INTERACTION OF PARTICLES WITH MATTER

LECTURE 3

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I. INTRODUCTION

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I.I HADRON COLLIDER ENVIRONMENT

- High energetic collisions occur at the interaction vertex
 - Heavy SM (or beyond-the-SM) particles cascade-decay
 - The interaction of secondary particles with the detector material is governed by the SM at low energies.
- Many aspects of detector technology and conceptual design are governed by the need to isolate leptons (e/μ) in the dense environment of the rest of the event.
 - E.g. leptons from e.g. the decay of W/Z bosons can be isolated and very energetic (~100 GeV)
- QCD multijet production and hadronic decay modes of t/W/Z/h produce highly energetic strongly interacting particles (quarks, gluons) that shower and 'hadronize', producing jets (σ(QCD)/σ(tt) ≈ 10¹⁰) of relatively low energetic particles.
 - At hadron colliders, simultaneous collisions happening at the same time ('pile-up) are always significant. LHC in 2016: <PU>≈15-40!
- Need to find the isolated, energetic particles in a very dense environment of low energetic background



I.2 JET CONSTITUENTS

physical evolution



simulated particle content of a hadronic jet



- I. The average jet composition does not depend on E
- 2. The interaction of the decay products with the detector is governed by Q, E (p), m and τ
- \rightarrow enormous simplification for data analysis

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I.3 A FEW OF THE KNOWN SM PARTICLES

This short table gives the name, the quantum numbers (where known), and the status of baryons in the Review. Only the baryons with 3- or

Many hundreds of particles are known

Good
 experimental
 understanding of
 decay branching
 ratios

p 1/2 ⁺¹ n 1/2 ⁺¹ N(1440) 1/2 ⁺¹ N(1520) 3/2 ⁻¹ N(1555) 1/2 ⁻¹ N(1655) 5/2 ⁻¹ N(1675) 5/2 ⁻¹ N(1685) N(1700) N(1700) 3/2 ⁻¹ N(1720) 3/2 ⁺¹	· ****	$\Delta(1232)$ $\Delta(1600)$ $\Delta(1620)$ $\Delta(1750)$ $\Delta(1900)$ $\Delta(1905)$ $\Delta(1910)$ $\Delta(1920)$ $\Delta(1920)$ $\Delta(1940)$ $\Delta(1940)$	$3/2^+ ***$ $3/2^- ***$ $1/2^- ***$ $1/2^+ *$ $1/2^- **$ $1/2^+ *$ $1/2^+ ***$ $3/2^+ ***$ $3/2^- ***$ $3/2^- ***$ $3/2^+ ***$	* Σ^+ Σ^0 * Σ^- * $\Sigma(1385)$ $\Sigma(1480)$ * $\Sigma(1560)$ * $\Sigma(1620)$ $\Sigma(1670)$ $\Sigma(1670)$ $\Sigma(1670)$ * $\Sigma(1670)$ * $\Sigma(1670)$ * $\Sigma(1750)$	1/2+ **** 1/2+ **** 3/2+ **** ** 1/2- * 1/2- * 1/2+ *** 3/2- * 1/2+ **** **	Ξ^{0} Ξ^{-} $\Xi(1530)$ $\Xi(1620)$ $\Xi(1690)$ $\Xi(1950)$ $\Xi(2030)$ $\Xi(2120)$ $\Xi(2250)$ $\Xi(2370)$ $\Xi(2500)$	$\frac{1/2^{+}}{1/2^{+}}$ $3/2^{-}$ $\geq \frac{5}{2}$?	**** * * * * * * * * * * * * * * * * *	$\begin{array}{c} \Lambda_{c}^{+} \\ \Lambda_{c}(2595)^{+} \\ \Lambda_{c}(2625)^{+} \\ \Lambda_{c}(2765)^{+} \\ \Lambda_{c}(2800)^{+} \\ \Sigma_{c}(2455) \\ \Sigma_{c}(2455) \\ \Sigma_{c}(2520) \\ \Sigma_{c}(2520) \\ \Sigma_{c}^{-}(2800) \\ \Xi_{c}^{+} \\ \Xi_{c}^{0} \\ \Xi_{c}^{-} \end{array}$	$1/2^+$ $1/2^-$ $3/2^-$ $5/2^+$ $1/2^+$ $3/2^+$ $1/2^+$ $1/2^+$ $1/2^+$ $1/2^+$	**** *** *** *** *** *** *** *** *** *	$\begin{array}{c} \pi^{\pm} \\ \pi^{0} \\ \eta \\ \epsilon_{0}(500) \\ \rho(770) \\ \omega(782) \\ \epsilon_{0}(980) \\ \epsilon_{0}(980) \\ \epsilon_{0}(980) \\ \epsilon_{0}(102) \\ \epsilon_{0}(111) \\ \epsilon_{0}(122) \\ \epsilon_{0}(12) \\ \epsilon_$
N(180) 5/2 ⁻¹ N(180) 5/2 ⁻¹ N(1875) 3/2 ⁻¹ N(1895) 1/2 ⁻¹ N(1990) 3/2 ⁻¹ N(1990) 7/2 ⁻¹ N(2000) 5/2 ⁻¹ N(2000) 5/2 ⁻¹ N(2000) 5/2 ⁻¹ N(2100) 1/2 ⁻¹ N(2120) 3/2 ⁻¹ N(2120) 3/2 ⁻¹ N(2250) 9/2 ⁻¹ N(2250) 9/2 ⁻¹ N(2300) 1/2 ⁻¹ N(2570) 5/2 ⁻¹ N(2570) 5/2 ⁻¹ N(2570) 13/2		A(1300) A(2000) A(2150) A(2200) A(2300) A(2300) A(2420) A(2420) A(2420) A(2420) A(2420) A(2420) A(2420) A(2420) A(2520) A(1652) A(1650) A(1660) A(1680) A(1810) A(1810) A(1830) A(1830) A(2000) A(2100) A(2110) A(2325)	$5/2^+$ ** $5/2^-$ * $7/2^-$ * $9/2^+$ ** $5/2^-$ * $1/2^+$ ** $1/2^+$ ** $1/2^+$ ** $1/2^+$ ** $1/2^+$ ** $1/2^+$ ** $1/2^+$ ** $1/2^-$ ** $3/2^-$ ** $1/2^+$ ** $3/2^-$	$\begin{array}{c} \Sigma(1770)\\ \Sigma(1770)\\ \Sigma(1775)\\ \Sigma(1880)\\ \Sigma(1915)\\ \Sigma(1940)\\ \Sigma(2000)\\ *\\ \Sigma(2000)\\ *\\ \Sigma(2000)\\ \times\\ \Sigma(3170)\\ \times\\ \times\\ \ast\\ \ast\\$	$1/2^+$ * $5/2^-$ **** $3/2^+$ * $1/2^+$ ** $3/2^-$ **** $1/2^-$ * $7/2^+$ **** $5/2^+$ * $3/2^+$ ** $7/2^-$ * ** ** * * * *	Ω ⁻ Ω(2250) ⁻ Ω(2380) ⁻ Ω(2470) ⁻	3/2+	****	$ \begin{array}{l} = c^{*}_{c} \\ \equiv c^{*}_{c} \\ \equiv c(2645) \\ \equiv c(2790) \\ \equiv c(2815) \\ \equiv c(2930) \\ \equiv c(2980) \\ \equiv c^{*}_{c}(3026) \\ \equiv c^{*}_{c}(30280) \\ \equiv c^{*}_{c}(30280) \\ \equiv c^{*}_{c}(30280) \\ \equiv c^{*}_{c}(3123) \\ \Omega^{0}_{c} \\ \Omega^{0}_{c} \\ (2770)^{0} \\ \equiv c^{*}_{c} \\ \Lambda^{0}_{b} \\ \Lambda^{0}_{b} \\ (5920)^{0} \\ \Sigma^{*}_{b} \\ \equiv c^{*}_{b} \\ \Xi^{0}_{b} \\ \equiv c^{*}_{b} \\ \Xi^{0}_{b} \\ \Xi^{$	1/2+ 1/2+ 3/2+ 1/2- 3/2- 1/2+ 3/2- 1/2+ 3/2- 1/2+ 3/2+ 1/2+ 3/2+ 1/2+ 3/2+ 1/2+ 1/2+ 1/2+ 1/2+ 1/2+ 1/2+ 1/2+ 1/2+ 1/2- 1/2+ 1/2- 1/2+ 1/2- 1/2+ 1/2- 1/2+		$ \begin{array}{c} & f_1(22) \\ & f_1(22) \\ & f_1(12) \\ & f_1(13) \\ & f_1(14) $

Baryon Summary Table

Meson Summary Table

See also the table of suggested $q\overline{q}$ quark-model assignments in the Quark Model section.

• Indicates particles that appear in the preceding Meson Summary Table. We do not regard the other entries as being established.

	LIGHT UNFLAVORED			STRAN	IGE	CHARMED, S	TRANGE	c c		
	(S = C =	= B = 0)		$(S = \pm 1, C =$	= B = 0)	(C = S =	±1)		$I^{G}(J^{PC})$	
	$I^{G}(J^{PC})$		$I^{G}(J^{PC})$		$I(J^{p})$		$I(J^p)$	• $\eta_c(1S)$	$0^{+}(0^{-+})$	
• π^{\pm}	$1^{-}(0^{-})$	• $\pi_2(1670)$	$1^{-}(2^{-+})$	• K [±]	$1/2(0^{-})$	 D[±]_s 	$0(0^{-})$	 J/ψ(1S) 	$0^{-}(1^{-})$	
 π⁰ 	$1^{-}(0^{-+})$	 \$\phi(1680) \$ \$ \$	$0^{-}(1^{-})$	• K ⁰	$1/2(0^{-})$	 D^{*±} 	0(? [?])	• $\chi_{c0}(1P)$	$0^{+}(0^{++})$	
• 7	$0^{+}(0^{-+})$	 ρ₃(1690) 	$1^{+}(3^{-})$	 K⁰_S 	$1/2(0^{-})$	 D[*]_{s0}(2317)[±] 	$0(0^{+})$	• $\chi_{c1}(1P)$	$0^{+}(1^{++})$	
 f₀ (500) 	$0^{+}(0^{+}+)$	 ρ(1700) 	$1^{+}(1^{-})$	 K⁰_L 	$1/2(0^{-})$	 D_{s1}(2460)[±] 	$0(1^+)$	 h_c(1P) 	? [?] (1 ⁺ ⁻)	
 ρ(770) 	$1^{+}(1^{-})$	$a_2(1700)$	$1^{-}(2^{++})$	K [*] ₀ (800)	$1/2(0^+)$	 D_{s1}(2536)[±] 	$0(1^+)$	• $\chi_{c2}(1P)$	$0^{+}(2^{++})$	
 ω(782) 	$0^{-}(1^{-})$	 f₀(1710) 	$0^{+}(0^{+}+)$	 K*(892) 	$1/2(1^{-})$	 D_{s2} (2573) 	$0(?^{?})$	• $\eta_c(2S)$	$0^{+}(0^{-+})$	
 η'(958) 	$0^{+}(0^{-+})$	$\eta(1760)$	$0^{+}(0^{-}+)$	 K₁(1270) 	$1/2(1^+)$	 D[*]₁(2700)[±] 	$0(1^{-})$	 ψ(2S) 	$0^{-}(1^{-})$	
 f₀(980) 	$0^{+}(0^{+}+)$	• $\pi(1800)$	$1^{-}(0^{-+})$	 K₁(1400) 	$1/2(1^+)$	D 1(2860) ±	0(??)	 ψ(3770) 	$0^{-}(1^{-})$	
 a₀(980) 	$1^{-}(0^{++})$	$f_2(1810)$	$0^+(2^{++})$	 K*(1410) 	$1/2(1^{-})$	$D_{sJ}(3040)^{\pm}$	$0(?^{?})$	 X(3872) 	$0^{+}(1^{++})$	
 \$\phi(1020)\$ 	$0^{-}(1^{-})$	X(1835)	? [?] (? ⁻ +)	 K[*]₀(1430) 	$1/2(0^+)$			• $\chi_{c0}(2P)$	$0^{+}(0^{++})$	
 h₁(1170) 	$0^{-}(1^{+})$	• $\phi_3(1850)$	0 (3 -)	 K[*]₂(1430) 	$1/2(2^+)$	BOTTO	M	• $\chi_{c2}(2P)$	$0^{+}(2^{++})$	
 b₁(1235) 	$1^{+}(1^{+})$	$\eta_2(1870)$	$0^{+}(2^{-+})$	K(1460)	$1/2(0^{-})$	$(B = \pm$	1)	X(3940)	5t(5tt)	
 a₁(1260) 	$1^{-}(1^{++})$	• $\pi_2(1880)$	$1^{-}(2^{-+})$	$K_2(1580)$	$1/2(2^{-})$	• B [±]	$1/2(0^{-})$	 ψ(4040) 	0-(1)	
 f₂(1270) 	$0^{+}(2^{++})$	$\rho(1900)$	$1^+(1^-)$	K(1630)	$1/2(?^{?})$	• B ⁰	$1/2(0^{-})$	$X(4050)^{\pm}$?(?*)	
 f₁(1285) 	$0^{+}(1^{++})$	$f_2(1910)$	$0^{+}(2^{++})$	$K_1(1650)$	$1/2(1^+)$	 B[±]/B⁰ ADM 	IXTURE	X(4140)	$0^+(?^{;+})$	
 η(1295) 	$0^{+}(0^{-+})$	 f₂(1950) 	$0^{+}(2^{++})$	 K*(1680) 	$1/2(1^{-})$	 B[±]/B⁰/B⁰_s/l 	-baryon	• $\psi(4160)$	0-(1)	
• $\pi(1300)$	$1^{-}(0^{-+})$	$\rho_3(1990)$	$1^{+}(3^{-})$	 K₂(1770) 	$1/2(2^{-})$	ADMIXTURE	KM Ma	X(4160)	21(211)	
 a₂(1320) 	$1^{-}(2^{++})$	 f₂(2010) 	$0^+(2^{++})$	 K[*]₃(1780) 	$1/2(3^{-})$	trix Elements	TYNE MIG.	$X(4250)^{\pm}$	2(2+)	
 f₀(1370) 	$0^+(0^++)$	$f_0(2020)$	$0^+(0^{++})$	 K₂(1820) 	$1/2(2^{-})$	• B*	$1/2(1^{-})$	• X(4260)	?*(1)	
$h_1(1380)$?=(1 + -)	 a₄(2040) 	$1^{-}(4^{+}^{+})$	K(1830)	$1/2(0^{-})$	B [*] _J (5732)	?(? [?])	X(4350)	$0^+(?^+)$	
• $\pi_1(1400)$	1 (1 ')	 f₄(2050) 	$0^+(4^++)$	$K_0^*(1950)$	$1/2(0^+)$	 B₁(5721)⁰ 	$1/2(1^+)$	• X (4360)	(1) a=(1 = -)	
 η(1405) 	0+(0-+)	$\pi_2(2100)$	$1^{-}(2^{-+})$	$K_{2}^{*}(1980)$	$1/2(2^+)$	 B[*]₂(5747)⁰ 	$1/2(2^+)$	 ψ(4415) 	0 (1)	
 f₁(1420) 	$0^+(1^+)$	$f_0(2100)$	$0^+(0^++)$	 K[*]₄(2045) 	$1/2(4^+)$	DOTTOM C	DANCE	X(4430)+	r(r)	
 ω(1420) ω(1420) 	$0^{-}(1^{-})$	$f_2(2150)$	$0^+(2^+)$	$K_2(2250)$	$1/2(2^{-})$	BUTTOM, S	- T1)	• X (4660)	(· (1)	
f2(1430)	$0^{-}(2^{+})$	$\rho(2150)$	1 (1)	$K_3(2320)$	$1/2(3^+)$	(0 = ±1, 0	- +1)	h	<u>b</u>	
 a₀(1450) a(1450) 	$1 (0 \cdot \cdot)$ $1 \pm (1)$	 φ(2170) ζ (2270) 	0 (1)	K [*] ₅ (2380)	$1/2(5^{-})$	• B ^o _S	0(0)	n (15)	$n \pm (n = \pm)$	
 ρ(1450) ρ(147E) 	$a \pm (a = \pm)$	$f_0(2200)$	$0^+(0^+)$	$K_4(2500)$	$1/2(4^{-})$	• B [*] ₅	0(1)	• T(15)	$0^{-}(1^{-})$	
• f (1600)	$a^{\pm}(a^{\pm}\pm)$	rj(2220)	0.(2	K(3100)	? [?] (? ^{??})	 B_{s1}(5830)^o D_s (5830)^o 	$0(1^{+})$	• Yw(1P)	$0^{+}(0^{+}+1)$	
• f(1500)	$0^+(1^++)$	m(200E)	$a \pm (a = \pm)$	CUADA	150	• B _{s2} (5840)*	0(21)	• X _{b1} (1P)	$0^{+}(1^{+}+)$	
f(1510)	$0^+(2^++)$	n(2225)	$1^{+}(2^{-})$	CHARIN	(ED	$B_{sJ}^{*}(5850)$?(?•)	• h _b (1 P)	$2^{?}(1 + -)$	
f (1666)	$a^{\pm}(2^{\pm}\pm)$	$p_3(2250)$	$a^{\pm}(3^{\pm}\pm)$	(c = 1)	- (0(0-)	BOTTOM, C	ARMED	• Yea(1P)	$0^+(2^{++})$	
o(1570)	1+(1)	$f_2(2300)$ $f_2(2300)$	$0^{+}(4^{+}+)$	• D [±]	1/2(0)	(B = C =	±1)	$n_b(2S)$	$0^{+}(0^{-}+)$	
p(1570) $b_{1}(1505)$	$0^{-}(1^{+})$	f ₄ (2330)	$0^{+}(0^{+}+)$	• D*	1/2(0)	• B [±]	$0(0^{-})$	 T(25) 	$0^{-}(1^{-}-)$	
 π₁(1595) π₁(1600) 	1 = (1 = +)	• f ₀ (2340)	$0^+(2^{++})$	• D*(2007)*	1/2(1)	c	-(-)	 T(1D) 	$0^{-}(2^{-})$	
a ₁ (1640)	1-(1++)	c (2350)	$1^{+}(5^{-})$	• D*(2010)*	1/2(1)			• Xh0(2P)	$0^{+}(0^{+}+)$	
£(1640)	$0^{+}(2^{+}+)$	$a_{2}(2450)$	1-(6++)	• D ₀ (2400)*	1/2(0 ')			• $\chi_{h1}(2P)$	$0^{+}(1^{+}^{+})$	
• n (1645)	$0^+(2^-+)$	£(2510)	$0^+(6^{++})$	$D_0^-(2400)^-$	1/2(0 ')			$h_b(2P)$	$?^{?}(1 + -)$	
• $\omega(1650)$	$0^{-}(1^{-})$.0(====)	- (-)	• D ₁ (2420) ⁺	1/2(11)			• $\chi_{b2}(2P)$	$0^{+}(2^{++})$	
• (a) (1670)	0-(3)	OTHER	LIGHT	$D_1(2420)^{-1}$	$1/2(f^{+})$			 <i>\(\Gamma\)</i> (35) 	$0^{-}(1^{-})$	
	- (-)	Further St	ates	$D_1(2430)^{\circ}$	$1/2(1^{+})$			 <i>χ</i>_b(3P) 	? [?] (? ^{?+})	
				• D ₂ (2460)*	1/2(2 ')			 <i>𝔅</i>(4<i></i>) 	$0^{-}(1^{-})$	
				 D[*]₂(2460)[±] D(25,50)⁰ 	1/2(21)			$X(10610)^{\pm}$	$?^{+}(1^{+})$	
				D(2550)"	1/2(0)			$X(10650)^{\pm}$	$?^{+}(1^{+})$	
				D(2600)	1/2(11)			 <i>𝔅</i>(10860) 	$0^{-}(1^{-})$	
				D*(2640)*	1/2(11)			 <i>𝔅</i>(11020) 	$0^{-}(1^{-})$	
				D(2750)	$1/2(2^{-})$					

I.4 WHEN CAN WE DETECT A PARTICLE?

- Particles can only be measured, if they
- A. live long enough after creation to reach the detector
 - The majority of particle states are short lived. From the hundreds of particles only few particles (and their antiparticles) have track lengths long enough to measure them:

	γ	р	n	e±	μ^{\pm}	π^{\pm}	K±	$\mathbf{K}_{0} \; (\mathbf{K}_{\mathrm{S}}/\mathbf{K}_{\mathrm{L}})$
τ ₀	8	8	8	8	2.2 μ s	26 ns	12 ns	89 ps / 51 ns
l _{track} (p=1GeV)	8	8	8	8	6.1 km	5.5 m	6.4 m	5 cm / 27.5 m

- Track length: $L_{track} = v\tau = c\beta\gamma\tau_0$ where τ_0 being the lifetime at rest.
- small but finite life-time: B-mesons have ps lifetimes
 → sub-mm impact parameter (IP) resolution allows tagging of b-quarks!
- B. interact with the detector
 - deposition of energy (dE/dx), transferred into a detector signal
 - neutrinos interact only weakly: need large detector volumes (e.g. IceCube) or giant fluxes (reactor experiments)



I.5 OVERVIEW OF INTERACTIONS IN THE DETECTOR

- Different type of interactions for charged and neutral particles
- Difference "scale" of processes:
 - atomic scale for the electromagnetic and nuclear scale for the strong interactions
- I. Detection of charged particles
 - Ionization, Bremsstrahlung, Cherenkov ...
- 2. Detection of γ -rays
 - Photo/Compton effect, pair production
- 3. Detection of neutrons
 - strong interaction
- 4. Detection of neutrinos
 - weak interaction

 In the following, a phenomenological treatment is given with a focus on the implications for detector design!

I.6 DETECTOR DESIGN CRITERIA

- I. little interaction with the measured particle
 - tracking detectors should trace the passage of a charged particle without disturbing it
- 2. high efficiency
 - probability of detection in case of a signal particle
- 3. high purity (high signal-to-noise ratio)
 - low probability of instrumental noise and unintended signal
- 4. high resolution
 - spatial, time, energy, momentum, angle, ...
- 5. fast signal processing
- 6. simple maintenance and detector control
- 7. radiation hardness
- 8. low cost

• there is no detector that satisfies 1.-8.

 K_0 .

• a well established compromise with existing technology:



I.7 ABSORPTION OF PARTICLES AND RAYS

- Particles can either be a) consumed by a single reaction or b) continuously loose energy until stopped
- a) Particle loss by absorption (dN/dx, typical for photons) leads to an exponential decay of particle intensity





 b) Energy loss through matter for heavy charged particles (dE/dx) results in a limited area with a strong drop in intensity Charged particles



Alpha particles, 1926



I.7 ABSORPTION OF PARTICLES AND RAYS

• Consider a thin detector volume $V = F \cdot l$ with a number N_T

particles calculated as
$$N_T = rac{
ho V}{M_T} N_A\,$$
 . The density $oldsymbol{
ho}$ is

given in g cm⁻³, the relative atomic mass M_T in mol⁻¹, and $N_A = 6.022 \cdot 10^{23} \text{mol}^{-1}$ is Avogadro's constant.

- The target density is $n = \frac{N_T}{V} = \frac{\rho}{A} N_A$ where A is the nuclear number.
- The beam sees a total cross section of $N_T \sigma$ such that the probability of a collision becomes

$$w = \frac{\dot{N}_R}{\dot{N}_{in}} = \frac{N_T \sigma}{F} = n\sigma l$$

• The cross-section can be calculated as $\sigma = \frac{N_R}{\dot{N}_{in}} \frac{1}{nl}$



• In a thick volume we have the differential equation $\frac{\mathrm{d}N}{N} = -n\sigma\mathrm{d}x$

with solution $N(x) = N_0 e^{-\mu x}$ (Beer-Lambert) and $\lambda = \mu^{-1} = (n\sigma)^{-1}$ is the mean free path.

• For cross-sections we have $(r_0 \approx 1.2 \text{fm}, a_0 = 0.5 \text{\AA})$ $\sigma_{\text{nucl}} \approx \pi r_0^2 A^{2/3} \approx 45 \text{mb} \cdot A^{2/3}$ $\sigma_{\text{atom}} \approx \pi a_0^2 \approx 10^8 \text{b}$

2. INTERACTION OF CHARGED PARTICLES WITH MATTER

2.1 OVERVIEW

- I. Ionization (or excitation) of detector material
 - The incoming particle loses energy by ionizing or exciting the atoms
- 2. Interaction with the nucleus
 - The particle is deflected (scattered) causing multiple scattering of the particle in the material. During this scattering a bremsstrahlung photon can be emitted.
- 3. Emission of Cherenkov light
 - In case the particle's velocity is larger than the velocity of light in the medium, the resulting EM shockwave manifests itself as Cherenkov Radiation.
- 4. Emission of transition radiation
 - When the particle crosses the boundary between two refractive media, there is a probability of the order of ≈1% to produced and X ray photon, called Transition radiation.



2.2 MAXIMUM ENERGY TRANSFER

• Simple rearrangements of energy momentum conservation $P + p_e = P' + p'_e \qquad P \cdot p'_e = P' \cdot p_e$ $(P - p'_e)^2 = (P' - p_e)^2 \qquad EE'_e - |\vec{P}||\vec{p}_e'|c^2\cos\theta = E'm_ec^2$ lead to $\cos\theta = \frac{EE'_e - E'm_ec^2}{c^2|\vec{P}||\vec{p}_e'|}$ where $\cos\theta$ is the angle of the scattered p_e

electron with respect to the direction of the heavy charged particle.

• From the definition of the kinetic energy $T = E'_e - m_e c^2$ it follows that E = E' + T. We express the electron momentum by using the kin. energy as $E'_e{}^2 = c^2 \vec{p_e}'^2 + m_e^2 c^4$

It follows that

$$c^{2}\vec{p_{e}}^{\prime 2} = E_{e}^{\prime 2} - m_{e}^{2}c^{4} = T^{2} + 2m_{e}c^{2}T$$

$$cos \theta = \frac{\gamma M c^{2}(T + m_{e}c^{2}) - (E - T)m_{e}c^{2}}{c^{2}\beta\gamma M\sqrt{T^{2} + 2m_{e}c^{2}T}} = \frac{T\left(\gamma + \frac{m_{e}}{M}\right)}{\beta\gamma\sqrt{T^{2} + 2m_{e}c^{2}T}}$$

$$c|\vec{p_{e}}'| = \sqrt{T^{2} + 2m_{e}c^{2}T}$$
or
$$T(\theta) = \frac{2m_{e}c^{2}\beta^{2}\gamma^{2}\cos^{2}\theta}{\gamma^{2}(1 - \beta^{2}\cos^{2}\theta) + 2\gamma\frac{m_{e}}{M} + \frac{m_{e}^{2}}{M^{2}}} \quad \text{with} \quad T_{\max} = \frac{2m_{e}c^{2}\beta^{2}\gamma^{2}}{1 + 2\gamma\frac{m_{e}}{M} + \frac{m_{e}^{2}}{M^{2}}} \approx \begin{cases} 2m_{e}c^{2}(\beta\gamma)^{2} & \gamma m_{e} \ll M \\ \gamma M c^{2} = E & \gamma \to \infty \\ m_{e}c^{2}(\gamma - 1) = E - m_{e}c^{2} M = m_{e} \end{cases}$$

Maximum energy transfer if all momenta aligned, $\cos \theta = 1!$

2.3 ENERGY LOSS BY IONIZATION

- The energy loss by ionization can be described by QM: Rutherford scattering in the electron rest frame (single photon exchange)
- In the following, however, we give Bohr's classical result
 - Particle with charge ze, mass M and velocity v moves through a medium with electron density n_e.
 - Electrons considered free and initially at rest.
 - The momentum transferred to a single electron is:

$$\Delta p_{\perp} = \int F_{\perp} dt = \int F_{\perp} \frac{dt}{dx} dx = \int F_{\perp} \frac{dx}{v} , \quad \Delta p_{\parallel} = 0$$

• Furthermore $F_{\perp}=eE_{\perp}$ and we can therefore use Gauss Law (SI units)

$$\int E_{\perp}(2\pi b) \mathrm{d}x = \frac{ze}{\varepsilon_0} \to \int E_{\perp} \mathrm{d}x = \frac{ze}{2\pi\varepsilon_0 b}$$

• transferred momentum:
$$\Delta p_{\perp} = e \int E_{\perp} \frac{\mathrm{d}x}{v} = \frac{ze^2}{2\pi\varepsilon_0 bv}$$
 and energy: $\Delta E(b) = \frac{\Delta p}{2m}$



2.3 ENERGY LOSS BY IONIZATION

- Next, we integrate the loss over the electron density
- For N electrons distributed on a barrel: N = $n_e \cdot (2\pi b) \cdot db dx$
- Energy loss per path length dx for distance between ٠ b and b + db in medium with electron density n_e :

$$\begin{aligned} -\mathrm{d}E(b) &= \frac{\Delta p^2}{2m_e} 2\pi n_e b \,\mathrm{d}b \,\mathrm{d}x = \frac{n_e z^2 e^4}{4\pi \varepsilon_0^2 m_e v^2} \frac{\mathrm{d}b}{b} \mathrm{d}x \\ &- \frac{\mathrm{d}E}{\mathrm{d}x} = \frac{n_e z^2 e^4}{4\pi \varepsilon_0^2 m_e v^2} \int_{b_{\min}}^{b_{\max}} \frac{\mathrm{d}b}{b} = \frac{4\pi n_e z^2 r_e^2 m_e c^2}{\beta^2} \ln \frac{b_{\max}}{b_{\min}} \end{aligned}$$

Here, we simplified the factors with the 'classical electron radius' ٠



 $\frac{e^2}{dt} = m_e c^2 \,.$

 $4\pi\varepsilon_0 r_e$

- b is the 'impact parameter' in a single-particle interaction.
- For a finite stopping power we need to determine the relevant range of interactions [b_{min}, b_{max}] ٠

2.3.1 CLASSICAL RESULT

- From conservation of energy it follows (see 2.2) that the maximum energy transfer happens for a 'head-on' collision and is given by $T_{\text{max}} = 2m_e c^2 \beta^2 \gamma^2$ (valid for $\gamma m_e \ll M$)
- For the maximal energy transfer: $T_{\max}(b_{\min}) = \frac{\Delta p^2}{2m_e} = \frac{(ze^2)^2}{8\pi^2 \varepsilon_0^2 m_e v^2 b_{\min}^2} = \frac{2r_e^2 m_e c^2 z^2}{\beta^2 b_{\min}^2} \propto \frac{1}{b_{\min}^2}$
- Insert in the logarithm of impact parameters: $\ln \frac{b_{\max}}{b_{\min}} = \frac{1}{2} \ln \frac{b_{\max}^2}{b_{\min}^2} = \frac{1}{2} \ln \frac{T_{\max}}{T_{\min}} = \frac{1}{2} \ln \frac{2m_e c^2 \beta^2 \gamma^2}{I}$
- What about T_{min}(b_{max})? Classically, the energy transfer to an electron can become arbitrarily small. However, below a minimal material-dependent ionization threshold I, there are only discrete energy levels. Moreover, if the movement of the electron during the interaction is important, interference effects need to be taken into account that further reduce the interaction to negligible levels.

• Substituting also the electron density
$$n_e = Z \frac{\rho}{A} N_A$$
 we find:

$$-\frac{\mathrm{d}E}{\mathrm{d}x} = \underbrace{4\pi N_A r_e^2 m_e c^2}_{K=0.307 \,\mathrm{MeV \, cm^2 \, mol^{-1}}} \rho \frac{1}{\beta^2} \frac{Z z^2}{A} \frac{1}{2} \ln \frac{2m_e c^2 \beta^2 \gamma^2}{I}$$
A \equiv atomic mass in g/mol

The QM treatment based on Rutherford scattering gives

$$-\left\langle \frac{dE}{dx}\right\rangle = K \frac{Z}{A} \rho \frac{z^2}{\beta^2} \left[\frac{1}{2} \ln \frac{2m_e c^2 \beta^2 \gamma^2 T_{max}}{I^2} - \beta_{\uparrow}^2 - \frac{\delta(\beta\gamma)}{\uparrow} - \frac{C(\beta\gamma, I)}{\uparrow} \right]$$

- fundamental constants
 - $K = 4\pi N_A r_e^2 m_e c^2 = 0.307 \text{ MeV } \text{cm}^2/\text{mol}$
 - $r_e = e^2/4\pi\epsilon_0 m_e c^2 \approx 2.8\,{
 m fm}$ (the classical radius of the electron)
- properties of the incident particle
 - z, β , γ are the charge (in units of e), the velocity, and the relativistic γ factor
- properties of the target material
 - Z atomic number, A the atomic mass in g/mol
 - I is the average energy required to ionize the medium
 - T_{max} is the maximum energy in a head-on collision
 - δ is the density correction, relevant for high $m{eta}\gamma$
 - C/Z is an additional 'shell' correction for low values of β

Shell corrections

The properties of the EM field in the medium govern the density correction

Spin-1/2 interactions (Mott scattering) instead of scalar Rutherford scattering is responsible for the β^2 term.

These are the important dependencies:

$$-\left\langle \frac{\mathrm{d}E}{\mathrm{d}x}\right\rangle \approx K\frac{Z}{\beta^2}\ln(\mathrm{const}\cdot\beta\gamma)$$

- Often, dE/dx is divided by *ρ* and then given in units MeV cm² g⁻¹.
- The Bethe-Bloch equation, with its various corrections, holds in a wide range of applications for charged particles heavier than electrons.
- It describes the 'stopping power' of matter on particles.
- The plot shows π[±] on silicon. At low energies, the β⁻² term dominates, at high energies the ln(βγ).
- There is a broad minimum at $\beta \gamma = 3-3.5$ or $\beta = 0.95$.
 - 'Minimum Ionizing particles' (MIP)



• Ionization is the universal detection principle of semi-conducting sensors, gaseous detectors, etcetcetc!

•	Often, dE/dx is divided by ρ and then given	material		A	ρ	Ι	$\frac{dE}{dx} _{min}$	
	in units MeV cm ² g ⁻¹ .				$\left(\frac{g}{cm^3}\right)$	(eV)	$\left(\frac{\text{MeV cm}^2}{\text{g}}\right)$	
•	The Bethe-Bloch equation, with its various	C graphite	6	12.01	2.21	78	1.74	
	corrections holds in a wide range of applications	AI	13	26.98	2.70	166	1.62	
	corrections, holds in a wide range of applications	Si	14	28.09	2.33	173	1.66	
	for charged particles heavier than electrons.	Fe	26	55.85	7.87	286	1.45	
		Pb	82	207.2	11.35	823	1.12	
•	It describes the stopping power of matter on particles.	Ar	18	39.95	0.00166	188	1.52	
•	The plot shows π^{\pm} on silicon. At low energies	Csl	108	259.8	4.51	553	1.24	
	The plot shows h on sincon. At low energies,	polysterene	56	104.2	1.06	68.7	1.94	
	the β^{-2} term dominates, at high energies the ln(β γ).	water	10	18	1.00	79.7	1.99	
	There is a buosd minimum of $\theta_{11} = 2.2$ F as $\theta = 0.0$ F	rock	11	22	2.65	136.4	1.69	
•	There is a broad minimum at $\beta \gamma = 3-3.5$ or $\beta = 0.95$.	nuclear emulsion			3.82	331.0	1.42	

- Ionization is the universal detection principle of semi-conducting sensors, gaseous detectors, etcetcetc!
- Example: Fe 7.87 g/cm³: A MIP looses ≈ 11.41 MeV/cm

• 'Minimum Ionizing particles' (MIP)

- mass stopping power of muon over 9 orders of magnitude in momentum
- Bethe-treatment is accurate to 1% down to β≈0.05, below that there is no accurate theory. Even further below, non-ionizing nuclear recoils dominate.
- At ultra-relativistic energies, radiative losses become important.
- At the muon critical energy $E_{\mu c}$, this effect equals the ionization loss described by the Bethe-Bloch equation.



http://pdg.lbl.gov/2017//

- The 'Bethe'-region is $0.1 < \beta \gamma < 100-1000$.
- The plot shows mass stopping power in liquid (bubble chamber) hydrogen, gaseous helium, carbon, aluminum, iron, tin, and lead.
- Typical value of 2 MeV cm²/g (except H_2) for a MIP.
 - liquid H_2 has Z/A=1 which otherwise is ≈ 0.5
- Radiative effects, relevant for muons and pions, are not included (significant for muons in iron for βγ ≥1000), and at lower momenta for muons in higher-Z absorbers.
- Note the qualitatively different behavior of the density correction in gaseous Helium.
- The minimum drops from βγ≈3 to 3.5 from Z=7 to 100.
 (or β ≈0.95).



EXAMPLE: MUONS TRAVERSING A LARGE DETECTOR

- CMS magnet: 12.5m length, 6.3m diameter, cooled to 4.7K
- 3.8T magnetic field inside, return yoke (steel) provides ~2T
- The magnet and the return yoke are large metallic structures

 What is the ionization loss of a 50 GeV muon in a 50cm iron layer?

$$-\Delta x \cdot \left\langle \frac{\mathrm{d}E}{\mathrm{d}x} \right\rangle \bigg|_{p_T(\mu) = 50 \mathrm{GeV}} \approx 2.2 \frac{\mathrm{MeV cm}^2}{\mathrm{g}} \cdot 7.87 \frac{\mathrm{g}}{\mathrm{cm}^3} \cdot 50 \mathrm{\,cm} \approx 0.86 \mathrm{\ GeV}$$

 \rightarrow few % losses in the whole detector

2.3.3 PARTICLE IDENTIFICATION WITH BETHE-BLOCH



- dE/dx can be used for identifying particles.
- Left: Time Projection Chamber (ALICE detector, Ne/CO₂ gas). Right: dE/dx measurement in Pb/Pb collisions.
- Note: The spread in dE/dx limits particle Identification.







2.3.4 STOCHASTICS OF THE ENERGY LOSS

- Bethe-Bloch equation describes the mean energy loss
 - "Few concepts in high-energy physics are as misused as <dE/dx>" [Pdg Live review of particles in matter, 2017]
- δ -electrons

For ultra-relativistic particles (2.2) gives $T(\theta)\Big|_{\gamma \to \infty} = \frac{2m_e c^2}{\tan^2 \theta}$

- rate of electrons in relatively central collisions (δ -electron) is significant and limited only by $T_{max} < \gamma Mc^2$
- These large fluctuations are described by 'Landau' distribution:

$$f_L(\lambda) = \frac{1}{\pi} \int_0^\infty e^{-t \ln t - \lambda t} \sin(\pi t) dt$$
$$\lambda = \lambda(\Delta E_w, \xi) = \frac{\Delta E - \Delta E_w}{\xi} - 0.22278$$

- Effects in the detector
 - limit detector resolution

 (can correct only the mean energy loss ΔE_w)
 - limit the separation capabilities in particle Id
 - spatial widening of secondary ionization and bremstrahlung



- The shape parameter is $\kappa = \frac{\xi}{T_{max}}$ where $\xi = \frac{1}{2} K \frac{Z}{A} \rho \frac{z^2}{\beta^2} \Delta x$
 - is the prefactor in the Bethe-Bloch equation multiplied with the thickness of the material and divided by T_{max} .
- For thick absorbers ($\kappa \gtrsim I$), the energy loss becomes Gaussian, for thin absorbers ($\kappa \ll I$) it has strong tails

2.3.5 DELTA ELECTRONS

Image source: Wikipedia



2.3.6 EXAMPLE: ATLAS MUON RECONSTRUCTION

- The ATLAS calorimetry for 100 GeV muons is equivalent to 175cm of Fe, but large T_{max}, so κ≈0.001.
- A small faction of high energetic muons deposits a significant energy fraction in the calorimetry.
- Because this is an instrumented region, the muon reconstruction can correct for the effect.

improved resolution for hypothetical signal:

2.3.7 PENETRATION DEPTH AND BRAGG PEAK

- Because <dE/dx> grows with β⁻² below the minimum (MIP) in the Bethe-Bloch regime, most heavy charged particles deposit their energy in a relatively narrow region.
 - Example (right): 100 MeV protons on water
 - The dislocation effect of e.g. δ -electrons is not taken into account
 - This 'Bragg-Peak' is exploited in e.g. medical cancer treatment.

energy deposits vs. depth in water

2.4 BREMSSTRAHLUNG

- So far, we have dealt with heavy charged particles where ionization of the atomic shell is the most important source of energy loss.
- Particles can also loose energy by interacting with the nucleus or the atomic shell.
 For electrons, this 'bremsstrahlung' is the dominant source of energy loss.
- In lecture I, we calculated the energy loss of an accelerated charge as $P = \frac{q^2 \dot{v}^2}{6\pi\varepsilon_0 c^3}$. Because $m\dot{v} \propto Ze$ in the field of a nucleus of charge Z, we find $P \propto \frac{Z^2 e^2}{m^2}$ (Note: $\propto Z^2$, because interaction is coherent).
- Therefore, the bremsstrahlung of the muon is suppressed by a factor $(m_{\mu}/m_{e})^{2} = 40000$.
- Careful QM calculations take into account dielectric suppression (polarization of the medium for low energies), coherent scattering on the shell, and energy dependent shielding of the nucleus by the shell.

2.4 BREMSSTRAHLUNG

• The fractional energy loss is nearly independent of energy, therefore, the loss per unit length is efficiently described by the 'radiation length' X_0 after which the energy of a highly energetic electron reduces to 1/e = 36.7%

$$\frac{\mathrm{d}E}{\mathrm{d}x} = -\frac{E}{X_0}$$

 The radiation length can be approximated to within 2.5% (Helium: 5%) by the empirical expression

$$\rho X_0 = \frac{716.408 \text{ g cm}^{-2} A}{Z(Z+1) \ln \frac{287}{\sqrt{Z}}}$$

- The critical energy E_c is defined as the energy where bremsstrahlung (dominant at high energies) and ionization losses (dominant at lower energies) agree.
 - Factor 10000 between electrons and muons!
- high-Z materials have low E_c (lon. \propto Z while brem. is \propto Z²)

	7	density	X	0	E_k	E^{μ}_{k}	
material	Z	(g/cm^3)	(g/cm^2)	(cm)	(MeV)	(GeV)	
Be	4	1.85	65.19	35.3	113.7	1328	
C (Graphit)	6	2.21	42.65	19.3	81.7	1060	
AI	13	2.70	24.01	8.9	42.7	612	
Si	14	2.33	21.82	9.36	40.2	582	
Fe	26	7.87	13.84	1.76	21.7	347	
Cu	29	8.96	12.86	1.43	19.4	317	
Ge	32	5.32	12.25	2.30	18.2	297	
W	74	19.30	6.76	0.35	8.0	150	
Pb	82	11.35	6.37	0.56	7.4	141	
U	92	18.95	6.00	0.32	6.7	128	
Szintillator							
Nsal	11, 53	3.66	9.49	2.59	13.4	228	
Csl	55, 53	4.53	8.39	1.85	11.2	198	
BaF_2	56, 9	4.89	9.91	2.03	13.8	233	
$PbWO_4$	82, 74, 8	8.30	7.39	0.89	9.64	170	
Polystyrol	1,6	1.06	43.79	41.3	93.1	1183	
Gases (20°C, I bar))						
H_2	1	$0.0838 \cdot 10^{-3}$	61.28	731000	344.8	3611	
He	2	$0.1249 \cdot 10^{-3}$	82.76	662610	257.1	2352	
Air	pprox 7.36	$1.205 \cdot 10^{-3}$	36.66	30423	87.9	1115	
Ar	18	$1.66 \cdot 10^{-3}$	19.55	11763	38.0	572	
Xe	54	5.48·10 ⁻³	8.48	1547	12.3	232	
other materials:							
water	1,8	1.0	36.1	36.1	78.3	1031	
rock	11	2.65	26.5	10.0	49.1	693	
fotoemulsion		3.82	11.33	2.97	17.4	286	

2.4 BREMSSTRAHLUNG

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$$\frac{\mathrm{d}E}{\mathrm{d}x} = -\frac{E}{X_0}$$

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- The critical energy E_c is defined as the energy where bremsstrahlung (dominant at high energies) and ionization losses (dominant at lower energies) agree.
 - Factor 10000 between electrons and muons!
- high-Z materials have low E_c (lon. \propto Z while brem. is \propto Z²)

2.5 MULTIPLE SCATTERING - MOLIÈRE THEORY

Charged particles can change their direction when interacting with the nucleus.
 Rutherford scattering formula (the two charges are z and Z in units of e)

$$\frac{d\sigma}{d\Omega}\Big|_{\text{Rutherford}} = z^2 Z^2 \alpha^2 \hbar^2 \frac{1}{\beta^2 p^2} \frac{1}{4 \sin^4 \theta/2}$$

 For many subsequent scattering events, Moliere-theory allows to predict the *angular spread*.
 Expect a Gaussian distribution: 1

$$f(\theta_{\text{plane}}) = \frac{1}{\sqrt{2\pi}\theta_0} e^{-\theta_{\text{plane}}^2/(2\theta_0^2)}$$

Left: Scattering of protons with T=2.18 MeV on an Aluminium foil of x = 13 μ m. The large-angle deviations predicted by Rutherford scattering make a 'non-Gaussian tail' for δ^2 >4.

Here, $\delta^2 = 1/2(\theta/\theta_0)^2$ where θ_0 is predicted by the 'Highland formula':

Rutherford scattering

Ze

ze

2.6 IMPACT PARAMETER RESOLUTION

- In order to identify b-quarks, the displaced decay of B mesons inside a jet is identified by reconstructing secondary, transversally displaced, vertices.
- B lifetime 1.5ps, <l> = βγcτ ≈ 2mm for e.g. a 10 GeV particle.
- The impact parameter is the most discriminating feature of a secondary vertex. Multiple scattering in the beampipe limits impact parameter resolution.
- Assume $r_B = 50 \text{ mm}$ and a beam-pipe of thickness Imm and $p(\pi)$ of 5 GeV.
- The resulting resolution is $\Delta d_0 = \theta_0 r_B$

Choose Be instead of AI for the beampipe to gain a factor 2 in impact parameter resolution!

2.7 ELECTROMAGNETIC FIELD IN DIELECTRIC MEDIUM

Energy can be lost to photons in a medium. Consider a dielectric medium.

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2.7 ELECTROMAGNETIC FIELD IN DIELECTRIC MEDIUM

• Let's write the 4-momentum of the incident charged particle:

$$p^{\mu} = M\gamma\begin{pmatrix} c\\ \vec{v} \end{pmatrix} = Mc\gamma\begin{pmatrix} 1\\ \vec{\beta} \end{pmatrix} = \begin{pmatrix} \frac{E}{c}\\ \vec{p} \end{pmatrix} \quad \text{(of course it satisfies } \frac{1}{c^2}\left(E^2 - c^2\vec{p}\,^2\right) = M^2c^2$$

• For the photon we make the ansatz $p_{\gamma} = \begin{pmatrix} \frac{\hbar\omega}{c}\\ \hbar\vec{k} \end{pmatrix}$.
Then we can use $p' = p - p_{\gamma}$ to find
 $M^2 = M^2 - \frac{2}{c^2}p \cdot p_{\gamma} \quad \text{or } 0 = p \cdot p_{\gamma} = \frac{E\hbar\omega}{c^2} - \hbar\vec{k} \cdot \vec{p}$.
• We can express the angular frequency as

$$\omega = \frac{\vec{k} \cdot \vec{p}}{E/c^2} = \vec{k} \cdot \vec{v} = |\vec{k}| |\vec{v}| \cos \theta_c = kv \cos \theta_c$$

$$\frac{|\vec{k}|^2 c^2}{E/c^2} = \vec{k} \cdot \vec{v} = |\vec{k}| |\vec{v}| \cos \theta_c = kv \cos \theta_c$$

and use the dispersion relation $\omega^2 = \frac{|k|^2 c^2}{\varepsilon} \rightarrow |\vec{k}| = \sqrt{\varepsilon} \frac{\omega}{c}$ to find $1 = \sqrt{\varepsilon} \frac{|\vec{v}|}{c} \cos \theta_c$ IF(!) we are in the optical regime, there is is a real solution $\cos \theta_c = (\beta n)^{-1}$

- IF(!) we are in the optical regime, there is is a real solution
 - real photon emission for v > c / n (Čerenkov)

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Close to resonance (ionisation), or in the hard UV, there is no real solution. •

2.8 TRANSVERSE REACH OF THE EM INTERACTION

- Next, we assume $k_x=0$ (no loss of generality) and use $\omega = vk_z = vk\cos\theta_c$ and $\frac{\omega^2\varepsilon}{c^2} = k_y^2 + k_z^2$ to express k_y as $k_y^2 = \left(\frac{\omega}{v}\right)^2 \left(\frac{v^2}{c^2}\varepsilon - 1\right) = \left(\frac{\omega}{v}\right)^2 \left(\frac{v^2}{(c/\sqrt{\varepsilon})^2} - 1\right)$.
- The phase velocity in the medium is $c_m = c/\sqrt{\varepsilon}, \ \beta' = v/c_m$ and therefore $k_y = \frac{\omega}{v}\sqrt{\beta'^2 1}$.
- For β '>I (optical!), the plane wave solution is real and the field extends infinitely in the y direction: $e^{i(\omega t \vec{k} \cdot \vec{r})}$
- For $\beta' < 1$, however, we have $e^{-i\frac{\omega}{v}(z-vt)}e^{-y/y_0}$ with

$$y_0 = \frac{v}{\omega} \underbrace{\frac{1}{\sqrt{1 - \beta'^2}}}_{\sqrt{1 - \beta'^2}} = \frac{v\sqrt{\varepsilon}}{\omega_0} \gamma' = \frac{c}{\omega_0} \beta' \gamma' = \frac{\beta' \gamma'}{k}$$

where ω_0 and $k_0^{\gamma'}$ are the vacuum angular frequency and wave vector: $\omega = ck/\sqrt{\varepsilon} = \omega_0/\sqrt{\varepsilon}$ $k_0 = \omega/c = k/\sqrt{\varepsilon}$

Take home message (1): The transverse reach of the electromagnetic interaction is governed by $y_0 = \beta' \gamma' / k$ of the incident particle! This explains the dependence of the density effect on $\beta \gamma$.

2.9 ČERENKOV RADIATION

- Below the excitation energies of the medium (optical regime) ε is real and $\varepsilon > 1$ so that $\cos \theta_c = (\beta n)^{-1}$ has a real solution if $\beta > n = \sqrt{\varepsilon}$. In this case, 'Čerenkov'-radiation is emitted. For $\beta \rightarrow 1$: $\theta_{\max} = \cos^{-1} n^{-1}$
- The minimum energy of the particle can be calculated from

$$\frac{E_{\rm th}}{mc^2} = \gamma_{\rm th} = \frac{1}{\sqrt{1 - \beta_{\rm th}^2}} = \frac{1}{\sqrt{1 - \varepsilon^{-1}}} = \frac{n}{\sqrt{n^2 - 1}} \ .$$

• The maximum angle is attained for $\beta = 1$ and satisfies

$$\sin\theta_{max} = \sqrt{1 - \cos^2\theta_{max}} = \sqrt{1 - \frac{1}{n^2}} = \frac{1}{\gamma_{th}} \cdot$$

• The angle of radiation for a given $\boldsymbol{\beta}$ as a function of $\boldsymbol{\theta}_{\max}$ is

$$\sin^2 \theta_c = 1 - \cos^2 \theta_c = 1 - \frac{1}{\beta^2 n^2} = 1 - \frac{\beta_{th}^2}{\beta^2} = \frac{1}{\gamma_{th}^2} \frac{\gamma^2 - \gamma_{th}^2}{\gamma^2 - 1}$$
$$= \sin^2 \theta_{max} \frac{\gamma^2 - \gamma_{th}^2}{\gamma^2 - 1} \quad \text{and therefore} \quad \frac{\theta_c}{\theta_{max}} = \frac{R_c}{R_{max}} \approx \sqrt{1 - \frac{\gamma_{th}^2}{\gamma^2}}$$

• For $\gamma \gg \gamma_{th}$, Čerenkov light is emitted in narrow rings. It is very weak: $\approx 0.1 - 1\%$ of Ionization.

• It also emitted promptly, in contrast to scintillation and ionization signals.

Kolanoski, Wermes 2015

Čerenkov ring:

flight path

2.10 TRANSITION RADIATION

- If a particle traverses a surface separating two materials $\$ with different refraction index $n_1 \neq n_2$, transition radiation is emitted.
- The most probable emission angle is $\theta_{mpv} \sim \gamma^{-1}$, and Intensity $\sim \gamma$.
- For ultra-relativistic particles (γ >1000), hard UV radiation is emitted, closely collimated with the particle.
- Measuring the intensity (γ) allows to identify charged particles when their momentum is known.

3. INTERACTION OF PHOTONS WITH MATTER

3.1 OVERVIEW

• High energetic photons typically interact by absorption. Ansatz: $N(x) = N_0 e^{-\mu x}$

electromagnetic shower (bottom-up) in the I5-foot Bubble Chamber at FNAL

- Photo effect: the photon looses its energy to an atom which emits an electron
 - secondary emission of characteristic X-rays and Auger electrons when the holes are re-filled
- Compton effect: (quasi-) elastic scattering on an electron in the atomic shell (>>E_{bind})
- Pair creation: the photon converts to an e⁺e⁻ pair in the electric field of the nucleus

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3.2 MEAN FREE PATH OF PHOTONS

- Photo-effect increases when the energy passes the energy thresholds of nuclear shells (K,L,M,..) (there is always a smaller contribution from elastic coherent Rayleigh scattering on the shell)
- The Compton effect and pair production dominate $\geq IMeV \approx 2m_ec^2$

• Pair creation(E>1.022 MeV) and electron bremsstrahlung, the two dominant processes at high energies, have the same (leading-order) matrix element and are thus related by $\lambda_{\gamma} = \frac{9}{7}X_0$. When e[±] brem 1/e= 63%, γ loose 54%.

3.3 MATERIAL DEPENDENCE

3.4 OVERVIEW OF ELECTROMAGNETIC INTERACTIONS

4. HADRONIC INTERACTIONS

4.1 OVERVIEW

- Hadrons interact inelastically with a single nuclei in the nucleus
- Secondary hadronic decay products interact with other nuclei, leading to shower formation.
- If neutral pions are produced, they decay to photons with subsequent electromagnetic showers.
- Between hadronic interactions, charged shower particle also ionize the material, leading to an EM cascade (also, nuclei can de-excite emitting γ)
- The fluctuation of the shower components is large and, depending on the charge and the momenta of the particles, each has a different reconstruction efficiency.
- This limits the resolution in reconstructing the energy of the incident particle.

4.2 HADRONIC AND ELECTROMAGNETIC SHOWERS

- The hadronic interaction is stronger, but has a short range.
- Again, an exponential law is an excellent description

$$N(x) = N_0 e^{-x/\lambda_a}$$
 with the absorption length λ_a
 $\lambda_a = \frac{A}{N_A \rho \sigma_{inel}} \approx 35 \frac{g}{cm^2} \frac{A^{\frac{1}{3}}}{\rho} \propto A^{-\frac{2}{3}}$ (assuming $\rho \propto A$)

- The processes for energy loss of photons and electrons (bremsstrahlung, pair production) lead to electromagnetic shower development, however, at a shorter length scale.
- The shower depth (electromagnetic and hadronic) grows logarithmically with energy.
- Neutral hadrons (e.g. neutrons) therefore penetrate deeply.

4.3 MATERIAL DEPENDENCE

•	It is beneficial	Material	Ζ	$ ho \ ({ m g/cm^3})$	X_0 (cm)	A	λ_a (cm)	λ_a/X_0	$\left. \frac{dE/dx}{dm_{min}} ight. \ { m (MeV/cm)}$
	to have similar response	H_2O	1,8	1.00	36.1	18	83.3	2.3	1.99
		Luft	7,8	$1.205 \cdot 10^{-3}$	$3.0\cdot 10^4$	14.3	$7.5\cdot 10^4$	2.5	$2.19\cdot10^{-3}$
	showers ('compensating')	Be	4	1.85	35.3	9	42.1	1.2	2.95
•	Low-Z materials	С	6	2.21	19.3	12	38.8	2.0	3.85
	have a lower (bonefitial)	AI	13	2.70	8.9	27	39.7	4.5	4.36
	nave a lower (Denential)	Fe	26	7.87	1.76	56	16.8	9.5	11.42
	ratio λ_a / X_0 .	Cu	29	8.96	1.43	64	15.3	10.7	12.57
		W	74	19.30	0.35	184	9.9	28.3	22.10
		Pb	82	11.35	0.56	207	17.6	31.4	12.73
		U	92	18.95	0.32	238	11.0	34.4	20.48

