

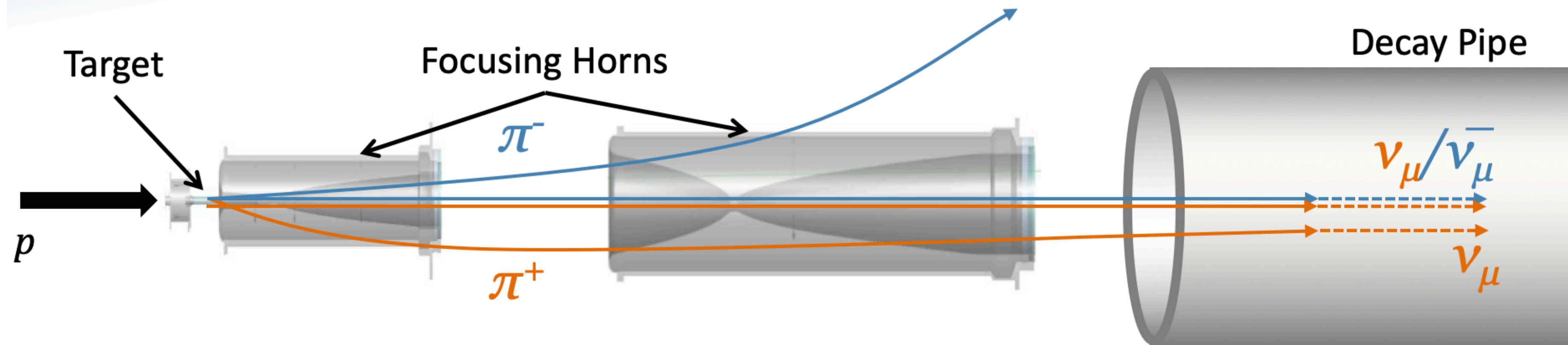
Fermilab long-baseline fluxes

Leo Aliaga

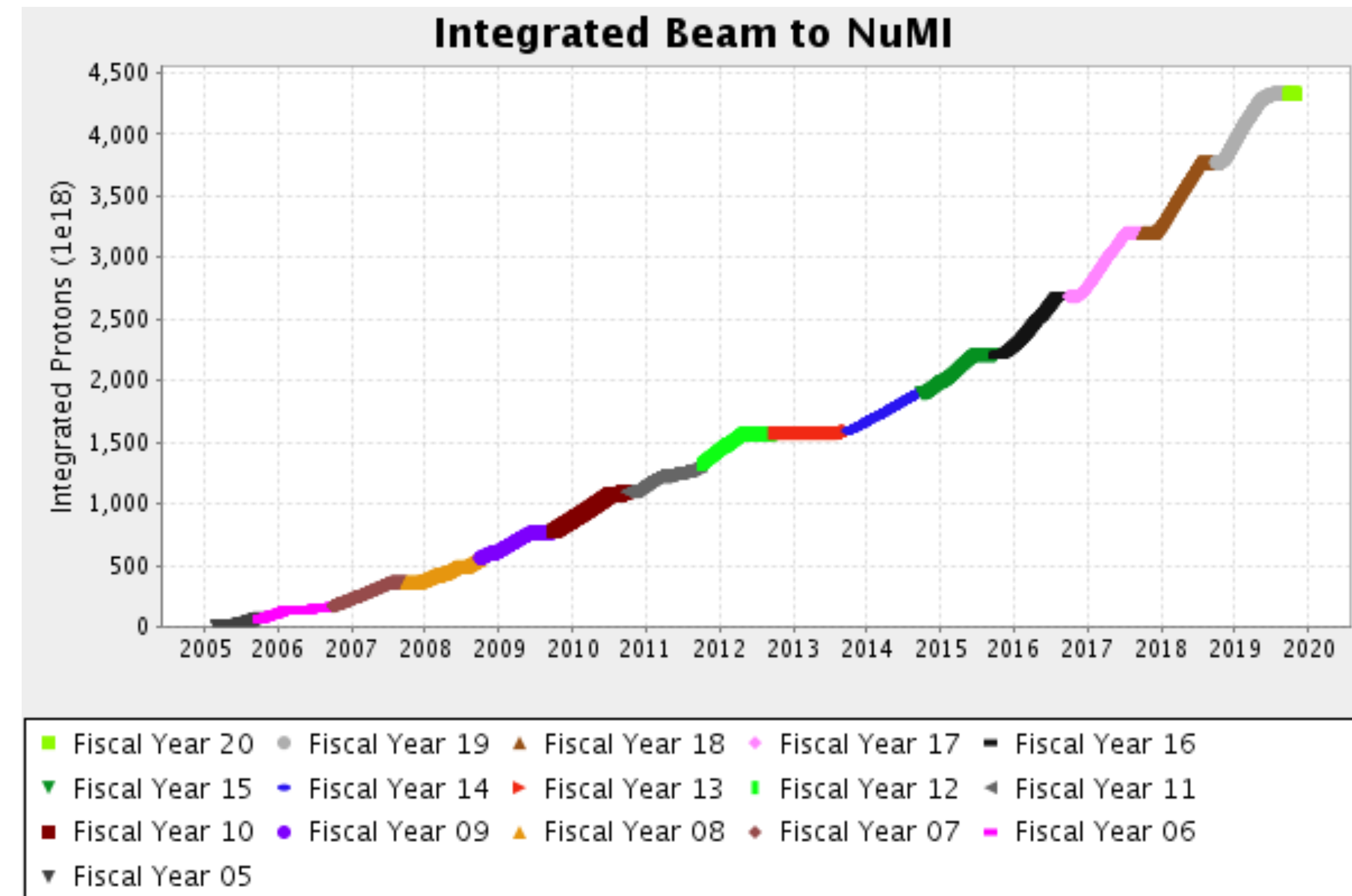
NA61/SHINE at Low Energy Workshop

December 10, 2020

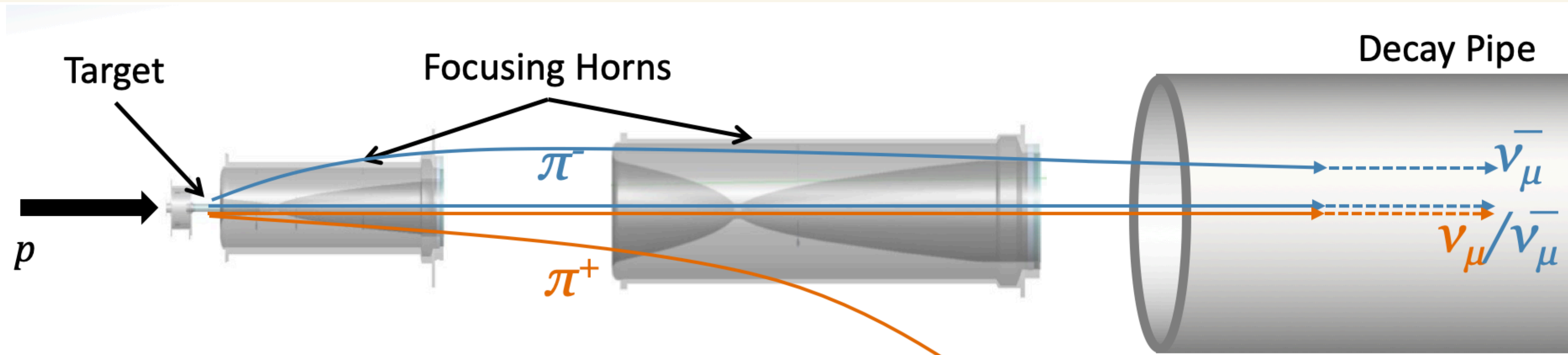
Neutrinos at the Main Injector



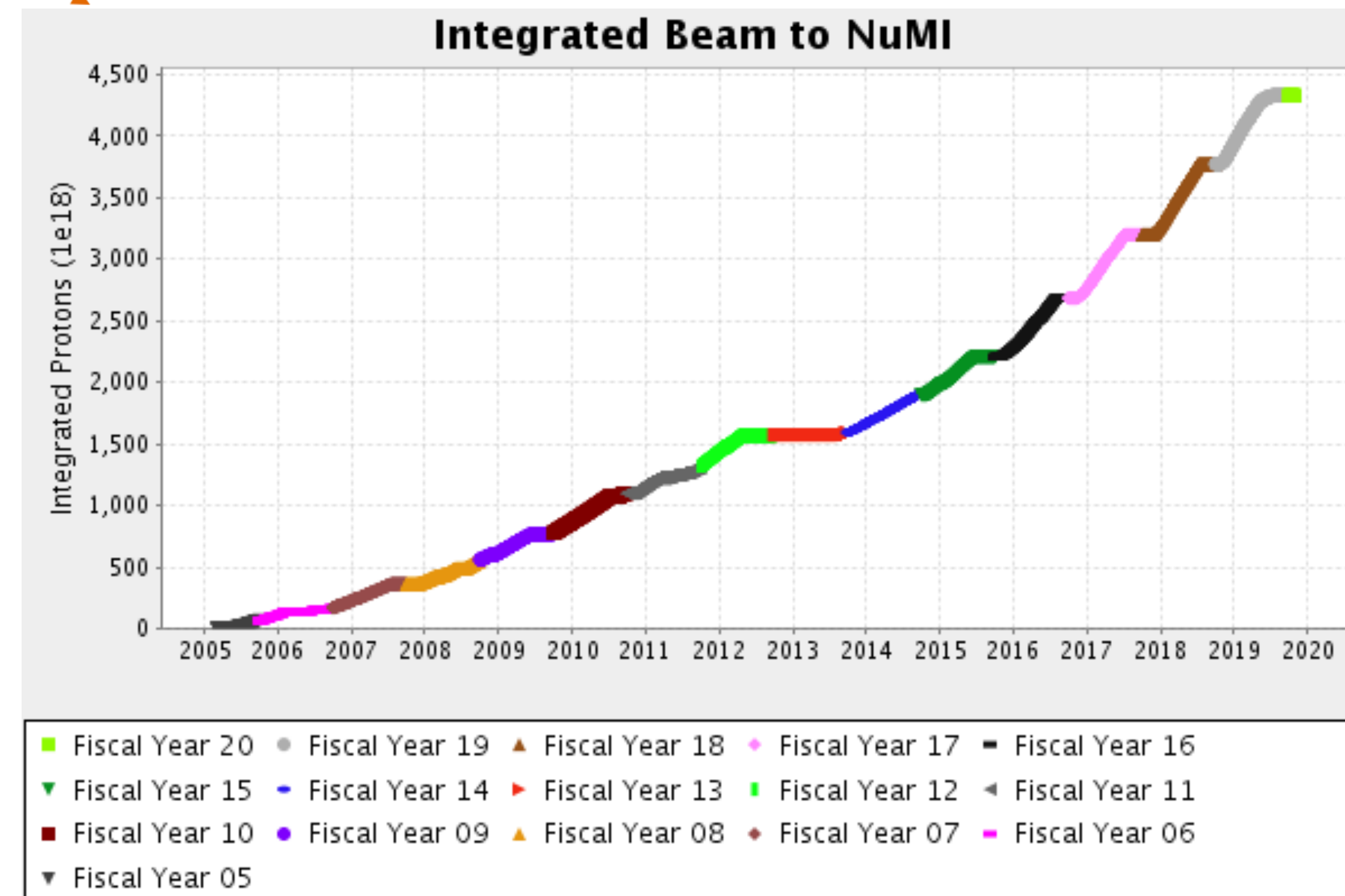
- » 120 GeV protons from Main Injector.
- » Graphite target
- » 2 Magnetic horns
- » Decay Pipe: 675 m filled with Helium.



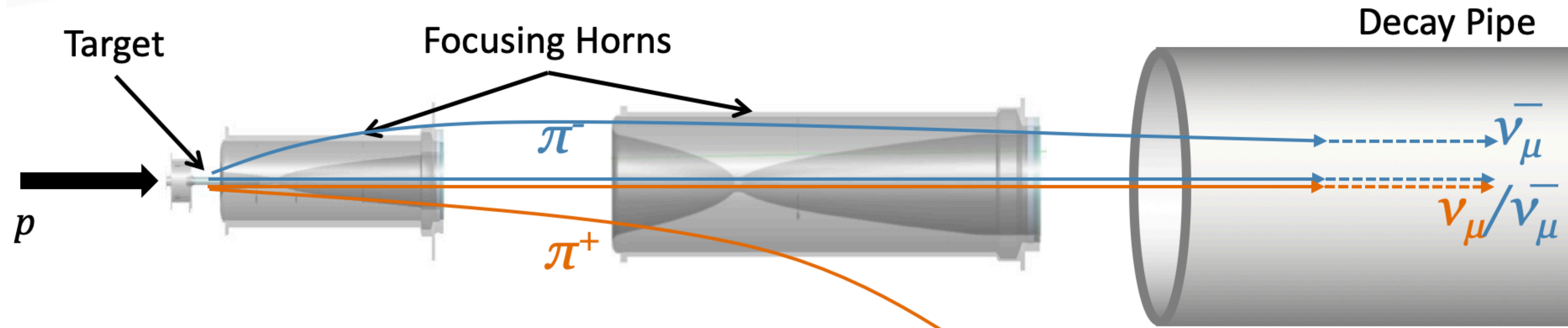
Neutrinos at the Main Injector



- » 120 GeV protons from Main Injector.
- » Graphite target
- » 2 Magnetic horns
- » Decay Pipe: 675 m filled with Helium.
- » It can run in neutrino or antineutrino modes



Neutrinos at the Main Injector

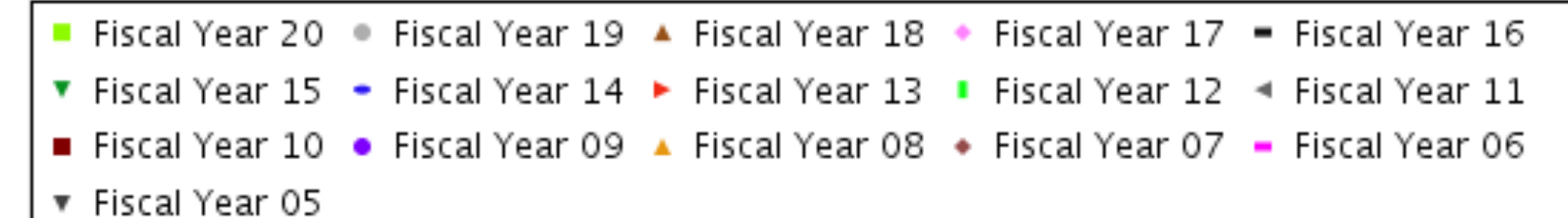
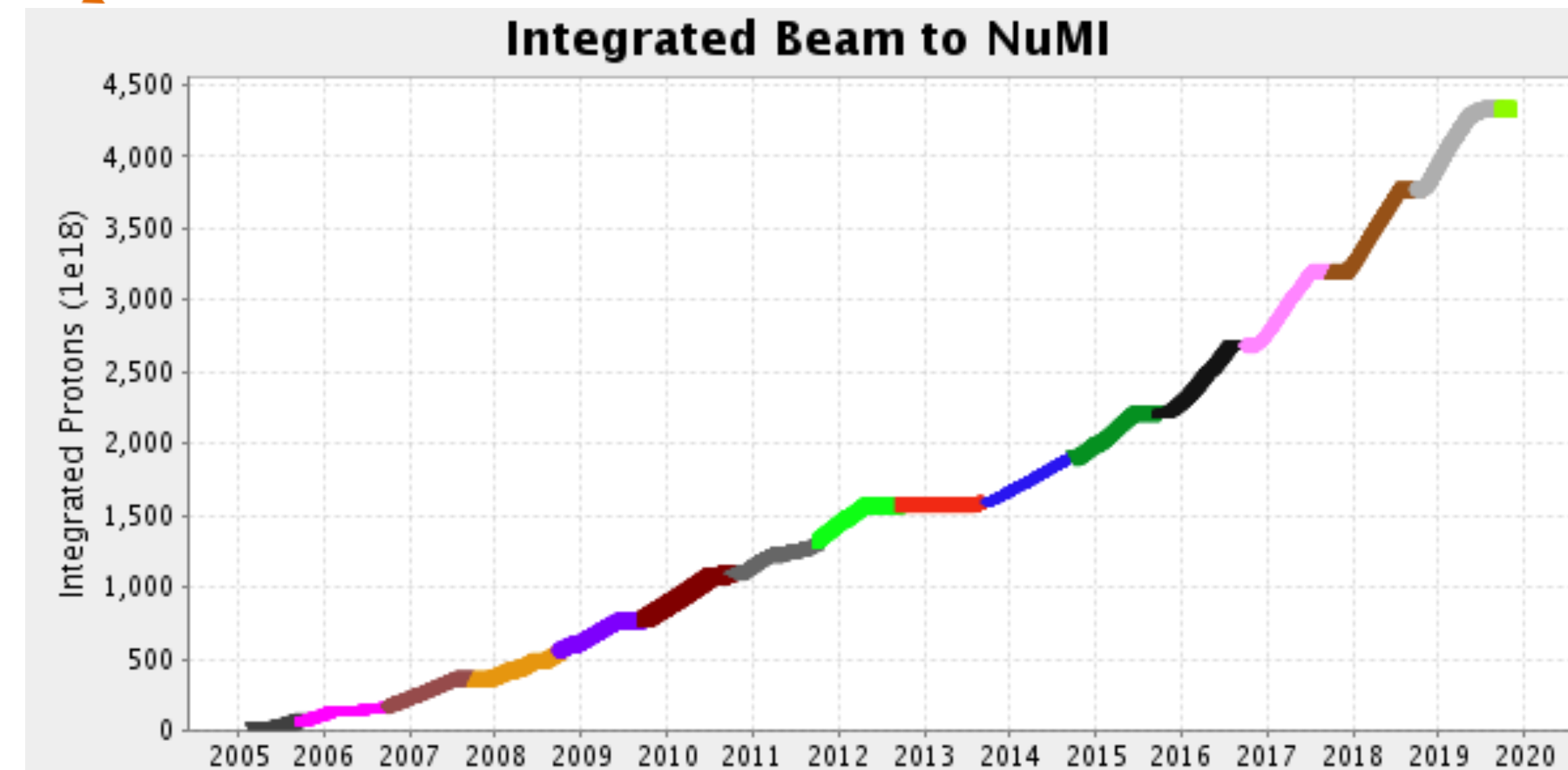


NuMI has run in two modes:

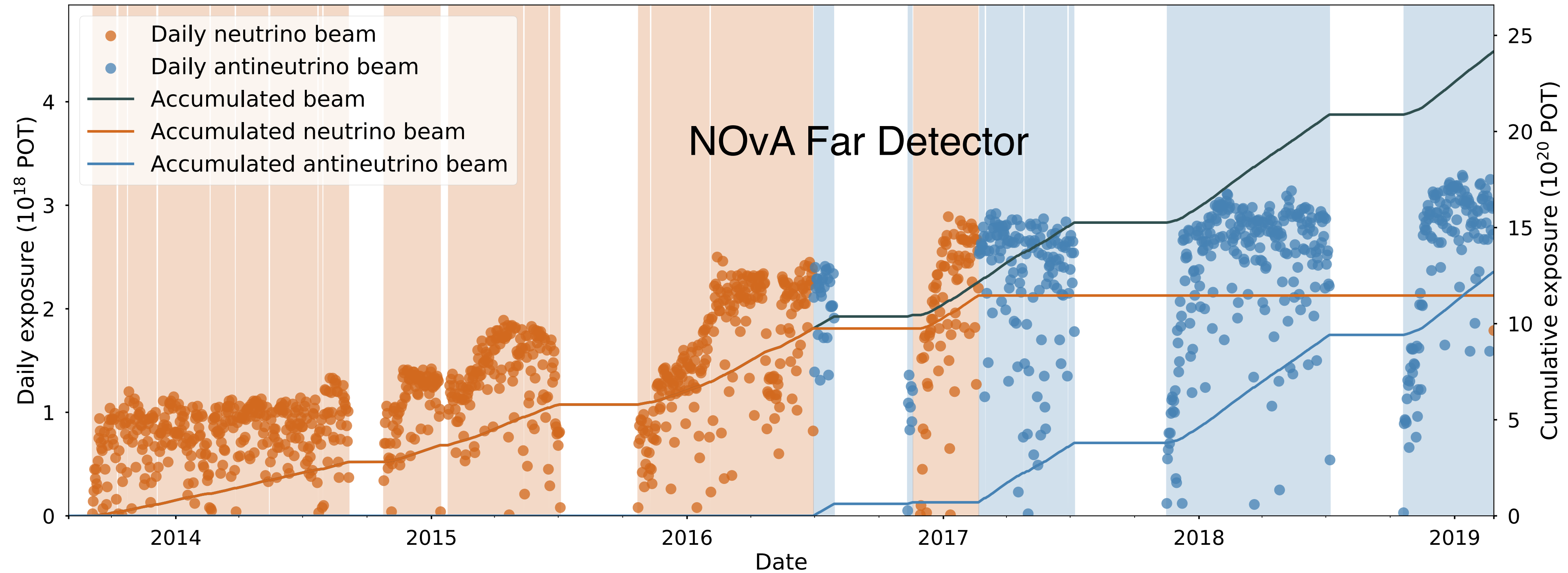
» Low energy	2005 - 2012	MINOS era
» Medium energy	2013 - present	NOvA era

» Currently running with 10 μ s spill every 1.33 s

» **Achieved 750 kW power since early 2019**



NOvA Exposure to NuMI Beam



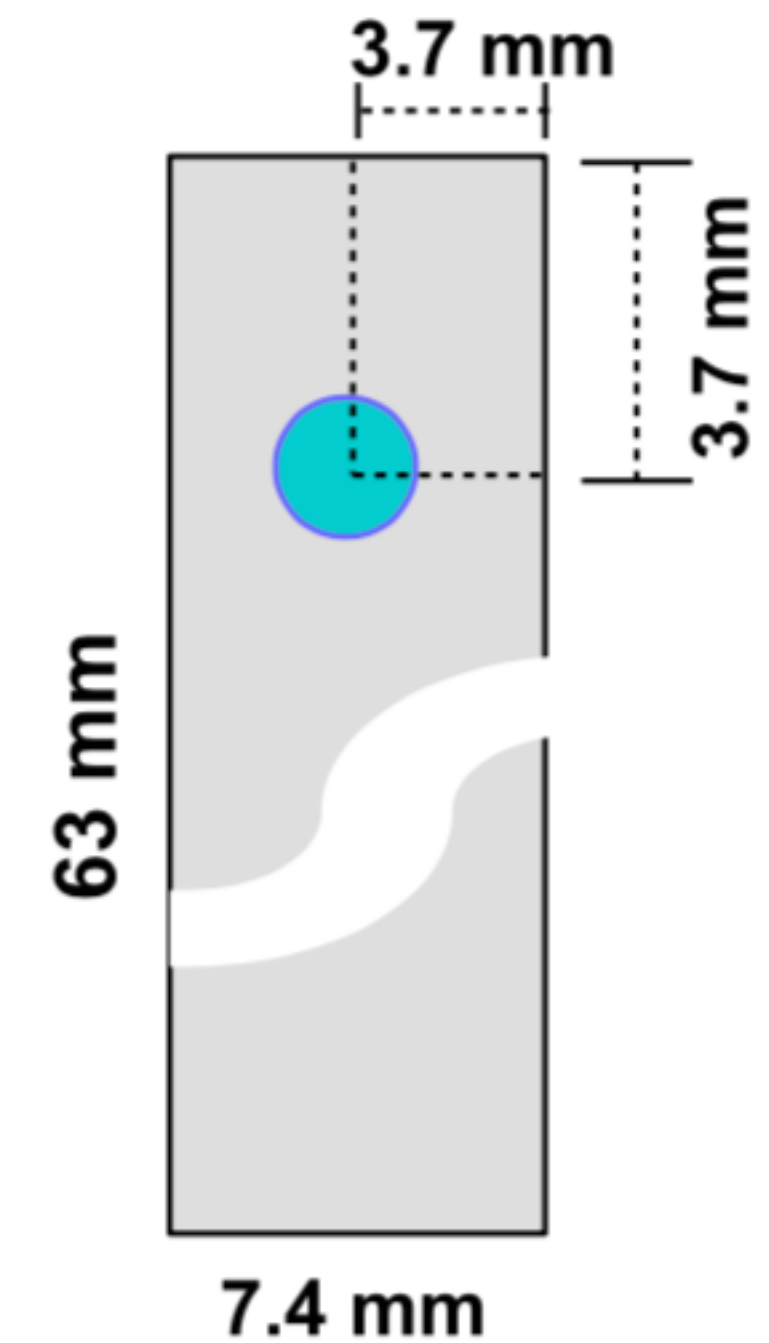
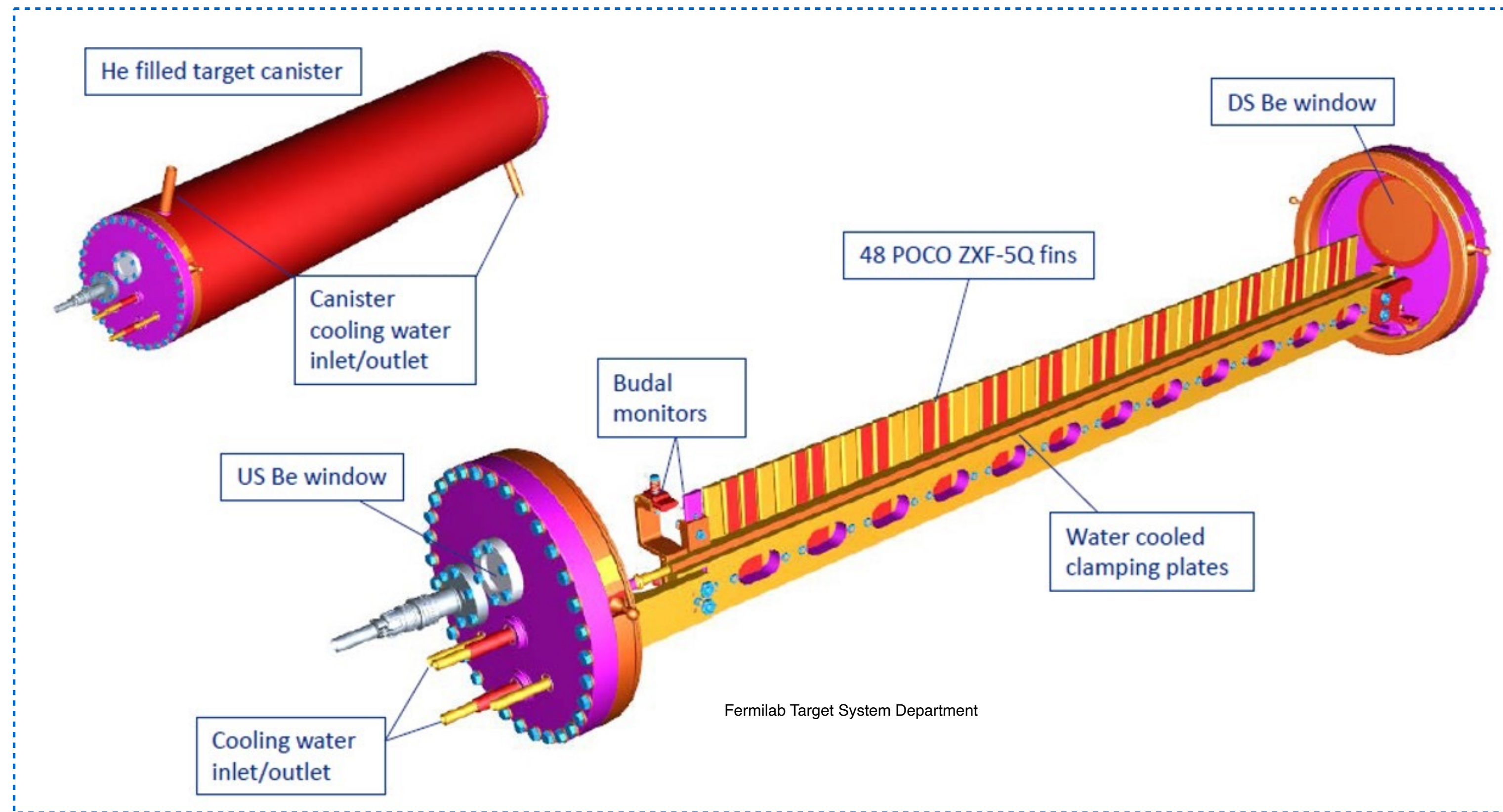
Collected data at NOvA ND to the present:

- » 11×10^{20} POT of neutrino-mode data
- » 11.8×10^{20} POT of antineutrino-mode data collected in the NOvA near detector so far.
- » More POT to come in the next years.

NuMI target

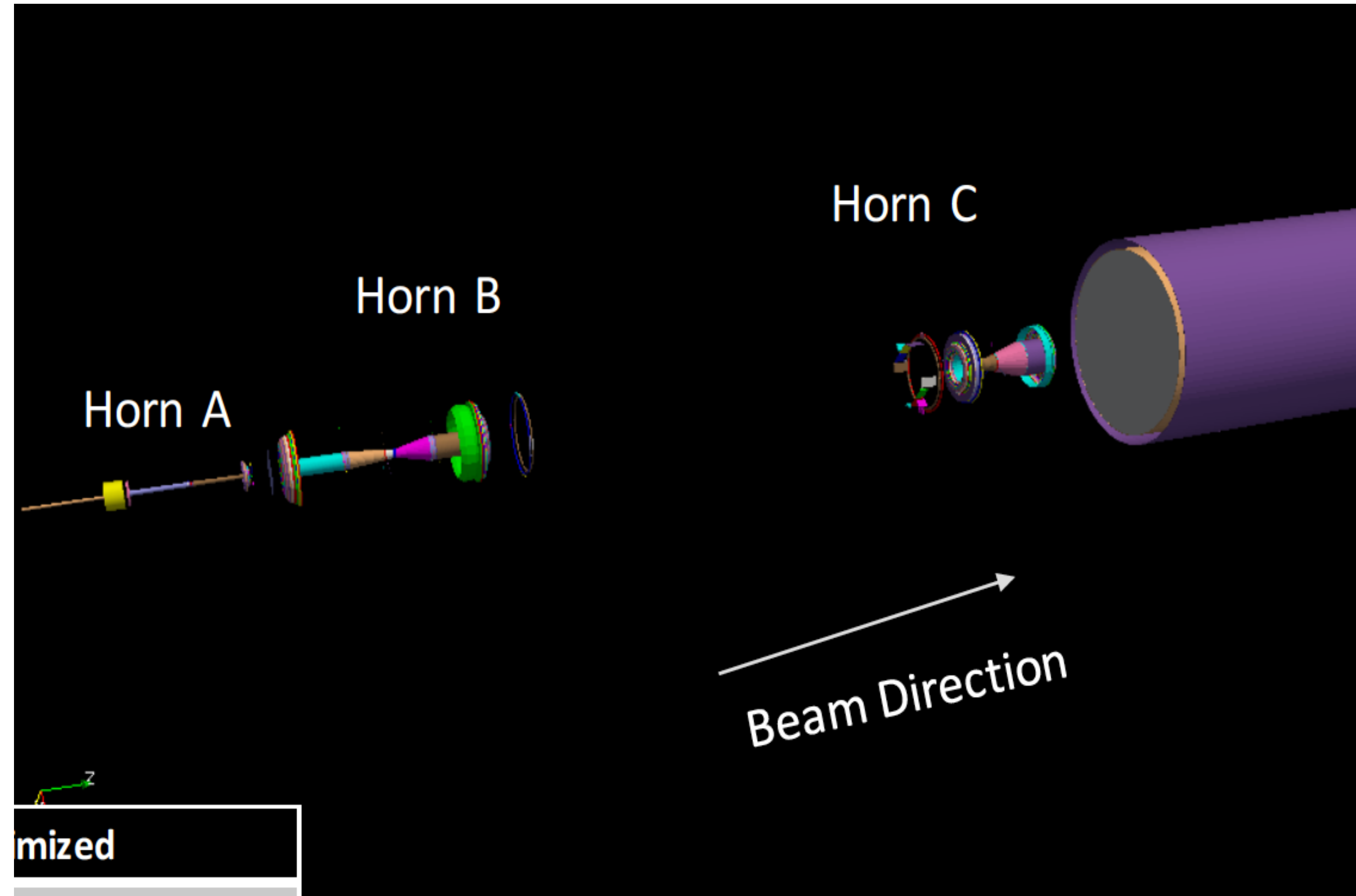
Rectangular graphite rod, segmented in rectangular pieces (“fins”).

50 fins in total: 1.2 m ($\sim 2.5 \lambda$).



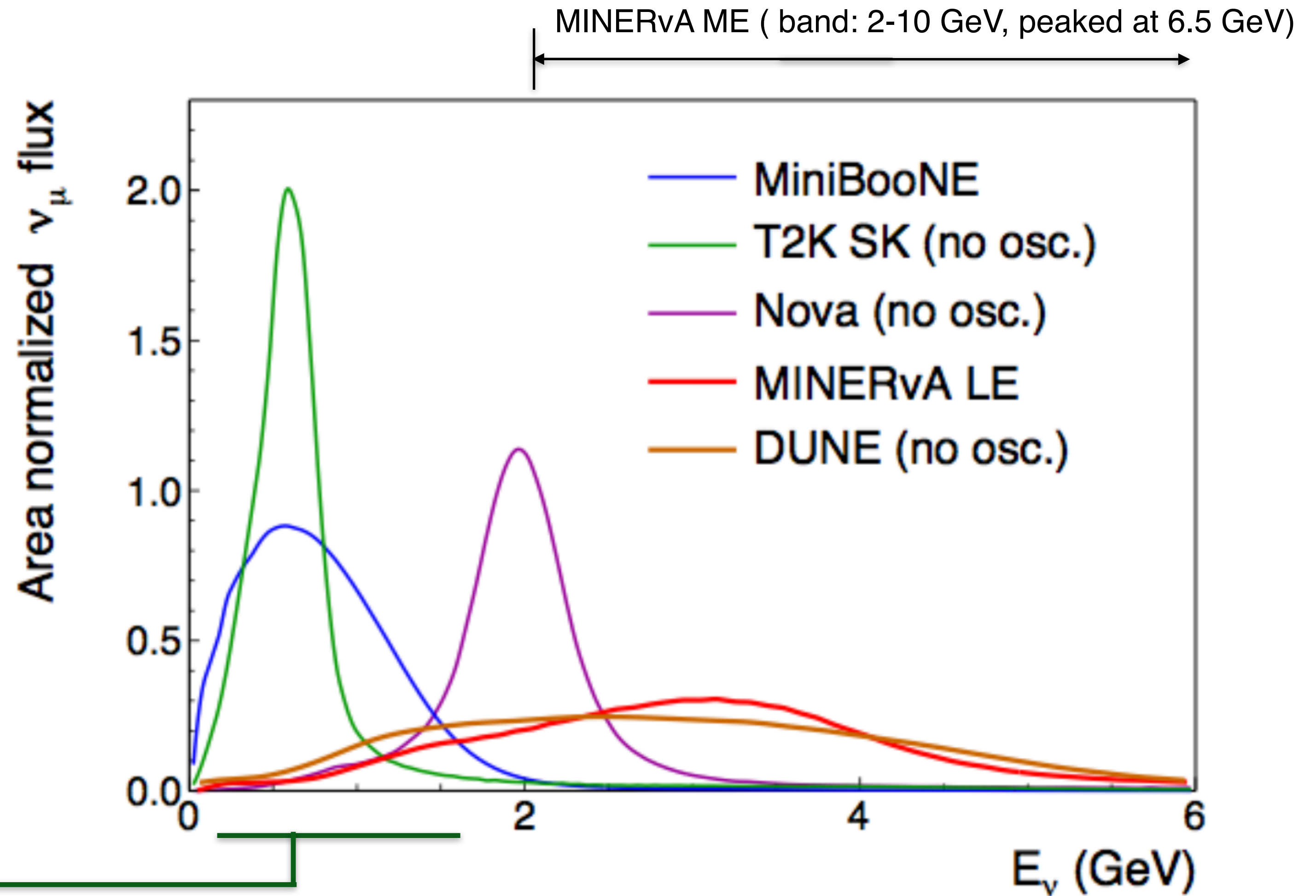
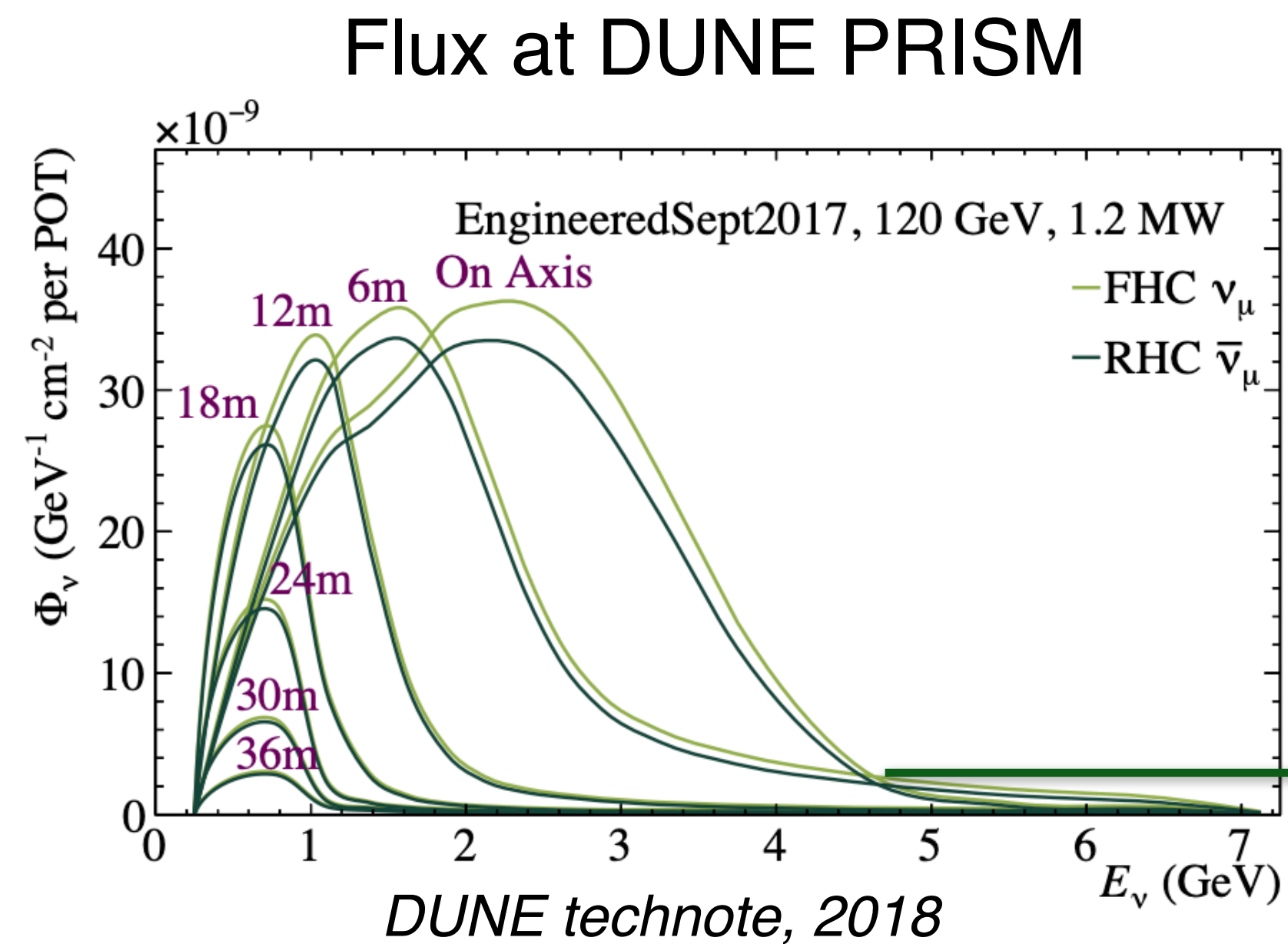
Long-Baseline Neutrino Facility

- » Primary proton beam in 60-120 GeV.
- » Initial 1.2 MW beam power, upgradable to 2.4 MW.
- » Graphite target
- » 3 Magnetic horns
- » Extended Decay pipe, He filled.
- » Experimental setup optimized for CP violation



Neutrino Fluxes

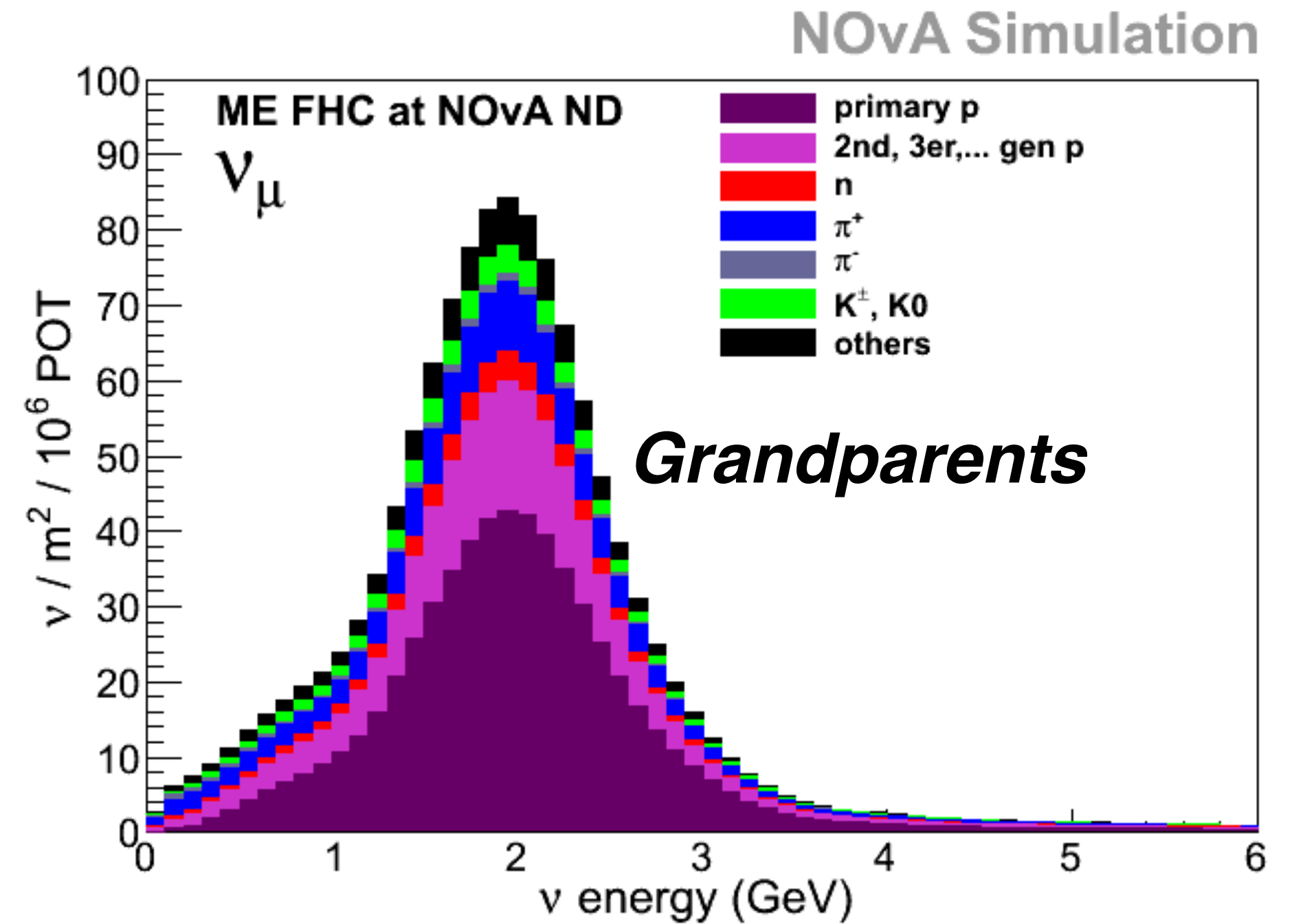
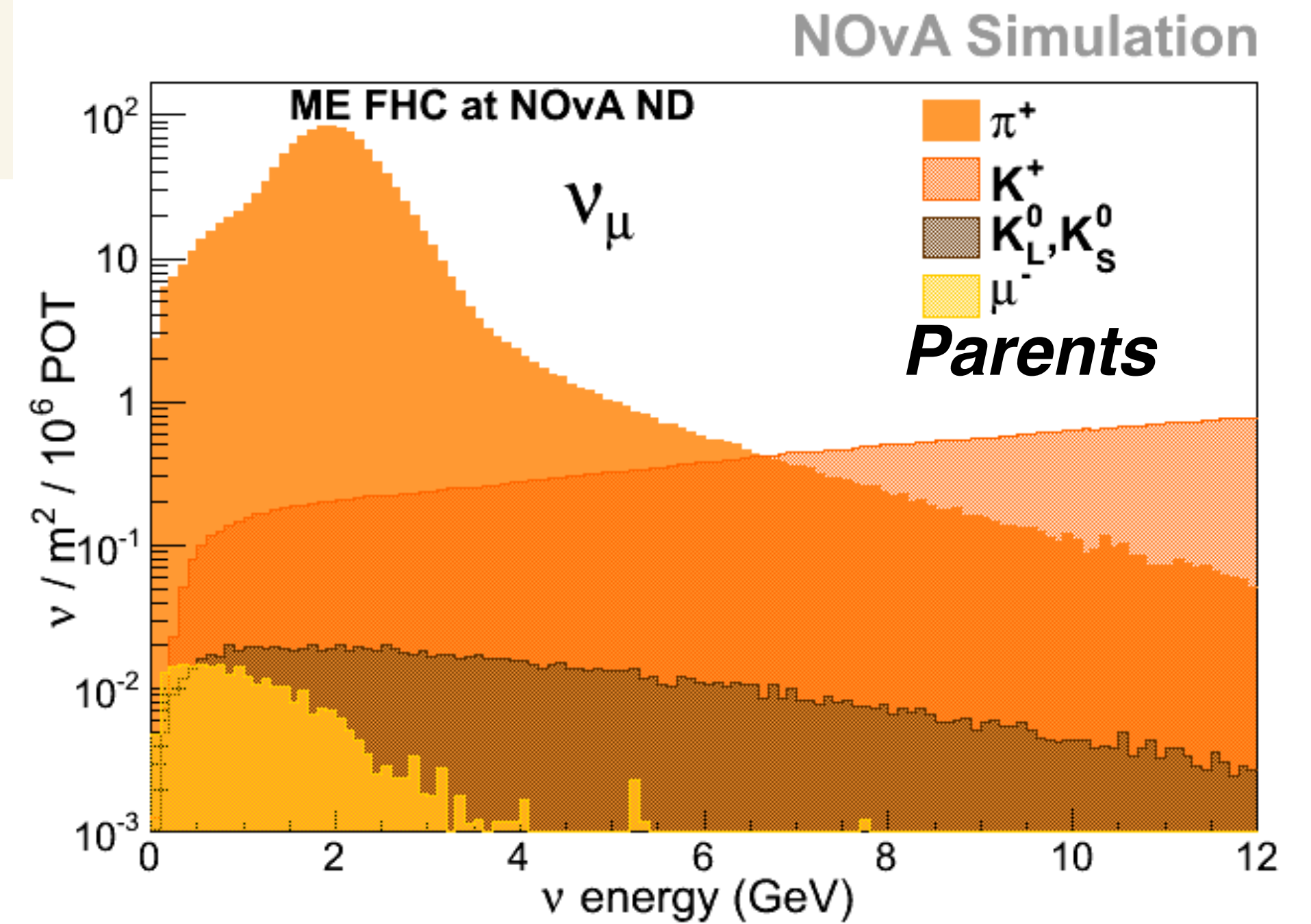
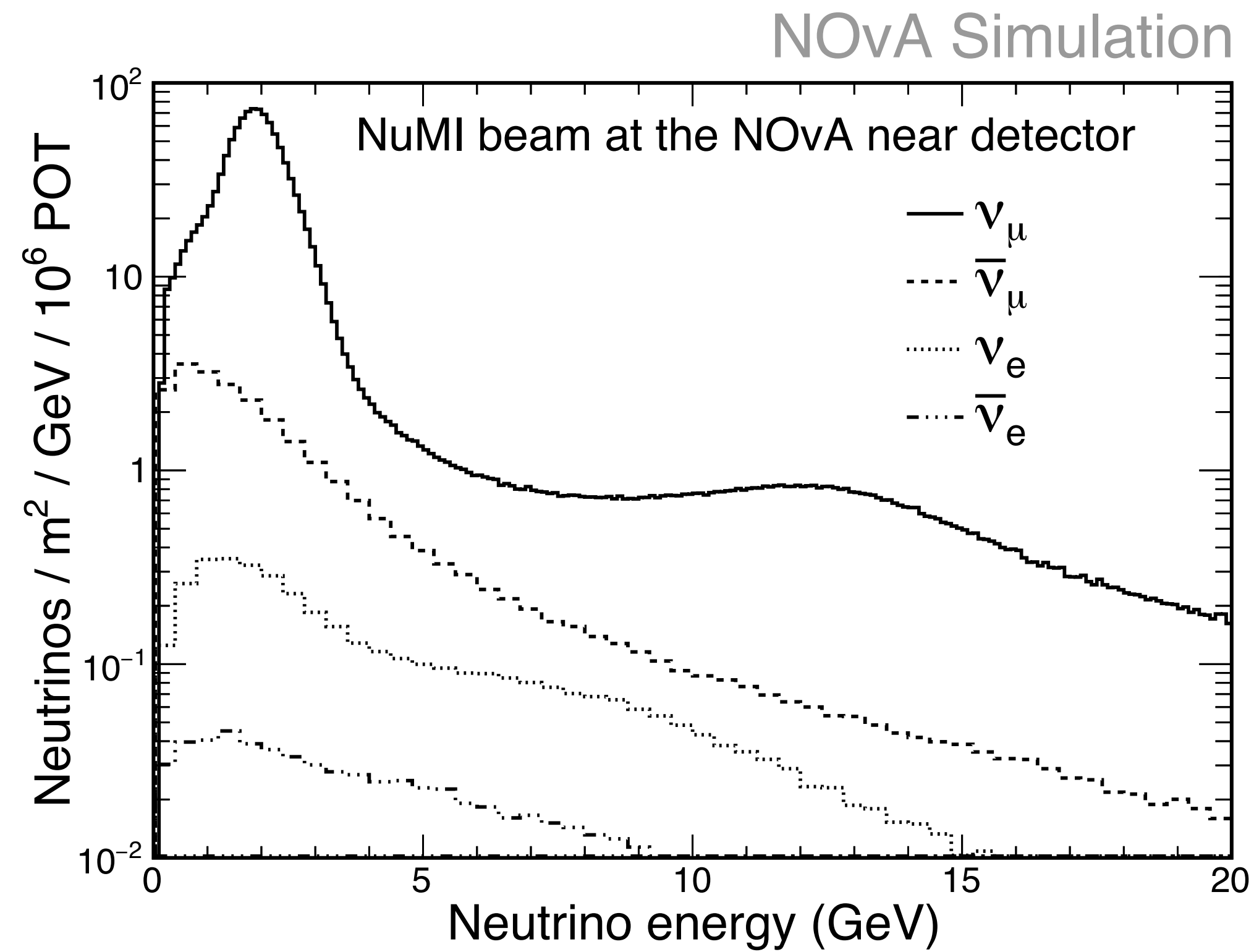
- » Not monochromatic: wide-band.
- » A-priori shape and normalization are determined from simulation and hadron production constraints.



Fermilab JETP Seminar, Phil Rodrigues 2015

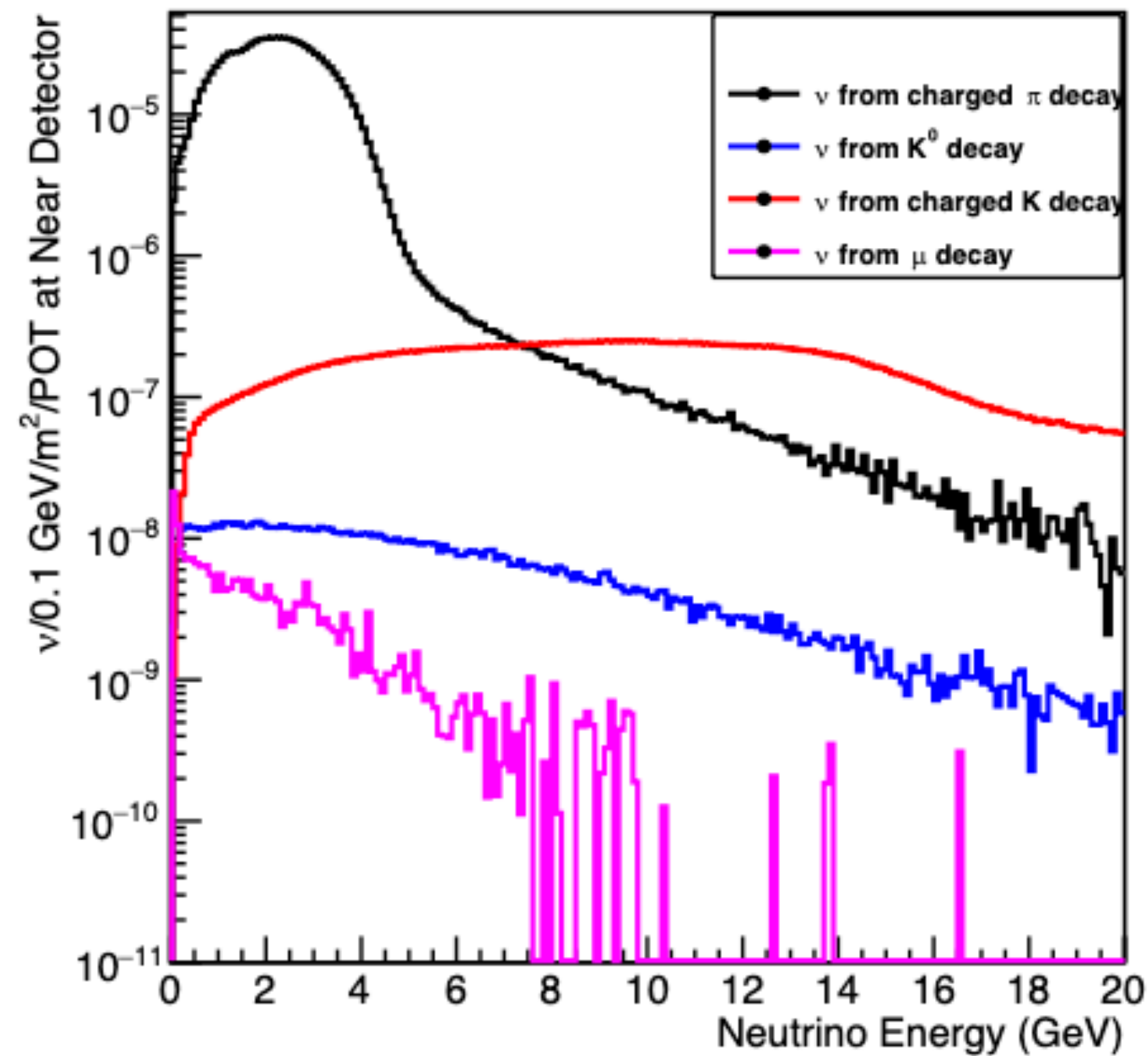
Neutrino flux at NOvA

» 96% pure ν_μ beam, 1% ν_e and $\bar{\nu}_e$

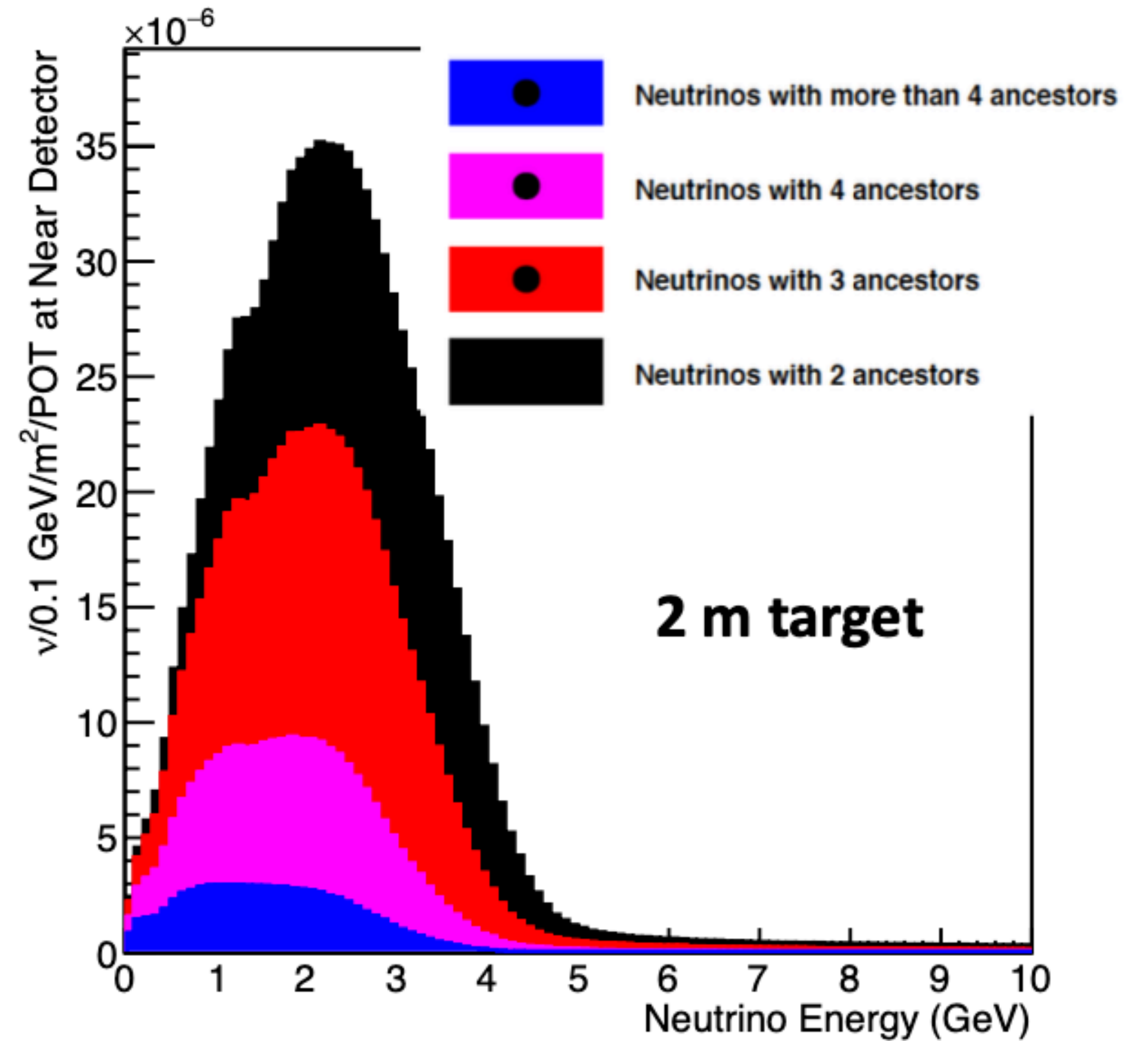


DUNE Flux

Muon Neutrino Flux at DUNE Near Detector (on axis)



Numu Flux at DUNE Near Detector

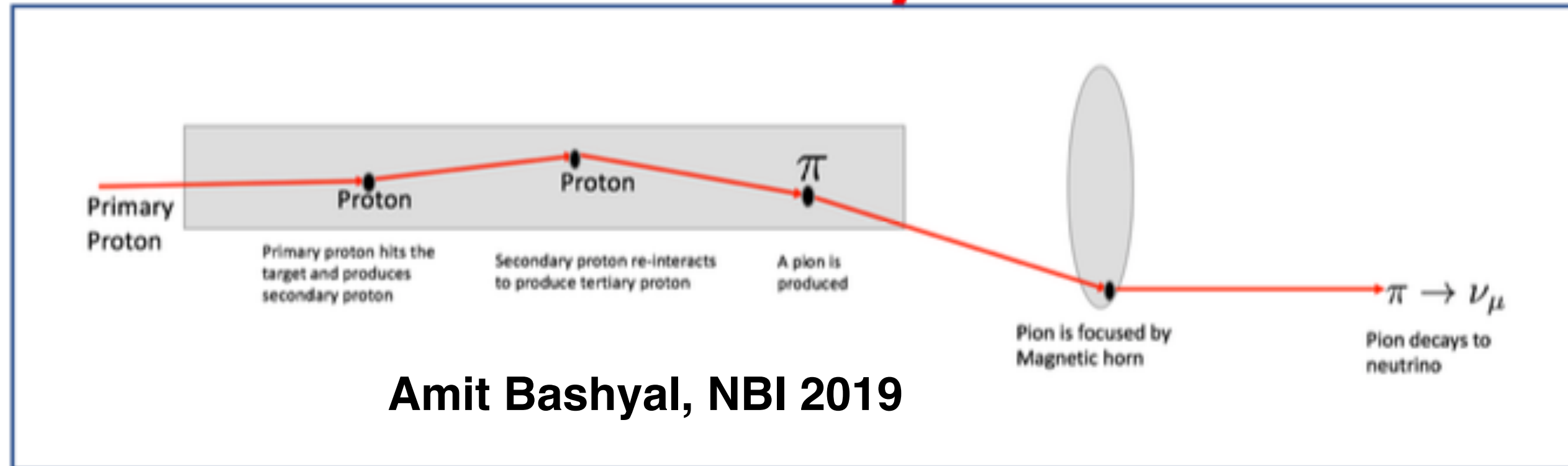


Amit Bashyal, NBI 2019

MINERvA strategy for predicting a-priori flux



Neutrino Flux Predictions for the NuMI Beam [Phys. Rev. D 94, 092005 (2016)]



Common framework is used for other long baseline neutrino experiments, such as NOvA and DUNE

- » The beamline simulation is G4NuMI (geant4 based).
- » The foundation is formed from **constraining** (correcting) the **interaction** and the **hadron production** with external measurements on **thick and thin targets**.
- » It corrects the yields given by the model to match the measured data. If not data is available we extend the data coverage or assign a well-educated guess.
- » It propagates the **underlying data correlations** to calculate the **neutrino interaction covariance** (“**multi-universe**” technique).
 - It requires the HP dataset bin-to-bin correlation

Hadron Production Corrections

Particle absorption

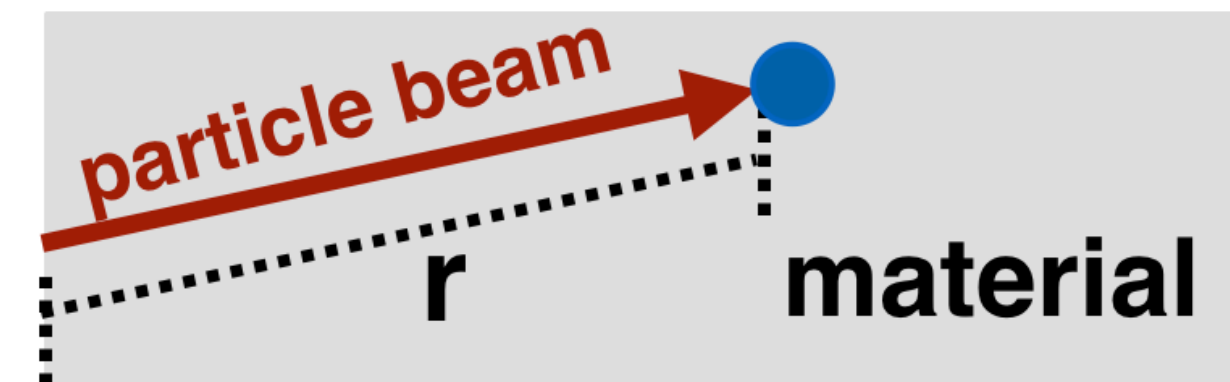
Neutrino Flux Predictions for the NuMI Beam [Phys. Rev. D 94, 092005 (2016)]

Particle absorption:

Interacting

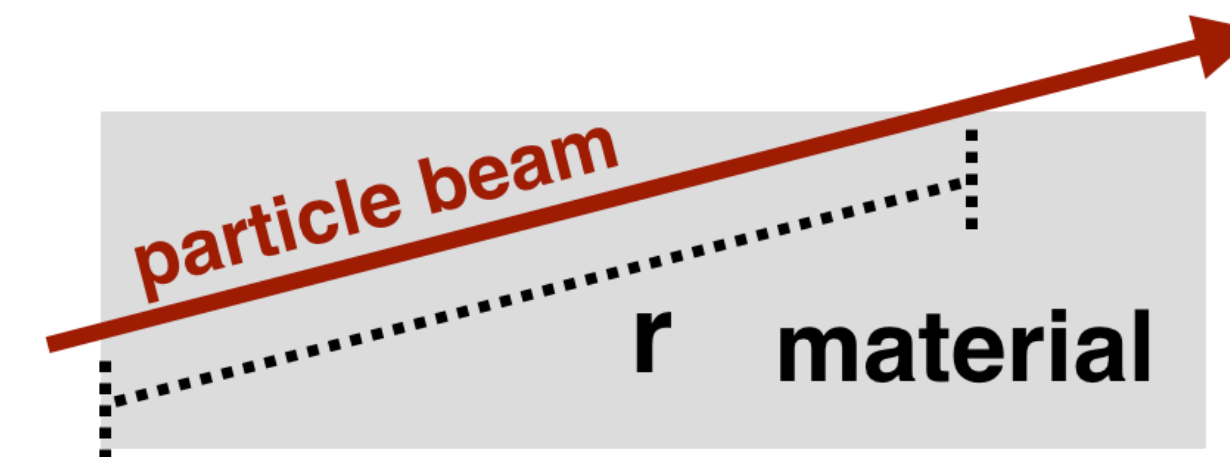
$$\text{correction}(r) = \frac{\sigma_{\text{Data}}}{\sigma_{\text{MC}}} e^{-r \frac{N_A \rho (\sigma_{\text{Data}} - \sigma_{\text{MC}})}{A}}$$

N_A : Avogadro Number, ρ : density, A : mass number



Not interacting

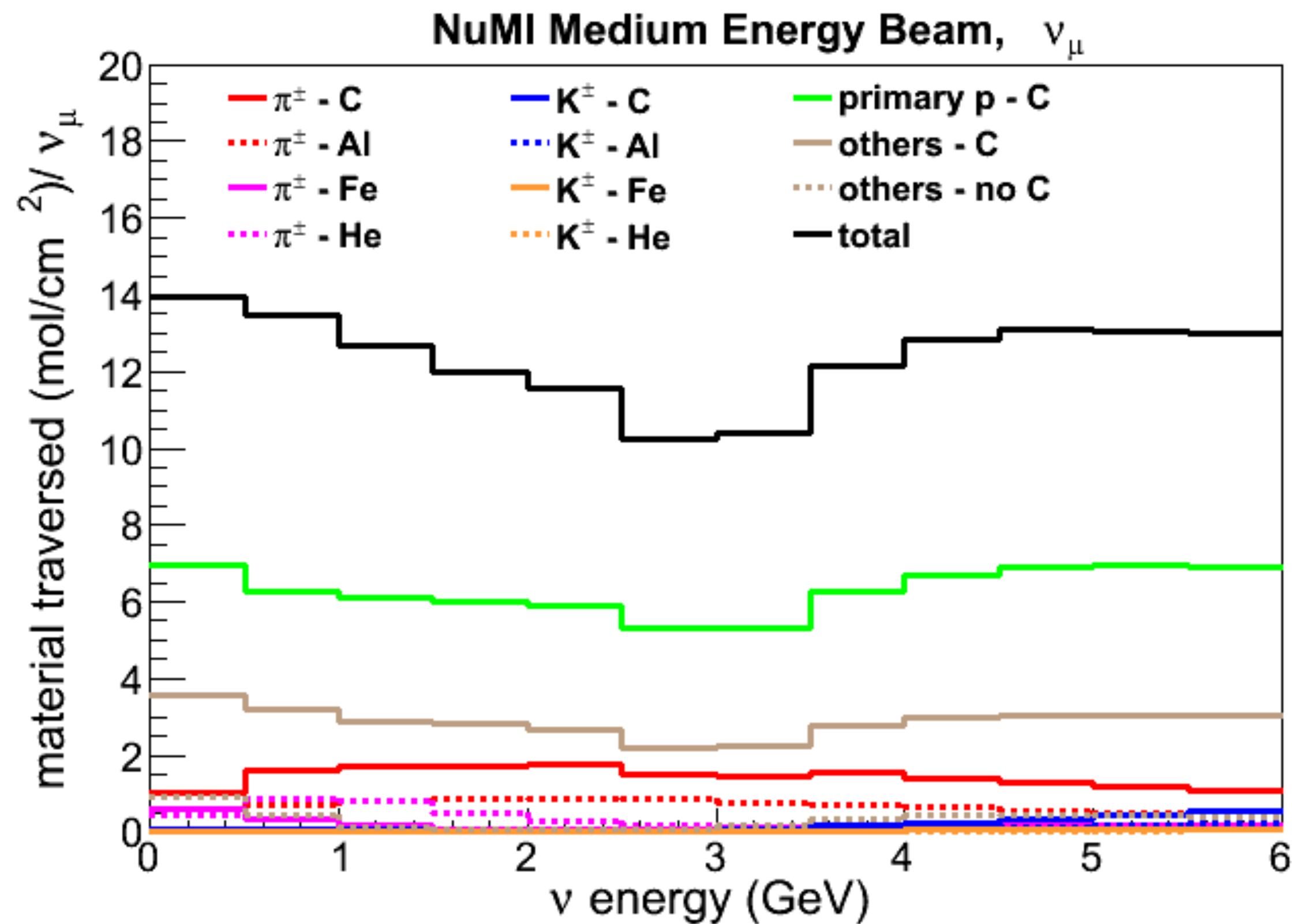
$$\text{correction}(r) = e^{-r \frac{N_A \rho (\sigma_{\text{Data}} - \sigma_{\text{MC}})}{A}}$$



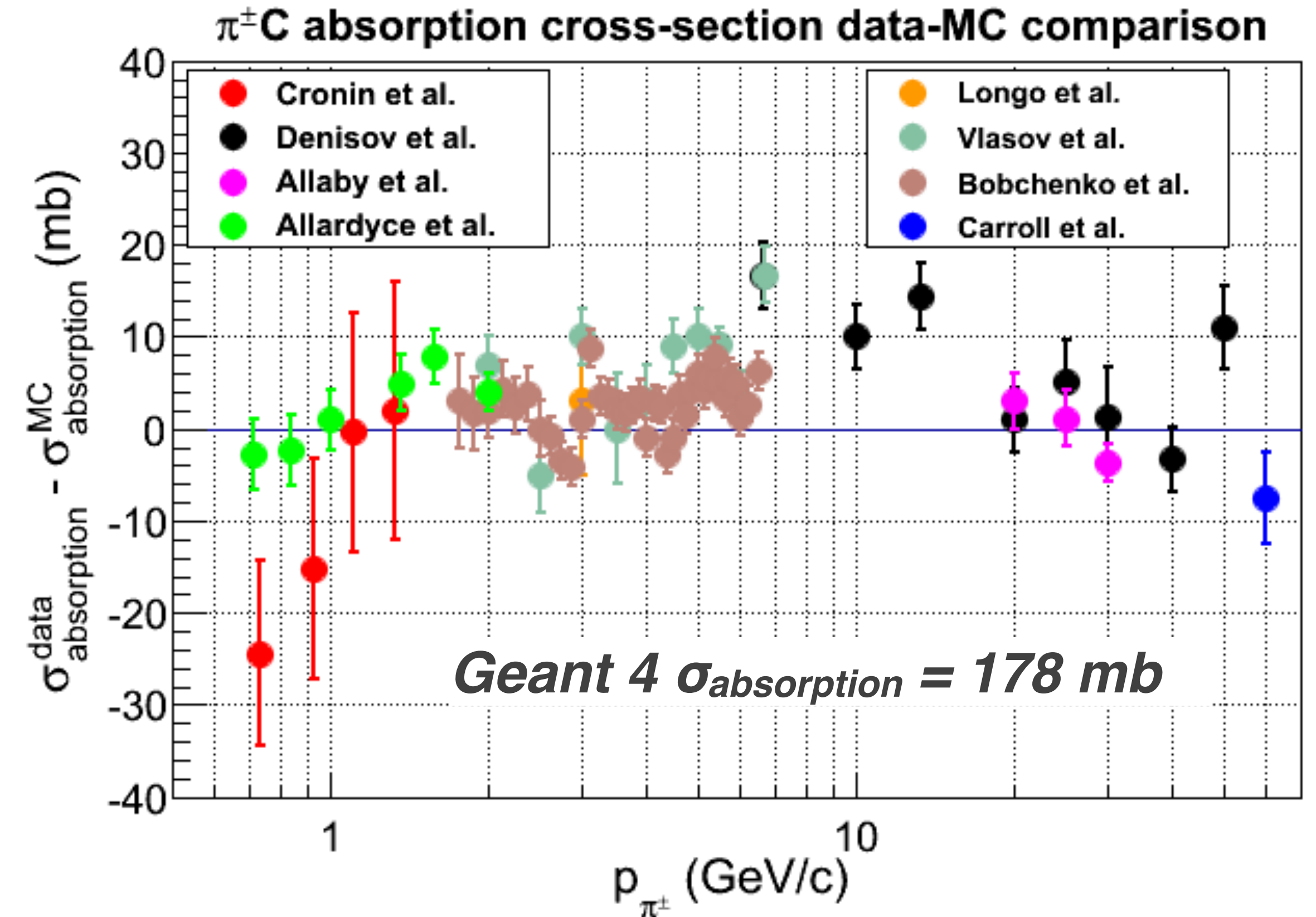
- Two variables are important here:
- The amount of material: $rN_A\rho/A$.
 - The σ_{Data} and σ_{MC} disagreement.

Absorption correction

$$\langle rN_{Ap}/A \rangle$$



- » C: $6 \text{ mol}/\text{cm}^2 \approx 40 \text{ cm}$
- » Al: $1 \text{ mol}/\text{cm}^2 \approx 10 \text{ cm}$
- » He: $1 \text{ mol}/\text{cm}^2 \approx 500 \text{ m}$

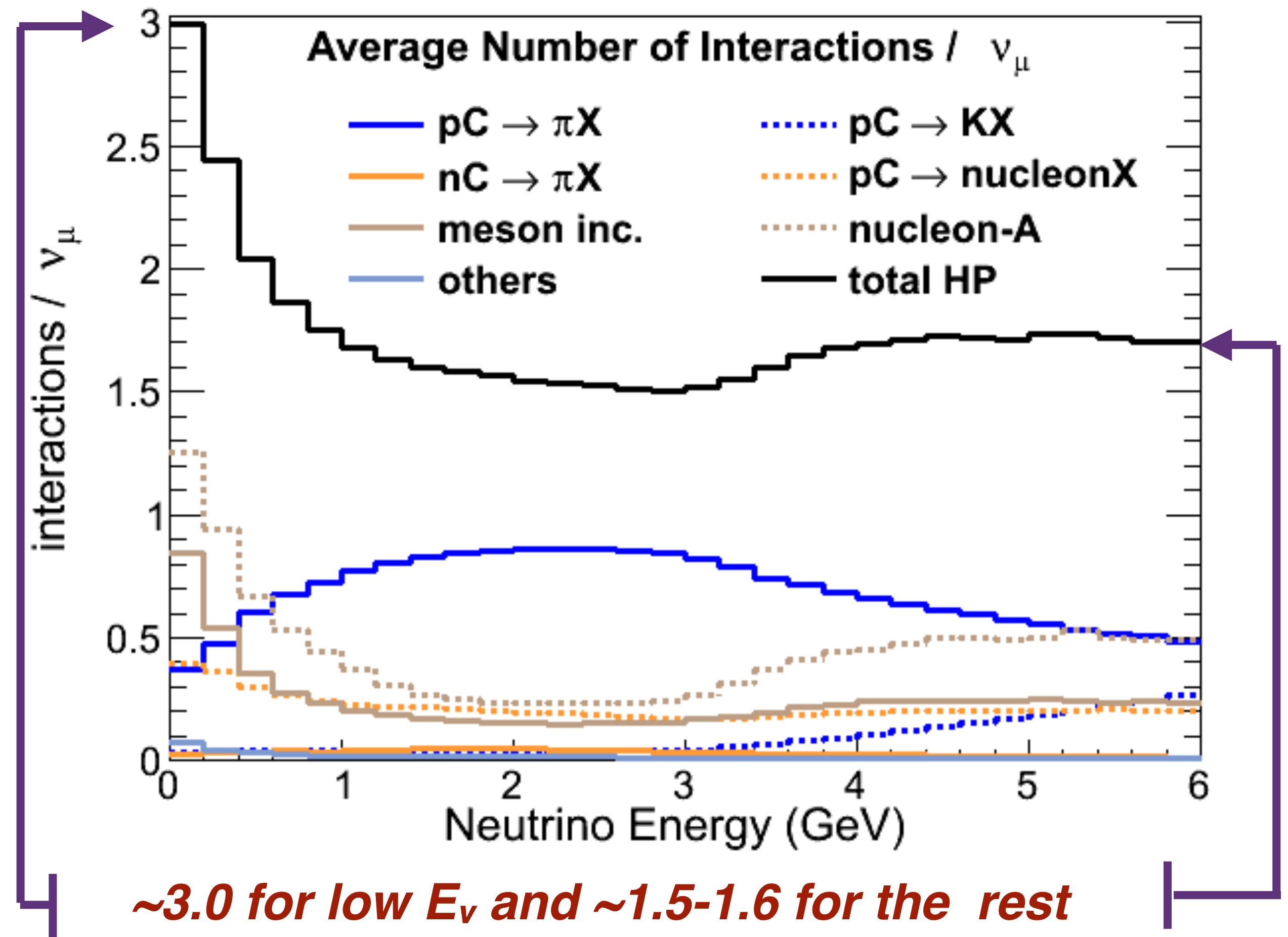


- » We apply 10 mb uncertainty ($\sim 5\%$)

Hadronic interactions

Average number of hadronic interactions per neutrino at NOvA ND

NOvA Simulation



Hadronic interactions

Average number of hadronic interactions per neutrino at NOvA ND

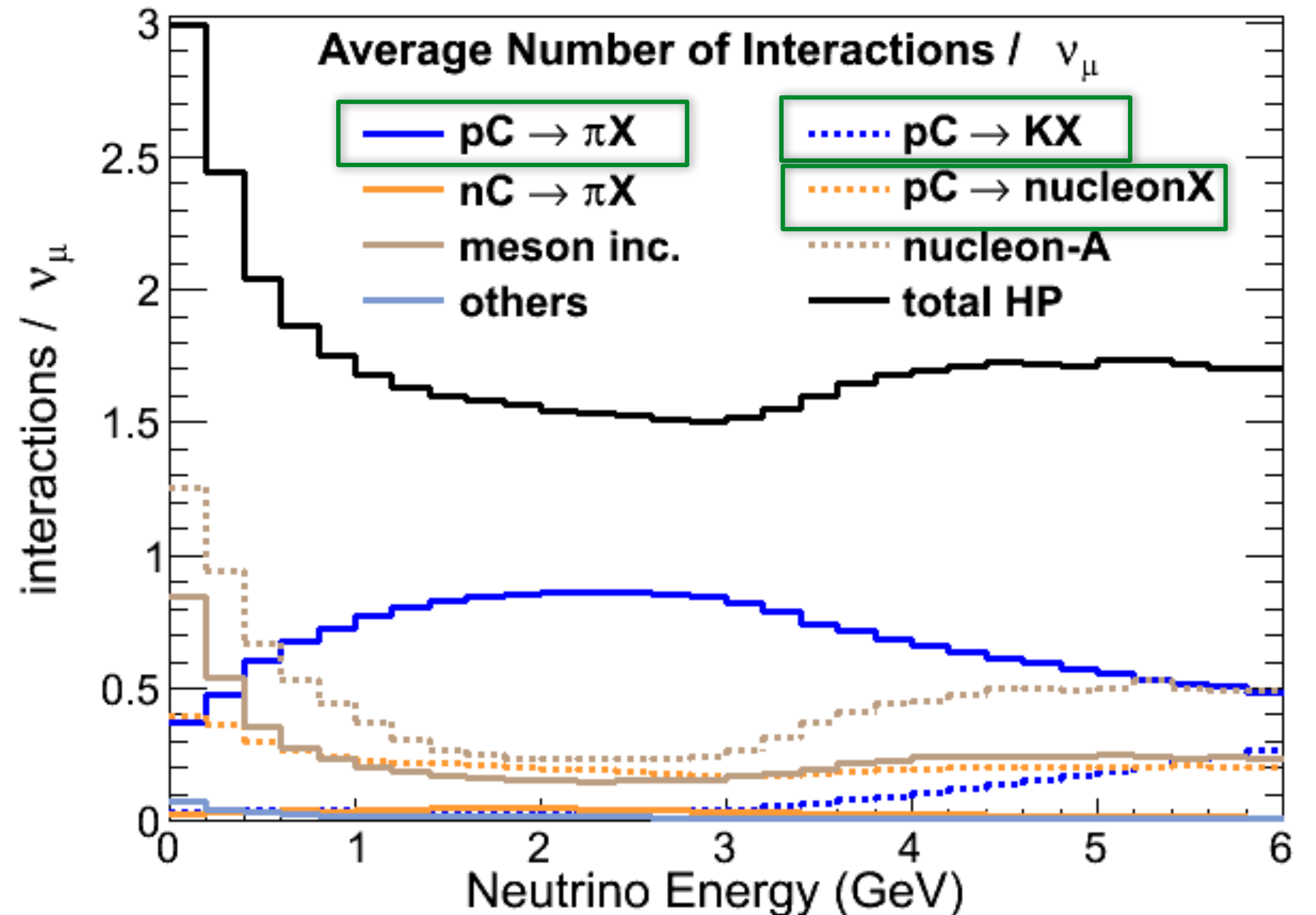
NOvA Simulation

Based directly on data:

- » Mostly NA49 proton on carbon producing charged pions, charged kaons, protons, neutrons,

Inclusive production of charged pions in p+C collisions at 158-GeV/c beam momentum

Eur.Phys.J.C 49 (2007) 897-917



Proton making pions

NA49 Data/MC comparison

Statistical:

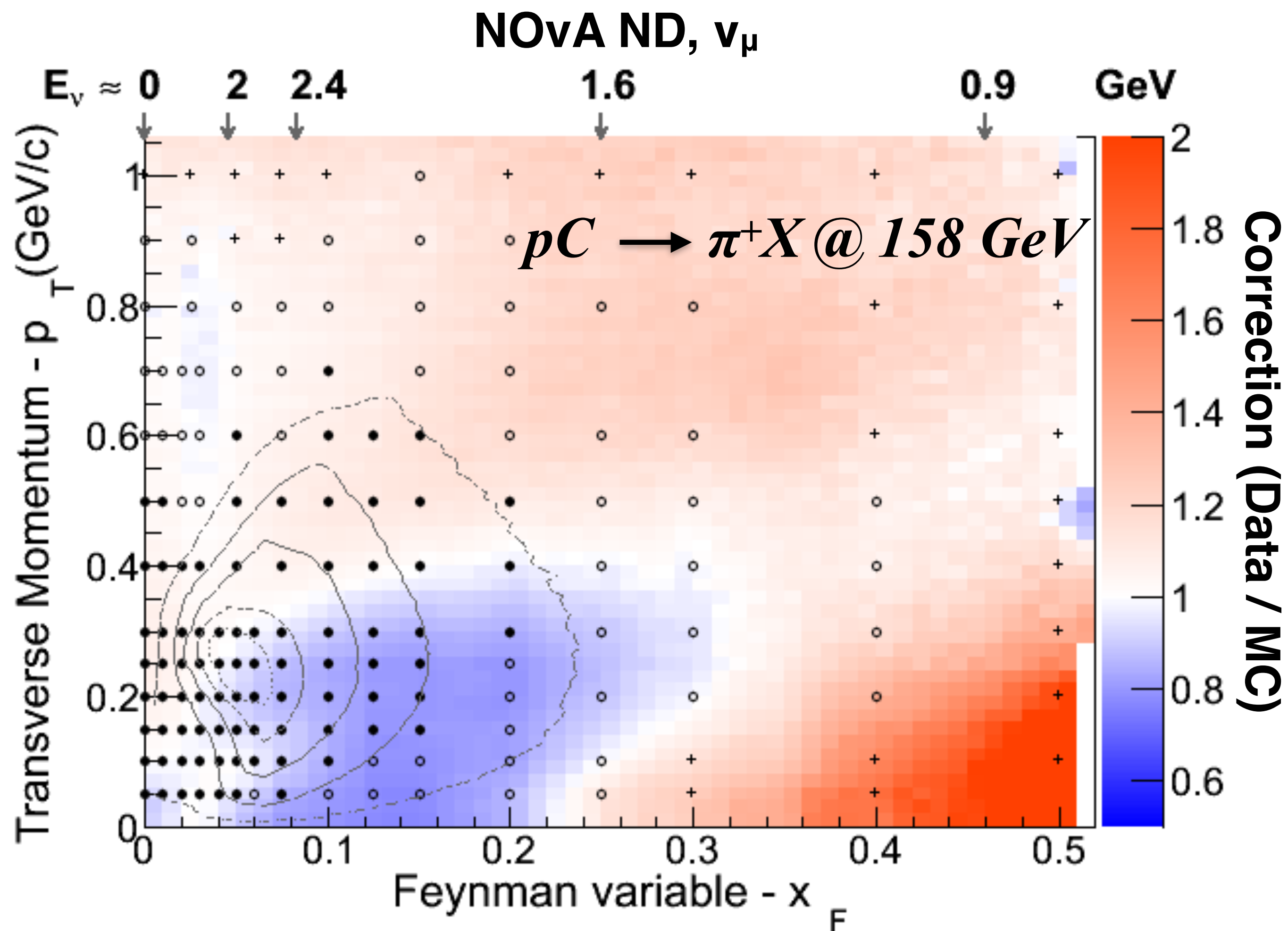
- » Closed circles: statistical error < 2.5%,
- » Open circles: statistical error 2.5-5.0%,
- » Crosses > 5%

Systematics:

- » 3.8% (added in quadrature).
- » Highly correlated bin-to-bin (assumed 100%)

Contours:

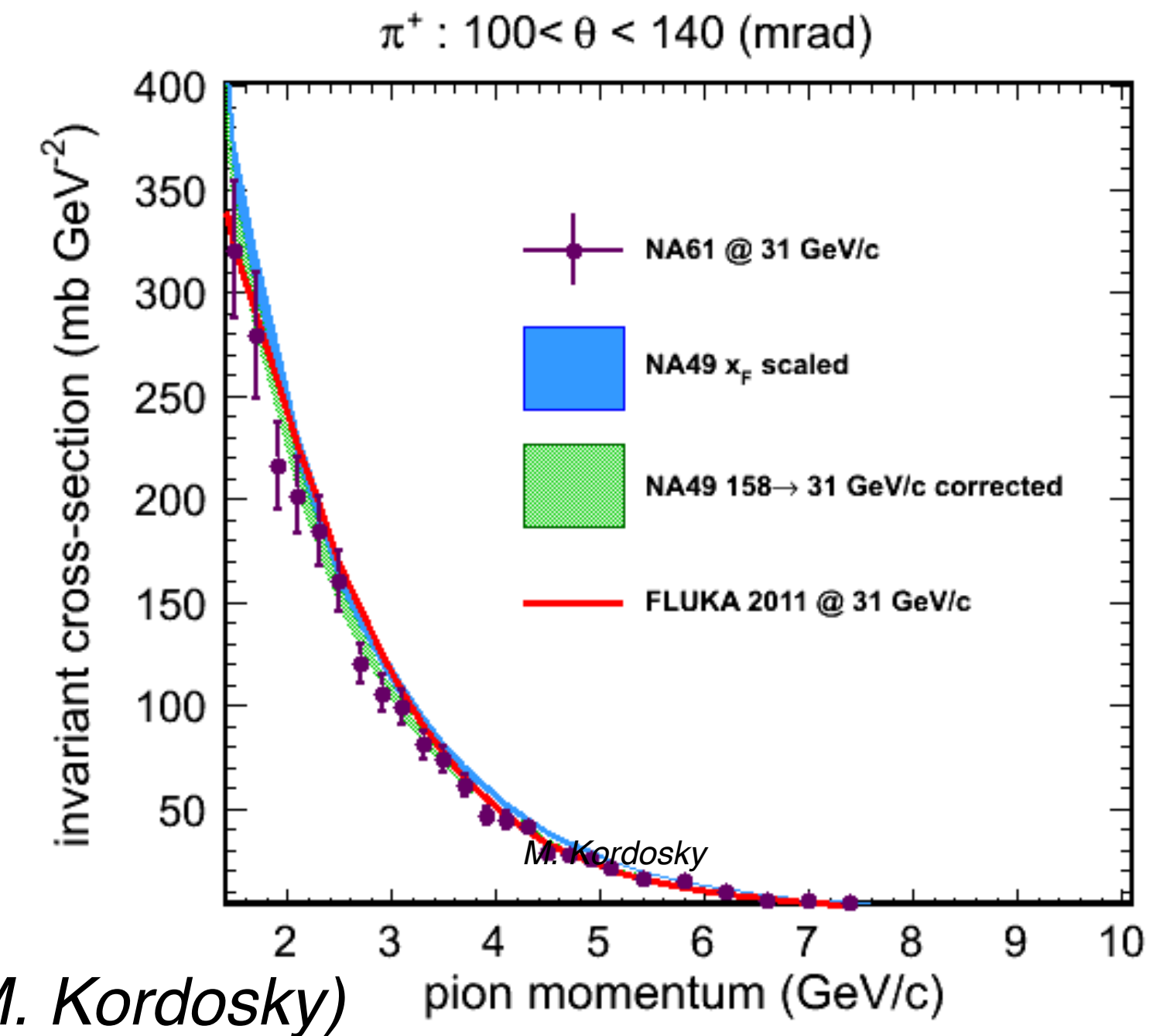
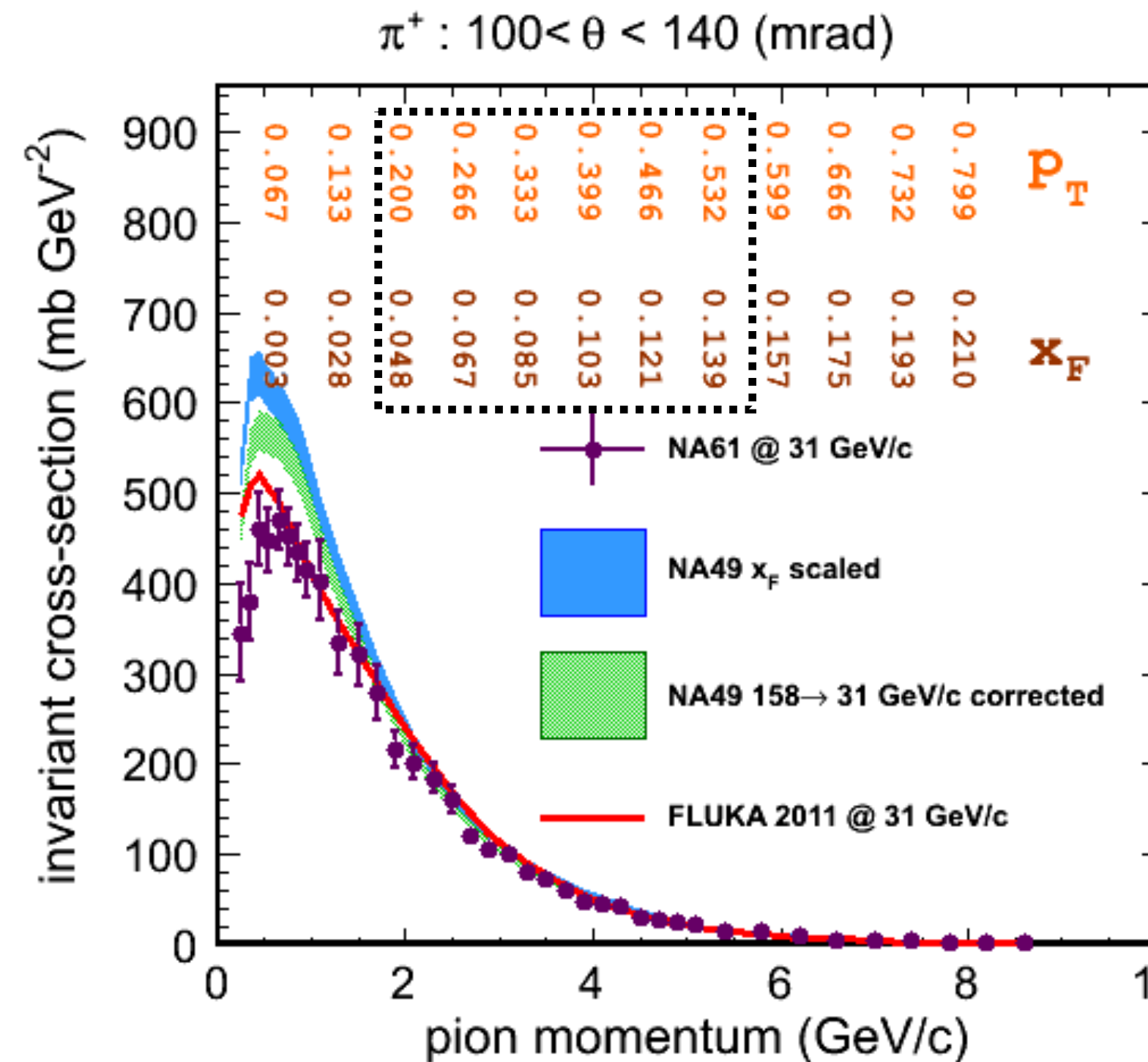
- » 2.5, 10, 25, 50 and 75 % of the π yields.



Using NA49 to correction proton - Carbon

$$correction(x_F, p_T, E) = \frac{f_{Data}(x_F, p_T, E = 158 GeV) \times scale(x_F, p_T, E)}{f_{MC}(x_F, p_T, E)}$$

- » We use NA49 for proton-Carbon interactions in 12-120 GeV using Feynman-x scaling
- » NA49 (158 GeV/c) and NA61 (31 GeV/c) look in good agreement in our region of interest.
- » The violation of the scaling is calculated using FLUKA.



Nucleon - A

Average number of hadronic interactions per neutrino at NOvA ND

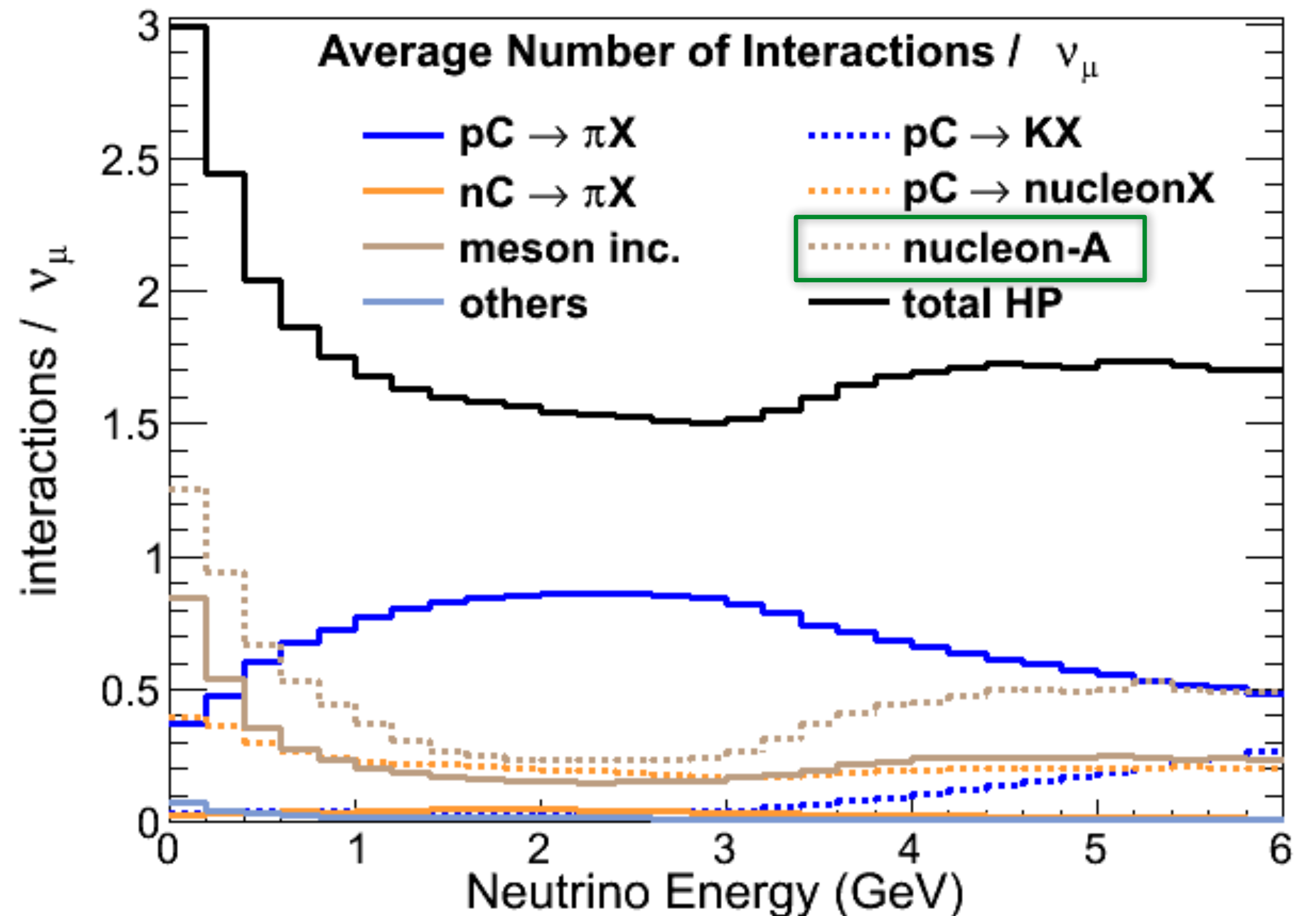
NOvA Simulation

» **Quasi-elastic interactions**, selected as $x_F > 0.95$

- We assign a 40% uncertainty

» **Extension of NA49.**

- We extend NA49 to other materials than carbon adding an additional uncertainty.
- Additional uncertainty is calculated by looking at: K^0 and Λ^0 production at 300 GeV in Skubic and check with Barton at 100 GeV.



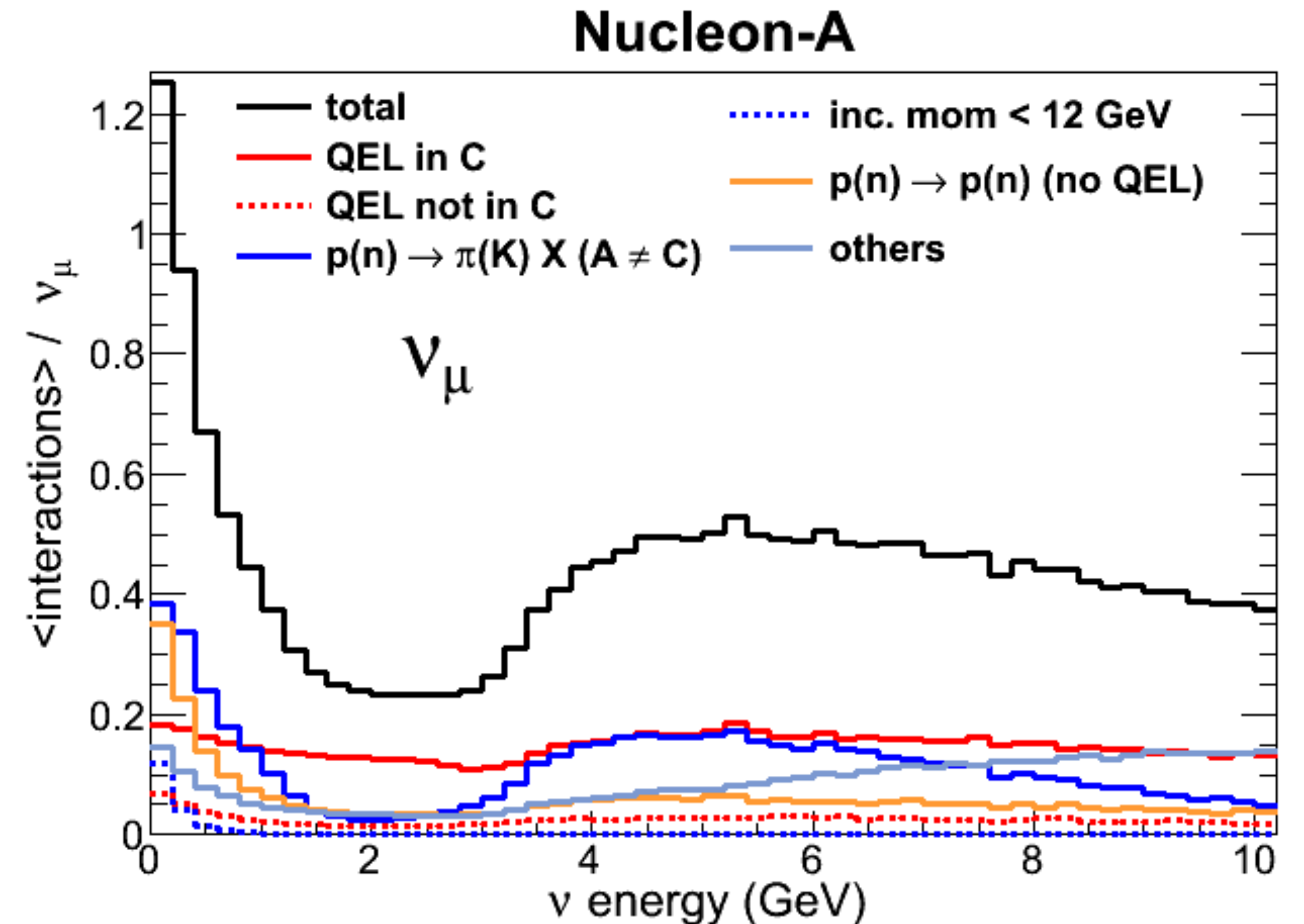
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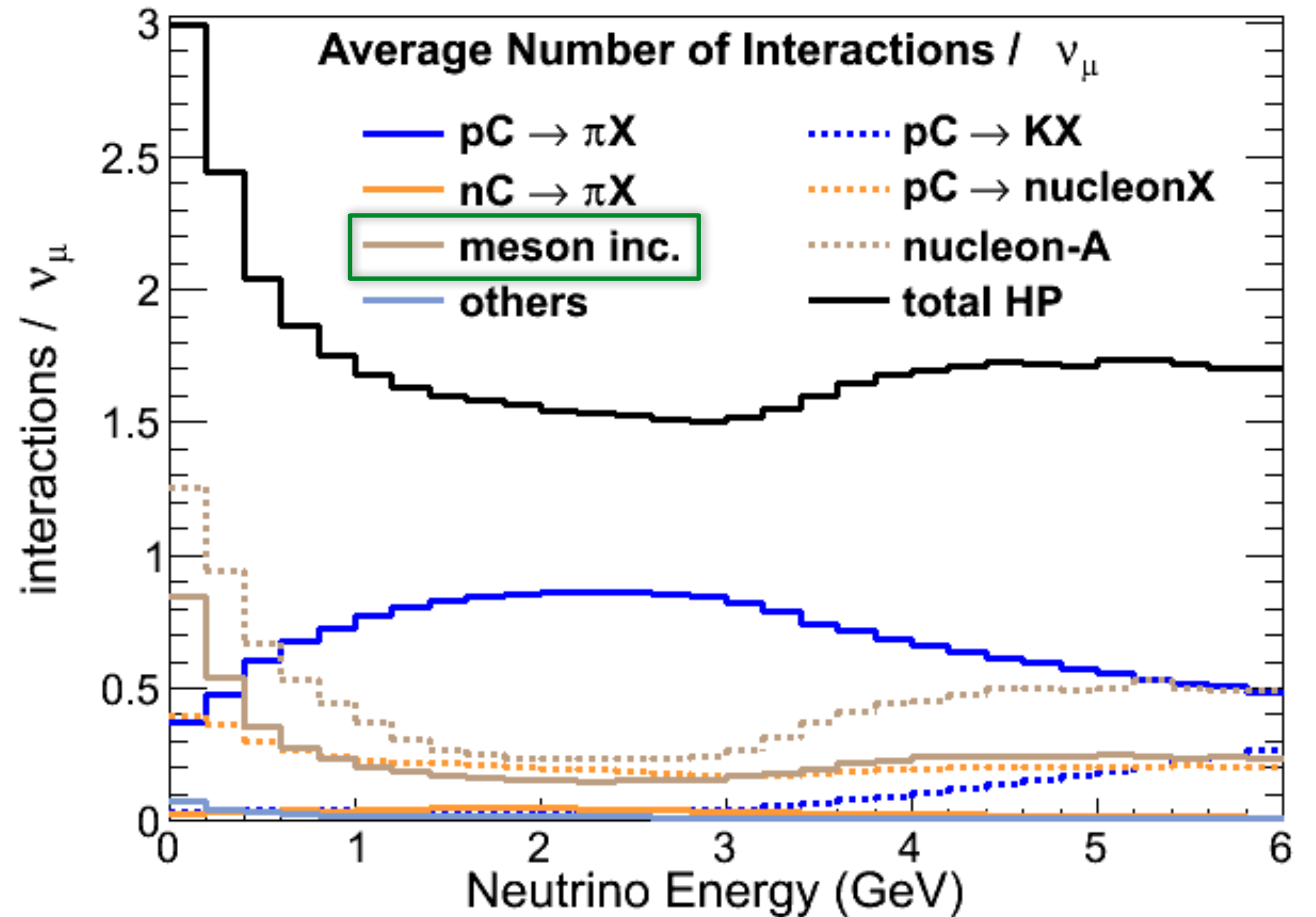


Incident meson

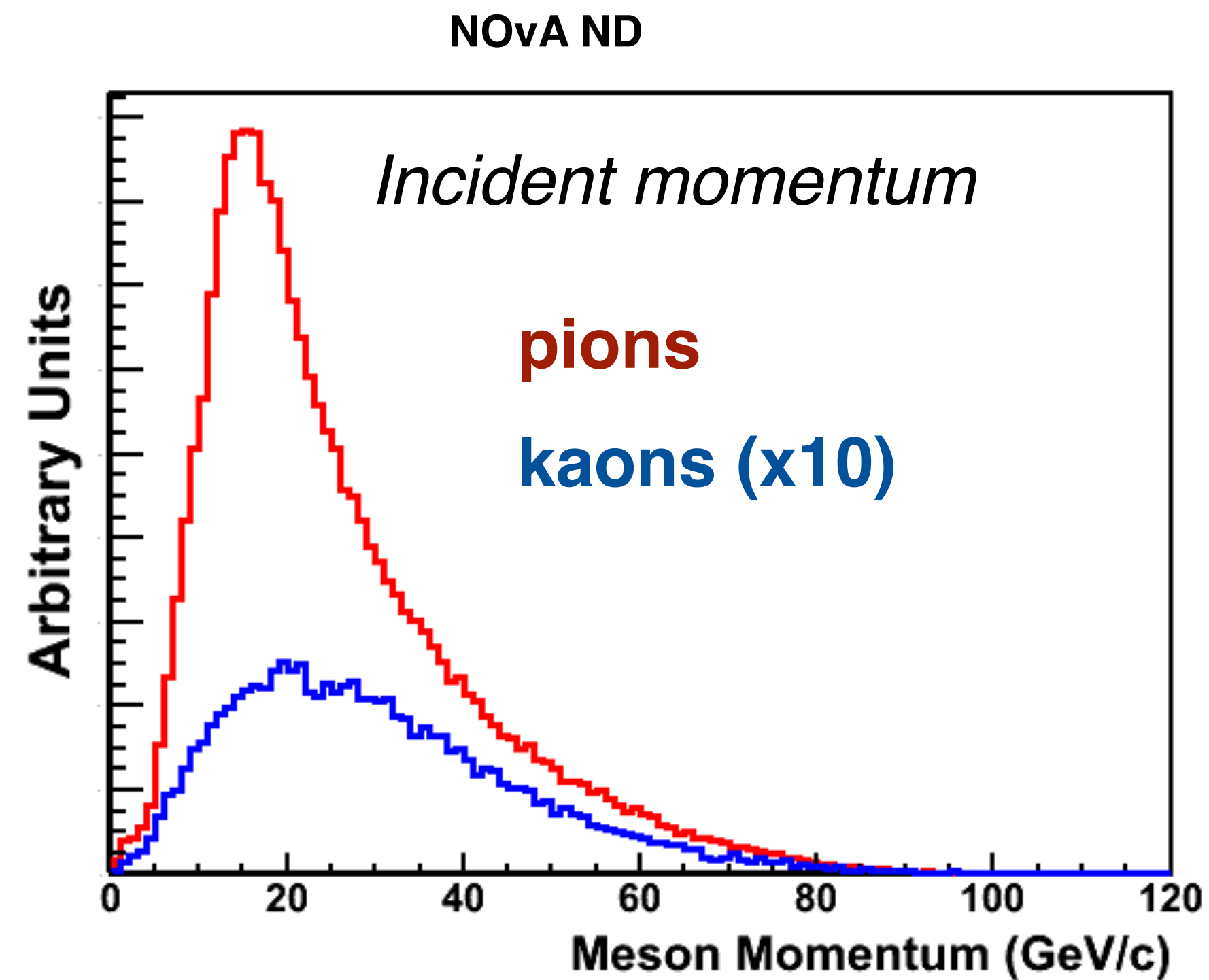
Average number of hadronic interactions per neutrino at NOvA ND

NOvA Simulation

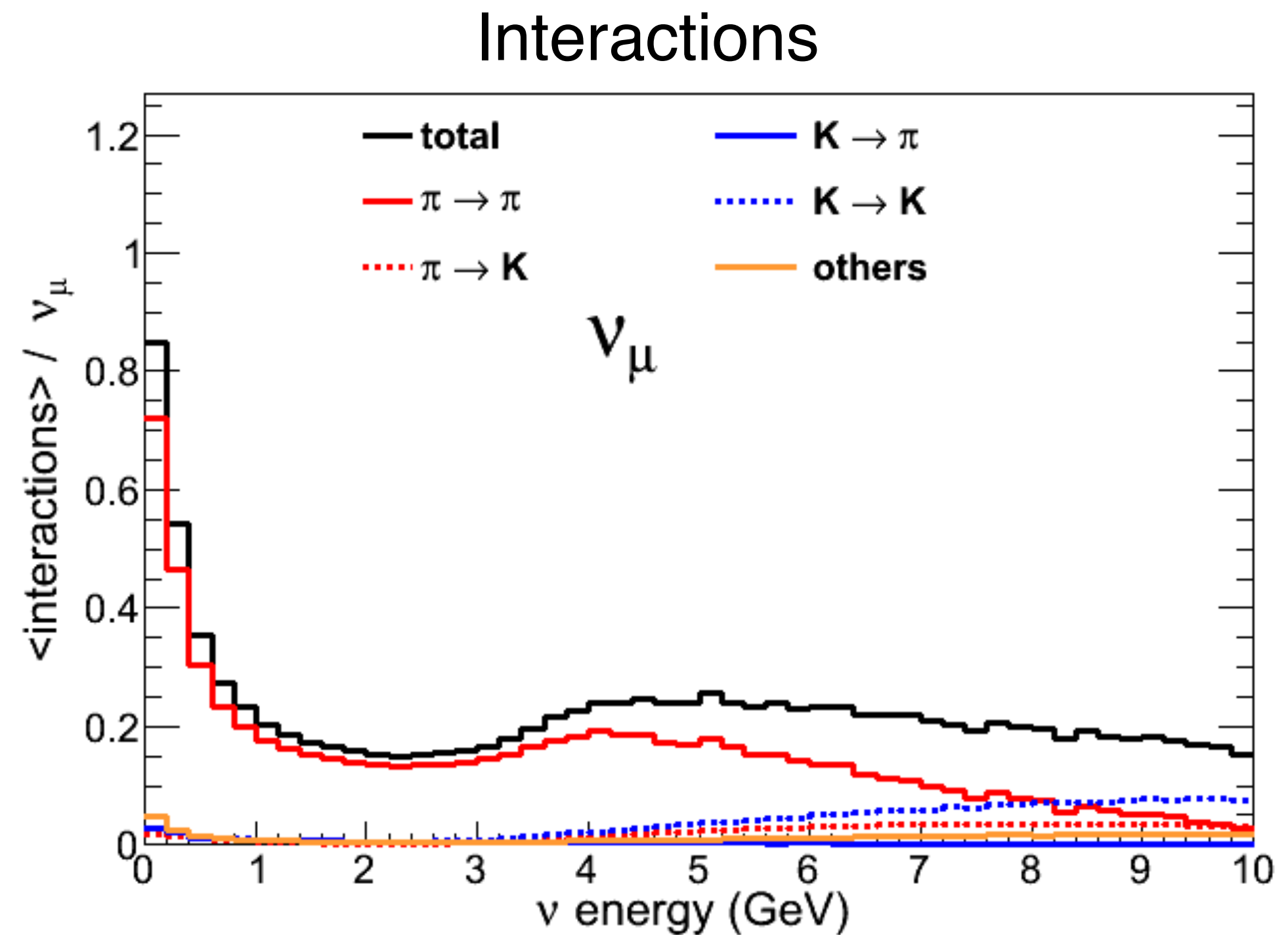
We assume large uncertainties for incident mesons: 40%



Incident meson



Main contribution from $\pi \rightarrow \pi$ in the whole neutrino energy spectrum

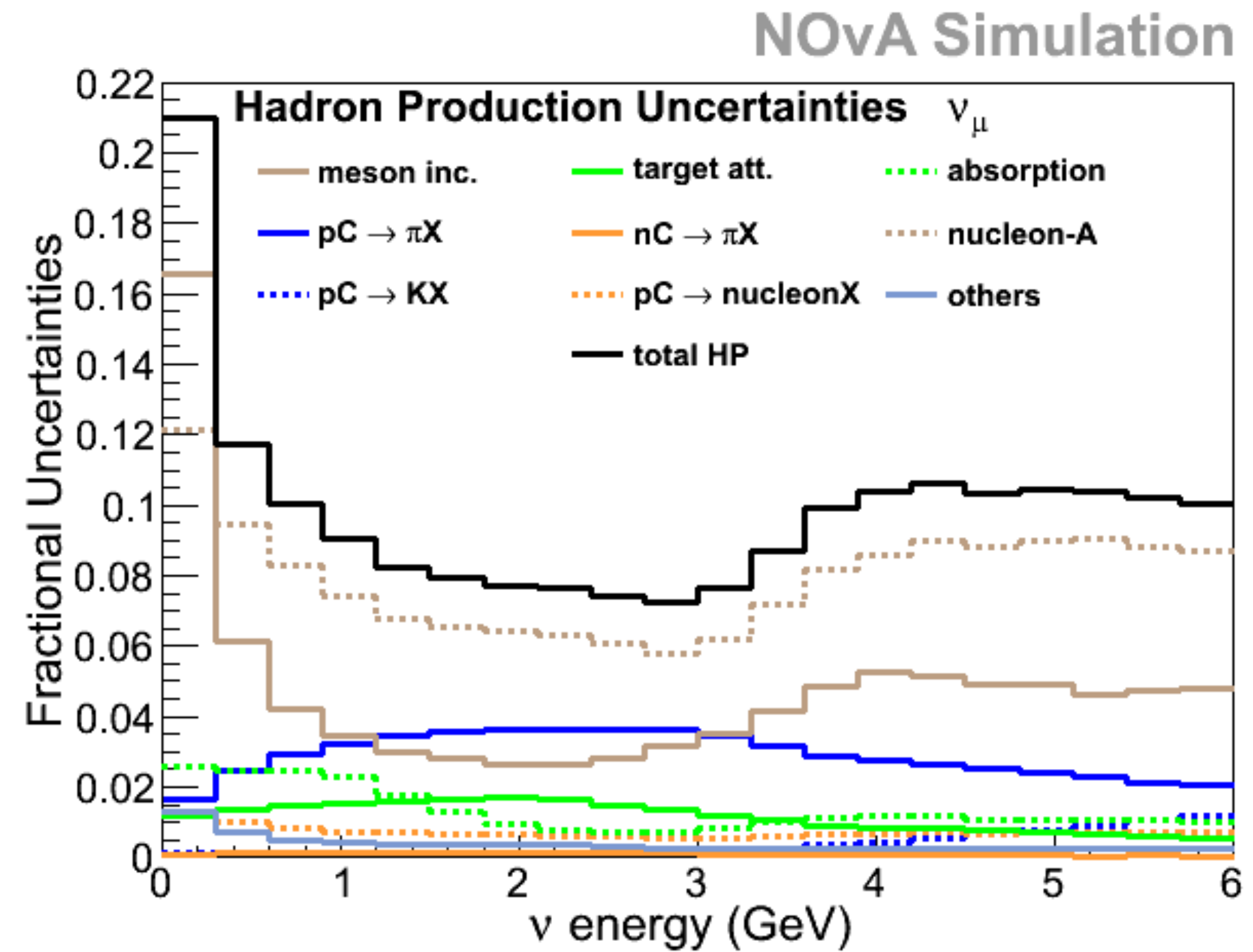
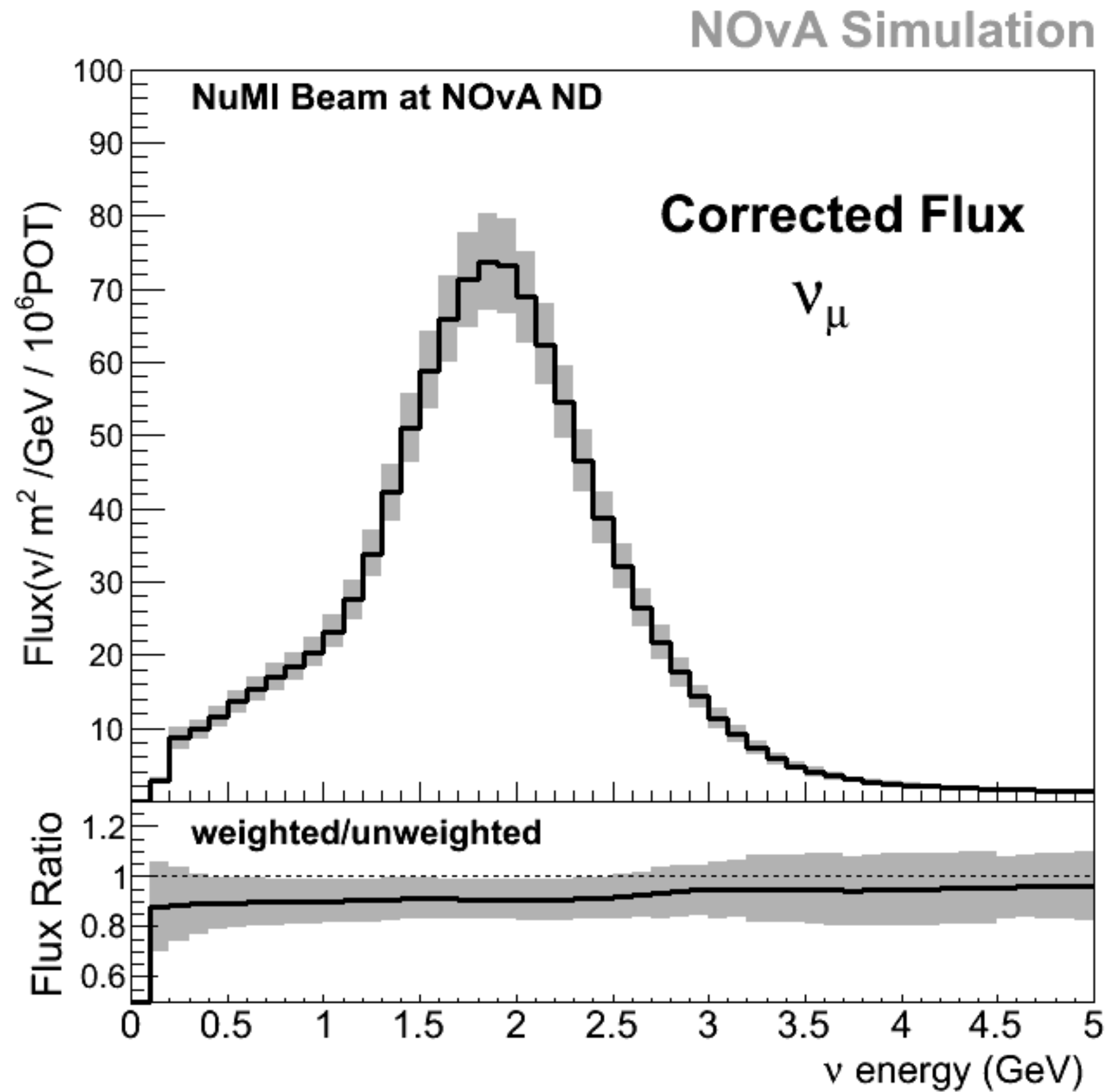


Currently we are working to include the NA61 incident pion on Be and C
Phys. Rev. D 100, 112004 (2019)

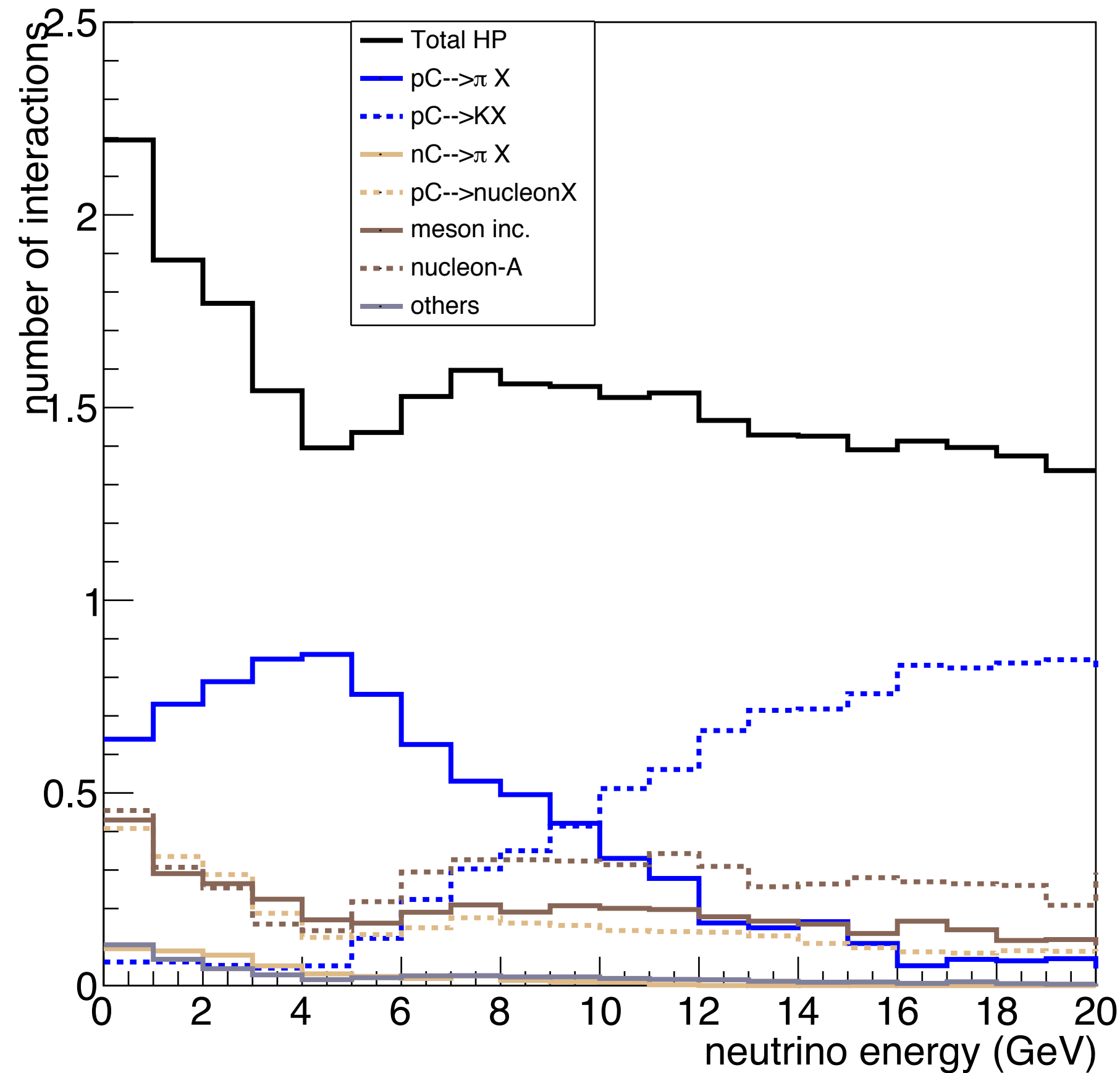
A Priori Flux Results for NOvA Near Detector



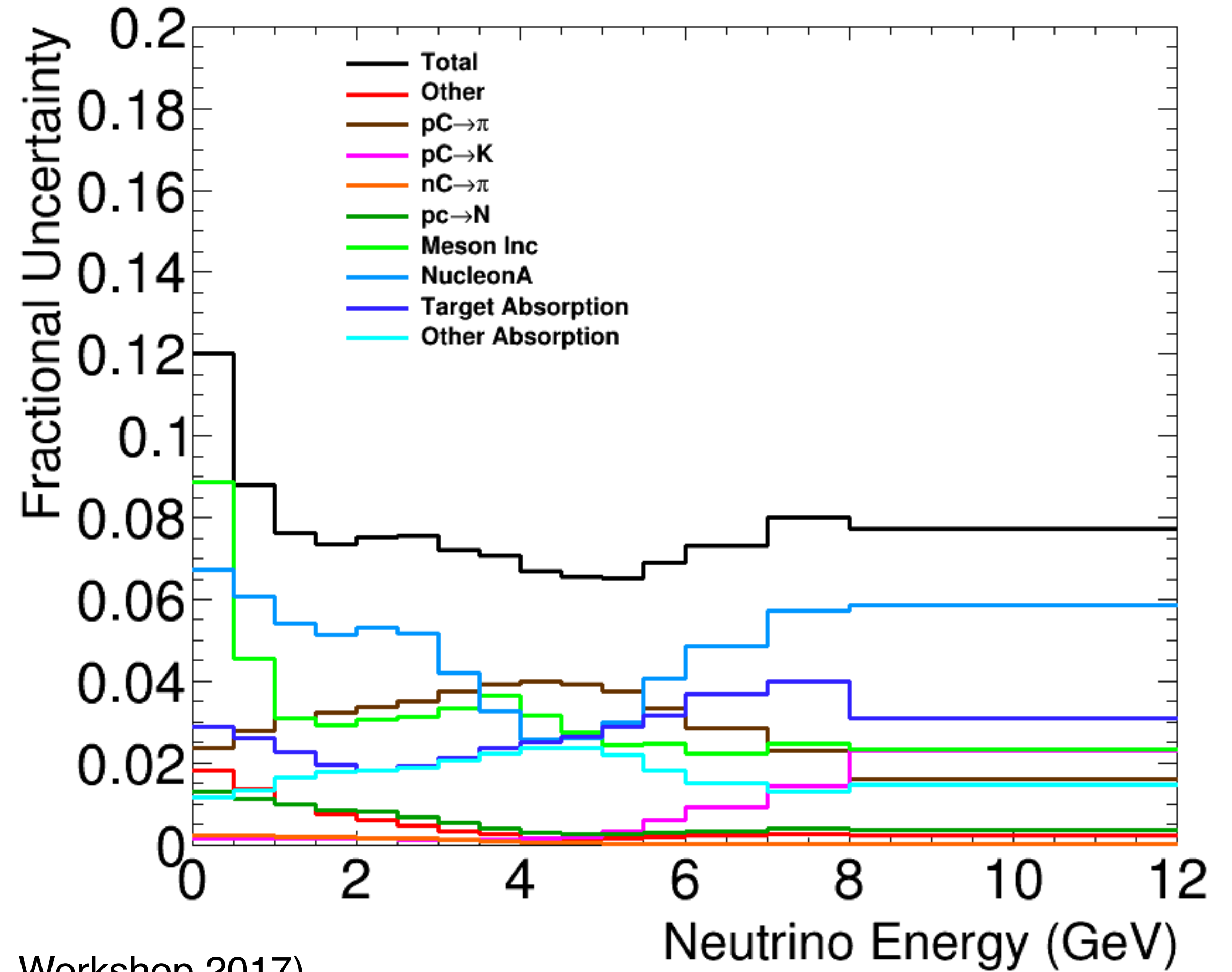
» A fully implemented a priori flux prediction in NOvA



A Priori Flux Results for DUNE Near Detector



L. Fields (NA61 Workshop 2017)



Summary

Summary

- » Fermilab long baseline neutrino experiments, such as MINERvA, NOvA and DUNE use external hadron production data to predict the flux in their detectors. A common framework has been pioneered by MINERvA that allows to extend to the upcoming new data
- » For interactions where we use HP data, the uncertainty associated to those interactions is relatively small. For instance, NA49 pion production: $\sim 4\%$.
- » New data is valuable (and crucial) for our physics program:
 - The flux simulation is the starting point of the simulation chain for all analysis.
 - The impact on the single detector analysis (such as cross section measurements) is direct
- » New data will be incorporated in our framework to be used for the Fermilab long baseline experiments.

Backup

NA49 vs NA61

