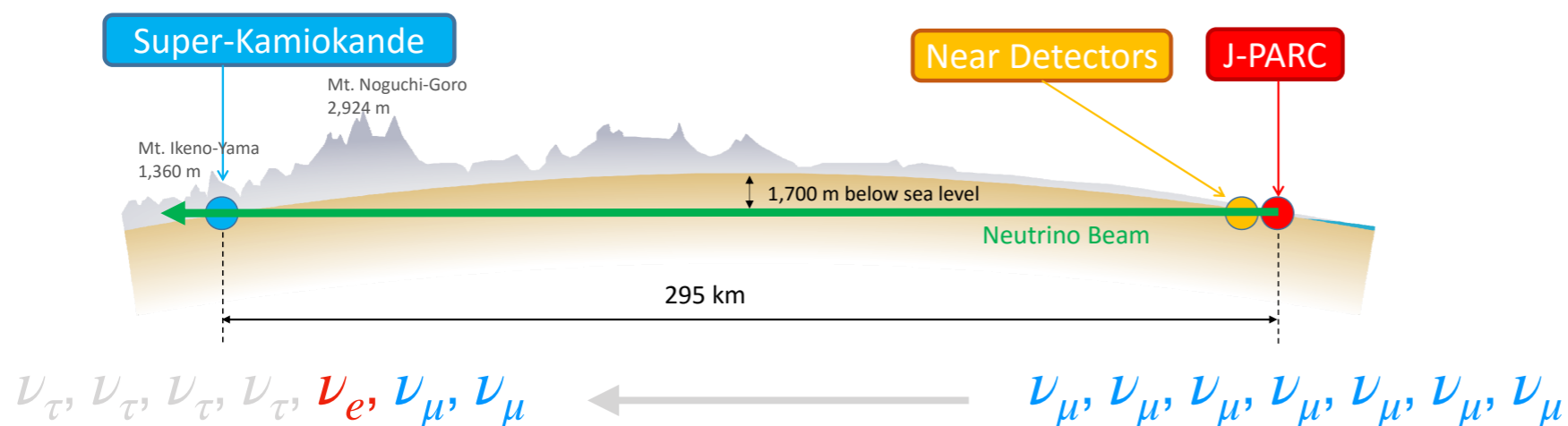
Wrong sign ν_e component entirely from 3-body decay of kaons, mostly K_L^0 .

For RHC contribution from K^+ increases.

Oscillation analysis

- ND280 constrains flux*xsec:
 flux: 5% → 3% (uncertainty on #evts)
 flux*xsec: 11% → 2%
If flux can be understood better than 3%,
 will be able to improve the xsec constraint through flux?
- Constraint is propagated to SK flux by constraining the SK/ND280 flux covariance matrix.

- Currently only μ -like samples fit at ND280, and e-like flux constraint relies on flux covariance matrix. Similarly when the near-detector has troubles running, the flux uncertainty relies on run-by-run correlations estimated with MC and INGRID monitoring data.
- Constraint on osc. params is stats-dominated, and the post-fit flux systematic plays a small role in the systematics.
- For HK expect to have larger contribution from non-hadronic systematics due to direction differences between HK and SK/ND280.

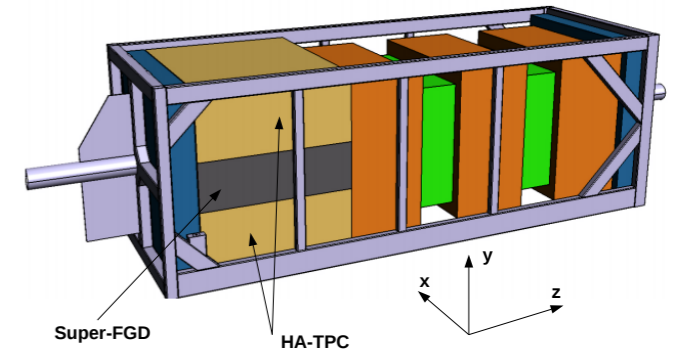


Near detectors

- Various near detectors perform wide variety of xsec measurements
- Flux uncertainty is dominant systematic on the xsec normalization
- Combined analyses at various off-axis angles allow studying energy dependence etc.

2.5° off-axis

ND280
magnetized
near detector
to be upgraded soon

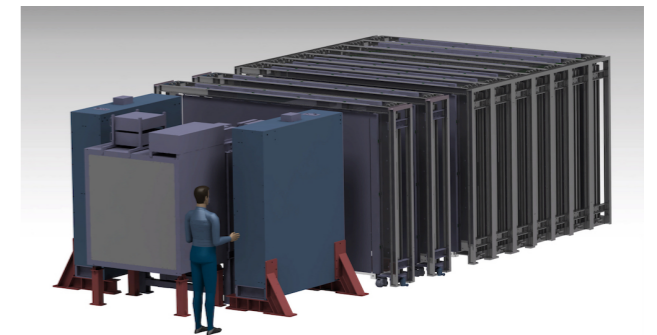


same off-axis angle as SK/HK

1.5° off-axis

WAGASCI + BabyMIND + NINJA
(not part of T2K)

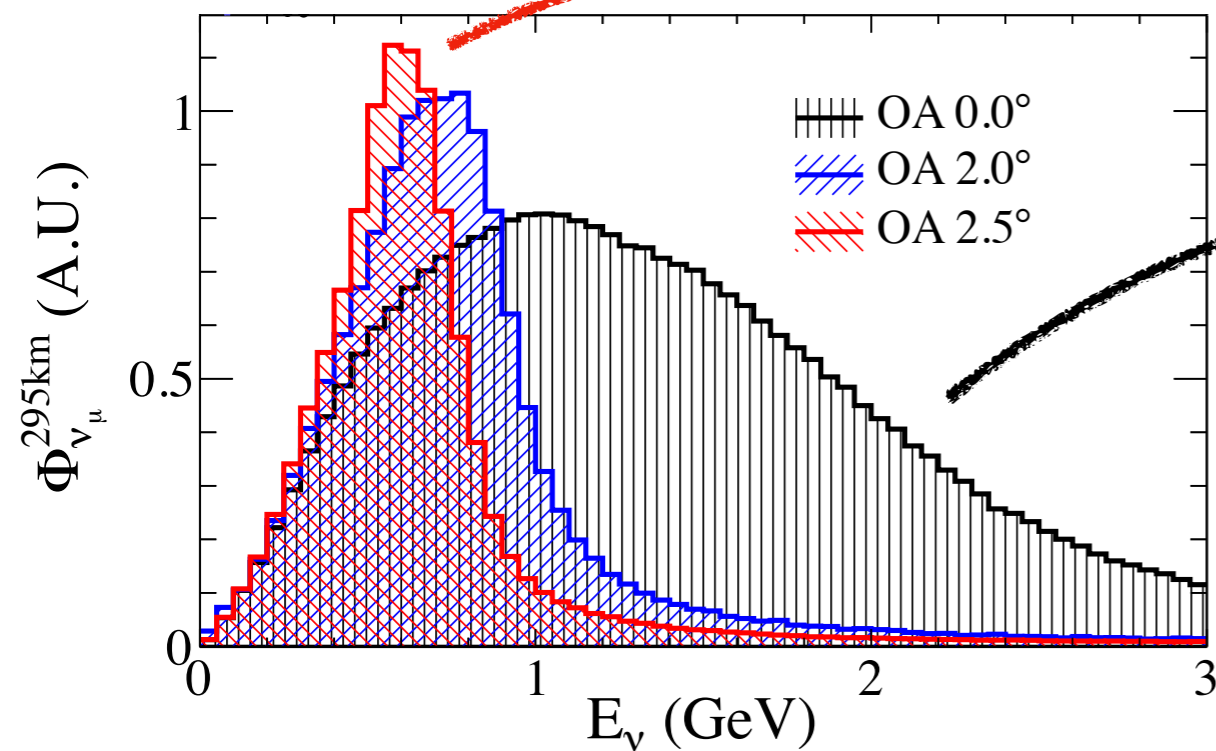
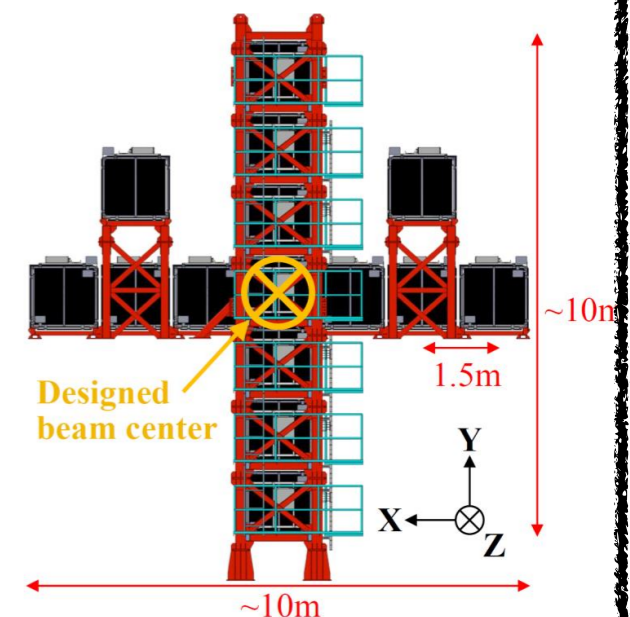
Water + scintillators
Compact magnetized
iron μ range detector
Emulsion detector



on-axis

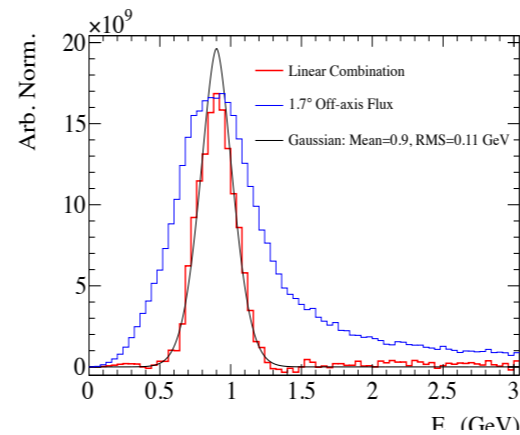
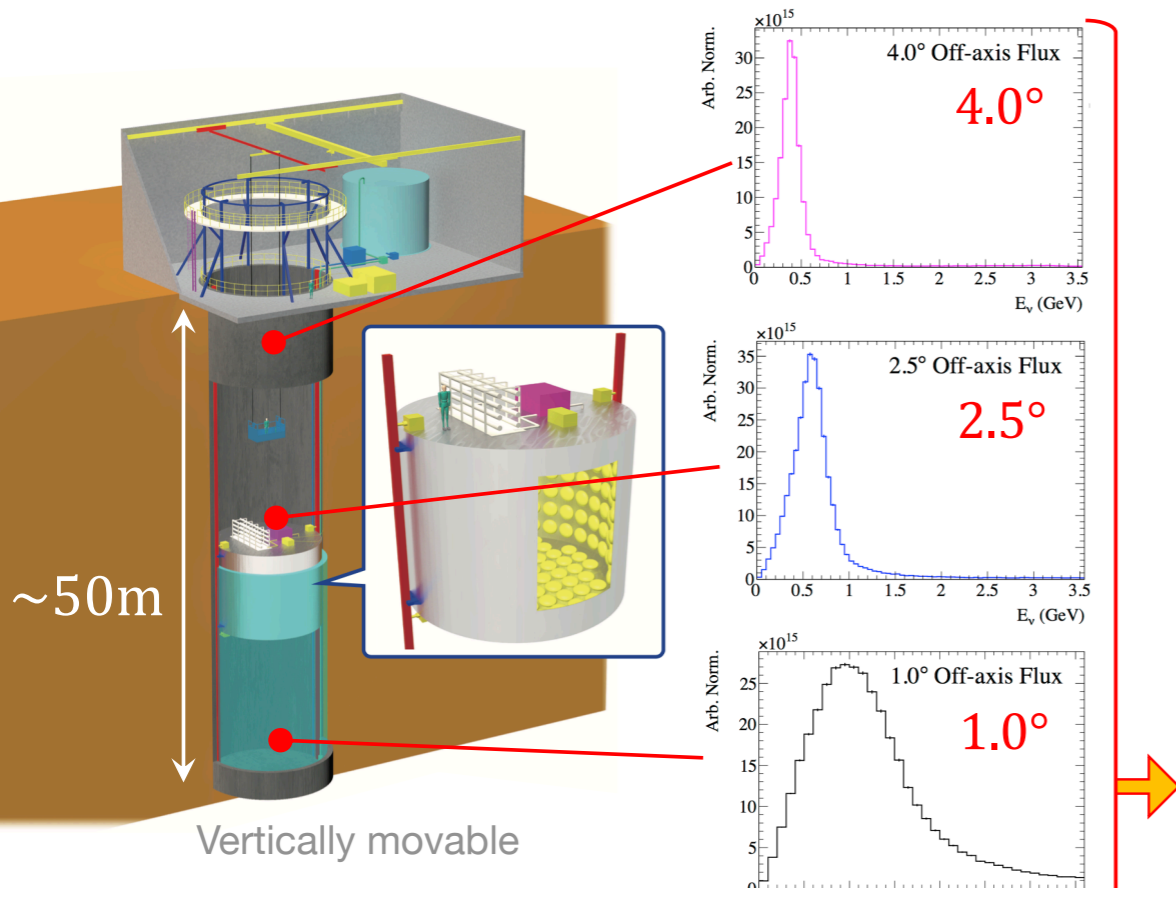
INGRID
monitor ν beam
rate + direction

14 modules of unmagnetized
iron-scintillator
sandwich detectors

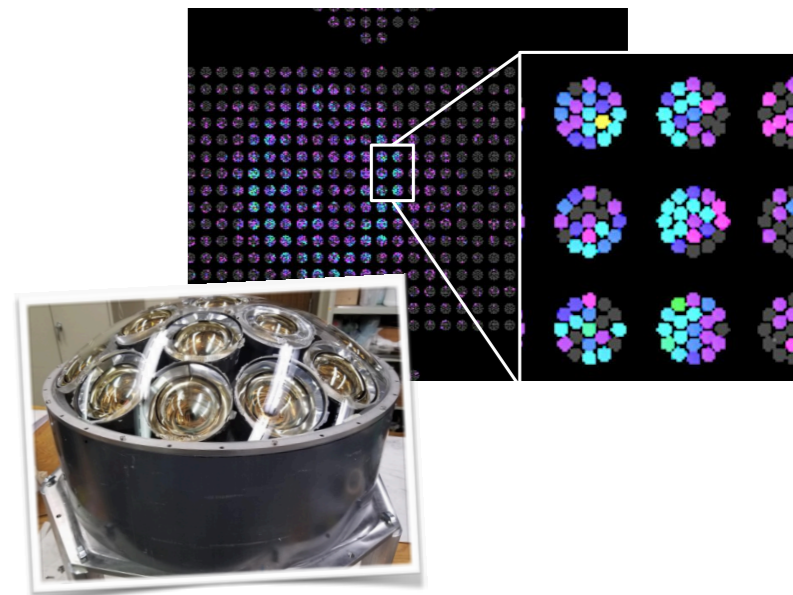
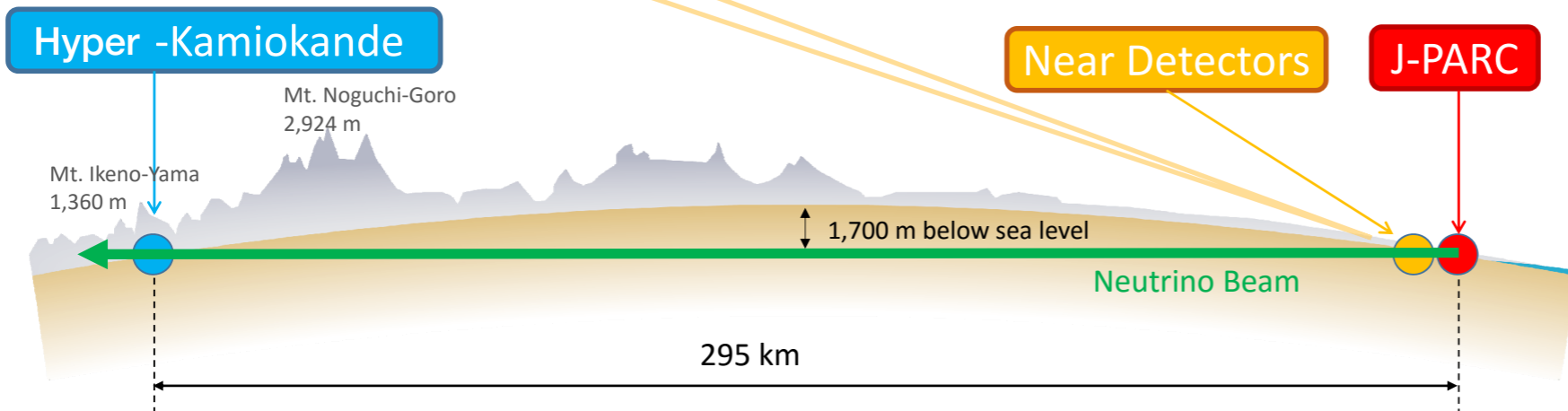


Intermediate Water Cherenkov Detector

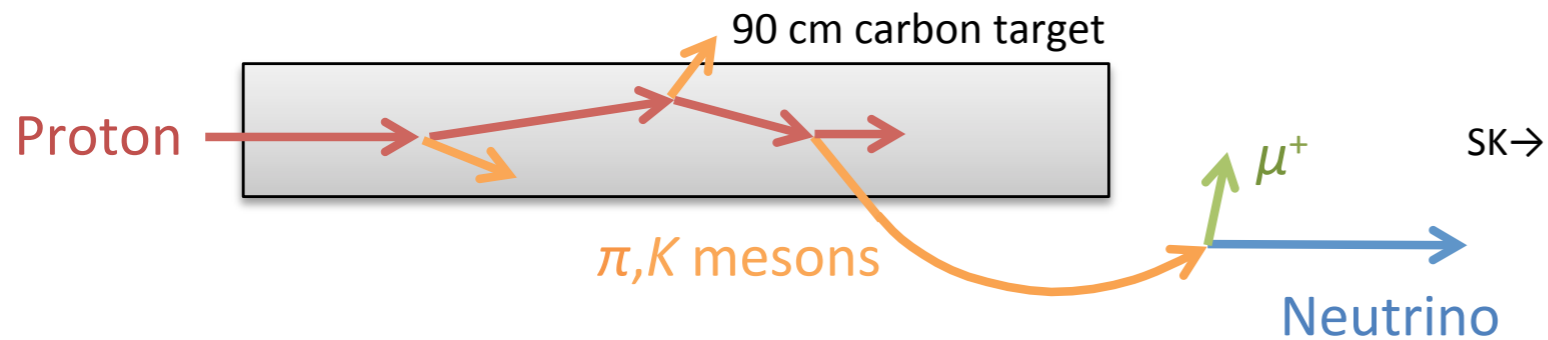
- New 1 kton scale water-Cherenkov detector at ~1 km baseline planned with goals:
 - ν -interaction with combined measurement of fluxes at various off-axis angles
 - ν_e cross section
 - Constrain NC and ν_e background
 - Neutron multiplicity with Gd loading
- Multi-PMTs for higher granularity + directionality



need precise understanding of flux



$\nu_\tau, \nu_\tau, \nu_\tau, \nu_\tau, \nu_e, \nu_\mu, \nu_\mu$ ← $\nu_\mu, \nu_\mu, \nu_\mu, \nu_\mu, \nu_\mu, \nu_\mu, \nu_\mu$

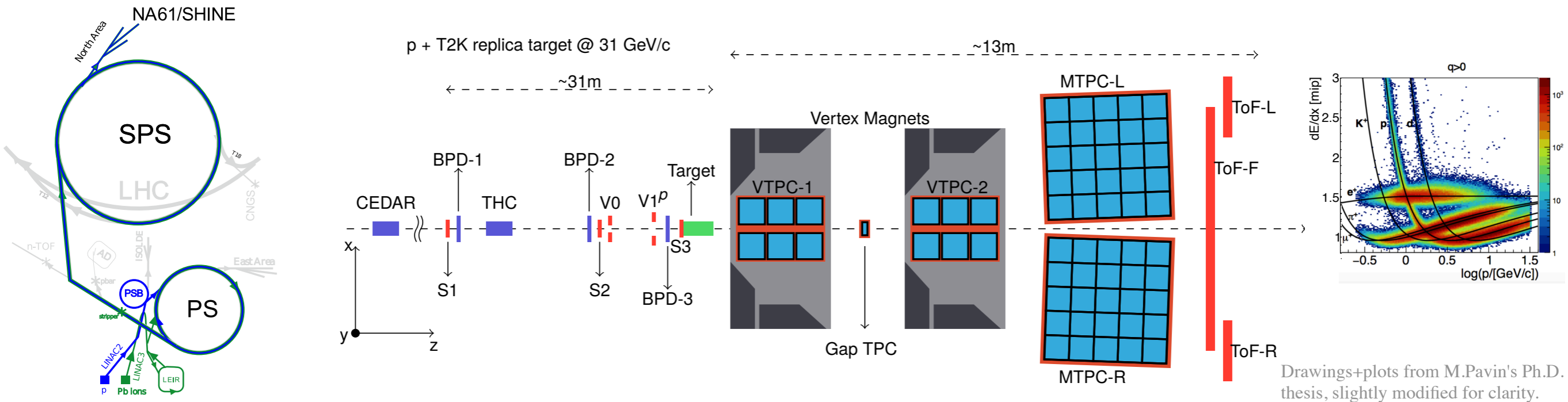


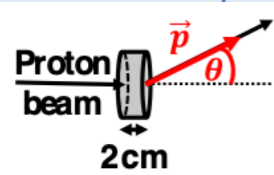
Flux Prediction

- Interactions of protons **inside target** simulated with **FLUKA** based on proton beam profile measured with upstream beam monitors.
- Horn **focusing** and **out-of-target** interactions (Al in horns and Fe in walls) of outgoing mesons using **Geant3 + GCalor**.
- Afterwards go through interaction chain and **apply weights** to **tune** output of MC generators to external hadron production data (mostly NA61/SHINE).
- **Covariance matrix** is used to constrain SuperK flux using near detector measurements.

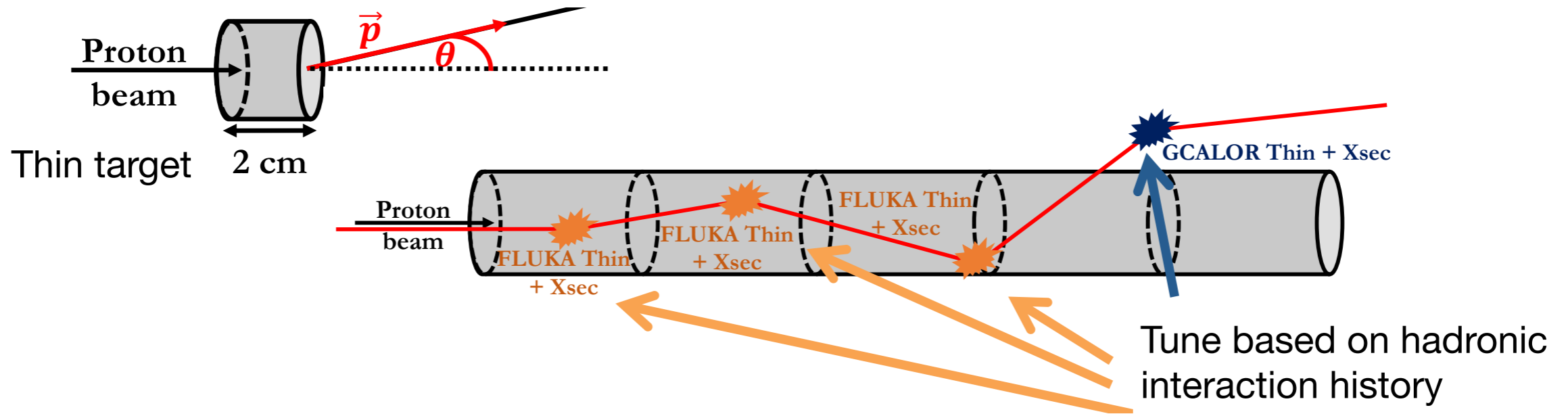
NA61 measurements for T2K

- Hadron production experiment, momentum measurement with TPCs in superconducting magnets, PID with dE/dx (Bethe-Bloch) and time of flight.



Beam	Target	Year	Stat (10^6)	Outgoing PID	Usage at T2K
 protons at 31 GeV/c	Thin (2cm)	2007	0.7	$\pi^\pm, K^\pm, K_S^0, \Lambda$	past
		2009	5.4	$\pi^\pm, K^\pm, p, K_S^0, \Lambda$	in use
	T2K replica (90cm)	2007	0.2	π^\pm	
		2009	2.8	π^\pm	next T2K results
		2010	10.	π^\pm, K^\pm, p	in development

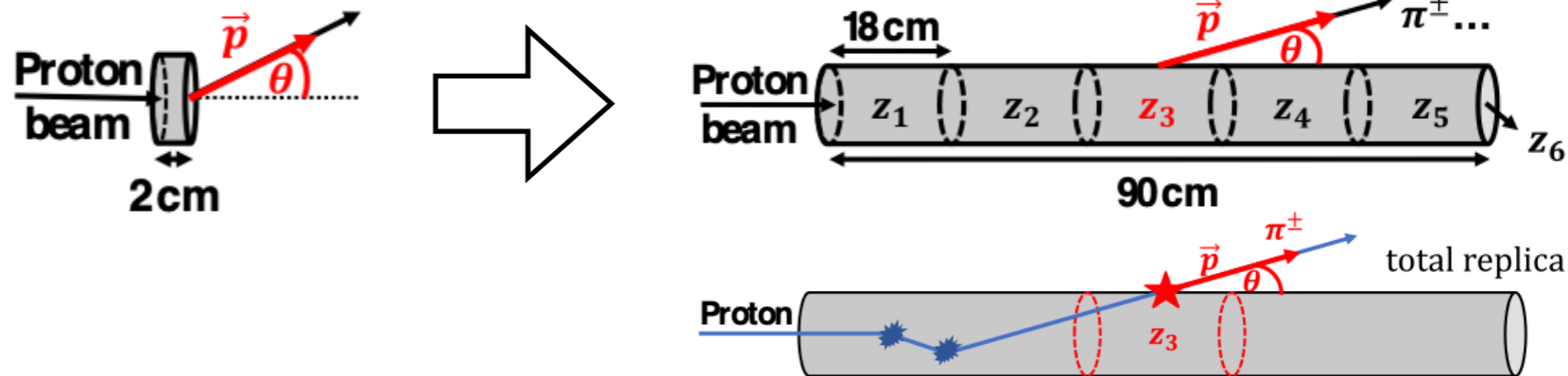
Thin target tuning



	Interaction length tune $\sigma_{\text{prod}}(p+C)$ to NA61 measurement	Multiplicity Mostly 30 GeV p+C data by NA61
At interaction 	"Vertex" weight $\sigma_{\text{DATA}} / \sigma_{\text{MC}}$	$\left(\frac{d^2n}{dp d\theta} \right)_{\text{DATA}} / \left(\frac{d^2n}{dp d\theta} \right)_{\text{MC}}$ <p>p, θ: outgoing particle kinematics</p>
For distance L traversed in matter 	"Attenuation" weight $e^{-(\sigma_{\text{DATA}} - \sigma_{\text{MC}}) \rho L}$	N.A.

Replica tuning

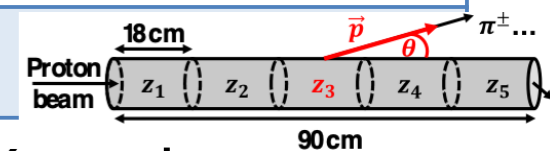
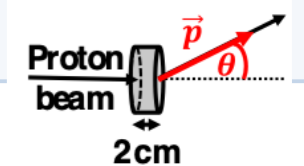
- NA61 took data with **full-sized replica** of T2K target, binned by (z, p, θ)
- Ignore interactions inside the target and apply **single DATA/MC weight** based on exiting particle.
- **Out-of-target** interaction and outgoing particles **not covered** by replica data are tuned with thin target data.



Reduced uncertainty from

- no interaction length uncertainty
- single weight per exiting particle

Beam	Target	Year	Stat (10^6)	Outgoing PID	Usage at T2K
protons at 31 GeV/c	Thin (2cm)	2007	0.7	$\pi^\pm, K^\pm, K^0_S, \Lambda$	past
		2009	5.4	$\pi^\pm, K^\pm, p, K^0_S, \Lambda$	in use
	T2K replica (90cm)	2007	0.2	π^\pm	
		2009	2.8	π^\pm	latest T2K results
		2010	10.	π^\pm, K^\pm, p	in development

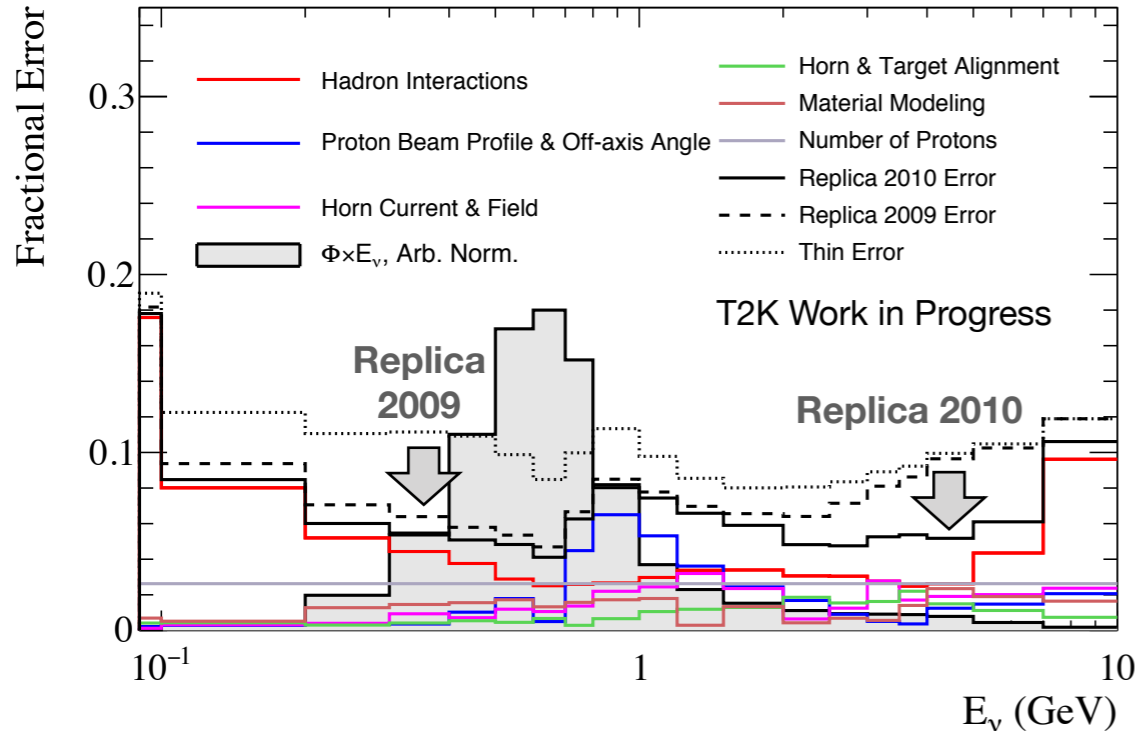


Replica tuning with 2010 data (π^\pm, K^\pm, p)

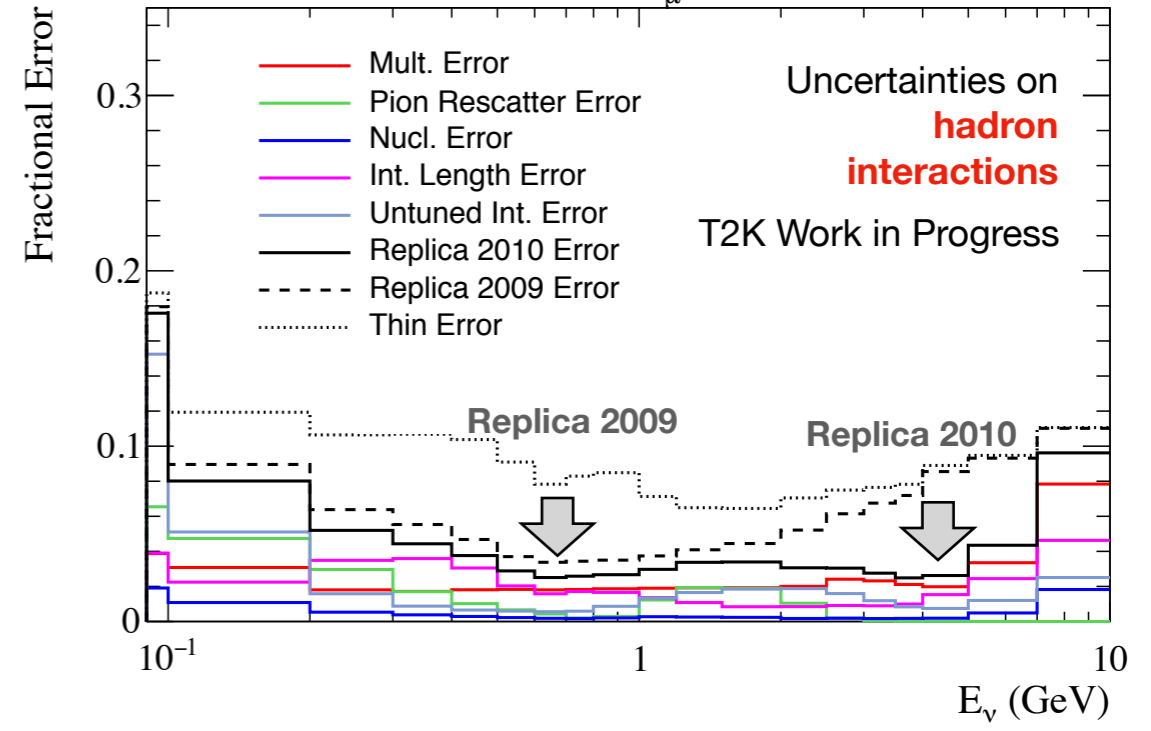
Adds K^\pm and proton yields + increased stats.
Achieve ~4% hadron interaction uncertainty over wide energy range.

- checked additional systematics → seems robust
- checking consistency with thin tuning

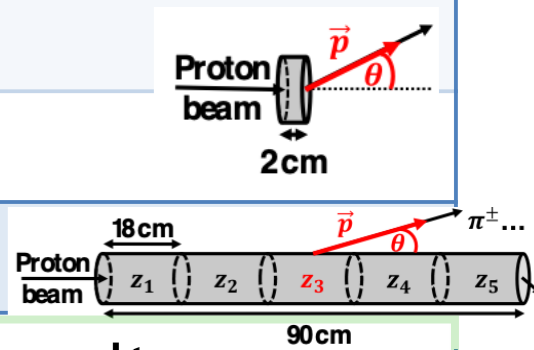
SK: Neutrino Mode, ν_μ



SK: Positive Focussing (ν) Mode, ν_μ



Beam	Target	Year	Stat (10^6)	Outgoing PID	Usage at T2K
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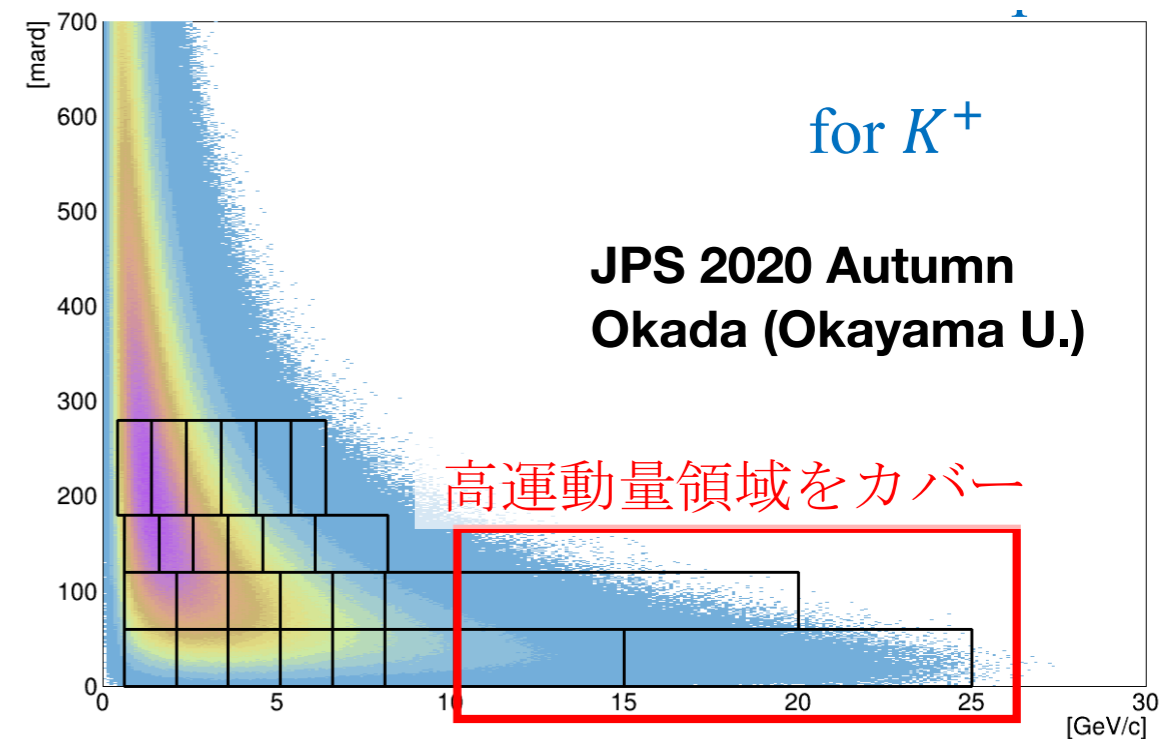


Future hadron production

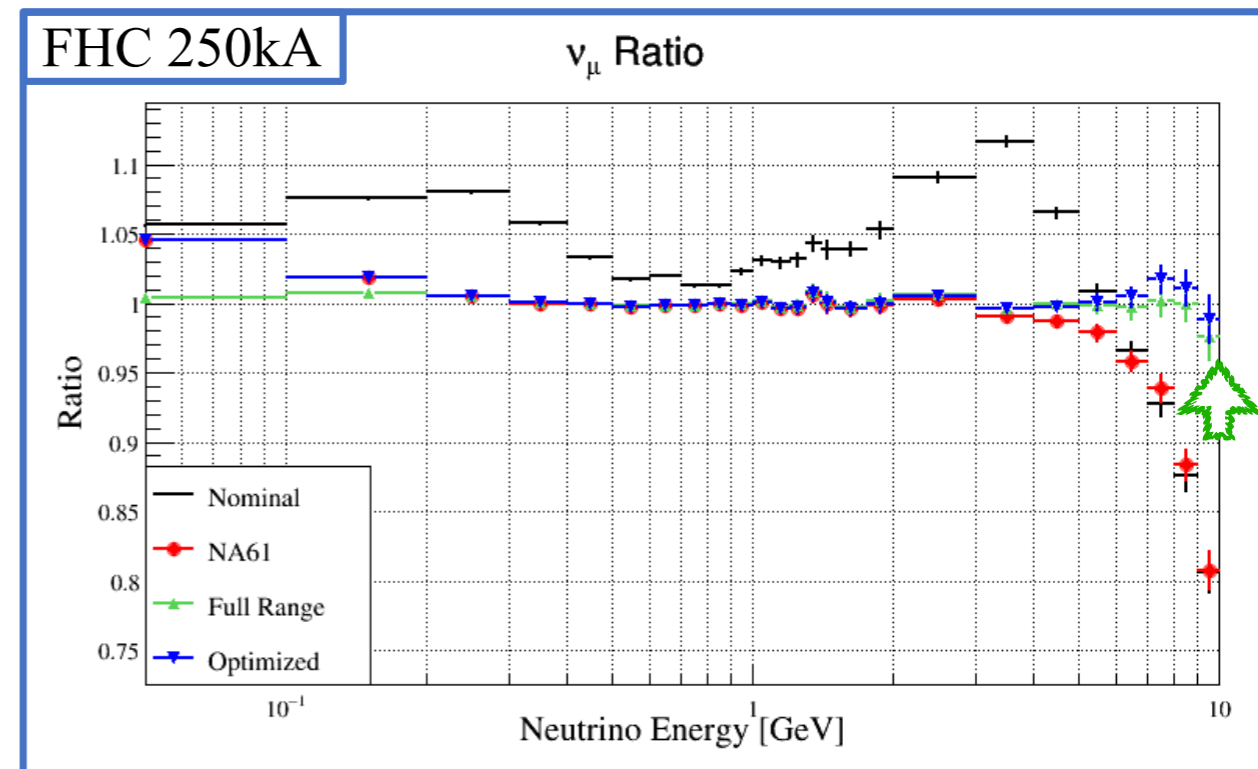
- NA61
 - many upgrades e.g. 10x higher trigger rate
 - more replica target data (SPSC approved) with extended Kaon coverage?
 - p + N,O for constraining atmospheric flux production

- if tertiary beamline can be built, $\pi \rightarrow \pi$ scattering at few GeV/c

today's topic



- EMPHATIC
 - focus on forward region (quasi-elastic)
 - subtraction of quasi-elastic (QE) xsec important for estimating the "production" cross section as $\text{prod} = \text{total} - \text{elastic} - \text{QE}$
 - plan to measure π, K, p scattering on various materials starting from few GeV



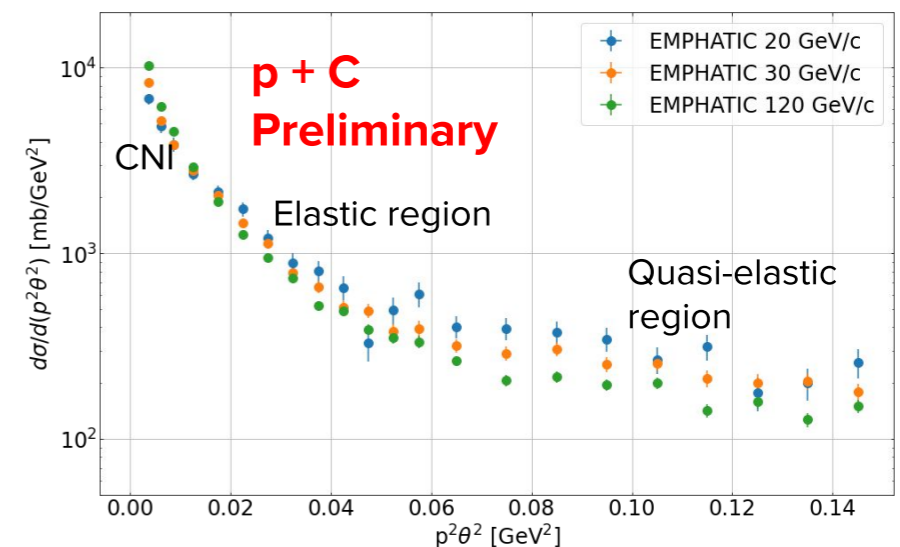
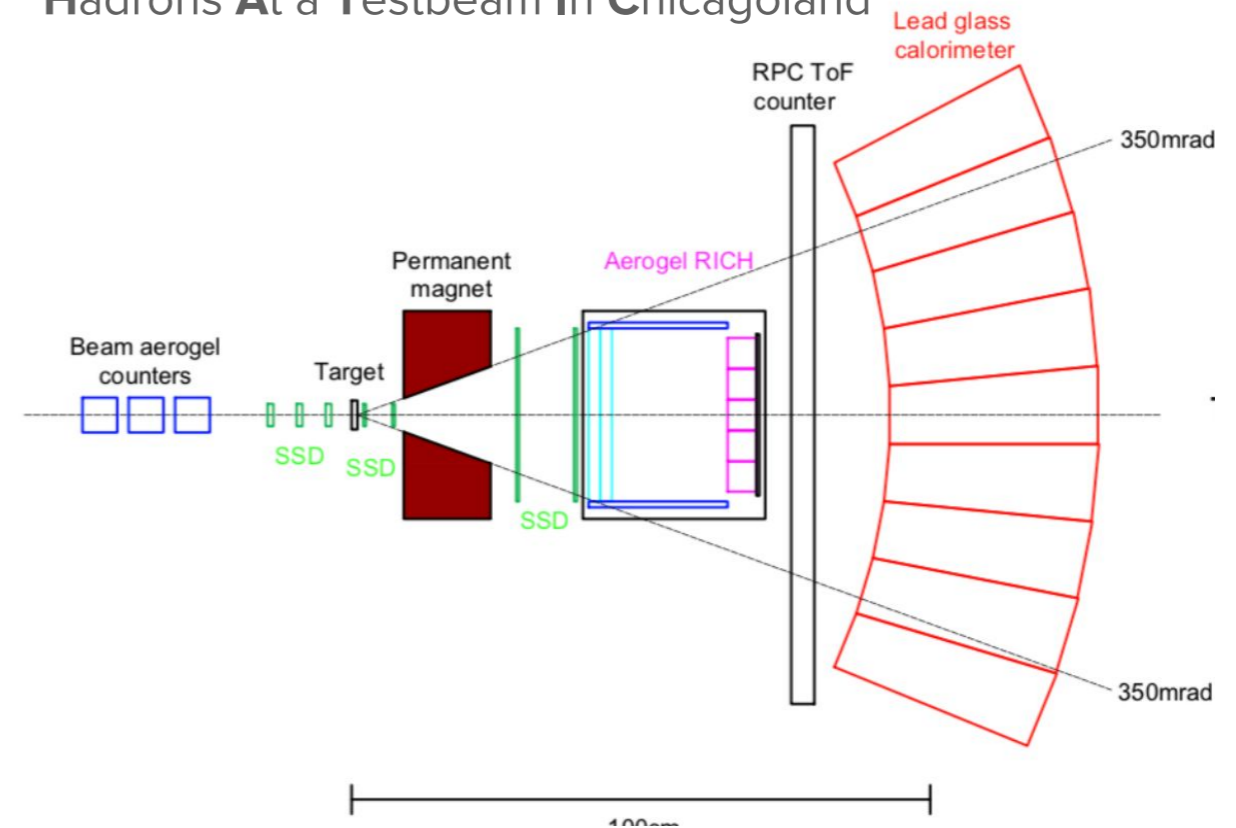
Future hadron production

- NA61
 - many upgrades e.g. 10x higher trigger rate
 - more replica target data (SPSC approved) with extended Kaon coverage?
 - $p + N, O$ for constraining atmospheric flux production
 - if tertiary beamline can be built, $\pi \rightarrow \pi$ scattering at few GeV/c

- EMPHATIC

- focus on forward region (quasi-elastic)
- subtraction of quasi-elastic (QE) xsec important for estimating the “production” cross section as $\text{prod} = \text{total} - \text{elastic} - \text{QE}$
- plan to measure π, K, p scattering on various materials at 2–15 GeV

Experiment to Measure the Production of Hadrons At a Testbeam In Chicagoland



Interactions not covered by replica data

Interaction in C	“Thin tuning”	Uncertainty
Proton prod xsec	NA61 2009 for $p_{in} > 20 \text{ GeV}/c$	QE xsec (Belletini model)
Meson prod xsec	FLUKA is treated as nominal, and GCALOR outside target is tuned to FLUKA xsec	QE xsec (Belletini model)
$p+C \rightarrow$ meson multiplicity	NA61 2007+2009 with Feynman scaling for $p_{in} < 31 \text{ GeV}/c$, BMPT extrapolation outside coverage	NA61 uncertainty Use other data with smaller p_{in} instead of Feynman scaling BMPT on/off
$p+C \rightarrow p$ multiplicity	NA61 2009 with outside region set by baryon number constraint on leading baryon. Only primary interaction is tuned.	NA61 uncertainty Secondary interaction tuning on/off (using Feynman scaling) NA49 $p+C \rightarrow n$ and NA61 $p+C \rightarrow \Lambda$ etc.
meson \rightarrow meson multiplicity	—	HARP tuning ($\pi \rightarrow \pi$) on/off. For others and outside coverage, a 50% correlated and 50% uncorrelated uncertainty in xF-pT space.
meson elastic+QE	—	multiplicity treated in cell above, xsec not treated
proton elastic+QE	—	—

relevant for this talk

Effect of additional tuning

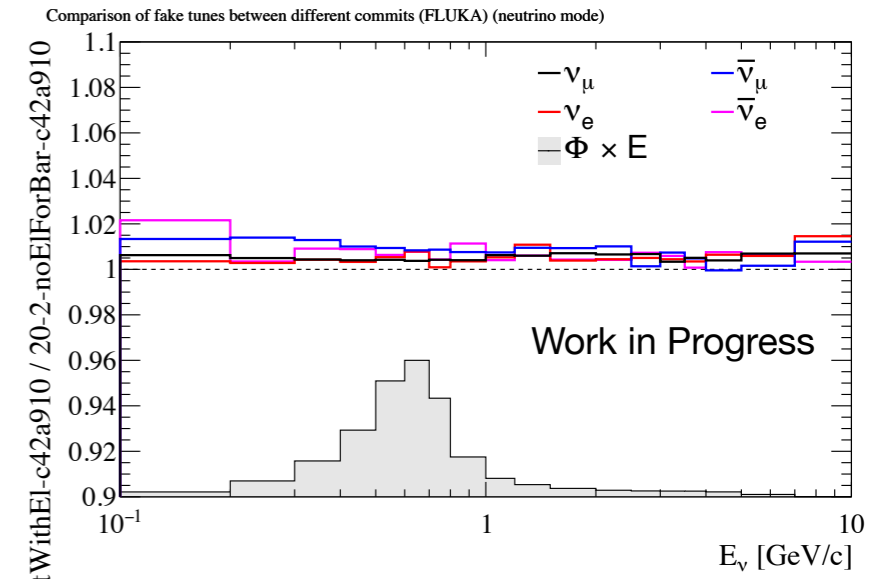
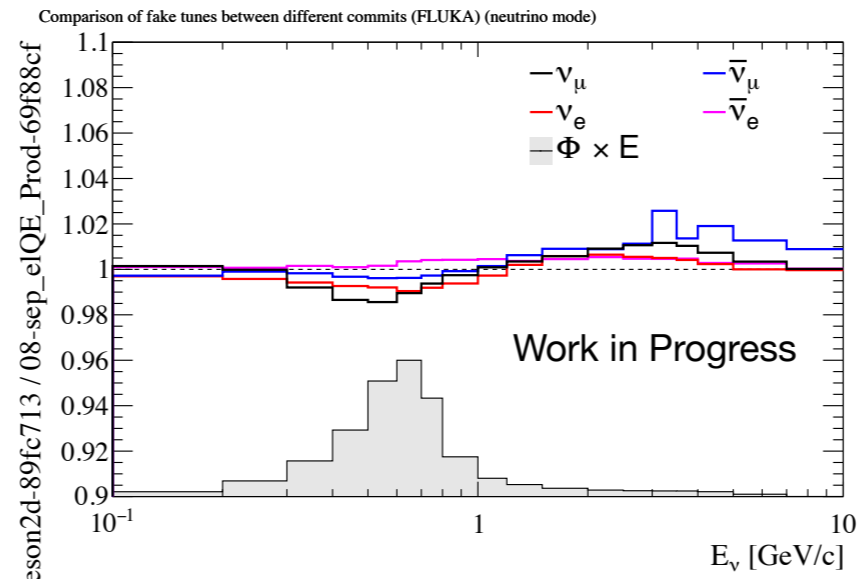
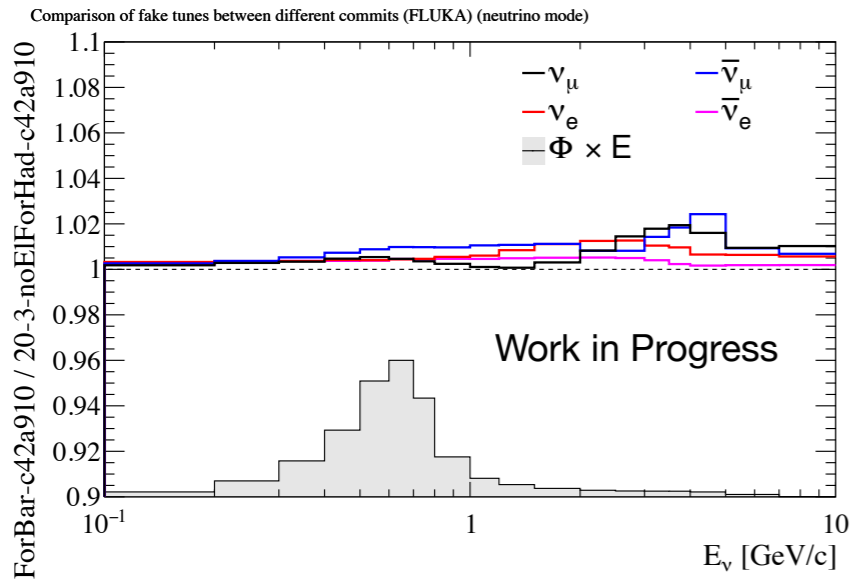
Here tuning from FLUKA to Geant4 FTFP_BERT
 — keep in mind that the physics are somewhat similar —

el+QE tuning of mesons

2d multiplicity tuning of mesons

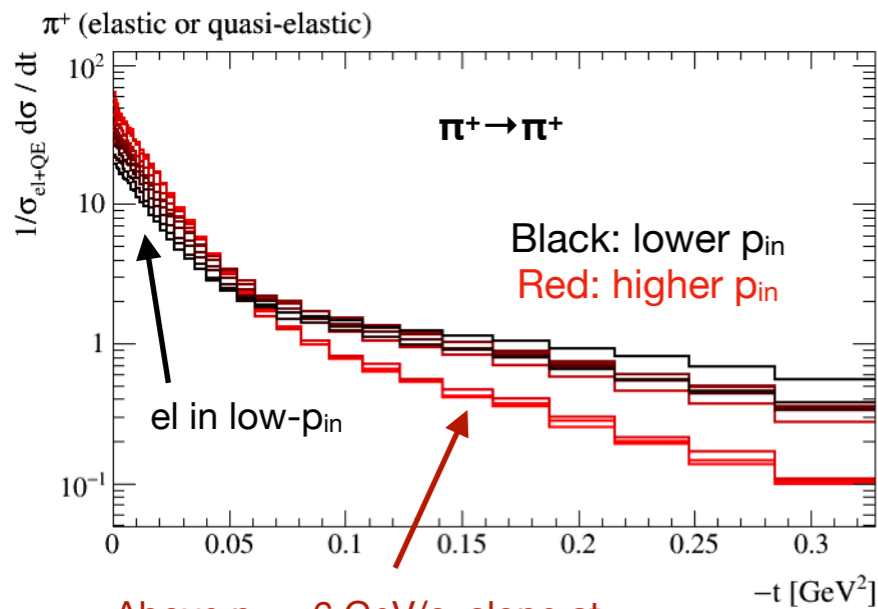
el+QE tuning of protons

FHC

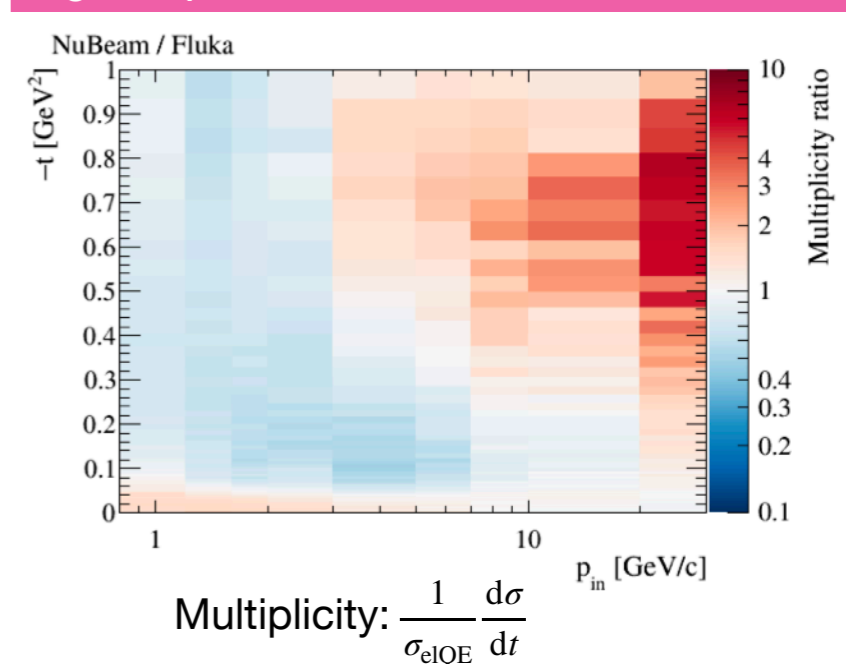
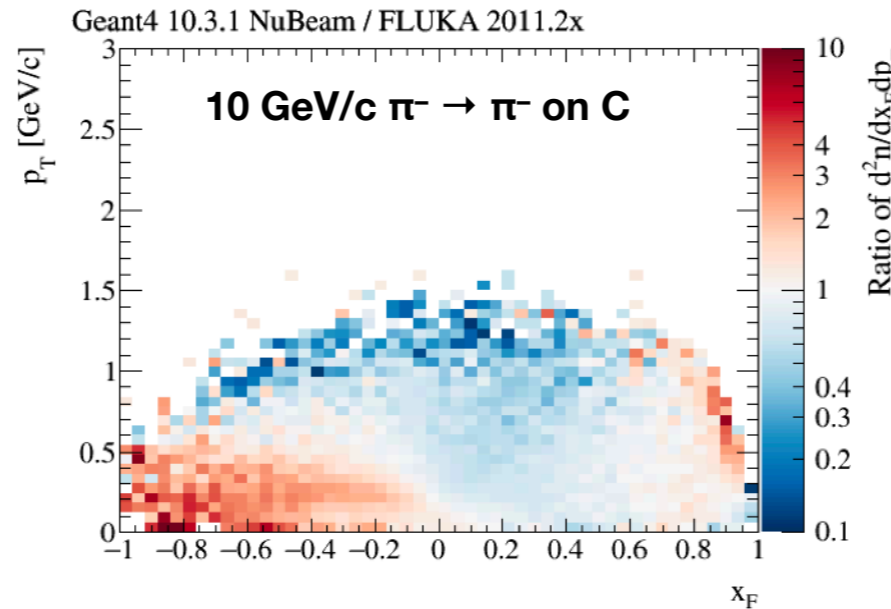


Need to check whether currently assigned uncertainty is large enough

This has no uncertainty assigned right now. Might be possible to tune with EMPHATIC data.

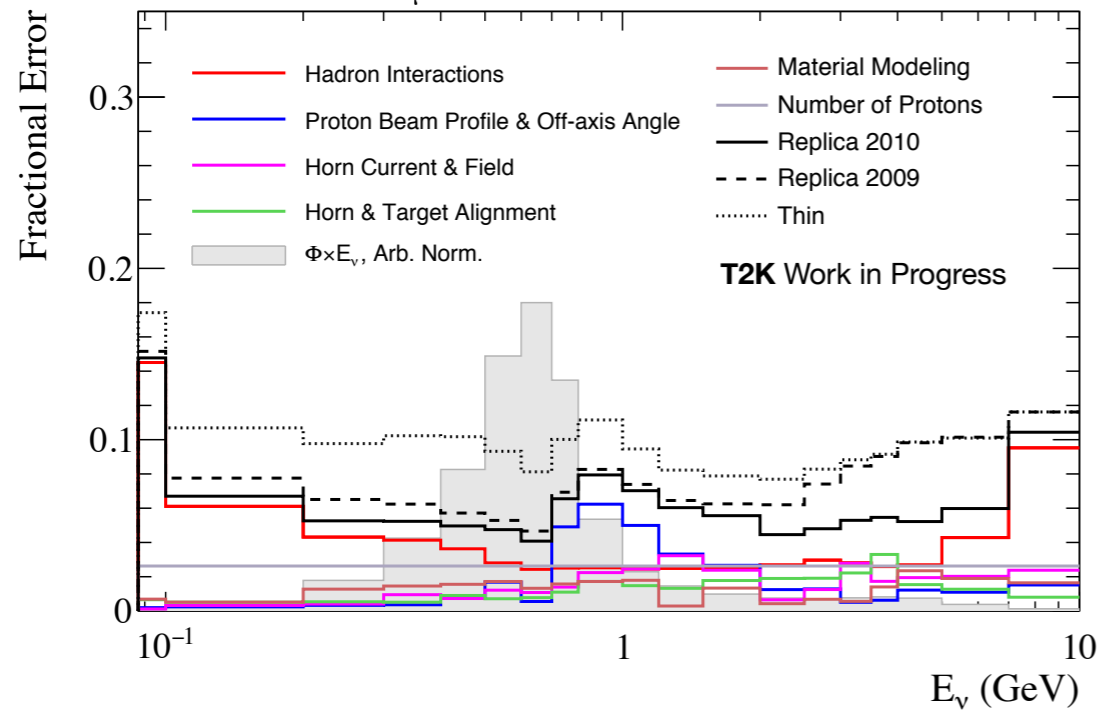


Above $p_{in} \sim 6$ GeV/c, slope at $-t > 0.06$ GeV² becomes steeper due to strange(?) modeling of QE

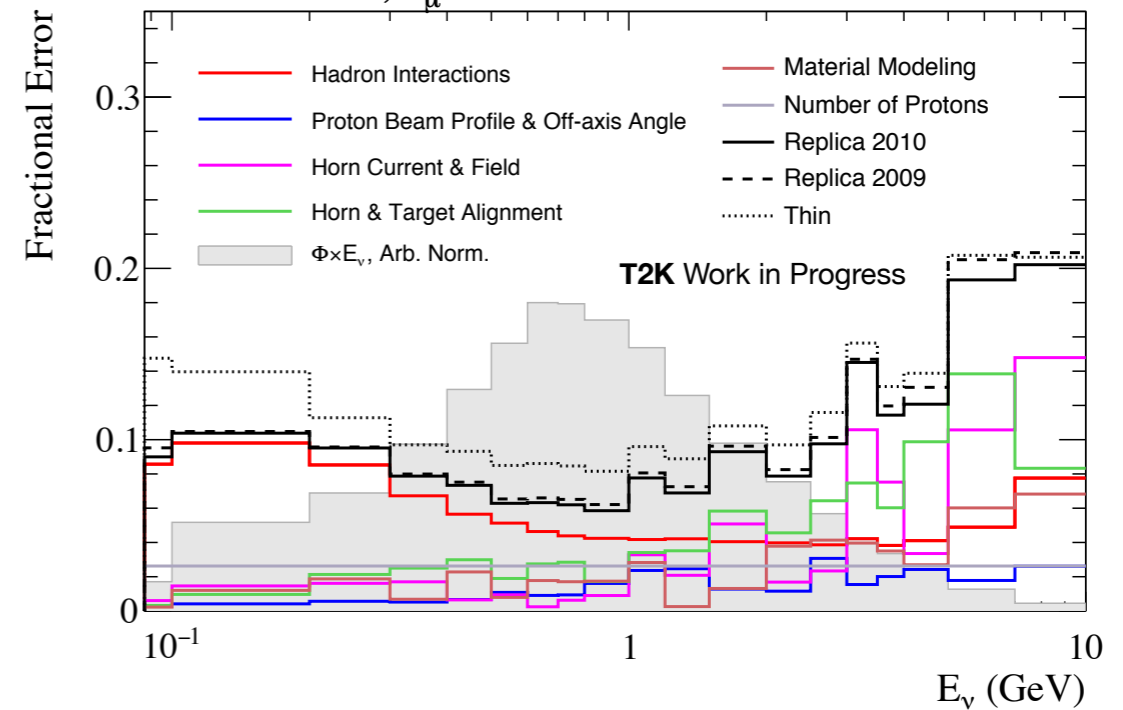


Total flux uncertainty

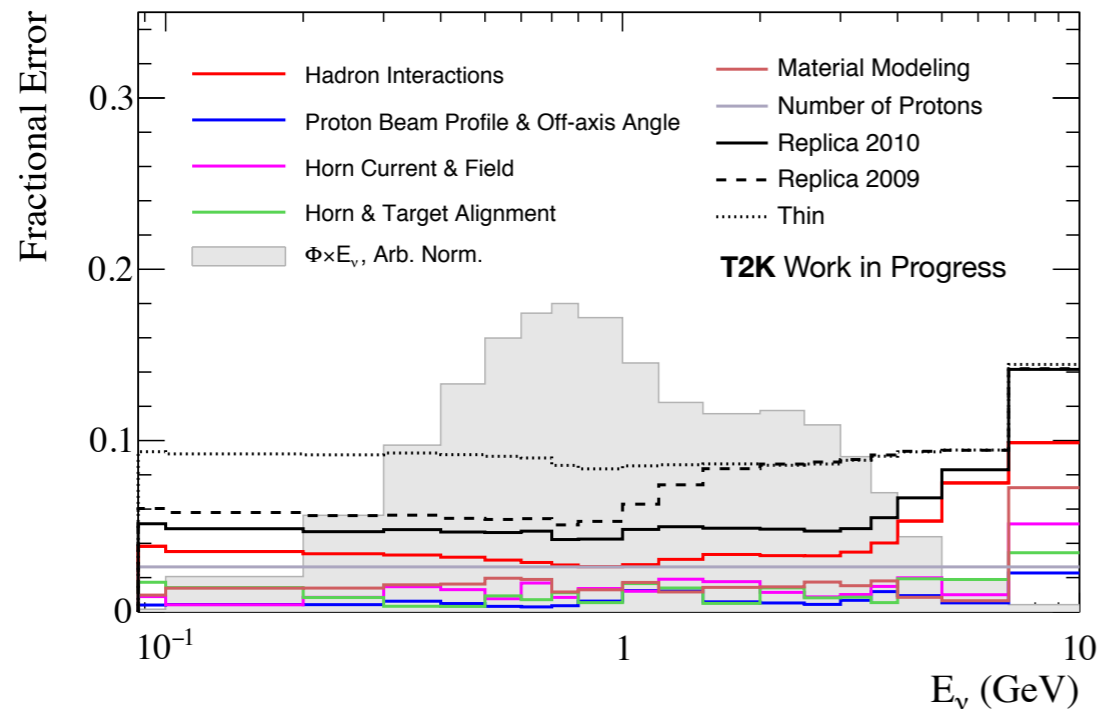
SK: Neutrino Mode, ν_μ



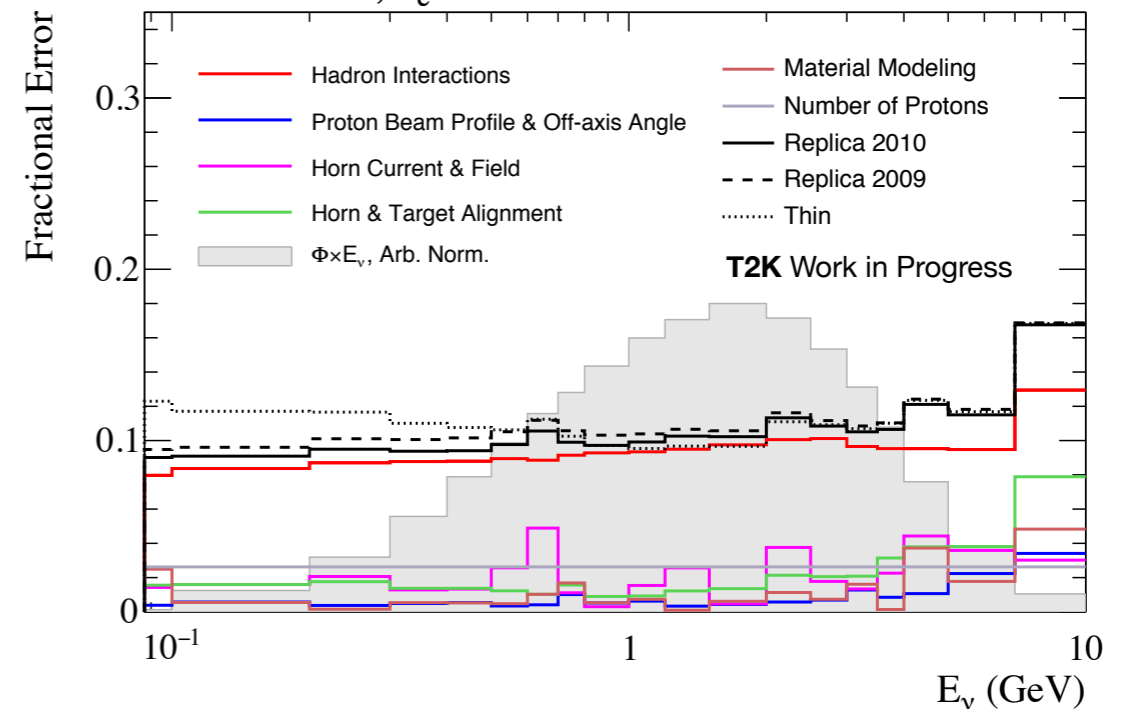
SK: Neutrino Mode, $\bar{\nu}_\mu$



SK: Neutrino Mode, ν_e

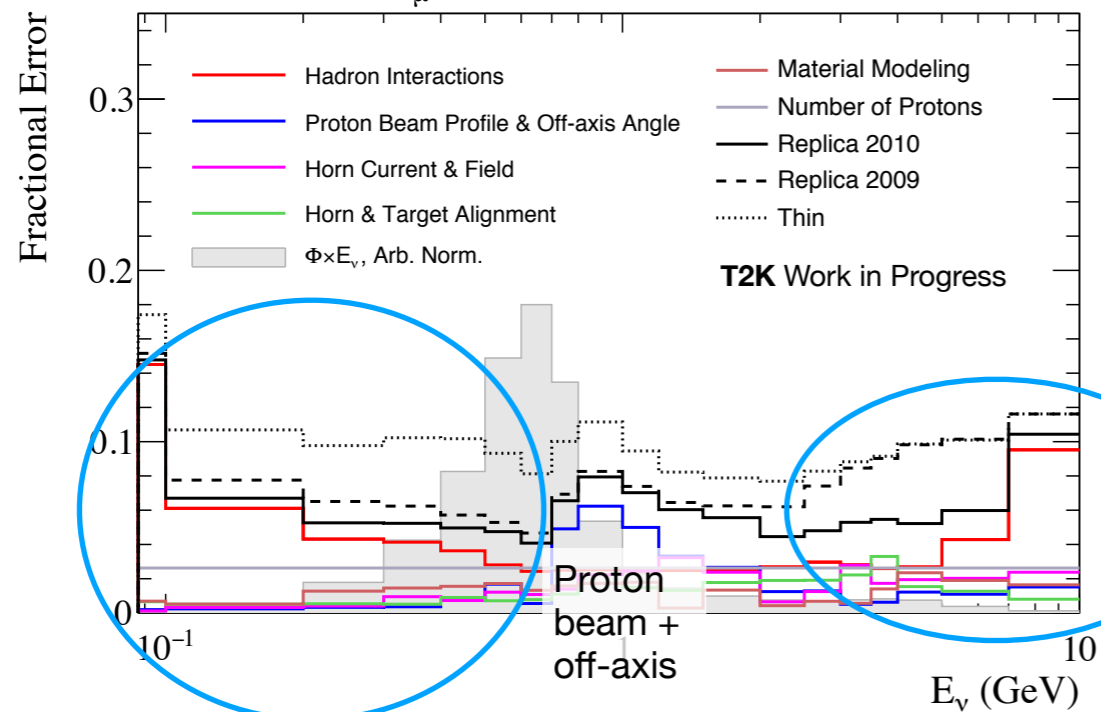


SK: Neutrino Mode, $\bar{\nu}_e$

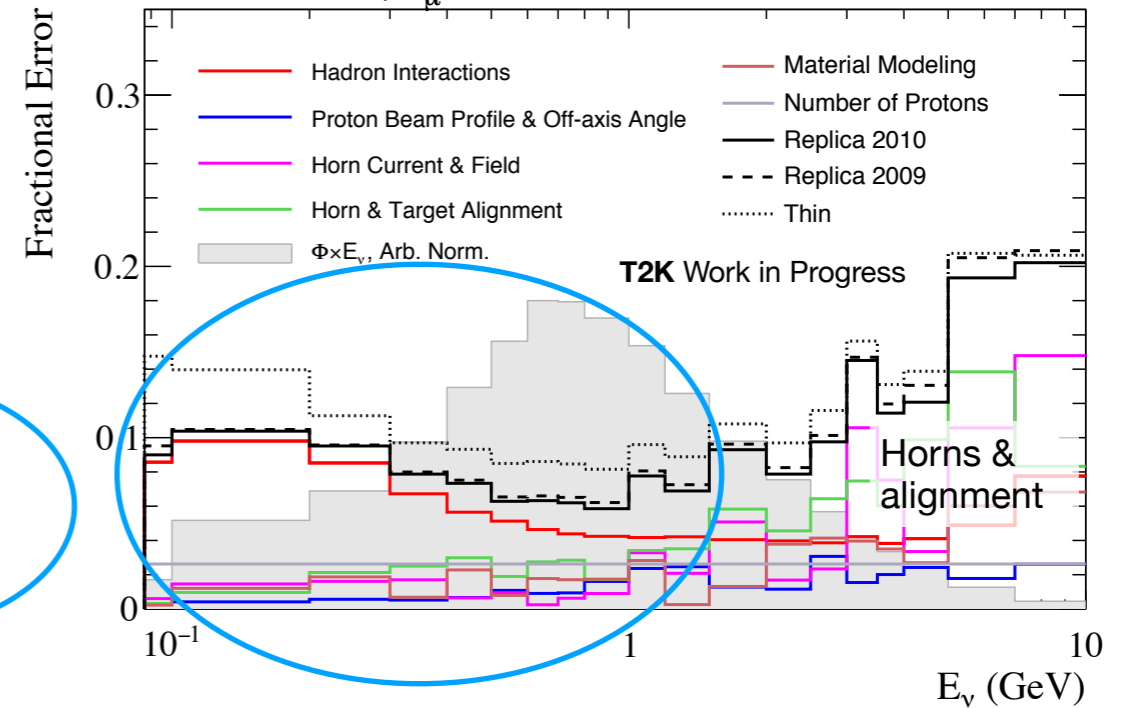


Total flux uncertainty

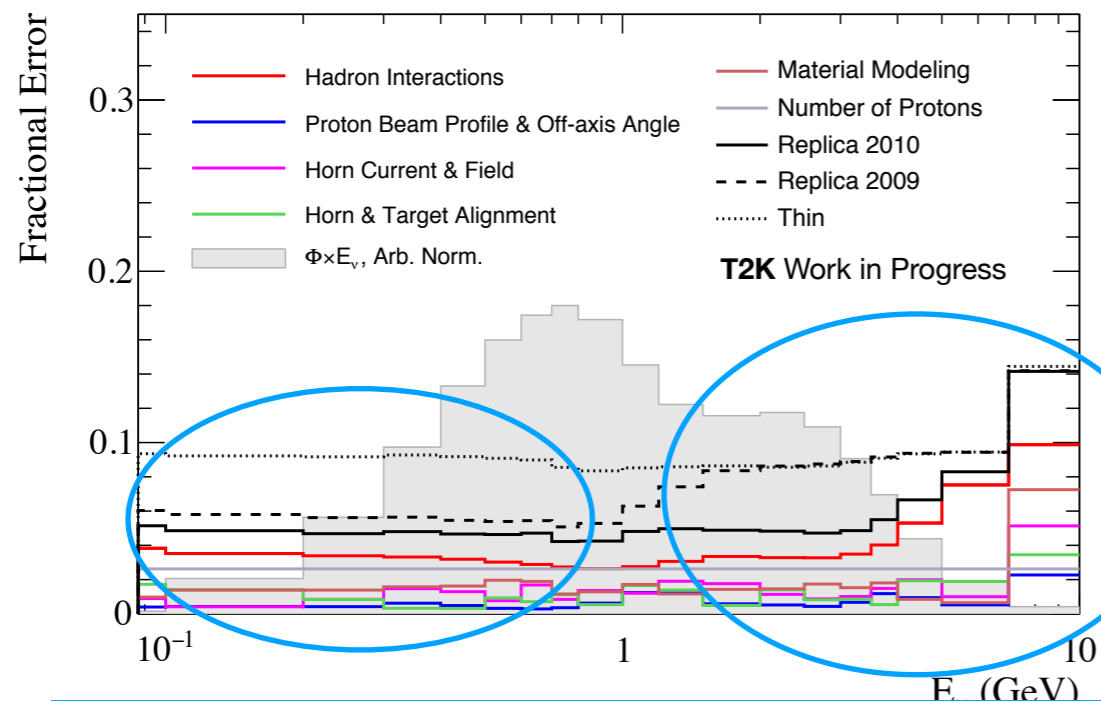
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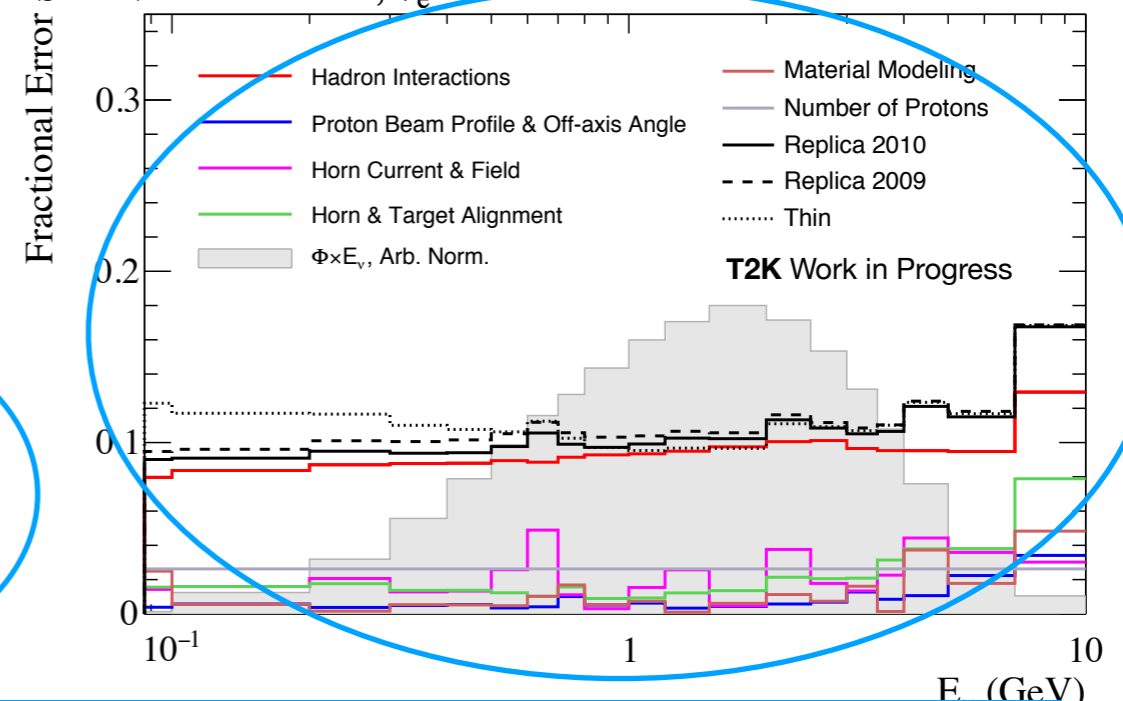
SK: Neutrino Mode, $\bar{\nu}_\mu$



SK: Neutrino Mode, ν_e

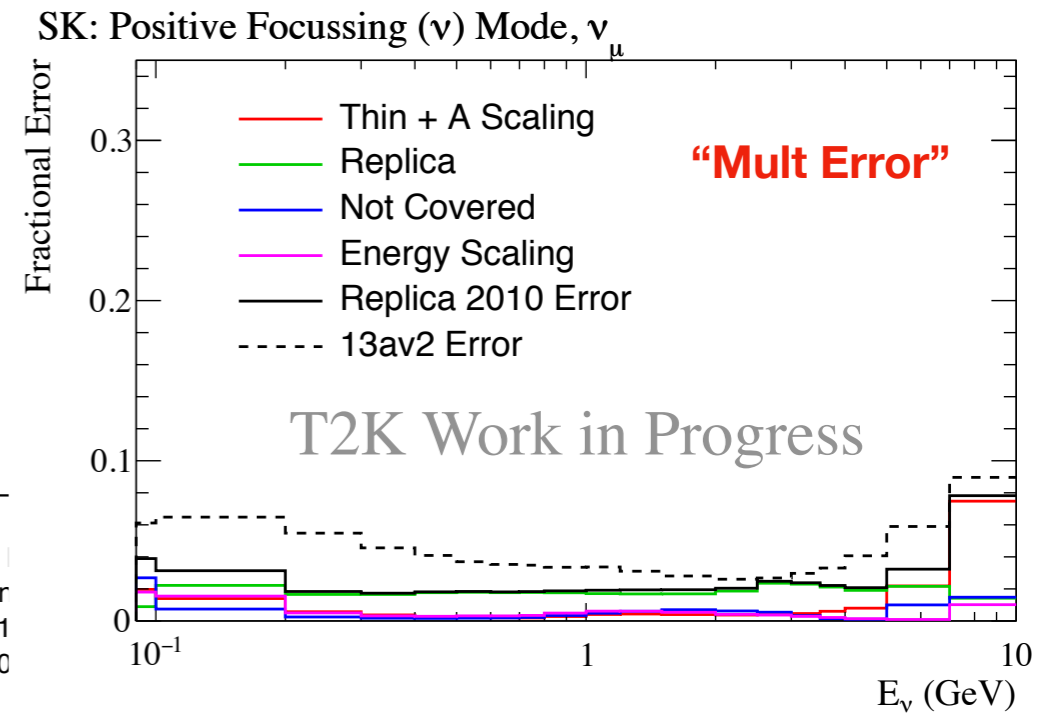
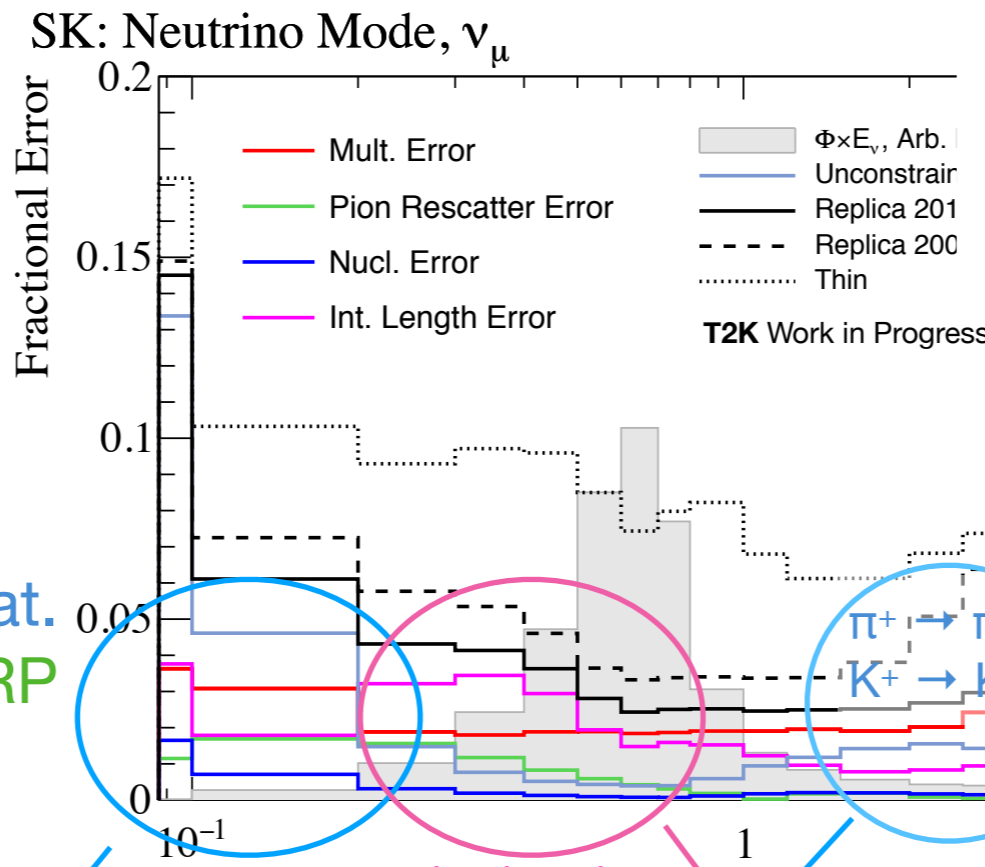


SK: Neutrino Mode, $\bar{\nu}_e$



After replica 2010 tuning, hadron int. uncertainty still dominant in most regions

Right-sign ν_μ

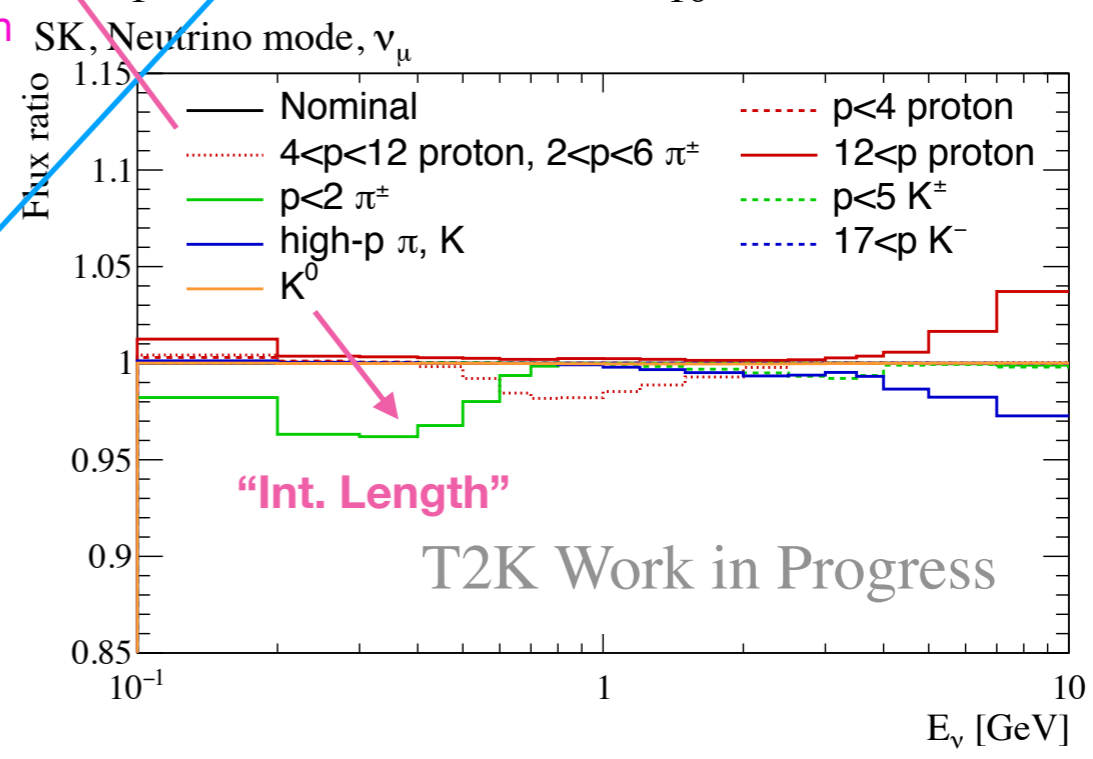
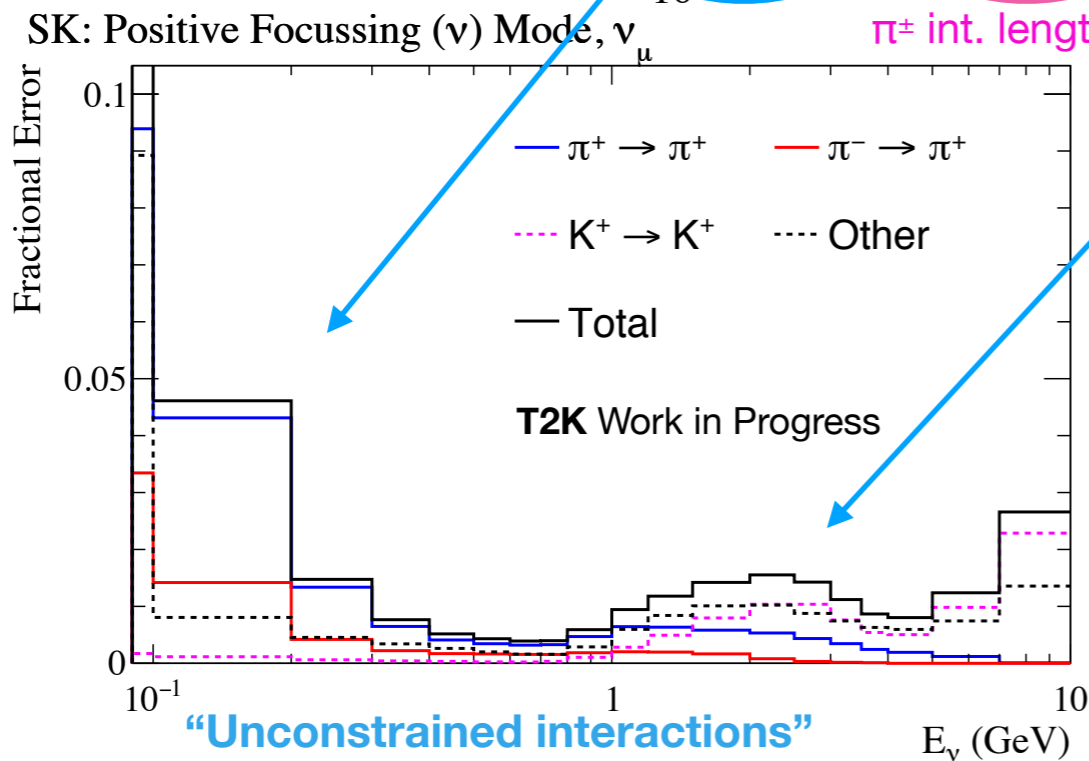


$\pi^+ \rightarrow \pi^+$ scat.
HARP

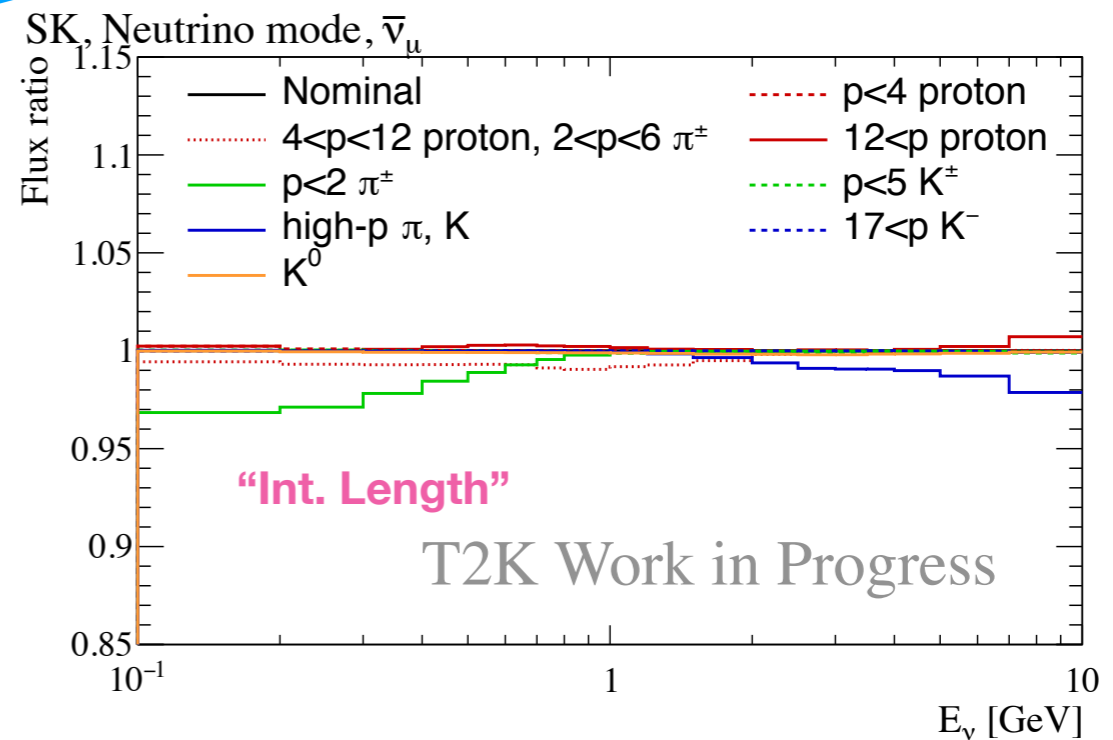
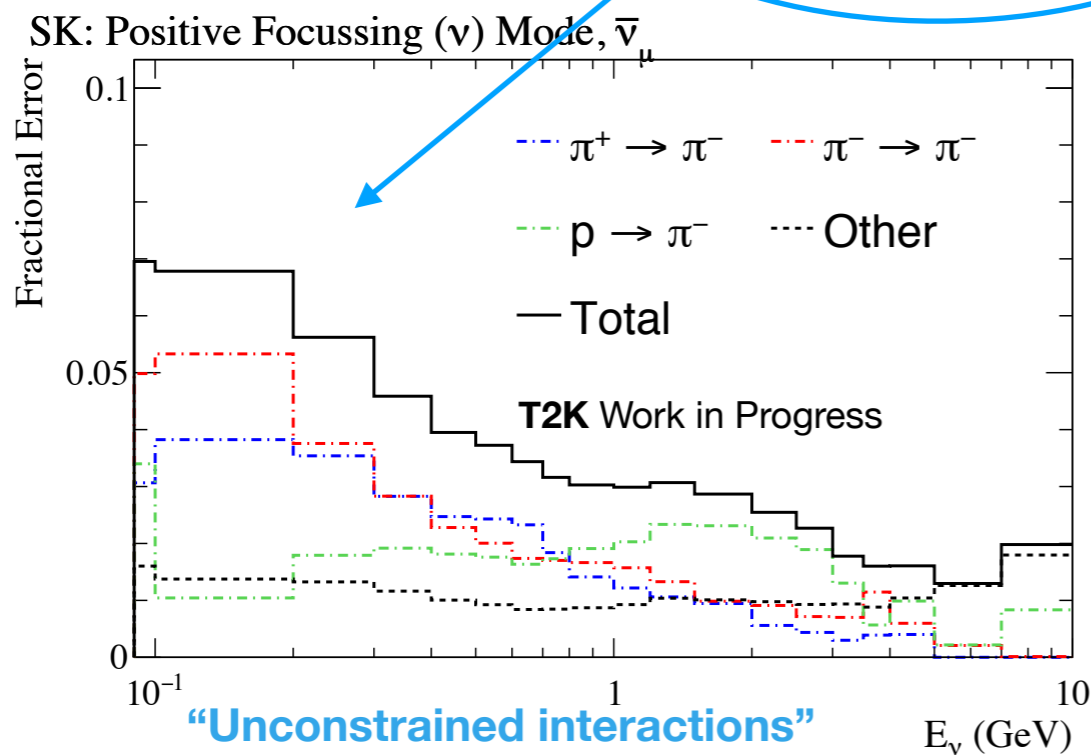
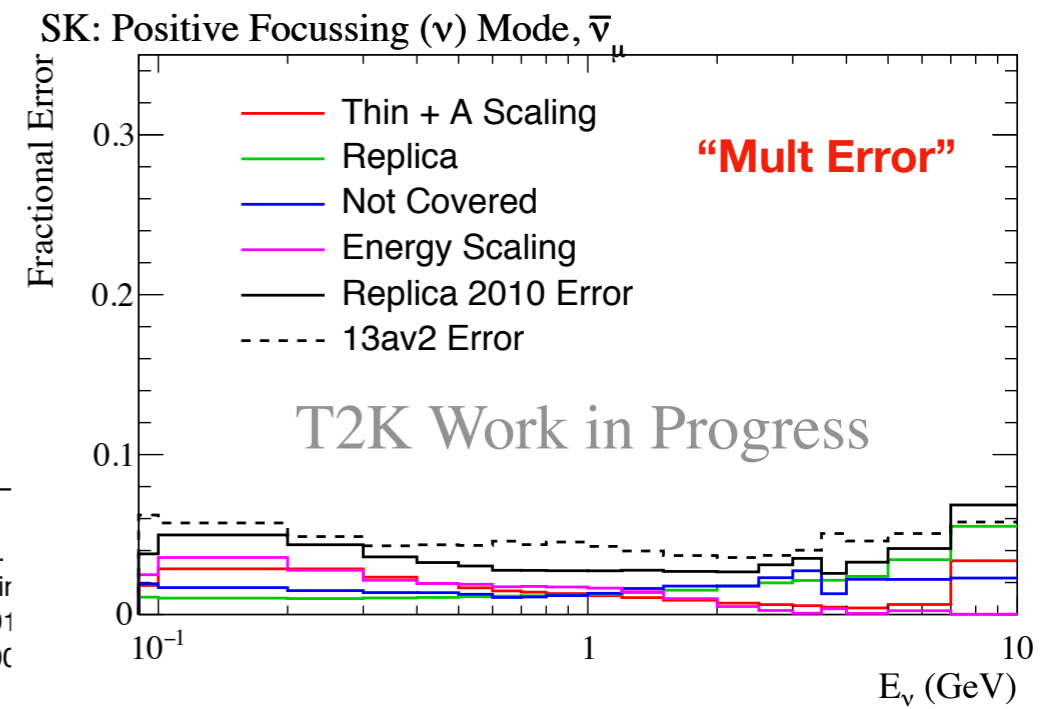
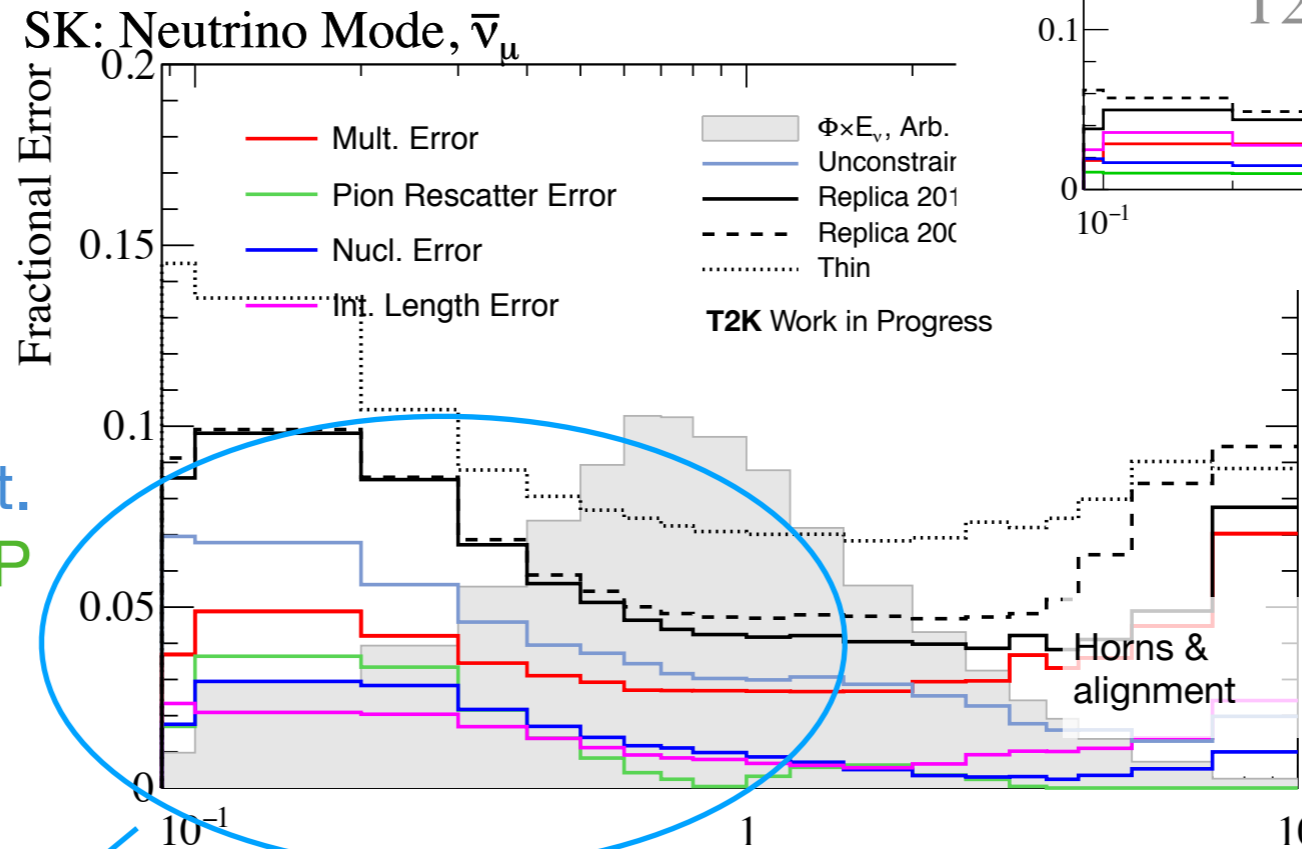
$\pi^+ \rightarrow \pi^+$
 $K^+ \rightarrow K^+$

p, K^\pm
prod.
xsec

π^\pm int. length

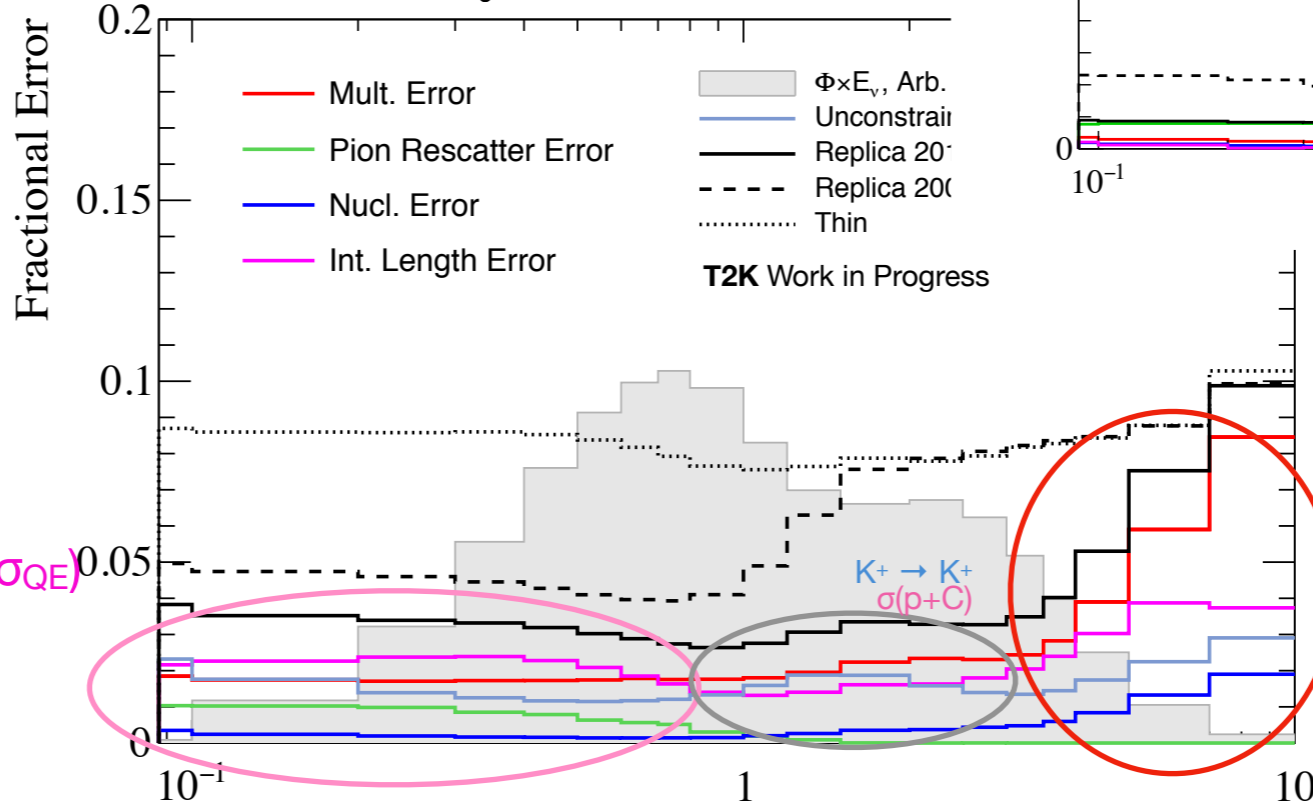


Wrong-sign ν_μ

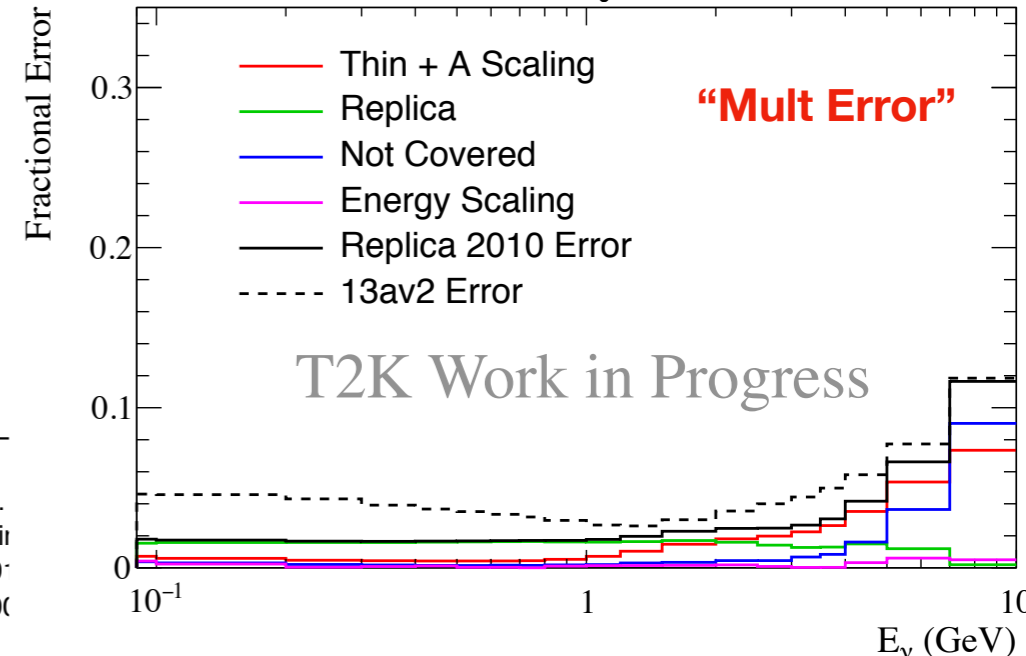


Right-sign ν_e

SK: Neutrino Mode, ν_e



SK: Positive Focussing (ν) Mode, ν_e

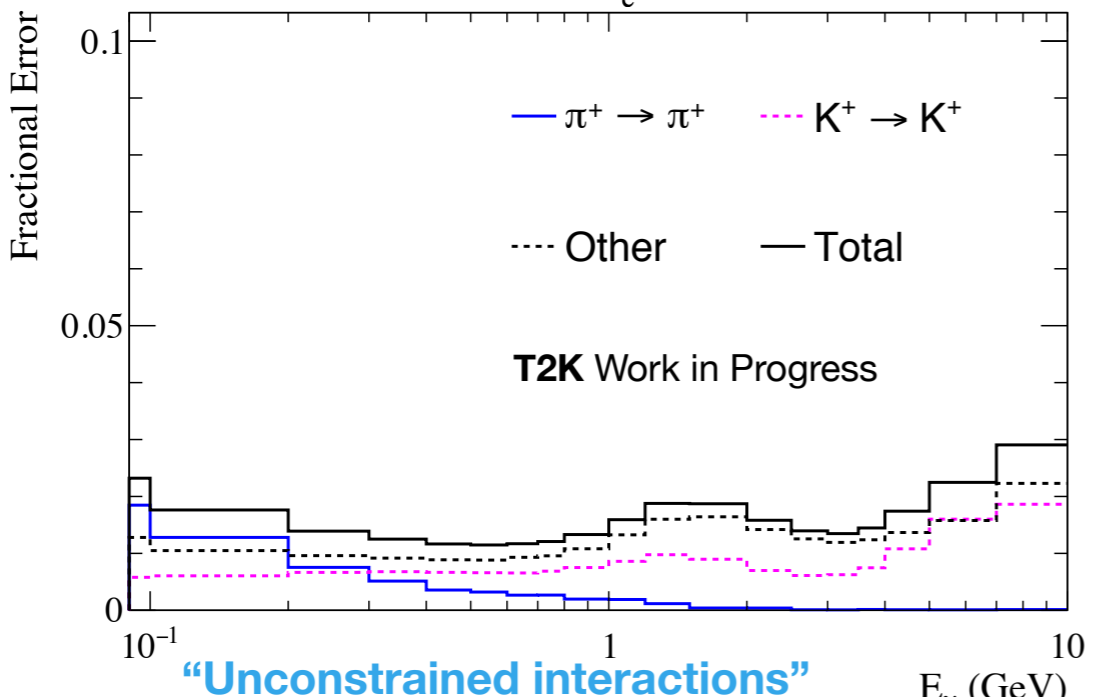


π^\pm prod xsec
uncertainty (taken σ_{QE})

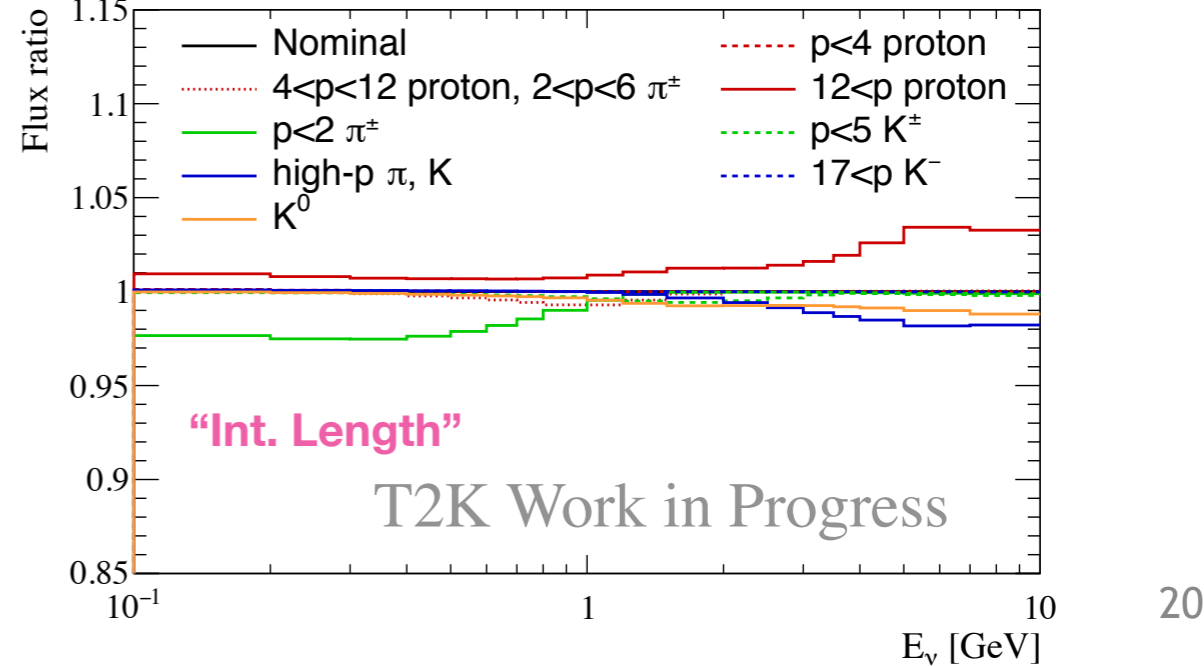
NA61 thin coverage
+ mult. error
(i.e. not covered by replica)

p, K⁺ prod.
xsec

SK: Positive Focussing (ν) Mode, ν_e

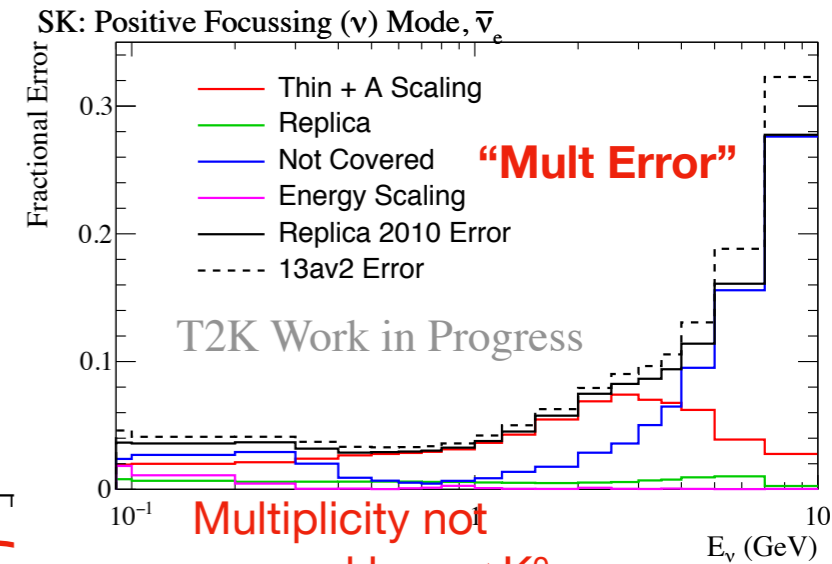


SK, Neutrino mode, ν_e

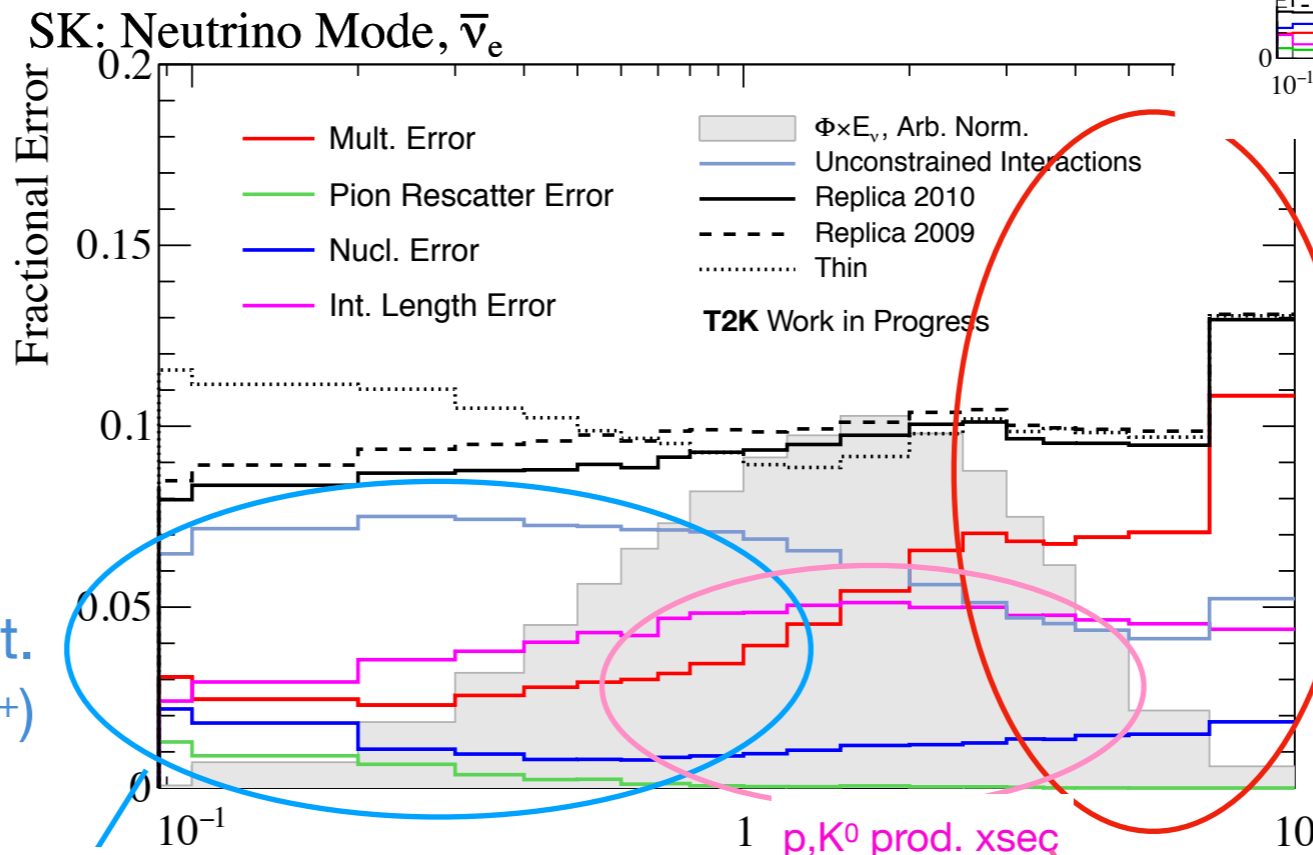


Wrong-sign ν_e are mostly due to K^0 , which are not covered by replica 2010 data

Wrong-sign ν_e

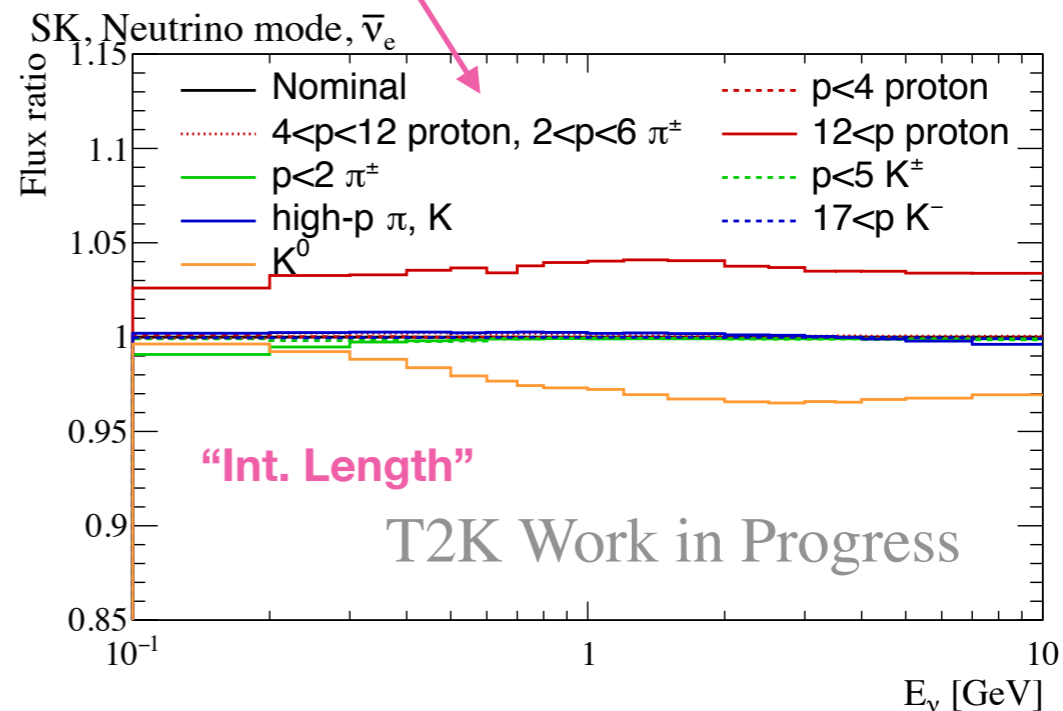
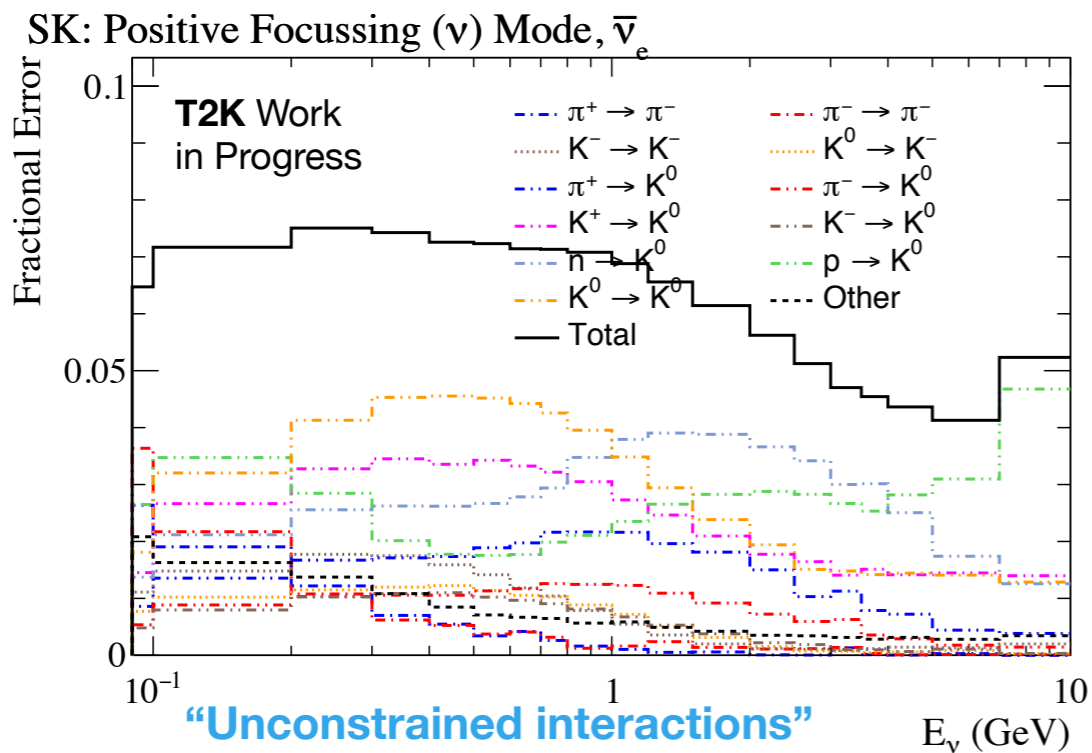


Multiplicity not covered by $p \rightarrow K^0$ thin data



$K^0 \rightarrow K^0$ scat.
(for RHC also $K^+ \rightarrow K^+$)

p, K^0 prod. xsec

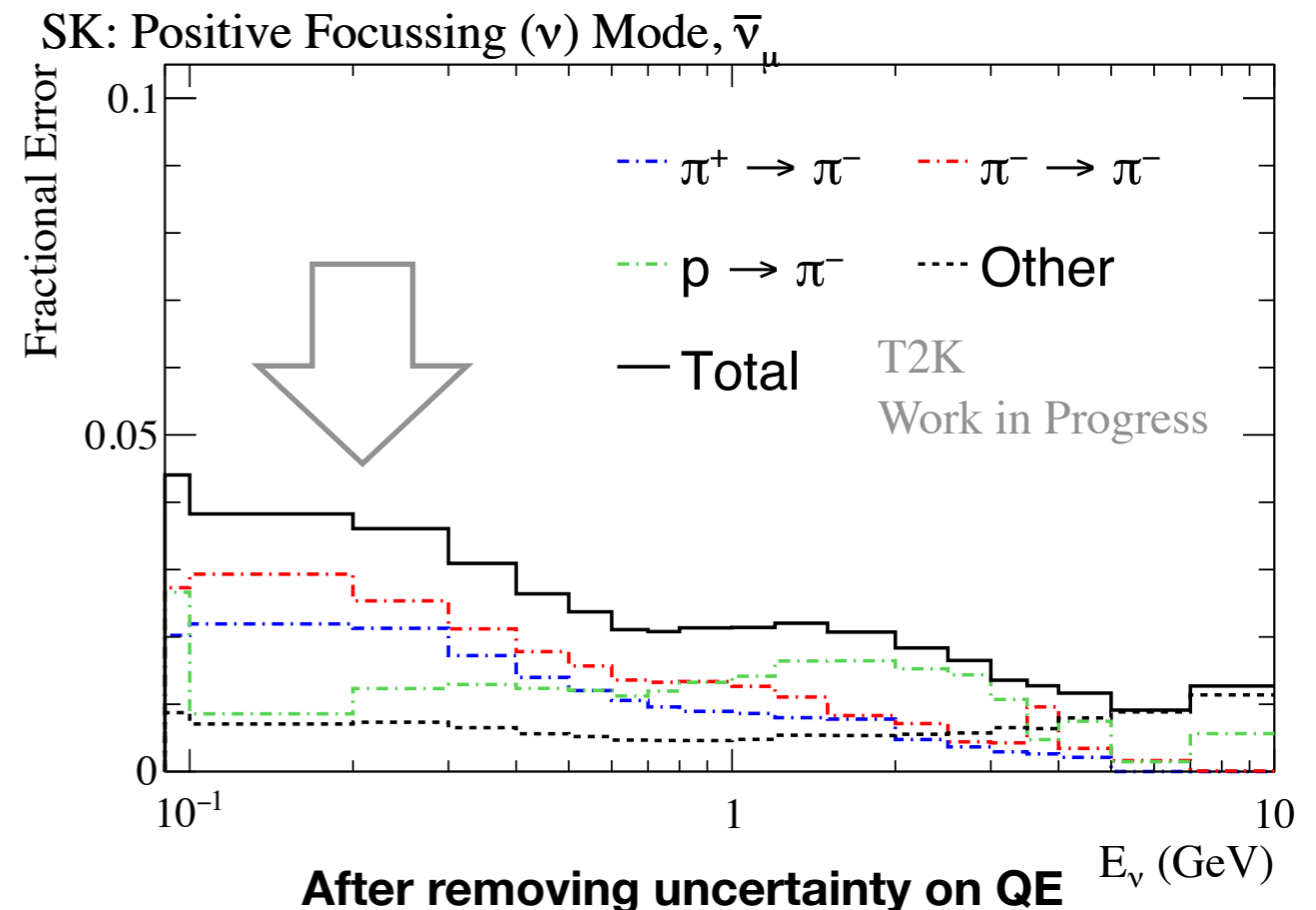
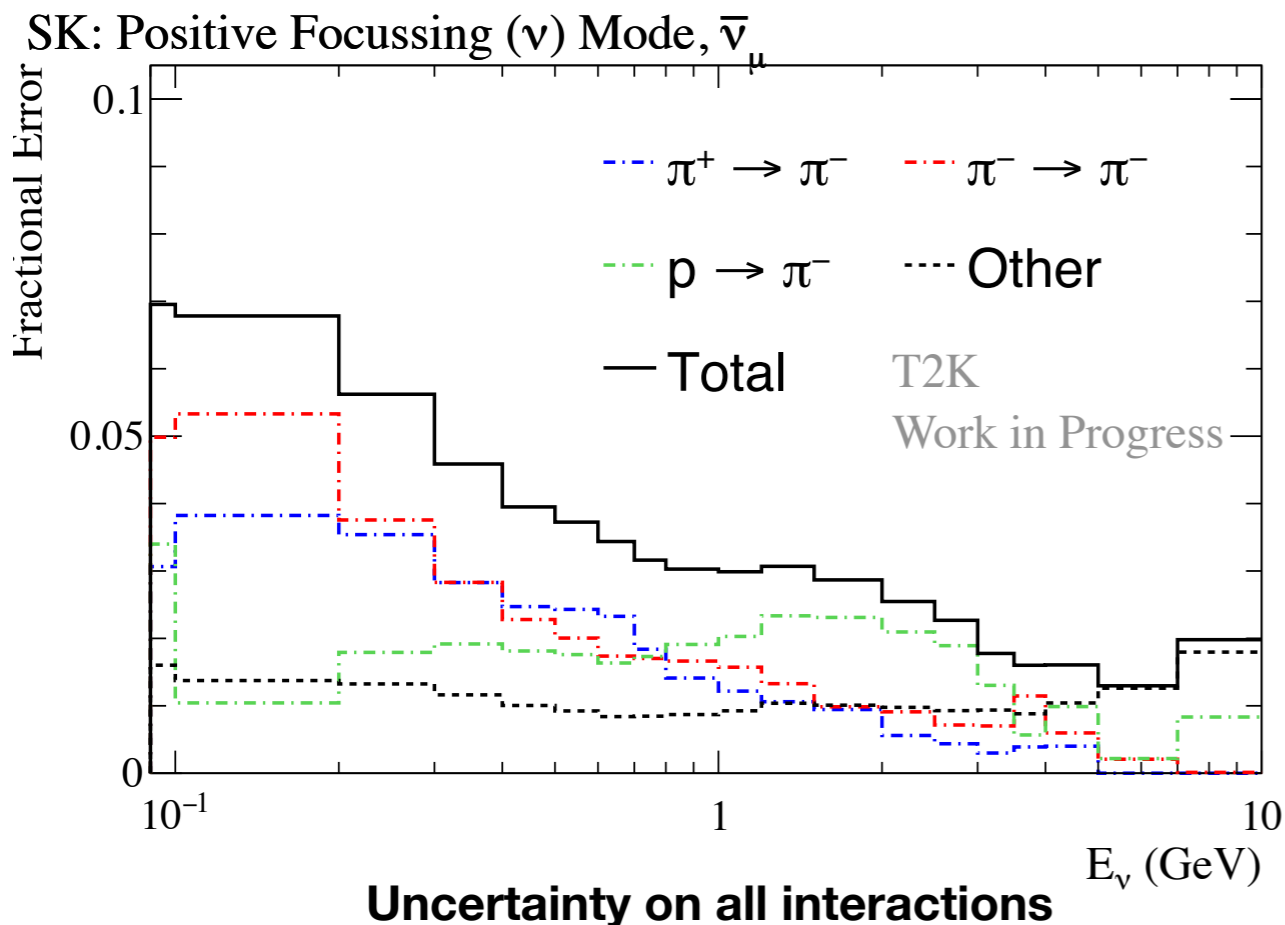


Pion scattering at low-E

- Since exiting pions are mostly covered by replica data, most important materials are Al (horns), Fe (decay volume walls)

note: for QE number of neutrinos doesn't change, so xsec is not as important, but t-distribution can matter

- Same sign scattering is to a large part QE:

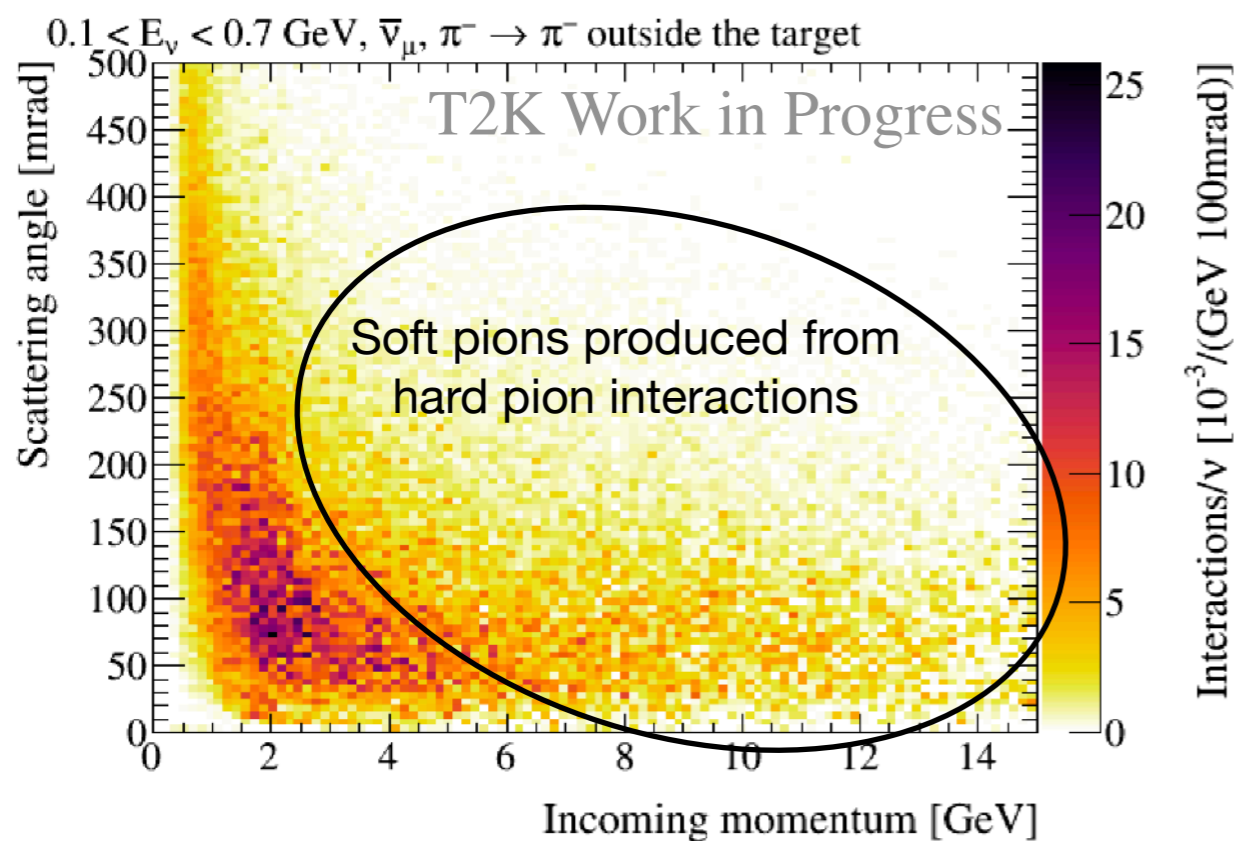


Pion scattering at low-E

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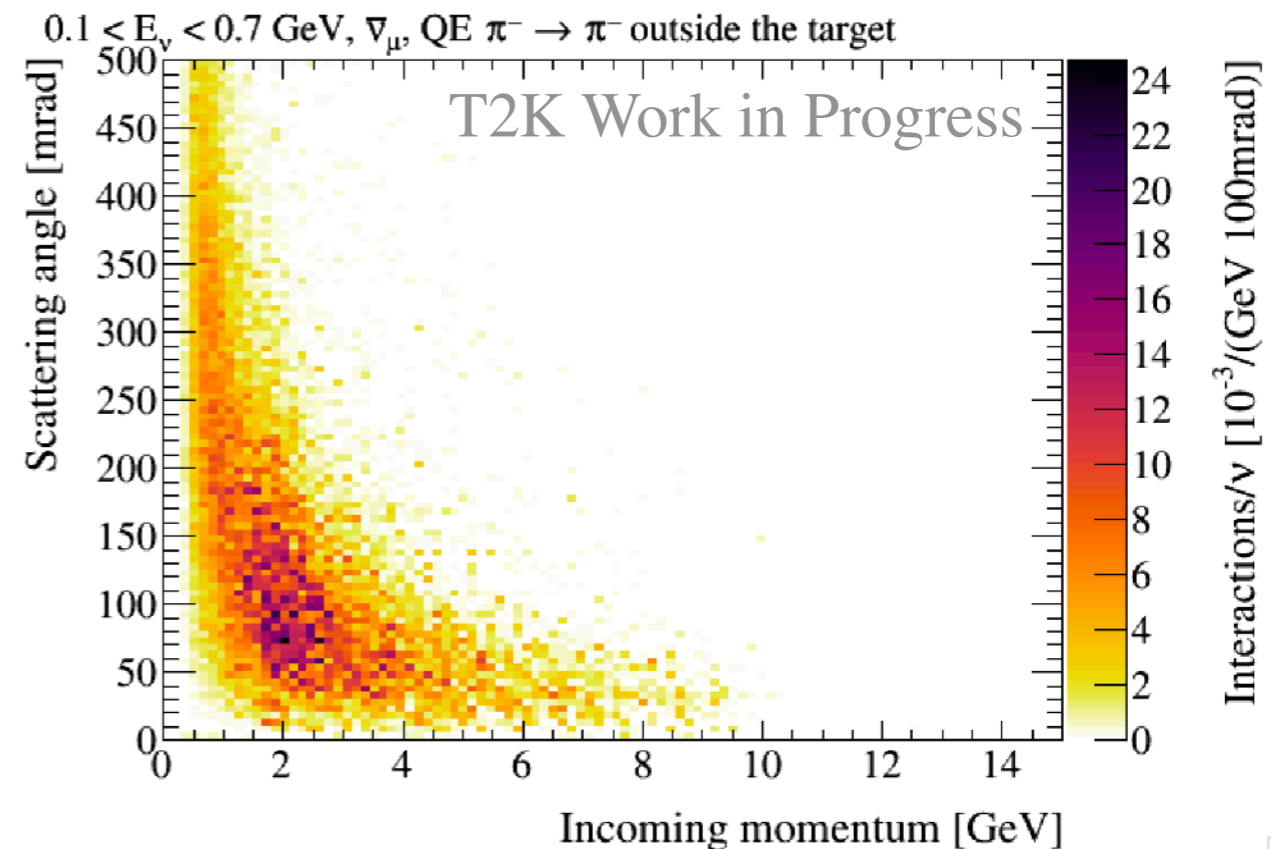
note: for QE number of neutrinos doesn't change, so xsec is not as important, but t-distribution can matter

- Same sign scattering is to a large part QE:



All interactions

$$\cos\theta := 1 - Q^2/(2|p_{\text{in}}| |p_{\text{out}}|)$$

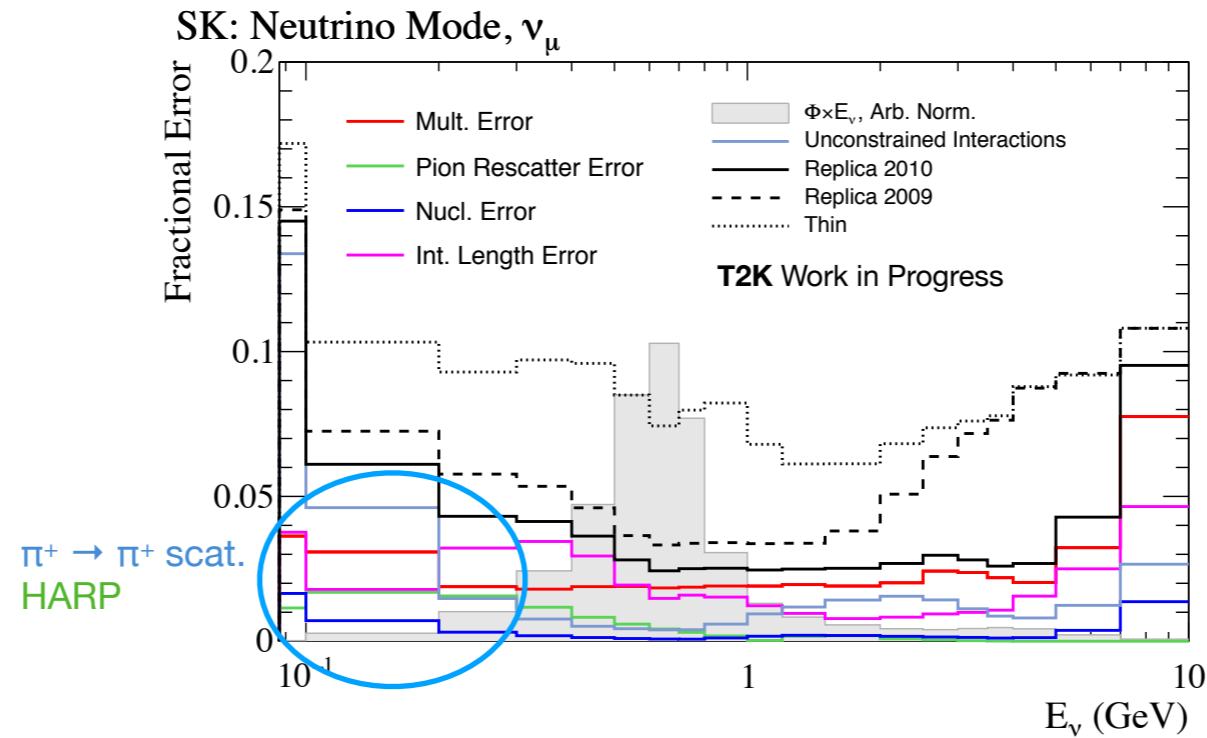
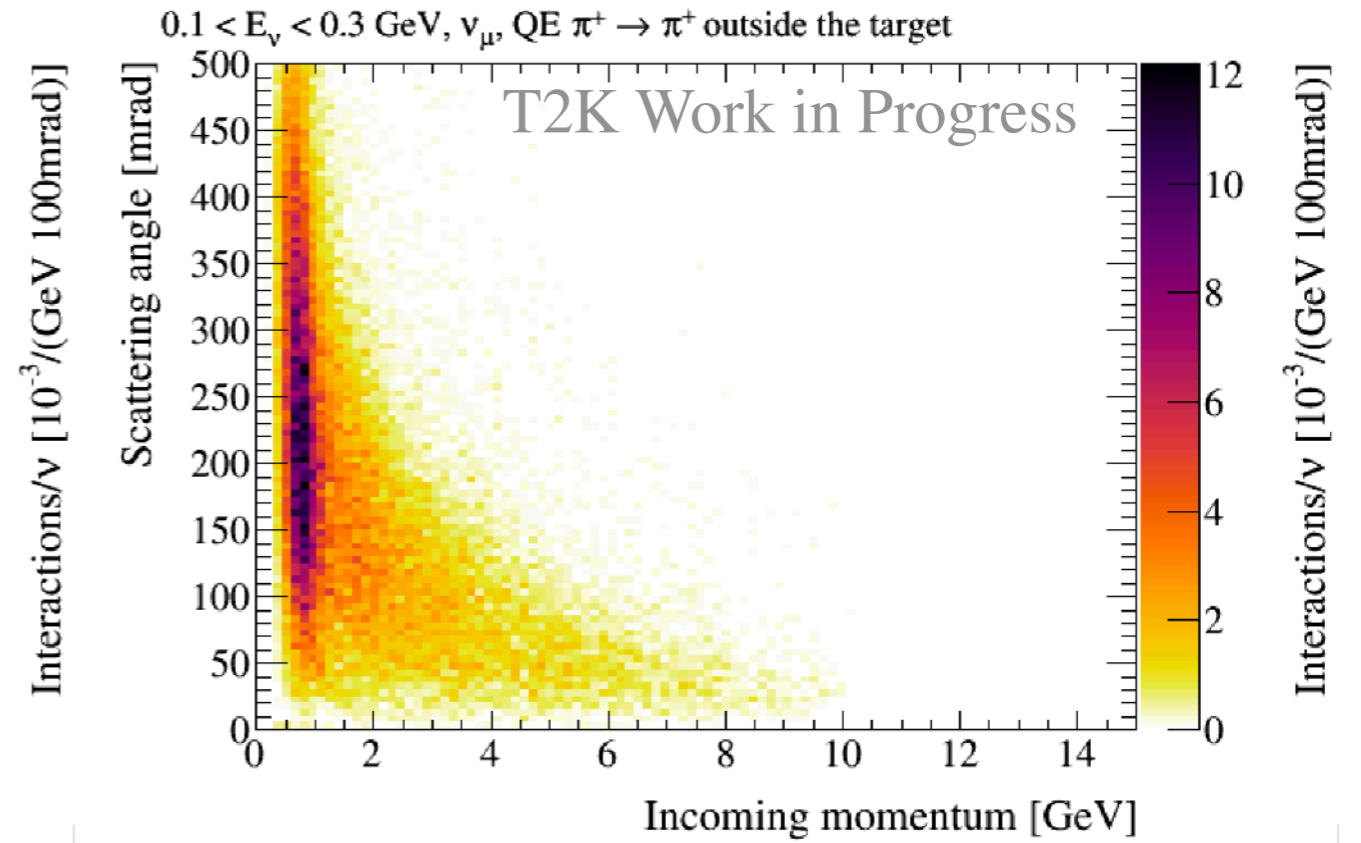
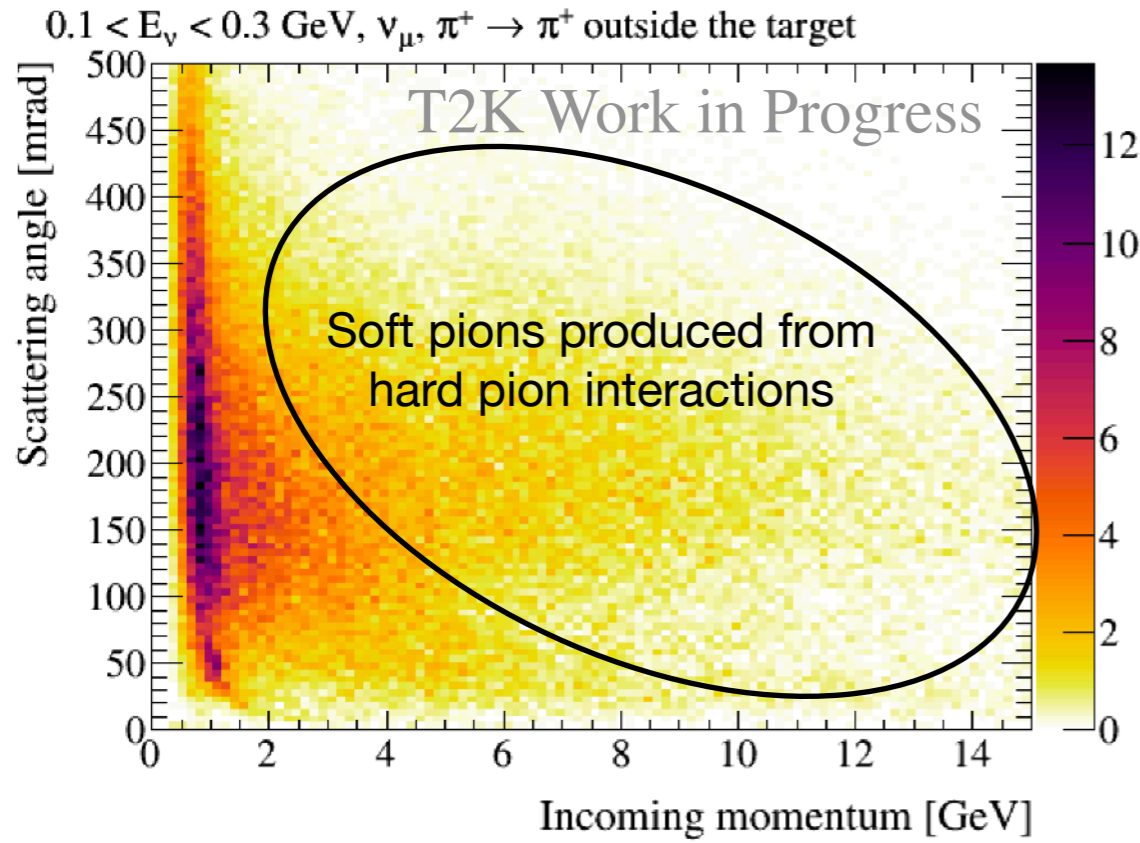


QE interactions

$$\cos\theta := 1 - Q^2/(2|p_{in}| |p_{out}|)$$

All interactions

QE interactions



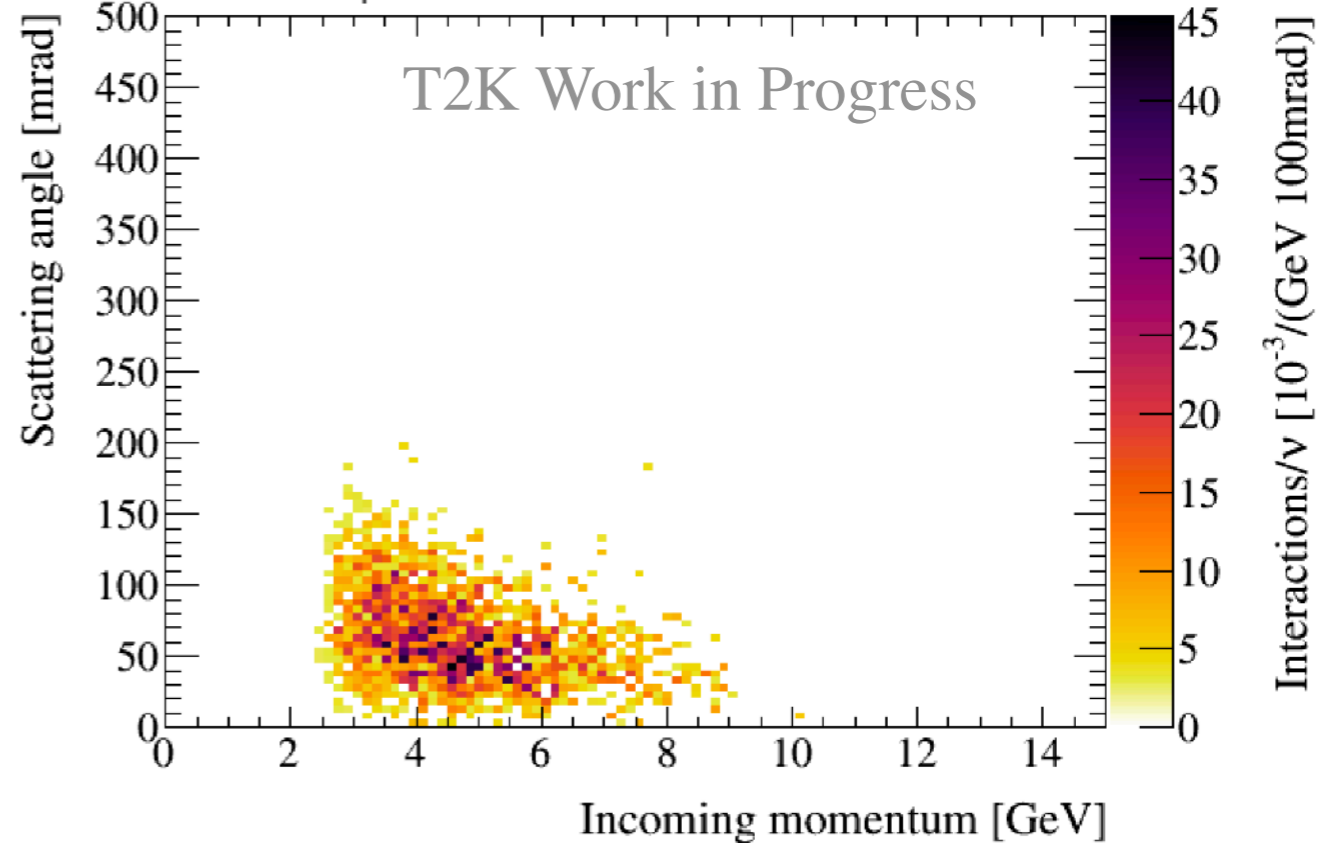
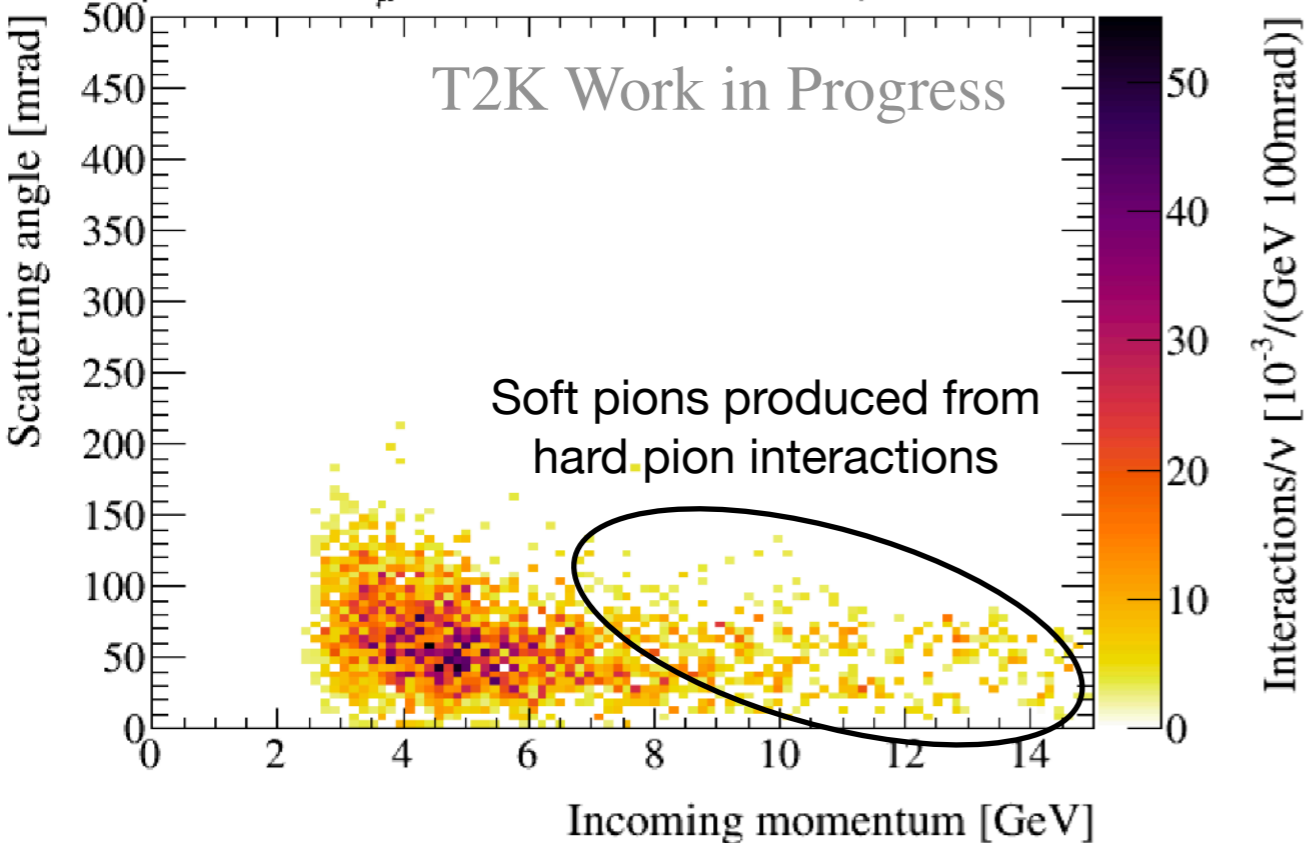
$$\cos\theta := 1 - Q^2/(2|p_{in}| |p_{out}|)$$

All interactions

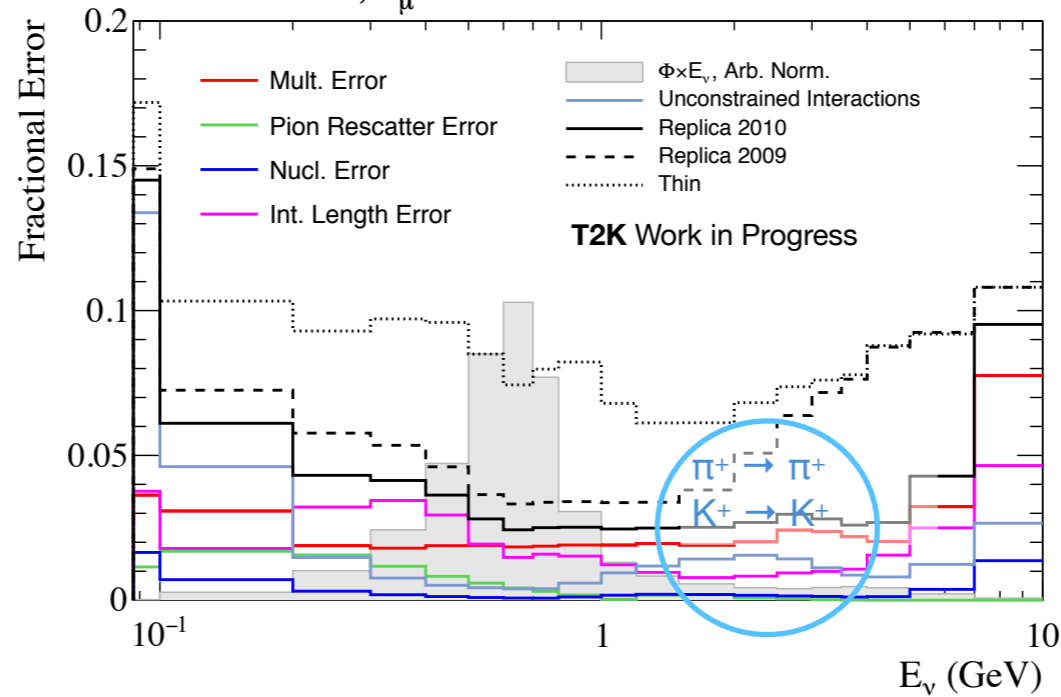
QE interactions

$1 < E_\nu < 2 \text{ GeV}$, ν_μ , $\pi^+ \rightarrow \pi^+$ outside the target

$1 < E_\nu < 2 \text{ GeV}$, ν_μ , QE $\pi^+ \rightarrow \pi^+$ outside the target



SK: Neutrino Mode, ν_μ



$$\cos\theta := 1 - Q^2/(2|p_{in}| |p_{out}|)$$

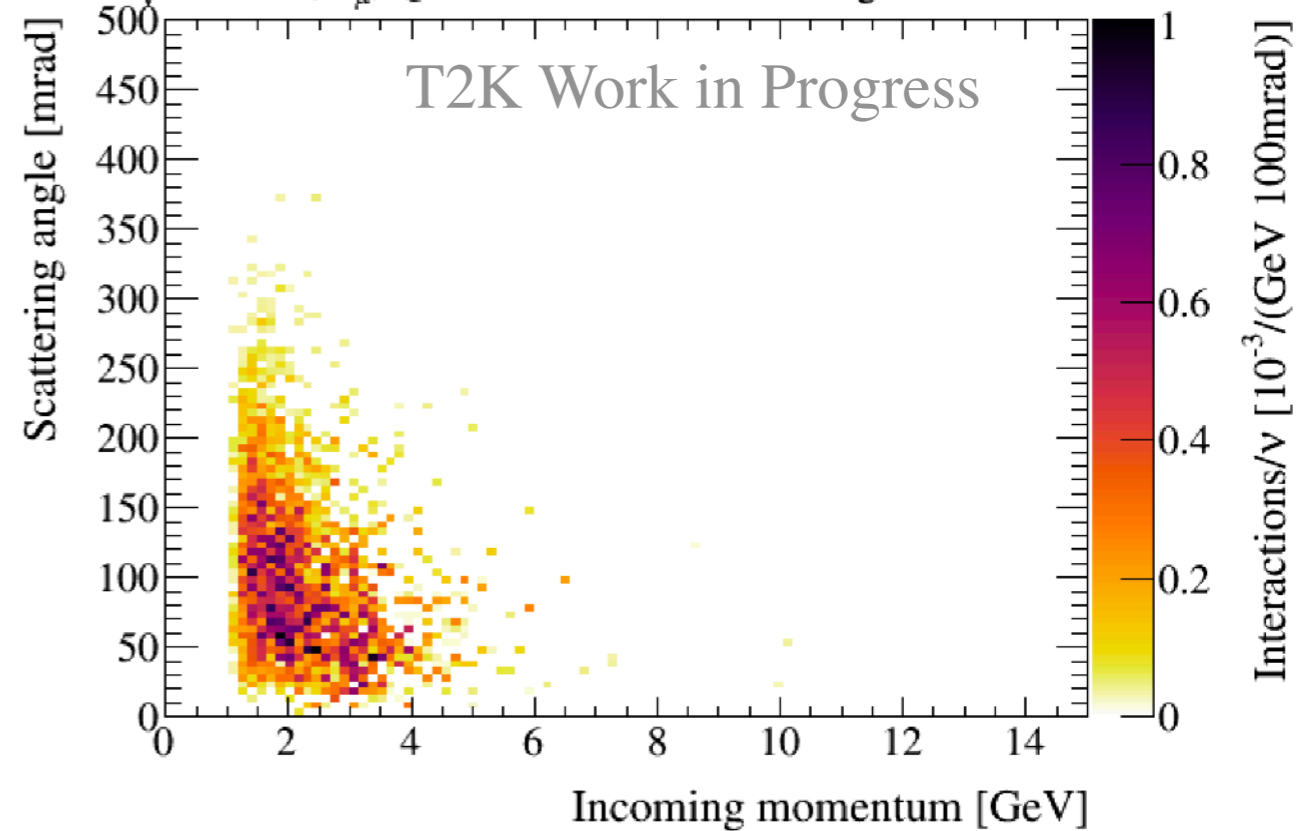
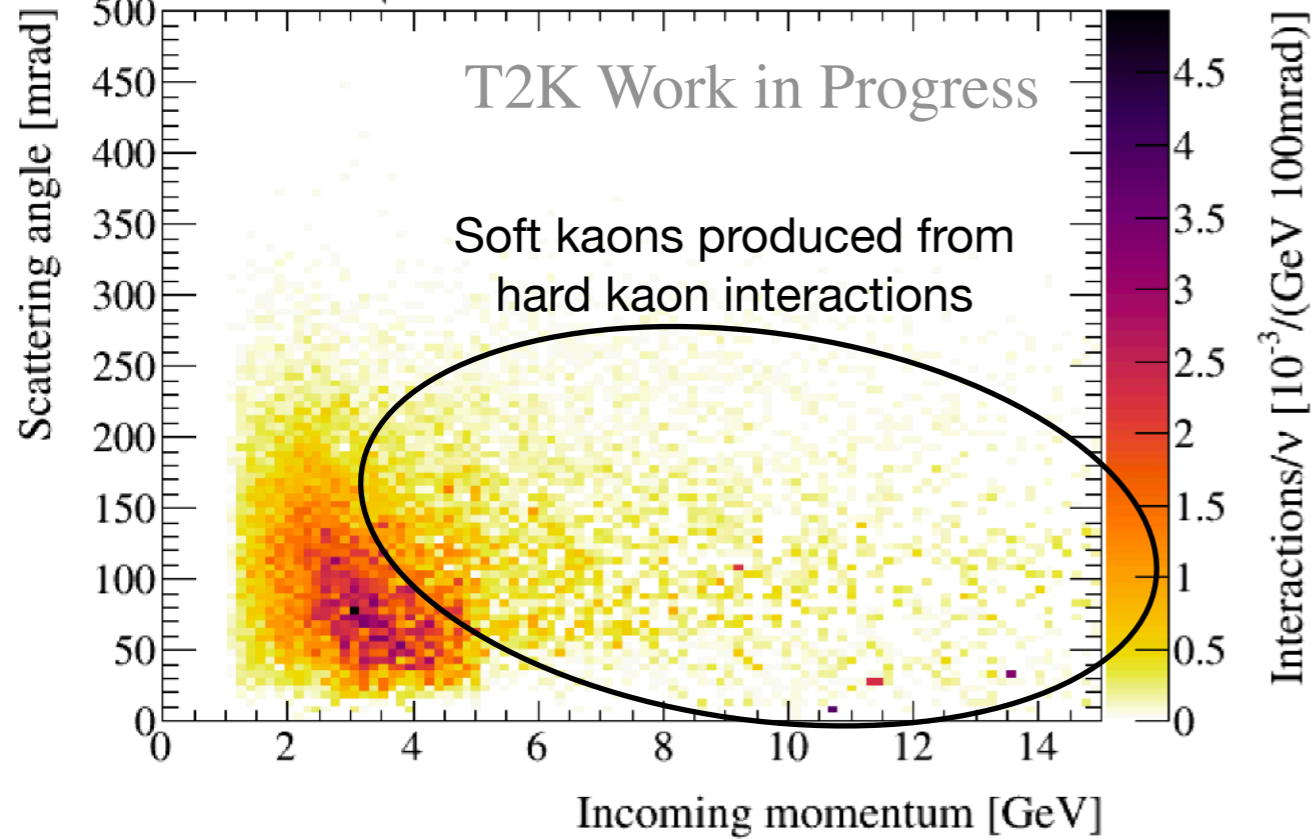
I think should take m_K into account for proper angle

All interactions

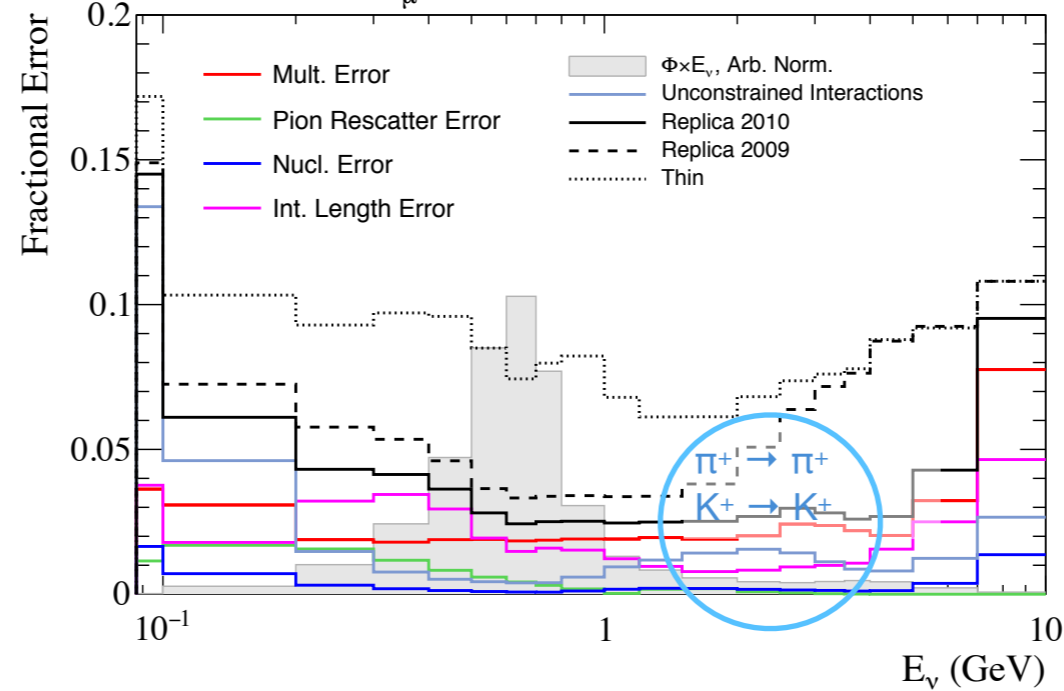
QE interactions

$1 < E_\nu < 3 \text{ GeV}, \nu_\mu, K^+ \rightarrow K^+$ outside the target

$1 < E_\nu < 3 \text{ GeV}, \nu_\mu, \text{QE } K^+ \rightarrow K^+$ outside the target



SK: Neutrino Mode, ν_μ



Note:

This K^\pm scattering may get covered by the proposed high-stats replica data taking with extended coverage, so need to check

Flux-weighted fractions of “unconstrained interactions”

SK, Neutrino mode

largest contributions:

	→ Carbon	→ Aluminum	→ Iron	→ Titanium	→ Helium	→ Other
$\pi^+ \rightarrow$	8.87	9.08	7.76	2.06	2.19	0.36
$\pi^- \rightarrow$	5.35	5.23	5.69	0.88	0.26	0.35
$K^0 \rightarrow$	4.22	3.33	1.71	0.24	0.05	0.11
$K^+ \rightarrow$	4.08	4.85	1.20	0.31	0.11	0.11
$K^- \rightarrow$	0.94	1.14	0.36	0.07	0.01	0.03
$n \rightarrow$	2.52	1.77	1.73	0.09	1.36	0.57
$p \rightarrow$	3.02	0.91	0.56	0.06	8.96	0.44
$\bar{p} \rightarrow$	0.21	0.33	0.09	0.02	0.01	0.01
$\Lambda \rightarrow$	2.78	1.21	0.11	0.14	0.00	0.01
$\Sigma \rightarrow$	0.94	0.22	0.03	0.03	0.00	0.00
$\bar{n} \rightarrow$	0.29	0.46	0.14	0.03	0.01	0.01

Channel	Flux-weighted fraction [%]
$\pi^+ \rightarrow$ Aluminum	9.08
$p \rightarrow$ Helium	8.96
$\pi^+ \rightarrow$ Carbon	8.87
$\pi^+ \rightarrow$ Iron	7.76
$\pi^- \rightarrow$ Iron	5.69
$\pi^- \rightarrow$ Carbon	5.35
$\pi^- \rightarrow$ Aluminum	5.23
$K^+ \rightarrow$ Aluminum	4.85
$K^0 \rightarrow$ Carbon	4.22
$K^+ \rightarrow$ Carbon	4.08
Other	35.90

SK, Anti-neutrino mode

largest contributions:

	→ Carbon	→ Aluminum	→ Iron	→ Titanium	→ Helium	→ Other
$\pi^+ \rightarrow$	7.26	7.09	8.02	1.25	0.39	0.42
$\pi^- \rightarrow$	6.17	6.47	5.79	1.45	1.59	0.30
$K^0 \rightarrow$	3.77	3.06	1.69	0.22	0.05	0.11
$K^+ \rightarrow$	2.75	2.81	1.35	0.17	0.04	0.13
$K^- \rightarrow$	1.35	1.95	0.35	0.13	0.04	0.03
$n \rightarrow$	4.49	1.60	1.67	0.07	1.35	0.53
$p \rightarrow$	2.78	0.88	0.68	0.06	7.23	0.61
$\bar{p} \rightarrow$	0.27	0.43	0.12	0.03	0.02	0.01
$\Lambda \rightarrow$	6.17	1.57	0.12	0.18	0.01	0.01
$\Sigma \rightarrow$	1.81	0.27	0.03	0.04	0.00	0.00
$\bar{n} \rightarrow$	0.23	0.37	0.14	0.02	0.01	0.01

Channel	Flux-weighted fraction [%]
$\pi^+ \rightarrow$ Iron	8.02
$\pi^+ \rightarrow$ Carbon	7.26
$p \rightarrow$ Helium	7.23
$\pi^+ \rightarrow$ Aluminum	7.09
$\pi^- \rightarrow$ Aluminum	6.47
$\pi^- \rightarrow$ Carbon	6.17
$\Lambda \rightarrow$ Carbon	6.17
$\pi^- \rightarrow$ Iron	5.79
$n \rightarrow$ Carbon	4.49
$K^0 \rightarrow$ Carbon	3.77
Other	37.54

Flux-weighted fractions of “unconstrained interactions”

SK, Neutrino mode

largest contributions:

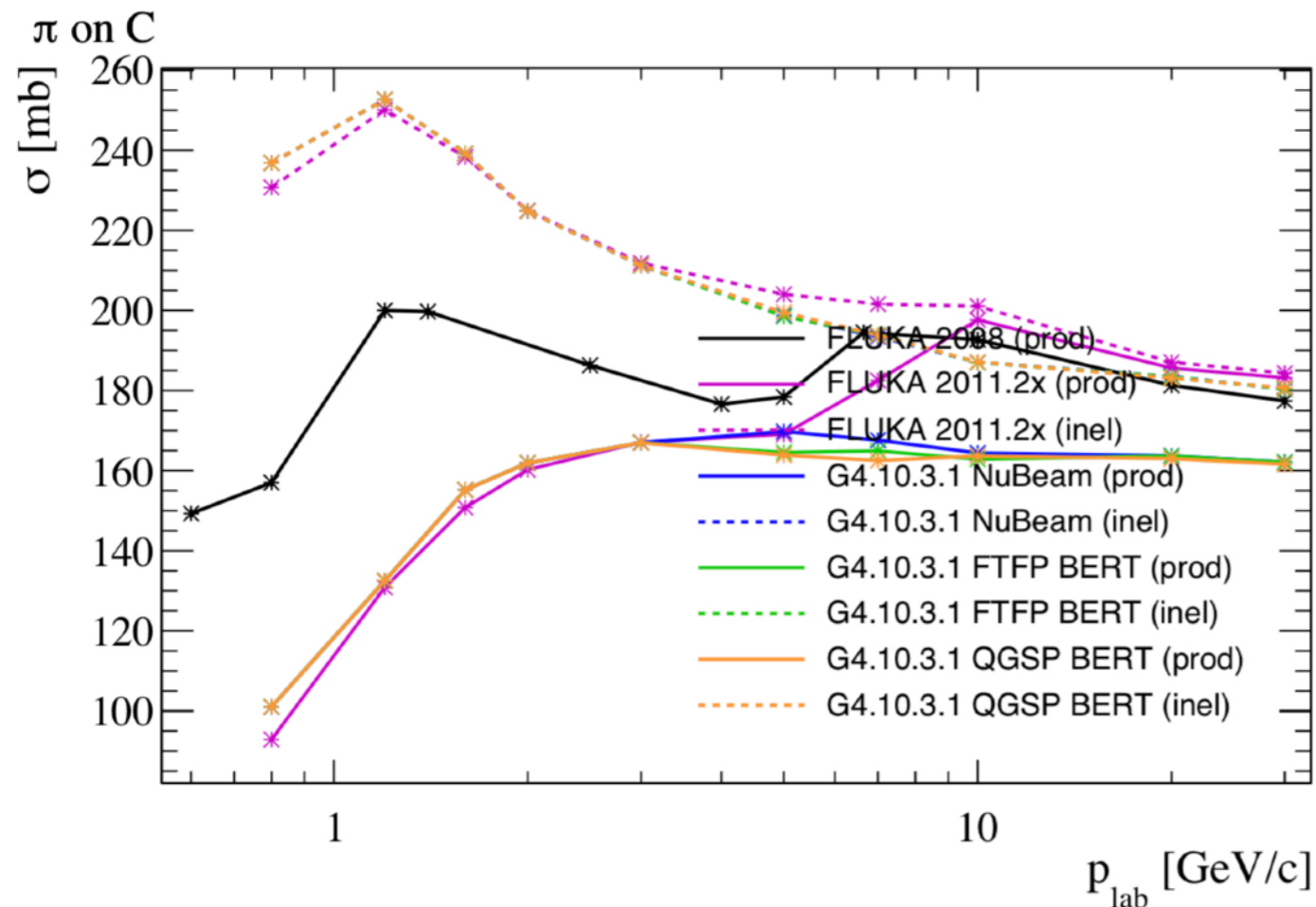
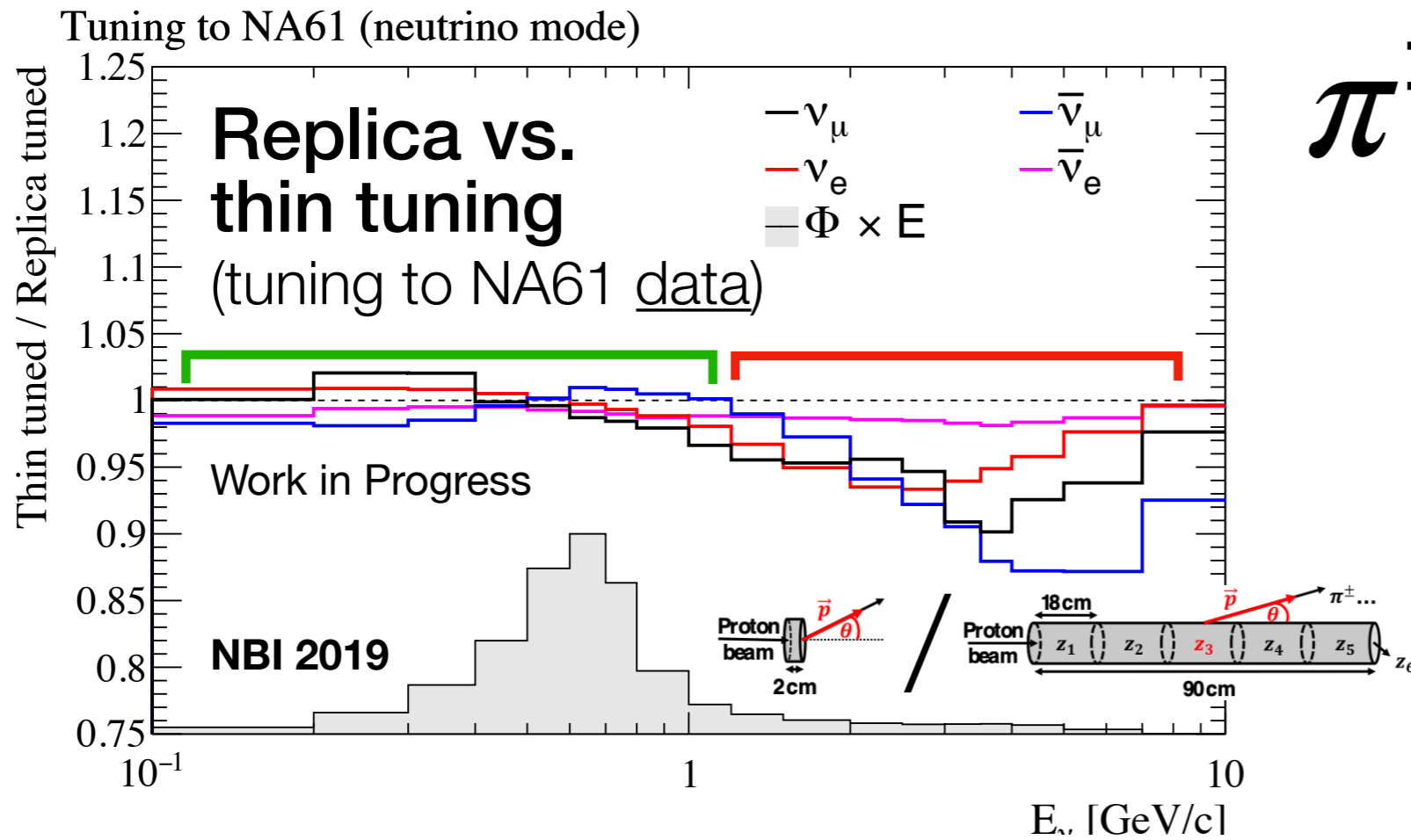
	π^+	π^-	K^0	K^+	K^-	n	p	\bar{p}	Λ	Σ	Channel	Flux-weighted fraction [%]
$\pi^+ \rightarrow$	20.74	5.15	1.98	1.28	0.26	0.29	0.35	0.00	0.11	0.16	$\pi^+ \rightarrow \pi^+$	20.74
$\pi^- \rightarrow$	8.99	6.00	1.38	0.58	0.21	0.30	0.15	0.00	0.07	0.08	$\pi^- \rightarrow \pi^+$	8.99
$K^0 \rightarrow$	2.49	0.64	3.69	2.08	0.65	0.03	0.02	0.00	0.03	0.04	$p \rightarrow \pi^+$	6.35
$K^+ \rightarrow$	2.70	0.42	2.73	4.73	0.02	0.02	0.02	0.00	0.01	0.01	$\pi^- \rightarrow \pi^-$	6.00
$K^- \rightarrow$	0.76	0.26	0.79	0.02	0.67	0.01	0.00	0.00	0.02	0.02	$\pi^+ \rightarrow \pi^-$	5.15
$n \rightarrow$	0.90	1.12	0.06	1.76	0.27	0.11	2.73	0.00	0.52	0.56	$K^+ \rightarrow K^+$	4.73
$p \rightarrow$	6.35	3.57	2.65	0.64	0.18	0.13	0.21	0.00	0.15	0.07	$K^0 \rightarrow K^0$	3.69
$\bar{p} \rightarrow$	0.47	0.15	0.02	0.01	0.00	0.00	0.00	0.00	0.00	0.00	$p \rightarrow \pi^-$	3.57
$\Lambda \rightarrow$	1.87	0.20	0.47	0.10	0.10	0.14	0.05	0.00	0.68	0.63	$n \rightarrow p$	2.73
$\Sigma \rightarrow$	0.44	0.04	0.15	0.03	0.02	0.03	0.02	0.00	0.16	0.33	$K^+ \rightarrow K^0$	2.73
$\bar{n} \rightarrow$	0.73	0.15	0.03	0.01	0.01	0.00	0.00	0.00	0.00	0.00	Other	35.31

SK, Anti-neutrino mode

largest contributions:

	π^+	π^-	K^0	K^+	K^-	n	p	\bar{p}	Λ	Σ	Channel	Flux-weighted fraction [%]
$\pi^+ \rightarrow$	8.15	12.06	1.87	0.68	0.43	0.33	0.26	0.00	0.39	0.27	$\pi^- \rightarrow \pi^-$	15.07
$\pi^- \rightarrow$	3.70	15.07	1.32	0.33	0.39	0.37	0.12	0.00	0.25	0.24	$\pi^+ \rightarrow \pi^-$	12.06
$K^0 \rightarrow$	0.63	2.31	3.51	1.14	1.08	0.03	0.01	0.00	0.10	0.08	$\pi^+ \rightarrow \pi^+$	8.15
$K^+ \rightarrow$	0.61	1.70	2.49	2.31	0.03	0.03	0.02	0.00	0.03	0.02	$p \rightarrow \pi^+$	4.29
$K^- \rightarrow$	0.17	1.30	0.81	0.01	1.41	0.01	0.00	0.00	0.07	0.06	$p \rightarrow \pi^-$	4.04
$n \rightarrow$	0.72	1.26	0.06	1.07	0.40	0.10	2.25	0.00	1.97	1.86	$\pi^- \rightarrow \pi^+$	3.70
$p \rightarrow$	4.29	4.04	2.65	0.53	0.17	0.12	0.20	0.00	0.15	0.08	$K^0 \rightarrow K^0$	3.51
$\bar{p} \rightarrow$	0.13	0.70	0.02	0.00	0.01	0.00	0.00	0.00	0.00	0.00	$\Lambda \rightarrow \Lambda$	3.48
$\Lambda \rightarrow$	0.19	1.80	0.45	0.03	0.28	0.18	0.03	0.00	3.48	1.61	$p \rightarrow K^0$	2.65
$\Sigma \rightarrow$	0.05	0.41	0.14	0.01	0.07	0.04	0.01	0.00	0.83	0.61	$K^+ \rightarrow K^0$	2.49
$\bar{n} \rightarrow$	0.19	0.54	0.02	0.00	0.01	0.00	0.00	0.00	0.01	0.00	Other	40.57

π^\pm prod. xsec



- Tuning using replica data and thin data only gives $\sim 10\%$ difference in flux prediction at high-energies
 - Possibly caused by π^\pm prod. xsec on C, the FLUKA 2011 values seem to be tuned to Denisov et al., which comparing to the recent NA61 30 GeV/c measurement is prod+QE, not prod xsec.
 - Fits to replica data actually suggest that one needs to move this more than just the Denisov/NA61 xsec difference at 30 GeV. Would be good to have measurements between 5~20 GeV
- (below 5 GeV the MC generators are tuned to Vlasov et al?)

Summary

- Better understanding of flux important for success of J-PARC neutrino program (T2K, many near detectors, IWCD, HK, ...)
- Dominant systematic at low energy wrong-sign ν_μ flux from few-GeV π^\pm scattering inside horns (Al) and decay volume (Fe)
- Similarly for right-sign ν_μ + few-GeV K^\pm scattering above flux peak
- For wrong-sign ν_e neutral Kaon scattering
- Having the option to run accurate thin target tuning will also be important toward HK era until HK-target replica tuning becomes available

backup

Alternative tunes

FHC, thin target tuning

