

<u>dzahini@</u>univ-grenoble-alpes.fr Xdigit is a start-up for specific ADC design Mainly for sensors' array like pixels

www.xdigit.fr

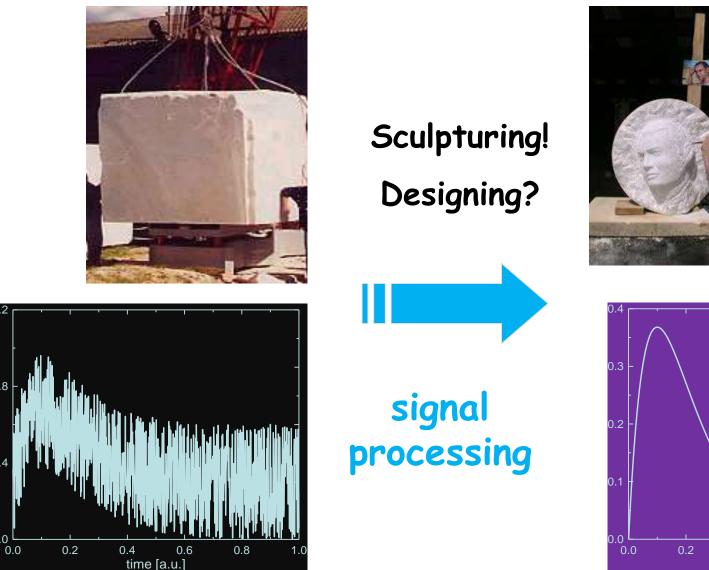


OVERVIEW FOR SIGNAL READOUT FOR <u>CAPACITIVE</u> DETECTOR <u>PULSE</u> PROCESSING

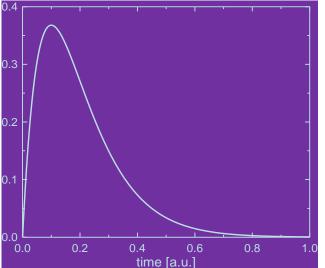
THANKS to many who provide slides and other documents online that I use for this lecturer

- Angelo Rivetti: Front end electronics for radiation sensors
- Emilio Gatti & Manfredi (INFN)
- Yan Kaplon (CERN)
- Christophe de la Taille (Omega lab)
- Glenn F. Knoll: Radiation detection & measurement
- Chiara Guazzoni; http://home.dei.polimi.it/guazzoni

What signal processing means?







How noisy are flowers in your garden

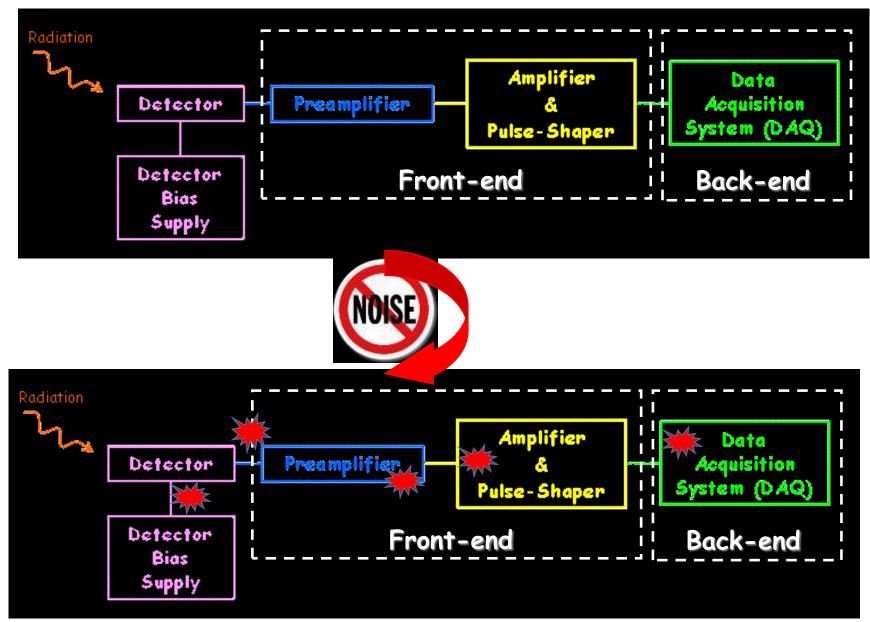


If plants can grow without ?? noise,

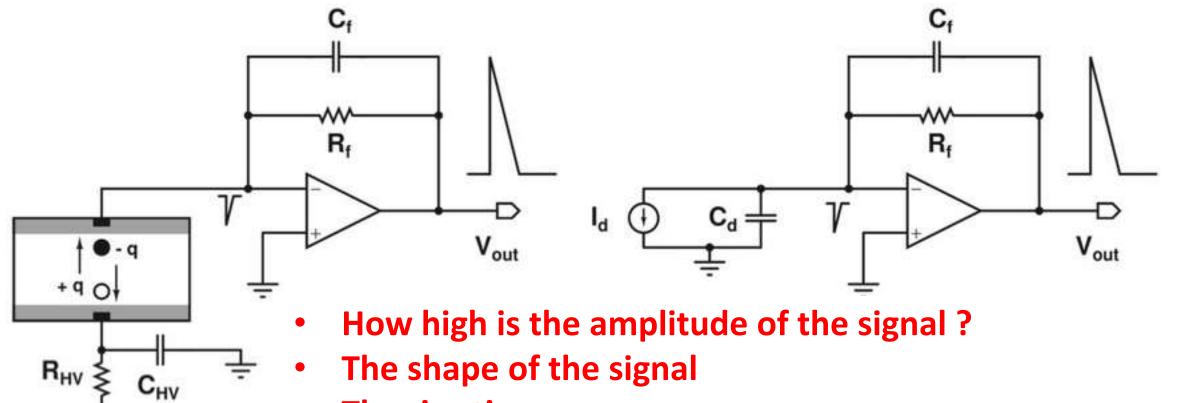
Why can't I amplify without adding noise

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Actual system is unfortunately noisy



Signal polarity with negative High voltage



- The rise time
- The duration

HV

• The fluctuation due to the process generating the signal

Detector signal duration: how short?

The signal generally is a **short** current pulse

- thin silicon detector (10 –300 μ m): 100 ps–30 ns
- thick (~cm) Si or Ge detector: 1 –10 µs
- proportional chamber:
- Microstrip Gas Chamber:
- Scintillator+ PMT/APD:

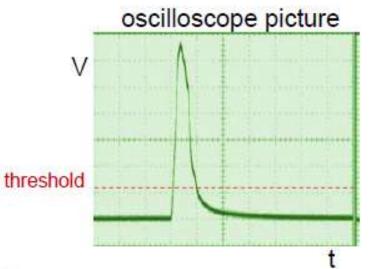
- 10 ns –10 µs
- 10 –50 ns
- 100 ps–10 µs

Energy ~
$$\int i(t) dt$$

What Information can one extract from a pulse?

Various measurements of this signal are possible Depending on information required:

- Signal above threshold digital response / event count
- Integral of current = charge
 → energy deposited
- Time of leading edge
 → time of arrival (ToA) or time of flight (ToF)
- Time of signal above threshold
 energy deposited by TOT

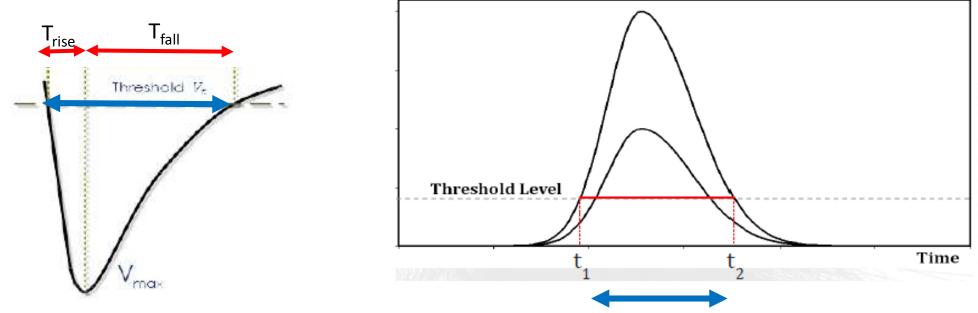


and many more ...

Counting and time over threshold

For **digital imaging** a counter is used after the comparator Then one proceed by counting the number of events in a frame rate.

The comparator system could be used also to quatify the amplitude of an incoming signal. The time spent over threshold by the amplifier output is somehow proportional to the amplitude of the incoming signal: TOT



Parameter impacting the pulse **amplitude**:

Ei = Minimum ionization energy (depends on the detector cristal, gaz, or liquide)
Ep (> Ei): average energy to generate a charges pair
E : Energy lost by an incoming particle =>
Np: Average Number of generated Pairs

Np = E/Ep => an *average* number

But instantly, the number follows a probabilistic law with a fluctuation from one event to another displaying a <u>standard</u>

deviation
$$\sigma_{Np} = \sqrt{F * Np}$$
; F is the Fano factor

In many material F<1 then σ_{Np} is better than one could expect from the Poisson statistics (\sqrt{Np});

Detector's equivalent circuit: C_D an I_L

0.1-10 pF

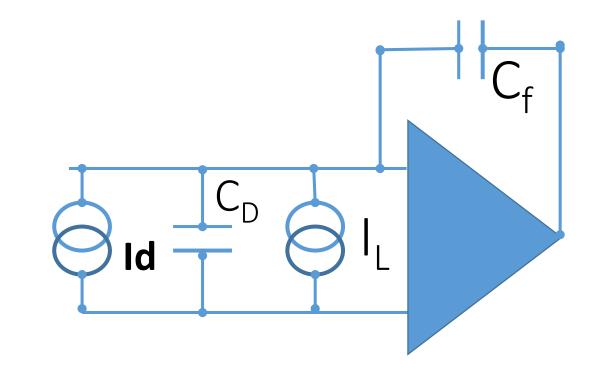
3-30 pF

Detector = capacitance Cd

- Pixels :
- PMs:
- Ionization chambers: 10-1000 pF
- Sometimes effect of transmission line

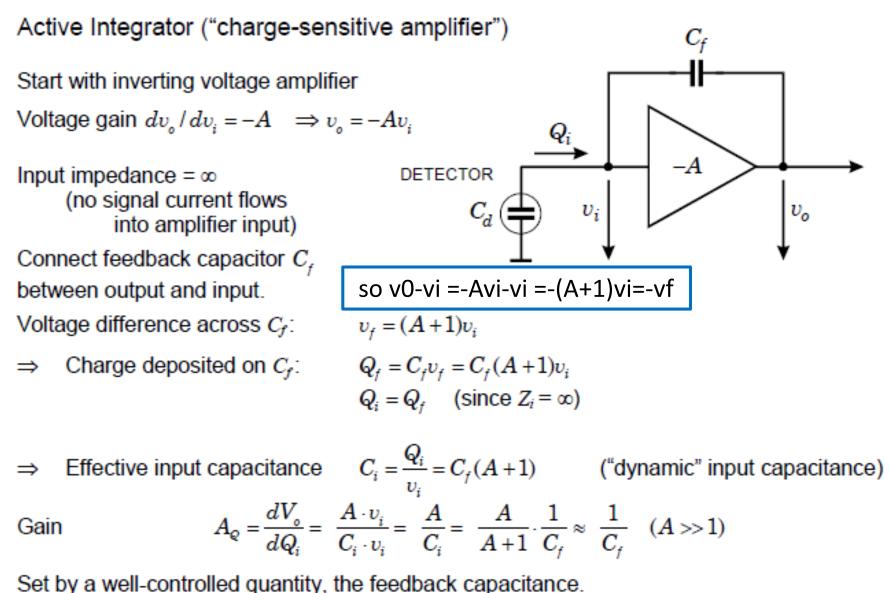
Signal : current source

- Pixels : ~100 e⁻/µm
- PMs : 1 photoelectron -> 105-107 e-
- Modeled as an impulse (Dirac) : $i(t)=Q_0 \overline{o}(t)$



- C_D impact on **speed** and **noise** figures
- I impact output DC level, and on noise

Charge sensitive preamplifier



Helmuth Spieler LBNL

Charge preamplifier: exemple of typical values

So finally the fraction of charge signal measured by the amplifier is:

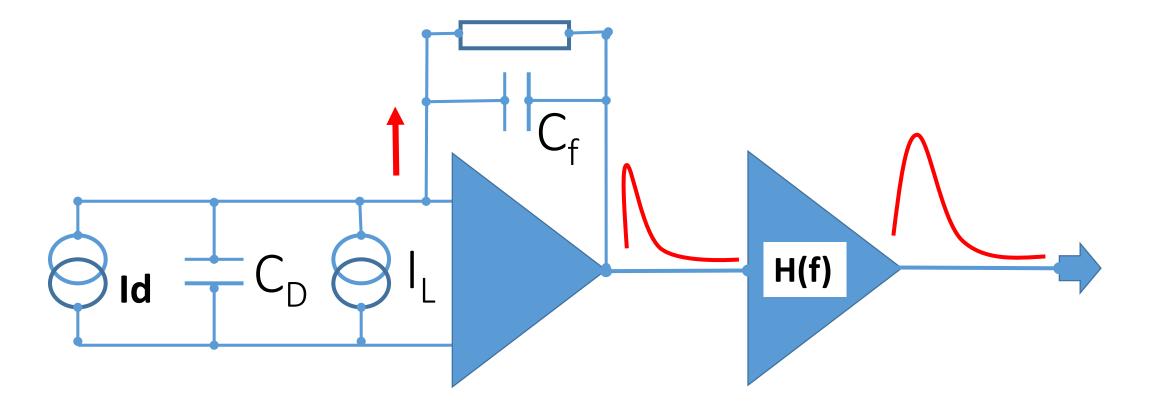
$$\frac{Q_i}{Q_s} = \frac{C_i v_i}{v_i (C_i + C_{det})} = \frac{1}{1 + C_{det} / C_i} \qquad C_f \approx \frac{A}{C_i} \quad (A >> 1)$$

Example:

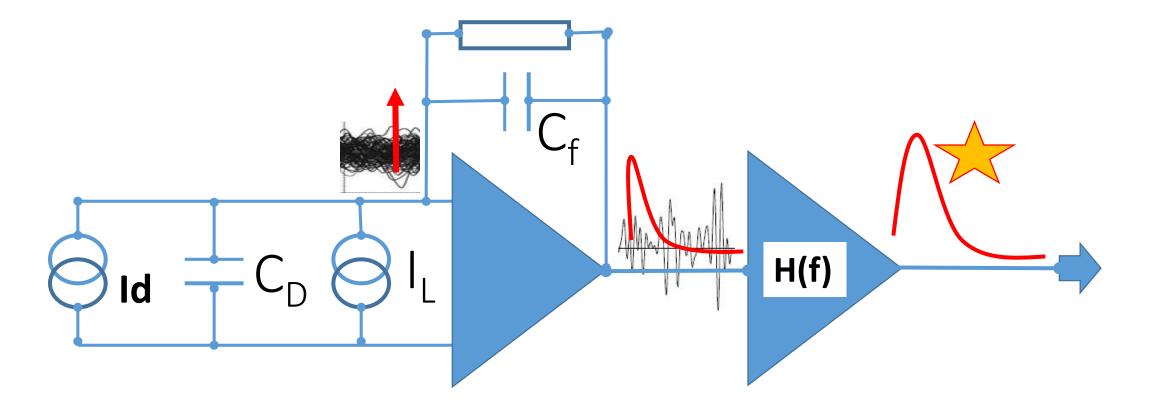
 $A = 10^{3}$ $C_{f} = 1pF \quad \Rightarrow \quad C_{i} = 1nF$ $C_{det} = 10pF \quad \Rightarrow \quad Q_{i}/Q_{s} = 0.99 \quad (C_{i} > > C_{det})$ $C_{det} = 500pF \quad \Rightarrow \quad Q_{i}/Q_{s} = 0.67 \quad (C_{i} \sim C_{det})$ \uparrow

Si det: 50um thick, 500mm² area

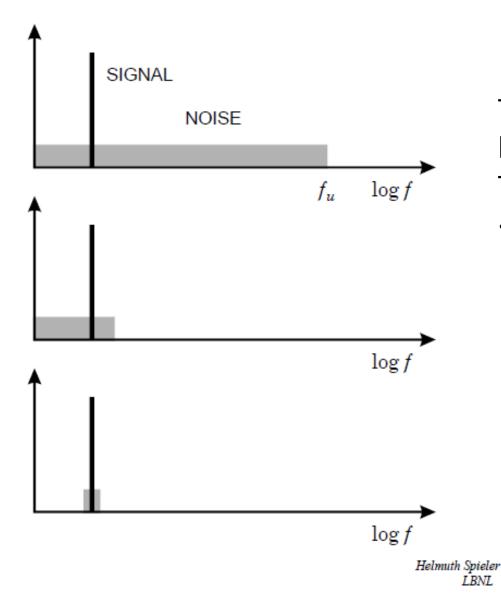
Front end amplifier and shaper circuit



Noise and Front end amplifiers

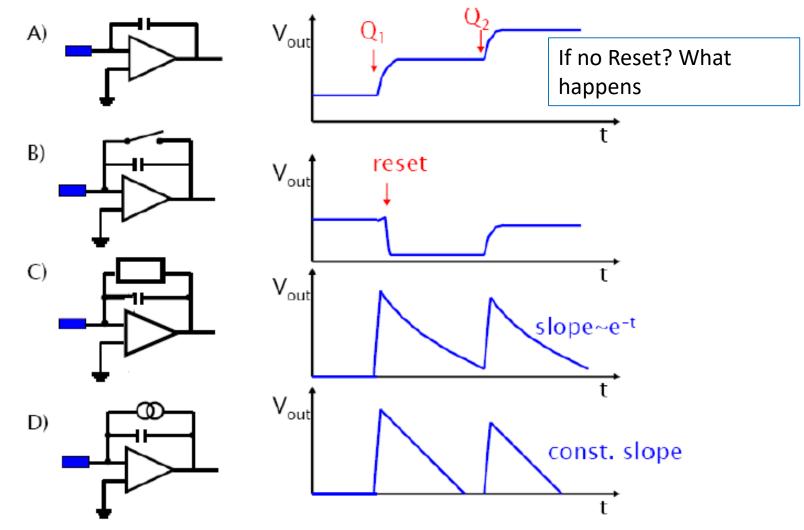


Signal & noise bandwidth



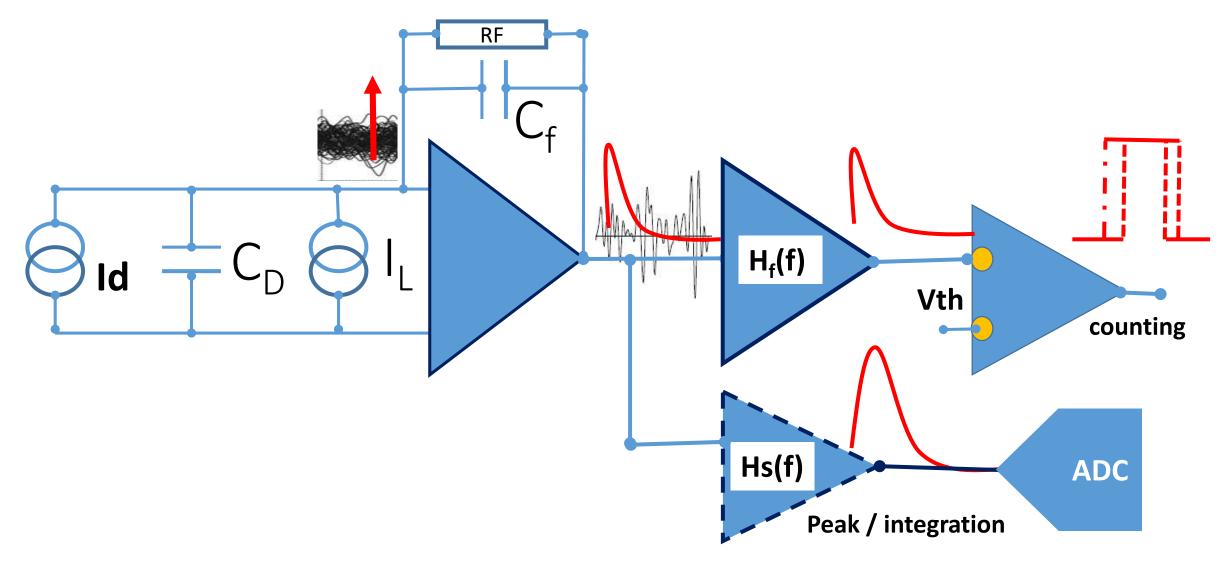
The total noise will be the integral over the bandwidth Then Signal/Noise could be improved by « optimizing » the noise bande close to the signal's.

Different options of reset systems

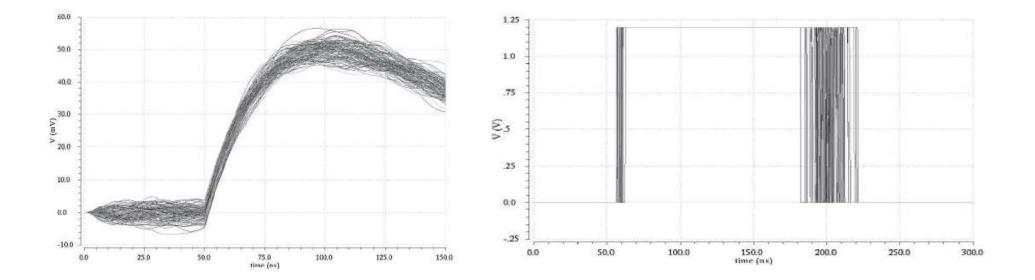


Ref LBL + de La Taille

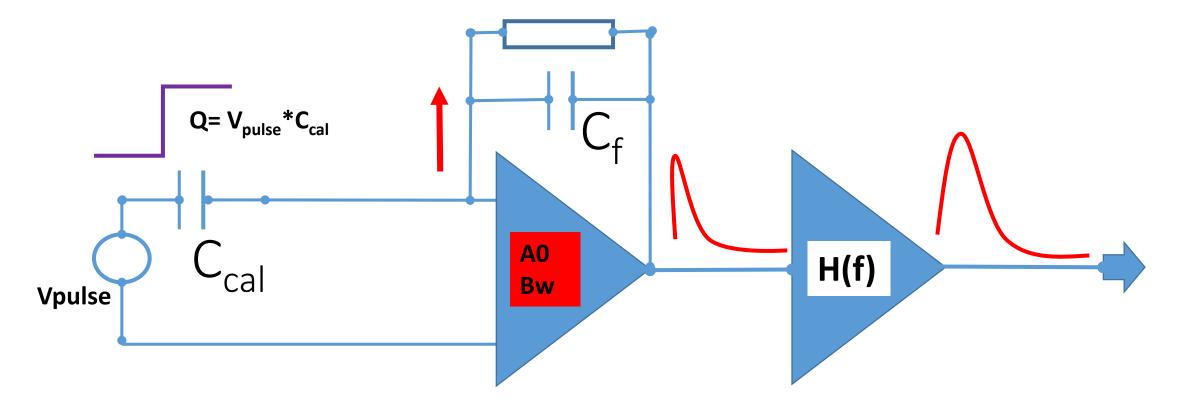
Noise and Front end amplifiers and comparators



Even for counting, beware of the noise: why using a shorter peaking time for counting readout?

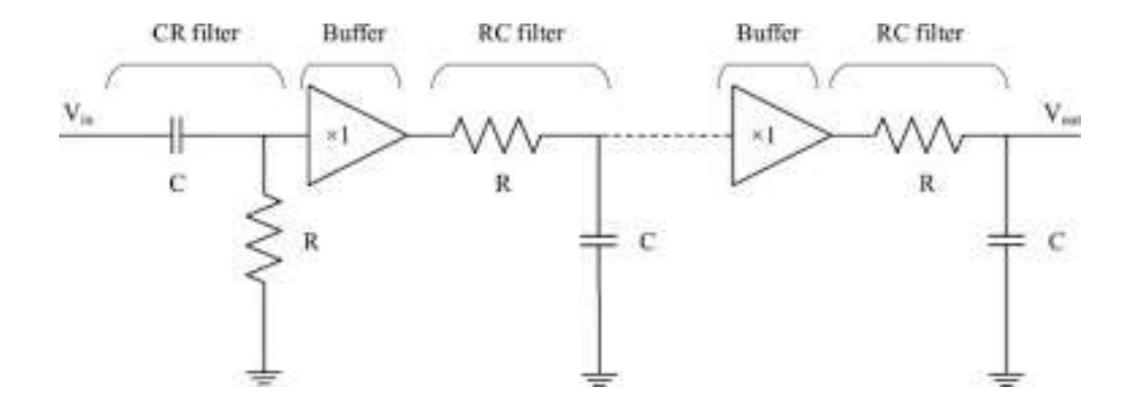


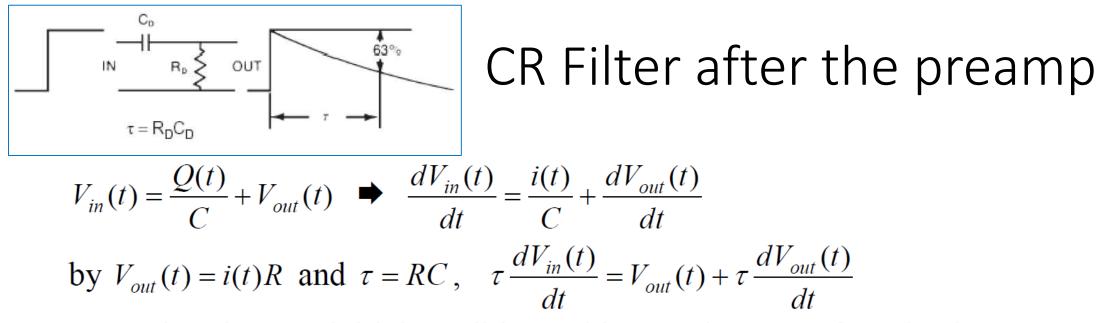
Calibration and simulation of Front end amplifiers



Before been so happy that you find a noise very low!!, make sure your circuit is still amplifying the signal

CR RC shapers (filters)





Assuming the zero initial condition, taking Laplace transform leads to

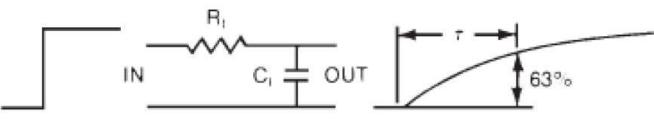
$$V_{out}(s) = \frac{\tau s}{1 + \tau s} V_{in}(s) = G_{CR}(s) V_{in}(s)$$

For the step function input

$$V_{in}(t) = \frac{V_0 \ (t > 0)}{0 \ (t \le 0)} \qquad \Rightarrow \quad V_{in}(s) = L[V_{in}(t)] = \frac{V_0}{s}$$

the output signal becomes

$$G_{CR}(i\omega) = \frac{i\omega\tau}{1+i\omega\tau} \implies |G_{CR}(i\omega)| = \frac{\omega\tau}{\sqrt{1+\omega^2\tau^2}}$$



RC stage of the shaper

 $\tau = R_I C_I$

$$V_{in}(t) = i(t)R + V_{out}(t) \text{ and } i(t) = \frac{dQ(t)}{dt} = C \frac{dV_{out}(t)}{dt}$$

$$\blacktriangleright V_{in}(t) = \tau \frac{dV_{out}(t)}{dt} + V_{out}(t) \quad \blacktriangleright \quad V_{out}(s) = \frac{1}{1 + \tau s} V_{in}(s) = G_{RC}(s) V_{in}(s)$$
Output signal for the step function input:

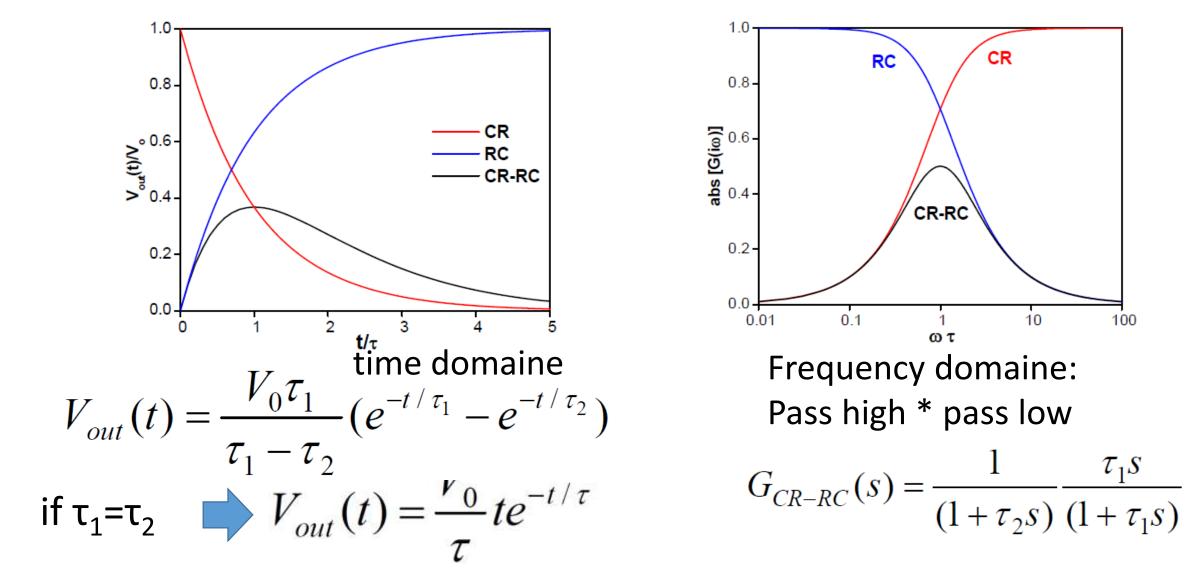
$$V_{out}(s) = \frac{1}{1 + \tau s} \frac{V_0}{s} \quad \blacktriangleright \quad V_{out}(t) = V_0(1 - e^{-t/\tau})$$
Frequency domain transfer function:

$$G_{RC}(i\omega) = \frac{1}{1 + i\omega\tau} \quad \bigstar \quad \left|G_{RC}(i\omega)\right| = \frac{1}{\sqrt{1 + \omega^2 \tau^2}}$$

CR+RC transfert functions for a step input This step stand for the integrator output signal

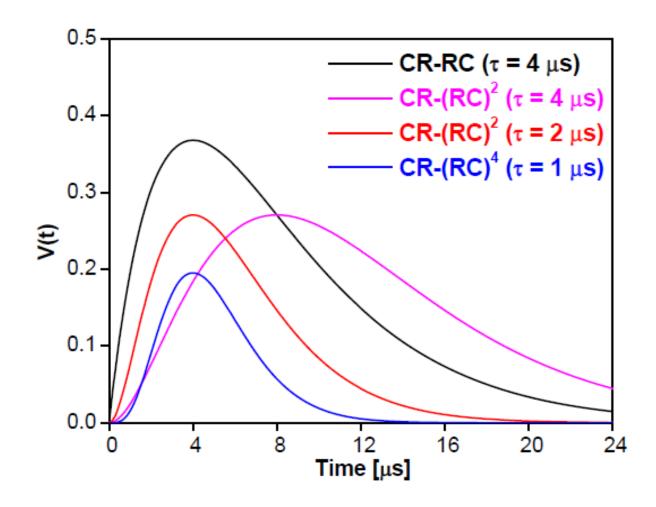
10

100



Filter / Shaper first order? second? Why

CR*RCⁿ filters or Semi-Gaussian pulse shaping

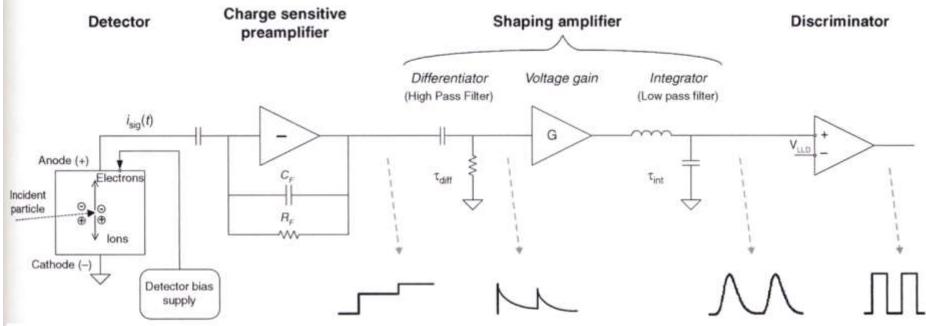


If a single CR high-pass filter is followed by several stages of RC integration, the output pulse shape becomes close to Gaussian amplifiers shaping, in this way are called **semi-Gaussian shaper**. Its output pulse is given by:

$$V_{out}(t) \propto \left(\frac{t}{\tau}\right)^n e^{-t/\tau}$$

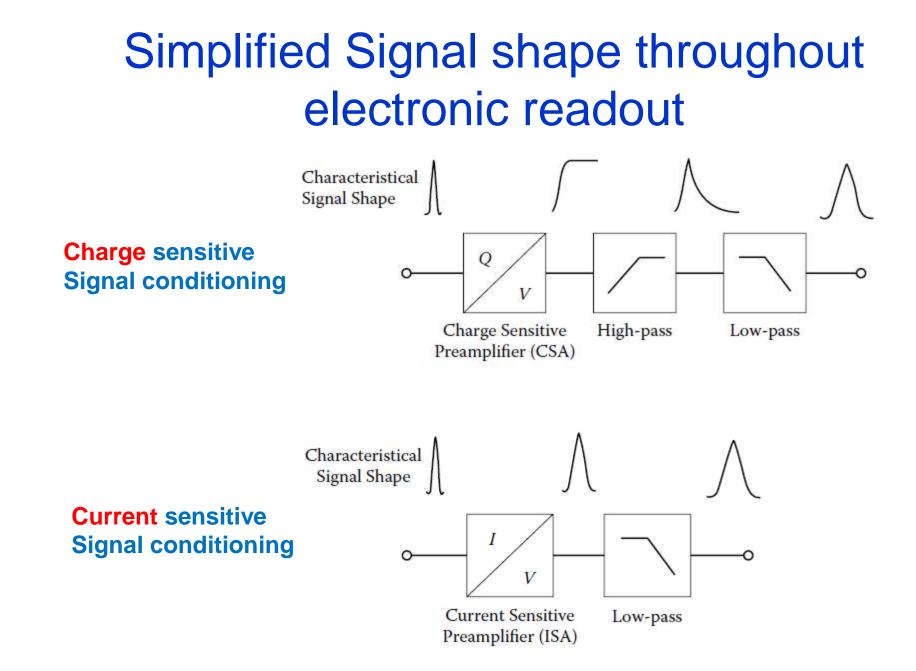
The peaking time in this case is equal to $n^*\tau$.

Signal shape following read out steps when 2 successive pulses (II)



Schematic of simple signal processing electronics. This circuit is suitable for use, as shown, in many applications and is conceptually similar to more complex circuits. These elements are discussed in greater detail in Chapter 17. (Courtesy of R. Redus, Amptek, Inc.)

Courtesy, Glenn F. « radiation detection »



WHAT IS NOISE ?

What is noise?

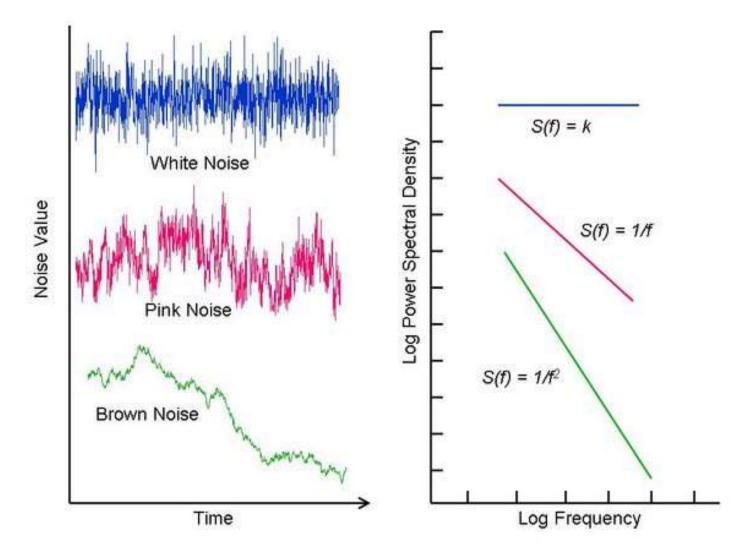
Noise is any undesired signal that masks the signal of interest.

- · Unwanted disturbance that interferes with a desired signal
- External: power supply & substrate coupling, crosstalk, EMI, etc.
- Internal: random fluctuations that result from the physics of the devices or materials
- Smallest detectable signal, signal-to-noise ratio (SNR), and dynamic range are determined by noise

$$SNR = \frac{P_{signal}}{P_{noise}} = \frac{V_{rms, signal}^2}{V_{rms, noise}^2}$$

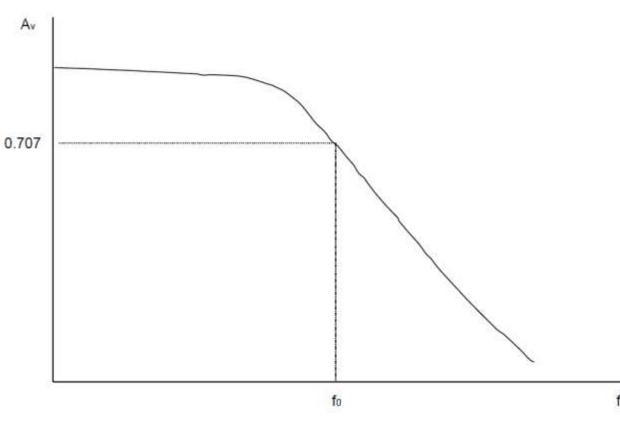
We will look at internal noise sources and how they affect key performance metrics.

Noise in general: time & frequency domain



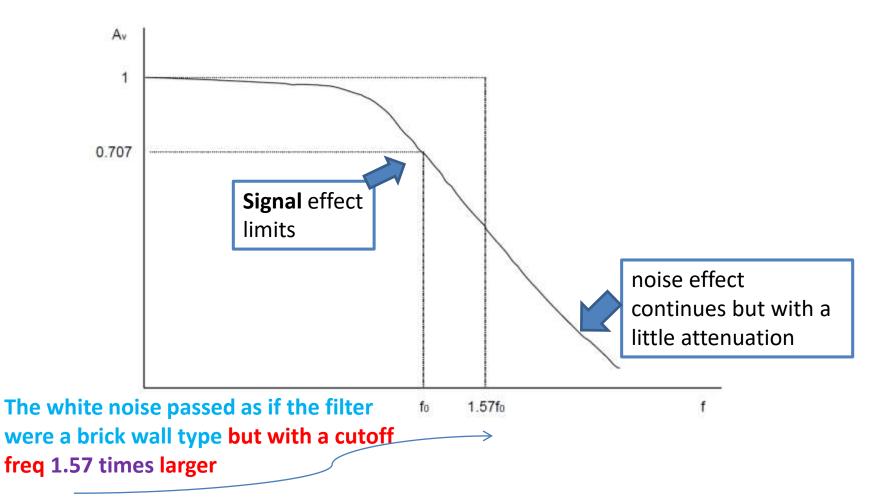
Noise Effective BW (NEB) of an amplifier (filter)

Lets consider a general *low pass* amplification (filter) system; What happen to a white noise located at the input of such amplifier? Is it amplified exactly as the signal?

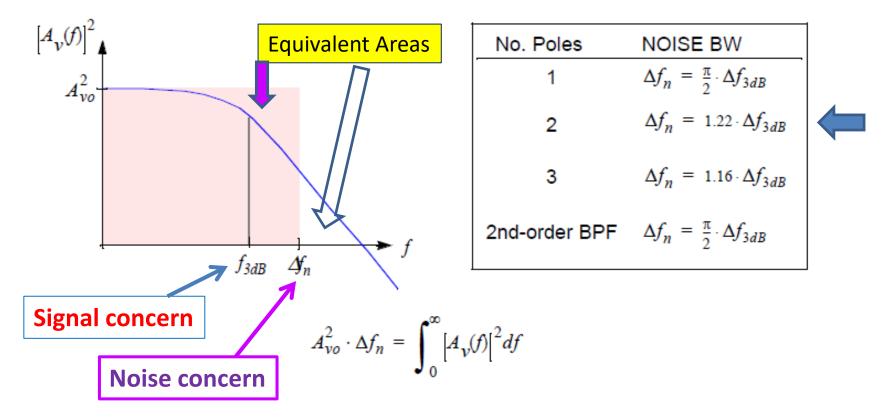


Noise bandwidth

Lets consider a low pass amplification system; What is the effect of a <u>larger bandwidth white noise</u> located at the input of such amplifier?



Noise Bandwidth # signal bandwidth



- · Noise bandwidth is defined for a brickwall transfer function
- Noise bandwidth is not the same as 3dB bandwidth

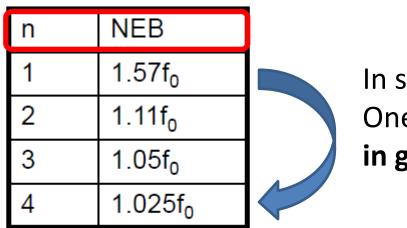
Noise Bandwidth improves when number of poles increase Keep it in mind and make the link later with CRRCn filtering

Optimizing Filtering <=> Optimize NEB

For Maximally flat (Butterworth) where f₀=f_{3dB}

$$NEB = \left(\int_{0}^{\infty} \frac{df}{1 + (f/f_0)^{2n}}\right)$$

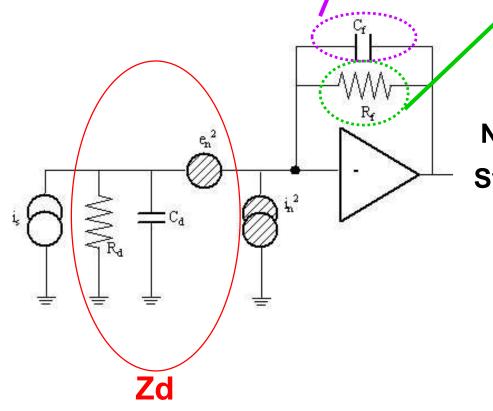
But higher order n means: More complicated and/or more power budget



In spectroscopie, One considers in general n=2

High order filter is good(0.5dB improvement)

Preamplifier: charges / Courant output noise



Noise spectrum at the output

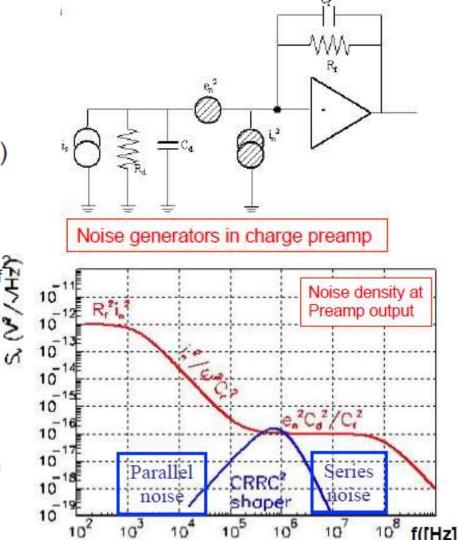
$$Sv(\omega) = (in^{2} + en^{2}/|Zd|^{2}) / \omega^{2}Cf^{2}$$

=
$$in^2 / \omega^2 C_f^2 + en^2 C_d^2 / Cf^2$$

If we neglect Rd and we translate en to a current By doing en/Zd

Noise issues for charge preamp: <u>frequency domaine</u>

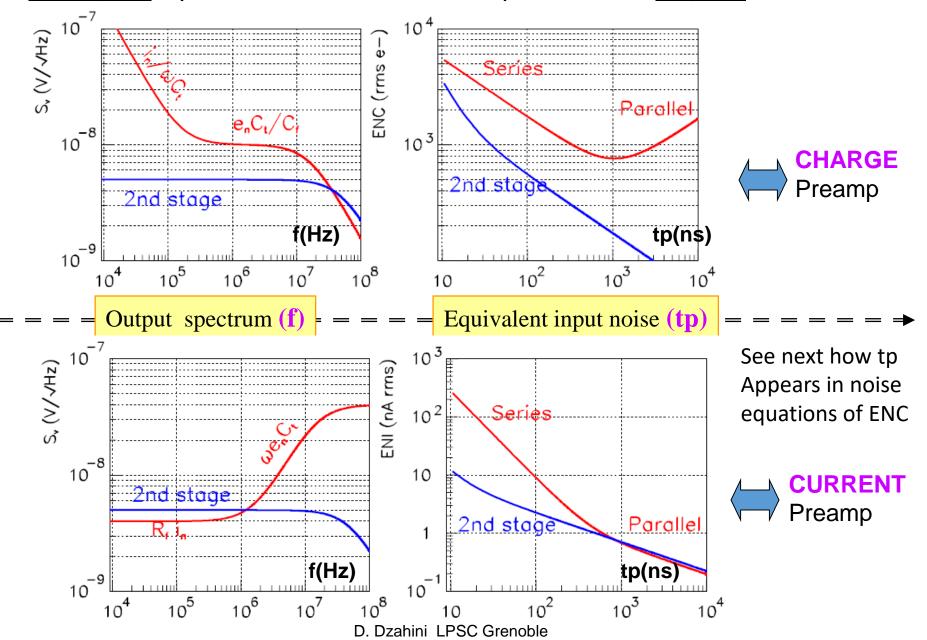
- 2 noise generators at the input
 - Parallel noise : (i_n²) (leakage currents)
 - Series noise : (e_n^2) (preamp)
- · Output noise spectral density :
 - $S_v(\omega) = (i_n^2 + e_n^2 / |Z_d|^2) / \omega^2 C_f^2$ $= i_n^2 / \omega^2 C_f^2 + e_n^2 C_d^2 / C_f^2$
 - Parallel noise in $1/\omega^2$
 - Series noise is flat, with a « noise gain » of C_d/C_f
- rms noise V_n
 - − $V_n^2 = \int Sv(\omega) d\omega/2\pi \rightarrow \infty$ (!)
 - Benefit of shaping...



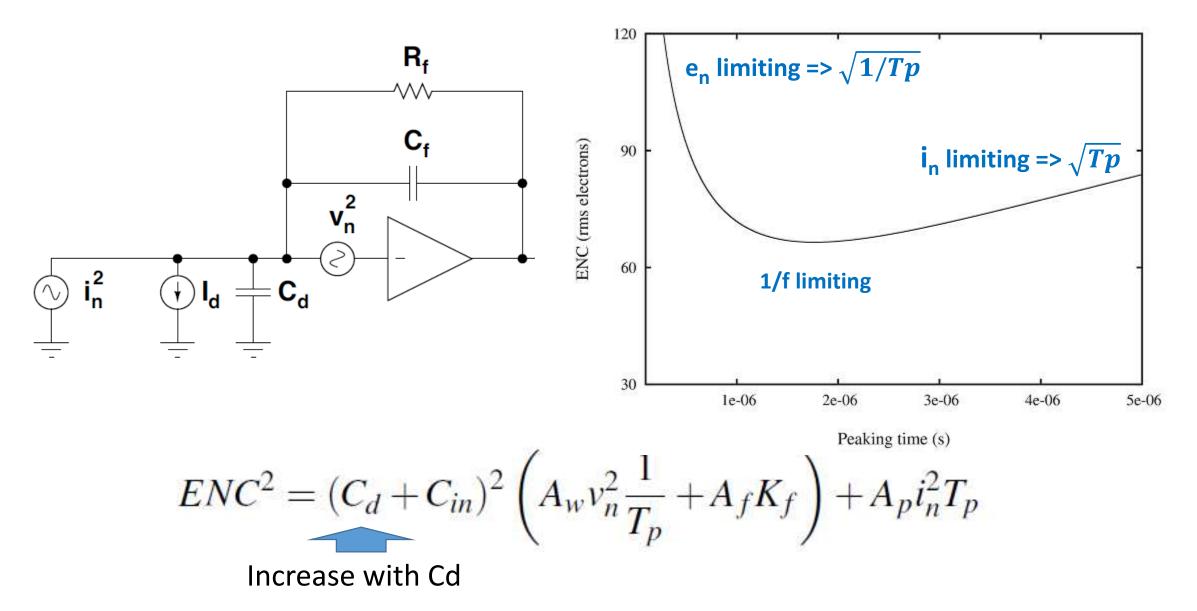
The methode here is to Transfert each input To current domain, then Multiply by the feedback



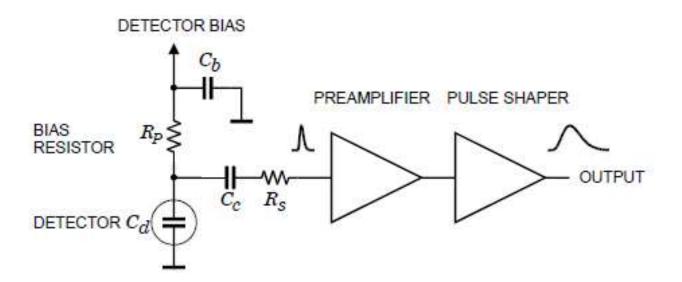
Output spectrum and it's equivalent input



Noise issues in charge preamp: time domaine

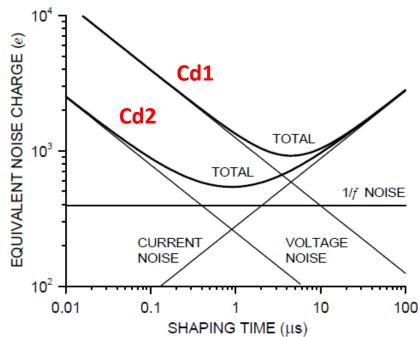


Practical summary about noise in charge integrator



ENC using a simple CR-RC shaper with peaking time T:

$$\begin{array}{c} Q_n^2 \approx \left[\left(2q_e I_d + \frac{4kT}{R_P} + i_{na}^2 \right) \cdot T \ + \ \left(4kTR_S + e_{na}^2 \right) \cdot \frac{C_d^2}{T} \ + \ 4A_f C_d^2 \right] \\ & \uparrow \qquad f \\ \text{current noise} \qquad \text{voltage noise} \qquad 1/f \text{ noise} \\ & \propto \tau \qquad & \propto 1/T \qquad \text{independent of } T \\ & \text{independent of } C_d \qquad & \propto C_d^2 \qquad & \propto C_d^2 \end{array}$$

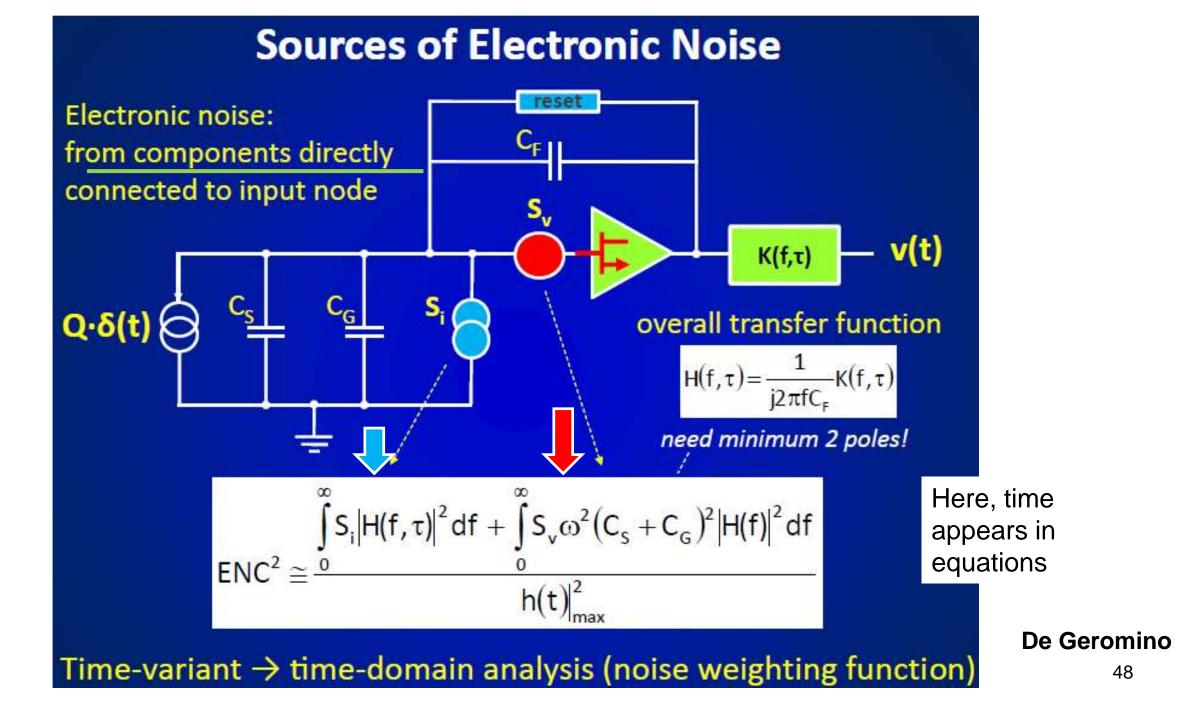


Helmuth Spieler LBNL

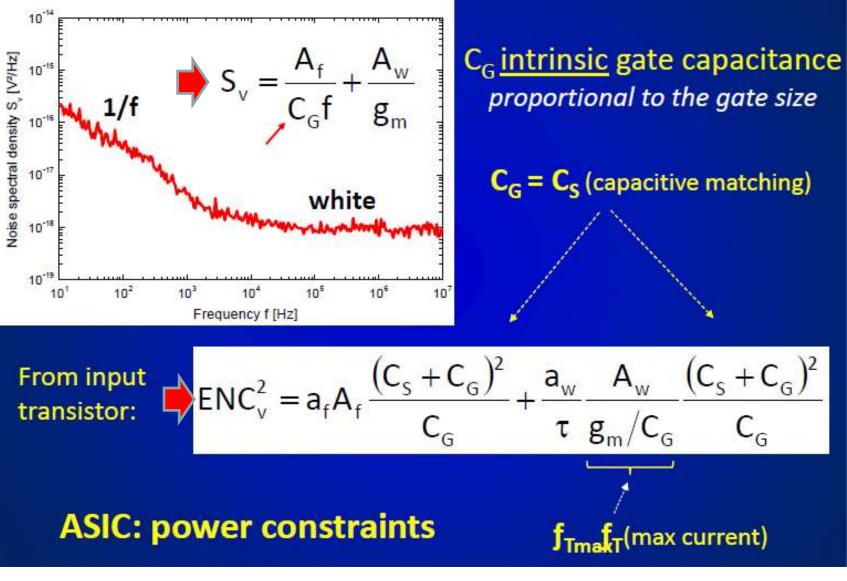
One stage preamplifier scheme

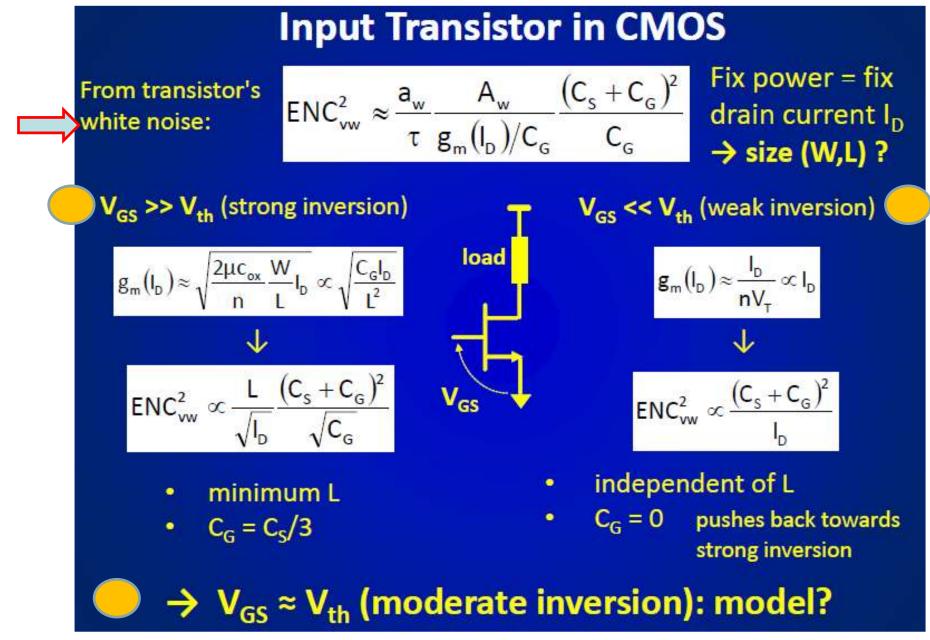
• The main contributor to the total noise is the **preamp input transistor**. We consider next the contribution of this transistor to the

equivalent noise $\mathbf{e_n}$ and $\mathbf{i_n}$



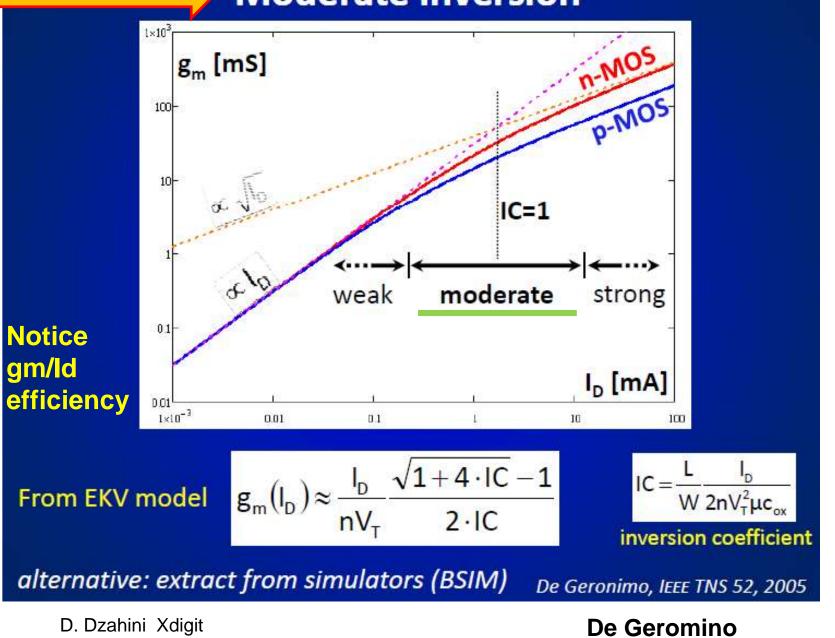
Noise from Input Transistor

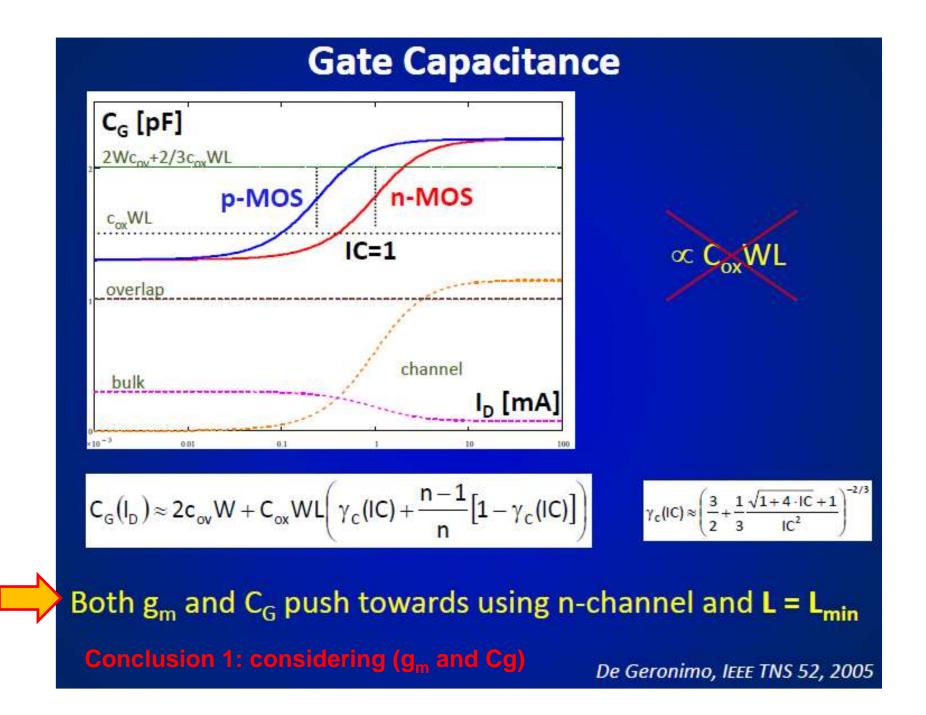




De Geromino

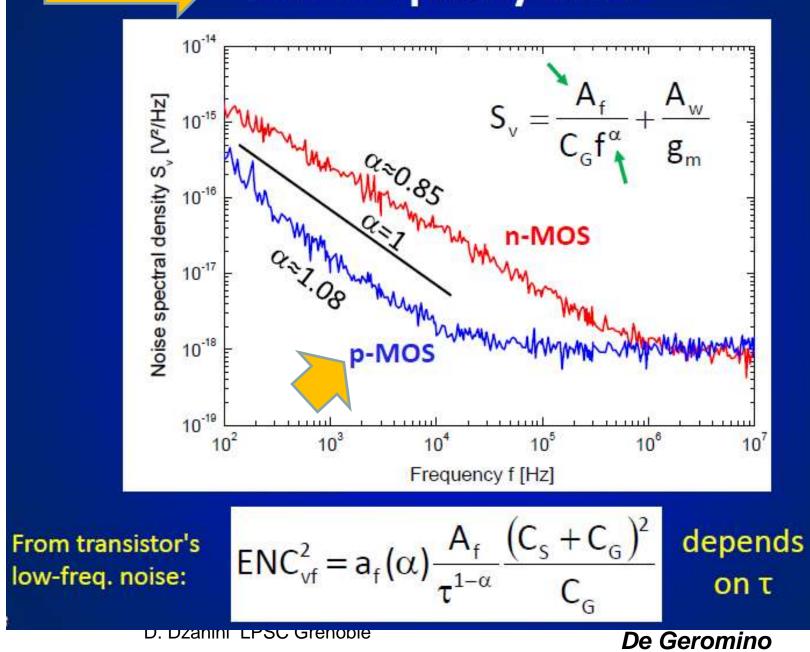
Moderate Inversion

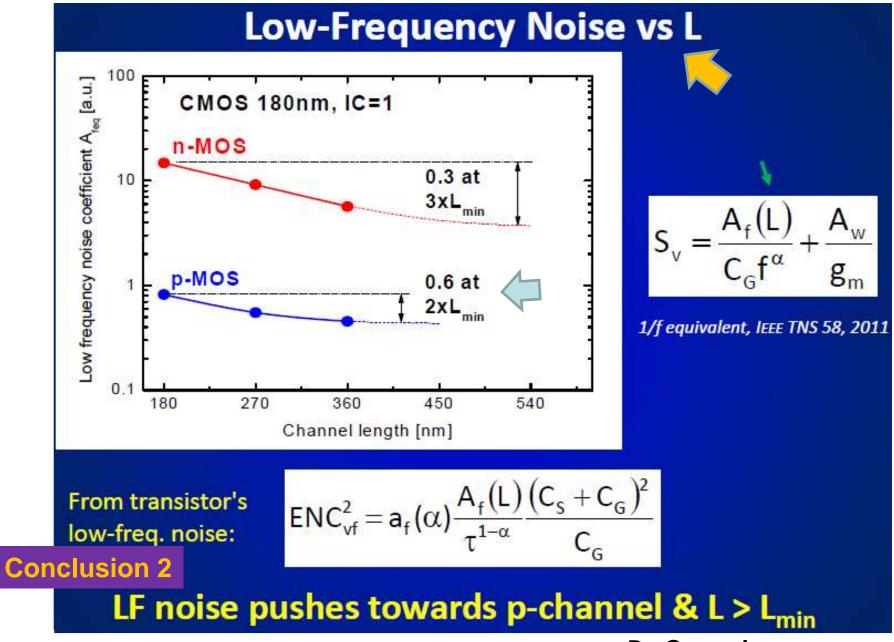




BUT, now let add the 1/f noise contribution

Low-Frequency Noise





D. Dzahini LPSC Grenoble

De Geromino

Preamp trends with agressive process

Preamp design & 'Scaling'

One may consider 2 prospection studies (old now !!)

1) Paul O'Connor => Brookhaven Lab, Upton , New York: 1,5µ au 180n

<u>At constant</u> power one save 23% in term of noise per generation: λ =0.7 ENC' = $\lambda^{3/4}$ * ENC

At constant noise: one save 60% of power per generation

 $\mathsf{P}' = \lambda^3 * \mathsf{P}$

But in dynamique range, one lose 10% of SNR per generation

SNR' = $\lambda^{1/4} *$ SNR

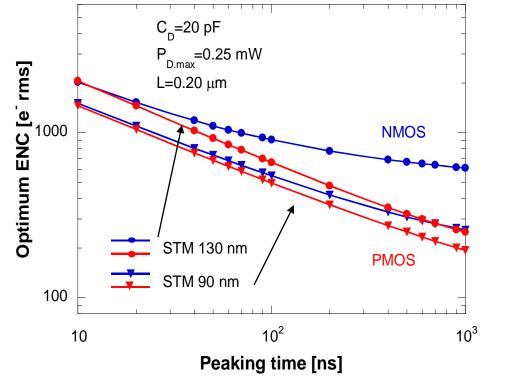
2) L. Rattia NSS 2007 => Università degli Studi di Pavia : 100n et 90n

Next slides show some of L. Rattia' work

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ENC vs peaking time, @ Pd=cte

L. Rattia NSS 2007 =>The 90nm leads to less ENC (noise) than the 130nm



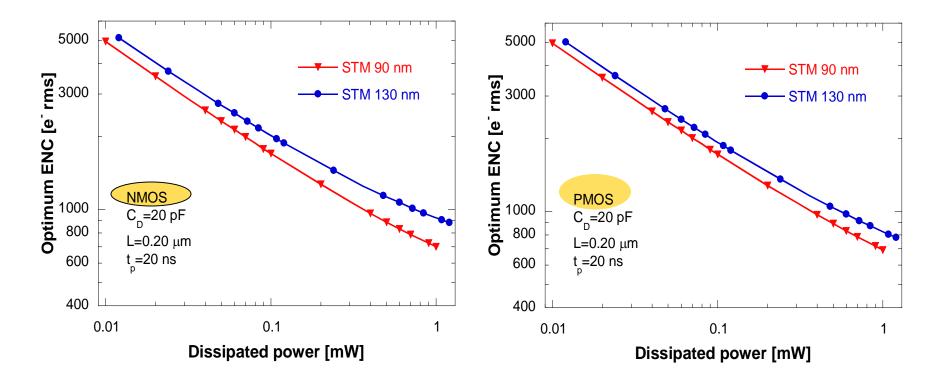
ENC was evaluated in the case of a second order, unipolar (RC²-CR) shaping processor

In the explored peaking time and power range, **PMOS input device** always provides better noise performances than NMOS input (*except for the 130 nm process at* t_p *close to 10 ns*) Using the 90 nm process may yield quite significant improvement with

respect to the 130 nm technology, especially when NMOS input charge preamplifiers are considered

ENC as function of power necessary (P et N) for given tp

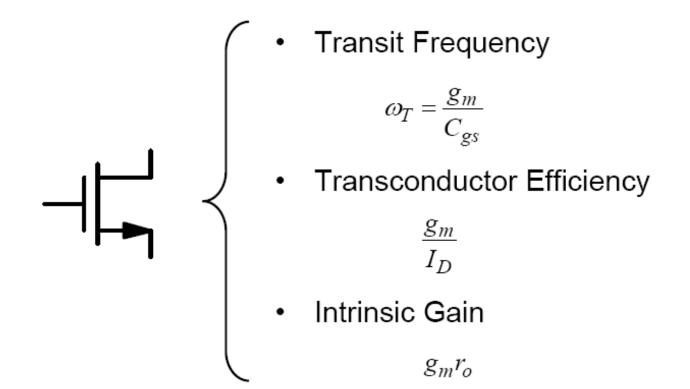
L. Rattia NSS 2007 => The 90nm needs less power at a constant noise level



At $t_p=20$ ns, noise performances provided by NMOS and PMOS input devices in the 90 nm technology are comparable

Better noise-power trade-off can be achieved by using the 90 nm technology

Figure of Merit for a MOS Process

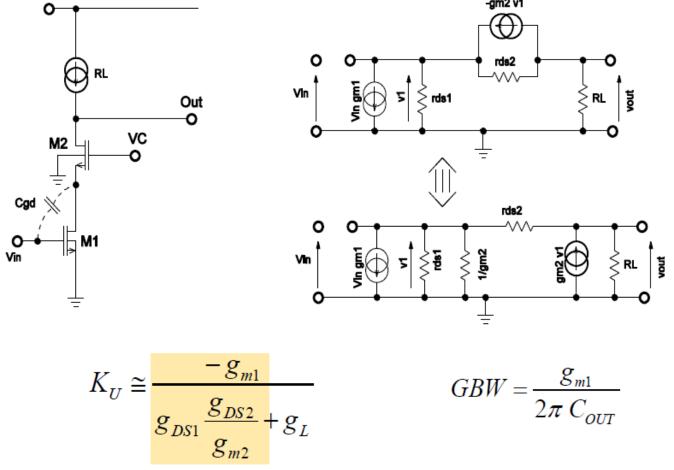


DO NOT FORGET: You may optimize the *bias point of input transistor* (moderate inversion, for a gm/ld efficiency)

HOW THE AMPLIFIER STAGE IS BUILT USING TRANSISTORS

Exemple of single stage amplifier

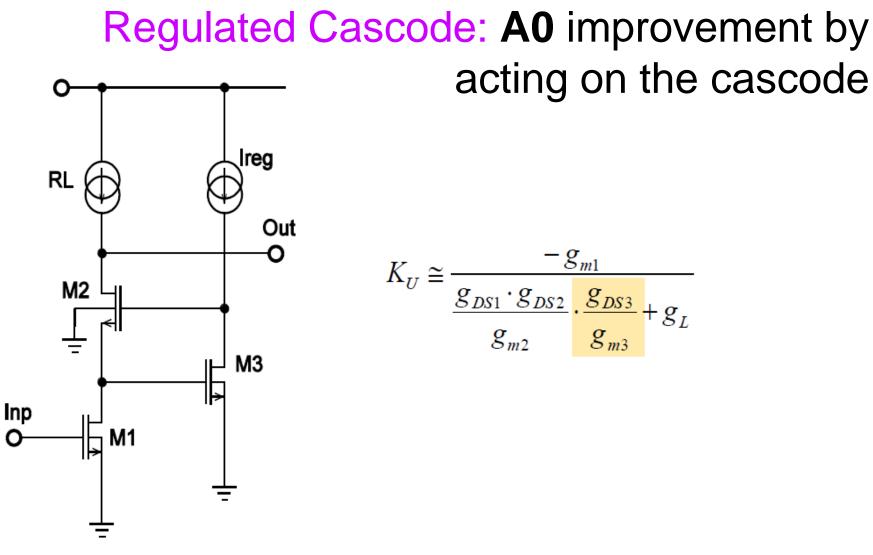
Telescopic Cascode: Common source-common gain



Courtesy Y. Kaplon, Cern

single stage amplifier; one dominant pole
 no Miller effect (low gain of common source stage)
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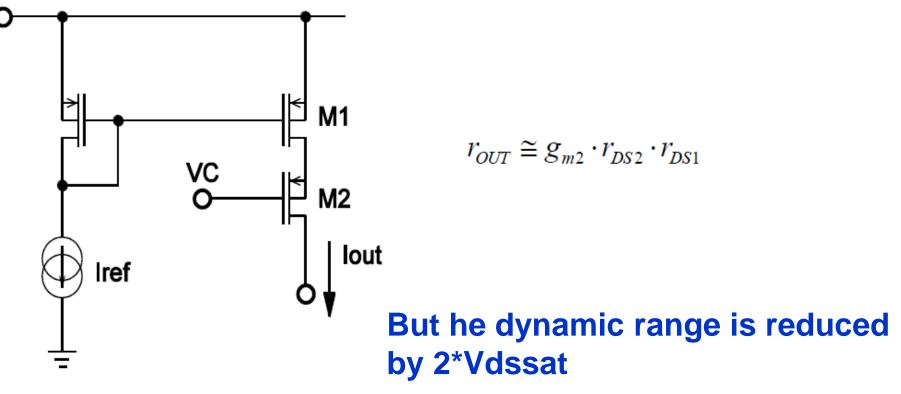
64



Cascode transistor controlled with common source amplifier
 Higher output conductance of cascode; possible higher gain
 GBW the same as for simple cascode

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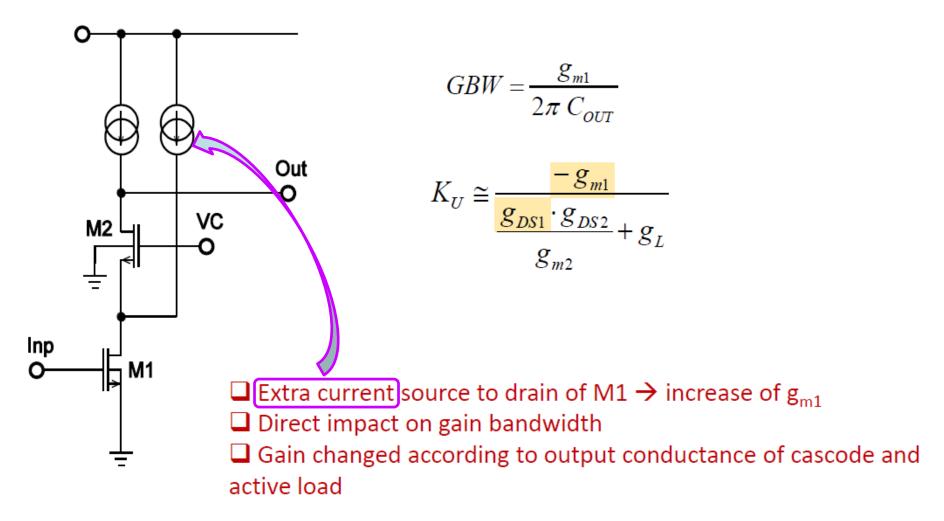
Cascode current: gain bosting via the load



Amplification of r_{DS1} by g_{m2}
 For short SSD application; OK for 250nm, not sufficient for 130 & 90nm

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Boosting both Gain & bandwidth



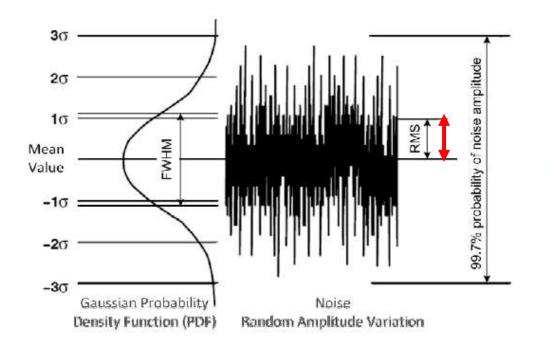
Courtesy Y. Kaplon, Cern

After the amplification schemes, let introduce how electronics noise looks like in the readout flow

NOISE EFFECT IN THE FLOW

Noise: Time domain characteristics

Noise: Time Domain Characteristics



- Noise amplitudes <u>vary randomly</u> with time.
- Noise can only be specified by a <u>Probability Density Function (PDF)</u>.
- Thermal noise and shot noise have <u>Gaussian PDFs</u>.
- Theoretically Gaussian noise amplitudes may have <u>values</u> <u>approaching infinity</u>.

Topics for later sessions

How can one characterize a comparator for counting accurately?

In couting flow: Frequency of noise hits (**fn**); Threshold **Vth**; input noise (**vn**) ...

S. O. Rice *Mathematical analysis of random noise* [1945] Bell System Technical journal, 24; 46-156

- How often are noisy events counted?
- Noise at your comparator input?
- Threshold value above the baseline?
- Counting rate and so your bandwidth or τ ?

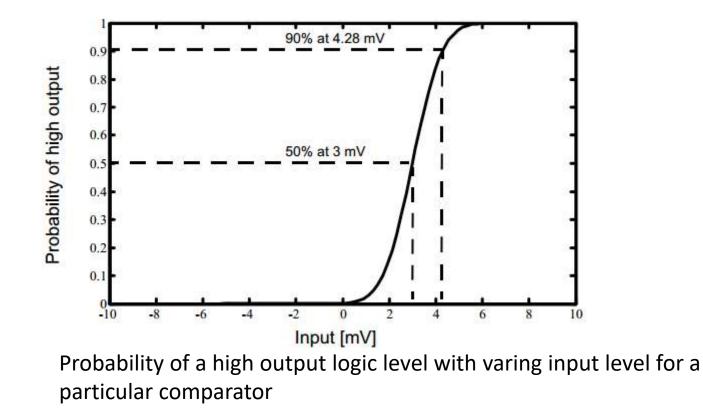
EXO: estimate f_n for V_{th}/V_n =7 for different τ

$$f_n = \frac{1}{2\pi\tau} e^{-(Vth^2/2Vn^2)}$$

Or Threshold over noise ratio

$$\frac{Vth}{Vn} = \sqrt{-2\ln(fn * 2\pi \tau)}$$

Cumulative distribution function (CDF) The comparator issue using S-curve



The 90% confidence interval for

A normal distribution is at **1.3** (or 1.2816) times its standard deviation (rms) In this exemple 90% correspond to 4.28mV; hence the rms input noise will be (4.28mV - 3mV)/1.2816 = 1mV rms

Cumulative Distribution Function

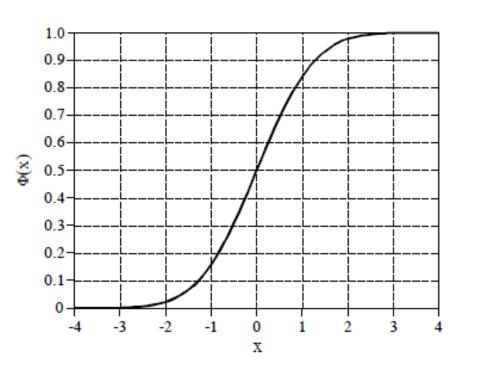


FIGURE 2-9 & TABLE 2-5

 $\Phi(x)$, the cumulative distribution function of the normal distribution (mean = 0, standard deviation = 1). These values are calculated by numerically integrating the normal distribution shown in Fig. 2-8b. In words, $\Phi(x)$ is the probability that the value of a normally distributed signal, at some randomly chosen time, will be less than *x*. In this table, the value of *x* is expressed in units of standard deviations referenced to the mean.

x	Ф(x)	х	Φ(x)	
-3.4 -3.3 -3.2 -3.1 -3.0 -2.9 -2.8 -2.7 -2.6 -2.5 -2.4 -2.3 -2.2 -2.1 -2.0 -1.9 -1.8 -1.7 -1.6 -1.5 -1.4 -1.3 -1.2 -1.1 -1.0 -0.9 -0.8 -0.7 -0.6 -0.5 -0.4 -0.3 -0.2 -0.1 0.0	.0003 .0005 .0007 .0010 .0013 .0019 .0026 .0035 .0047 .0062 .0082 .0107 .0139 .0179 .0228 .0287 .0359 .0446 .0548 .0287 .0359 .0446 .0548 .0668 .0808 .0968 .1151 .1357 .1587 .1841 .1357 .1587 .1841 .2119 .2420 .2743 .3085 .3446 .3821 .4207 .4602 .5000	$\begin{array}{c} 0.0\\ 0.1\\ 0.2\\ 0.3\\ 0.4\\ 0.5\\ 0.6\\ 0.7\\ 0.8\\ 0.9\\ 1.0\\ 1.1\\ 1.2\\ 1.3\\ 1.4\\ 1.5\\ 1.6\\ 1.7\\ 1.8\\ 1.9\\ 2.0\\ 2.1\\ 2.2\\ 2.3\\ 2.4\\ 2.5\\ 2.6\\ 2.7\\ 2.8\\ 2.9\\ 3.0\\ 3.1\\ 3.2\\ 3.3\\ 3.4\end{array}$.5000 .5398 .5793 .6179 .6554 .6915 .7257 .7580 .7881 .8159 .8413 .8643 .8849 .9032 .9192 .9332 .9192 .9332 .9452 .9554 .9641 .9713 .9772 .9821 .9641 .9713 .9772 .9821 .9861 .9893 .9918 .9938 .9953 .9955 .9974 .9981 .9987 .9990 .9993 .9995 .9997	

Cumulative **D**istribution Function (CDF) The comparator issue using S-curve

The mean value μ could correspond to: The Vth ± offset

1.0 $\mu = 0, \sigma^2 = 0.2, -$ 0.8 $\mu = -2, \sigma^2 = 0.5.$ $\Phi_{\mu,\sigma^2}(\chi)$ Very approximative 0.2 0.0 -1 0 2 3 5 Х

EXO: Can you calculate a more accurate value using 90% ->1.3