At a glance: Electroweak (EW) symmetry non-restoration (SNR) at high temperature (T) and vacua trapped in a metastable minimum at zero T are possible phenomena within the N2HDM thermal history that further constrain its parameter space



Fate of the electroweak symmetry in the early universe: Non-restoration and trapped vacua in the N2HDM T. Biekötter, S. Heinemeyer, J. M. No, <u>M. Olea</u>, G. Weiglein

1. <u>INTRODUCTION</u>: The first-order EW phase transition (FOEWPT) has been extensively studied in the N2HDM...

## but what other finite T effects could occur within this model?

**Trapped vacua:** Numerous studies take the existence of a critical temperature (T<sub>c</sub>) as a sufficient condition for a FOEWPT. We show that this is not always the case and leads to incorrect specifications of the parameter space. **EW SNR**: It is commonly assumed that the EW symmetry gets restored at high T. This is not always the case. We define **3 coefficients to identify N2HDM points with EWSNR**.

2.1. <u>MODEL</u>: N2HDM = 2HDM + real scalar singlet → 3 CP-even,
1 CP-odd Higgs boson and 2 charged Higgs bosons. 3 real vevs.
2.2. <u>METHODS</u>:

**1)** We apply **theoretical and experimental constraints** to the N2HDM parameter space (perturbative unitarity, flavor physics...)

2) Numerical analysis (studying trapped vacua). Calculation of the critical and the transition temperature (Tn) for N2HDM points.

**3) Analytical analysis (studying EW SNR)**. Study of the curvature of the N2HDM potential at the origin of field space in its **high-T approximation.** We defined **three coefficients c**<sub>ii</sub> that encode the sign of the curvature of the EW preserving minima at high T:

 $H^0_{ij} = \left. \frac{\partial^2 V}{\partial \rho_i \partial \rho_j} \right|_{(0,0,0)}$ 

$$c_{ii} \equiv \lim_{T \to \infty} H_{ii}^0/T^2 > 0$$

# 3. <u>RESULTS:</u>

Trapped vacua: Singlet component of one of the CP-even Higgs vs. its mass. Color bar: Tc (right), Tn (left). Black points are unphysical.



**EW SNR:** For all the points in the scan  $c_{11} < 0 \rightarrow$  the EW symmetry at high T can't be restored at the origin of field space and, under certain conditions, also outside.

### 4. CONCLUSIONS:

- With the coefficients **c**<sub>ii</sub> one can easily find regions of the N2HDM parameter space **where EW SNR happens** at high T.

- The calculation of the **Tn** is **needed** to specify the allowed parameter space.

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#### 1. **One-loop effective potential**

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$$H_{11}^{S}(\rho_{3},T) = \frac{\partial^{2} V}{\partial \rho_{1}^{2}}\Big|_{(0,0,\rho_{3})} \quad c_{11}^{S} = c_{11} + \mathcal{O}\left(\frac{v_{S}(T)^{2}}{T^{2}}\right) \quad c_{11}^{S} = \lim_{T \to \infty} \frac{H_{11}^{S}(v_{S}(T),T)}{T^{2}}$$

Under certain conditions,  $|\lambda_6|, |\lambda_7|, |\lambda_8| < 1$  the coefficients C<sub>ii</sub> also control the EW symmetry restoration at high T independently of the Z<sub>2</sub> symmetry restoration.



#### 3. Trapped vacua

Input parameters of the scan appearing in the figures of the poster

$m_{h_a}$	$m_{h_b}$	$m_{h_c}$	$m_A$	$m_{H^{\pm}}$	aneta	$C^2_{h_a t \bar{t}}$	$C_{h_aVV}^2$	$R_{b3}$	$m_{12}^2$	$v_S$
125.09	[30, 1000]	400	650	650	2	1	1	[-1, 1]	65000	[1, 1000]
Computation of the transition temperature Tn						Bounce action				
$\int_{T_n}^{T_c} \frac{T^4}{H^4} \frac{A(T)}{T} e^{-S_3(T)/T} dT \approx 1$				$\sim 1$	L.	$S_{3} = 4\pi \int r^{2} \mathrm{d}r  \left[ \frac{1}{2} \left( \frac{\mathrm{d}\phi_{\mathrm{B}}}{\mathrm{d}r} \right)^{2} + V \left( \phi_{\mathrm{B}}, T \right) \right]$				