

Neutrino transition in dark matter

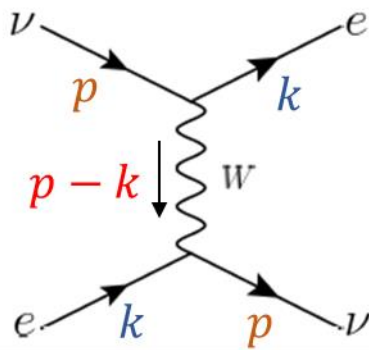
2112.05057

1909.10478, 2012.09474

Eung Jin Chun

Neutrino oscillations in matter Wolfenstein 1978

- “The effect of coherent forward scattering (leaving the medium unchanged) must be taken into account when considering the oscillations of neutrinos traveling through matter.”



$$\mathcal{H}_{eff} = 2\sqrt{2} G_F \bar{\nu}_{eL} \gamma^\mu e_L \bar{e}_L \gamma_\mu \nu_{eL}$$

$$\Rightarrow \sqrt{2} G_F N_e \bar{\nu}_{eL} \gamma^0 \nu_{eL}$$

$$\rightarrow E \approx p + \frac{m_\nu^2}{2p} + \sqrt{2} G_F N_e \quad \text{“Wolfenstein potential”}$$

Neutrino oscillations in Dark Matter

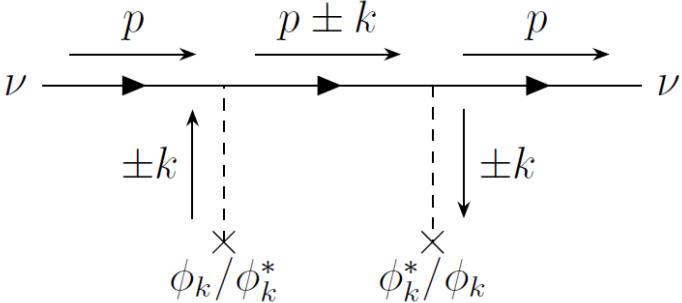
- Consider neutrinos propagating in a classical DM background

$$\mathcal{L}' = \frac{1}{2}g\hat{\phi}\nu\nu + \frac{1}{2}g^*\hat{\phi}^+\bar{\nu}\bar{\nu}$$

$$\mathcal{L}' = \frac{1}{2}g\hat{\phi}\nu f + \frac{1}{2}g^*\hat{\phi}^+\bar{\nu}f$$

$$\hat{\phi}(x) \rightarrow \phi_c(x) = \phi_k e^{-ik \cdot x} + \bar{\phi}_k e^{ik \cdot x} \quad \text{LNV: } \bar{\phi}_k = \phi_k^*$$

- Coherent forward scattering \rightarrow medium contribution to the propagator \rightarrow modified dispersion relation



In the relativistic limit:

$$E \approx p + \frac{m_\nu^2 + |g\phi_k|^2 + |g\bar{\phi}_k|^2}{2p}$$

Choi, EJC, Kim, 1909.10478
2012.09474

“Neutrino oscillations even with $m_\nu = 0$ ”
(Wolfenstein)

Neutrino Propagation in Dark Matter

- Medium effect in neutrino propagation in a classical DM background

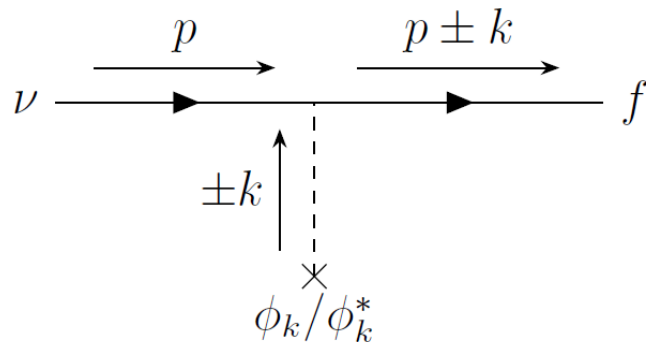
$$\mathcal{L}' = \frac{1}{2}g\hat{\phi}\nu\nu + \frac{1}{2}g^*\hat{\phi}^+\bar{\nu}\bar{\nu}$$

$$\mathcal{L}' = \frac{1}{2}g\hat{\phi}\nu f + \frac{1}{2}g^*\hat{\phi}^+\bar{\nu}\bar{f}$$

$$\hat{\phi}(x) \rightarrow \phi_c(x) = \phi_k e^{-ik \cdot x} + \bar{\phi}_k e^{ik \cdot x}$$

$$\text{LNV: } \bar{\phi}_k = \phi_k^*$$

- Yet unexplored coherent process of transition



$$A_{\nu \rightarrow f} \propto g\phi_k \bar{u}_{p+k} P_L u_p + g\bar{\phi}_k \bar{u}_{p-k} P_L u_p$$

EJC, 2112.05057

Coherent state=classical field

Glauber, Sudarshan, 1963

- Eigenstate of annihilation operator:

$$|\phi_c\rangle \propto e^{\int_k \phi_k a_k^\dagger} |0\rangle$$

$$\langle \phi_c | a_k | \phi_c \rangle = \phi_k; \quad \langle \phi_c | a_k^\dagger | \phi_c \rangle = \phi_k^*$$

- Energy density:

$$\rho_\phi V = \langle \phi_c | \int_k E_k a_k^\dagger a_k | \phi_c \rangle = \int_k E_k |\phi_k|^2$$

$$\hat{\phi}(x) = \int \frac{d^3k}{(2\pi)^3 2E_k} [a_k e^{-ik \cdot x} + a_k^\dagger e^{ik \cdot x}]$$

$$\begin{aligned} \phi_c(x) &\equiv \langle \phi_c | \hat{\phi}(x) | \phi_c \rangle \\ &= \int \frac{d^3k}{(2\pi)^3 2E_k} [\phi_k e^{-ik \cdot x} + \phi_k^* e^{ik \cdot x}] \end{aligned}$$

Monochromatic + non-relativistic

$$\phi_k = (2\pi)^3 2E_0 \delta^3(k - k_0) \phi_0, \quad k_0 \approx (m_\phi, m_\phi \vec{v})$$

$$\rho_\phi \approx 2E_0^2 |\phi_0|^2 \approx 2m_\phi^2 |\phi_0|^2$$

$$|\phi_0| \approx 10^8 \text{ GeV} \left(\frac{10^{-20} \text{ eV}}{m_\phi} \right)$$

Analysis setup

- Work in the mass basis of neutrinos coupling to a classical background:

$$\mathcal{L}' = \frac{1}{2} g_{ij} \hat{\phi}(x) \bar{\nu}_i^c P_L \nu_j + \frac{1}{2} g_{ij}^* \hat{\phi}^\dagger(x) \bar{\nu}_j P_R \nu_i^c$$

$$\mathcal{L}_{\text{mass}} = \frac{1}{2} m_{\nu_i} \bar{\nu}_i^c P_L \nu_i + \frac{1}{2} m_{\nu_i}^* \bar{\nu}_i P_R \nu_i^c$$

$$|g\phi_0| \sim 0.01\text{eV}$$
$$\Rightarrow g \sim 10^{-19} \left(\frac{m_\phi}{10^{-20}\text{eV}} \right)$$

“Flavour structure of the coupling”

$$g_{ij} = g_{\alpha\beta} U_{\alpha i} U_{\beta j}$$

Two-flavour case

- Consider two-flavor neutrinos:

$$\nu_e = c_\theta \nu_1 - s_\theta \nu_2$$

$$\nu_\mu = s_\theta \nu_1 + c_\theta \nu_2$$

- The $\nu_e \rightarrow \nu_\mu$ transition amplitude:

$$A_{\mu e} = c_\theta s_\theta (A_{11} - A_{22}) + c_\theta^2 A_{21} - s_\theta^2 A_{12}$$

Standard oscillation +
Medium effect in transition
with $g_{11,22} \neq 0$

Medium effect with $g_{12} \neq 0$

Transition amplitudes in medium

- $\nu_i \rightarrow \nu_j$ transition amplitude:

$$A_{ji} = \langle \phi_c; \nu_j, p_2 | e^{-iH_0 t_2} U(t_2, t_1) e^{iH_0 t_1} | \nu_i, p_1; \phi_c \rangle$$

$$U(t_2, t_1) = T e^{i \int_{t_1}^{t_2} dt \mathcal{L}'(t)} = I + i \int_{t_1}^{t_2} dt \int d^3x \mathcal{L}'(x) + \dots$$

- $U = I$: Standard neutrino oscillation

$$A_{ji}^0 = \langle \phi_c; \nu_j, p_2 | e^{-iH_0(t_2-t_1)} | \nu_i, p_1; \phi_c \rangle \propto e^{-i E_i L} \delta_{ji} \text{ with } L = t_2 - t_1$$

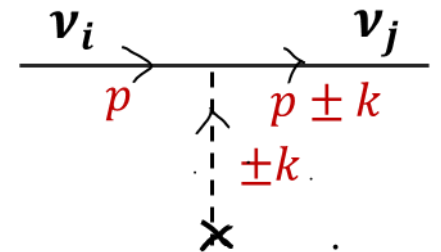
Transition amplitudes in medium

- $U = i \int_{t_1}^{t_2} dt \int d^3x \mathcal{L}'(x)$: Medium-induced transition $\nu_i(p) \rightarrow \nu_j(p \pm k)$

$$A_{ji}^{1\pm} = \langle \phi_c; \nu_j, p_2 | e^{-iH_0 t_2} i \int_{t_1}^{t_2} dt \int d^3x \mathcal{L}'(x) e^{iH_0 t_1} | \nu_i, p_1; \phi_c \rangle$$

$$\mathcal{L}' = \frac{1}{2} g_{ji} \bar{\nu}_j^c P_L \nu_i \hat{\phi}(x) + \frac{1}{2} g_{ji}^* \bar{\nu}_i P_R \nu_j^c \hat{\phi}^\dagger(x)$$

$$\phi_c(x) = \phi_0 e^{-ik_0 \cdot x} + \phi_0^* e^{ik_0 \cdot x} \quad \nu(x) = \int_p a_p^s u_p^s e^{-ip \cdot x} + a_p^{s+} v_p^s e^{ip \cdot x}$$



$$A_{ji}^{1\pm} \propto g_{ji}^{(*)} \int_{t_1}^{t_2} dt e^{i(E_{p \pm k_0} - E_p \mp E_{k_0})t} \phi_0^{(*)} \frac{[\overline{u_{p \pm k_0}^s} P_L u_p^s - \overline{v_p^s} P_R v_{p \pm k_0}^s]}{2E_{p \pm k_0} 2E_p}$$

Transition probability

- Three independent contributions: $P_{\mu e} \propto |A_{\mu e}^0|^2 + |A_{\mu e}^{1+}|^2 + |A_{\mu e}^{1-}|^2$

- Assuming $|g\phi_0| \ll m_\nu$, consider two couplings independently g_{12}, g_{22} :

$$E_p \approx E_\nu + \frac{m_\nu^2}{2E_\nu}, \quad E_{k_0} \approx m_\phi$$

- Additional contribution to be smaller:

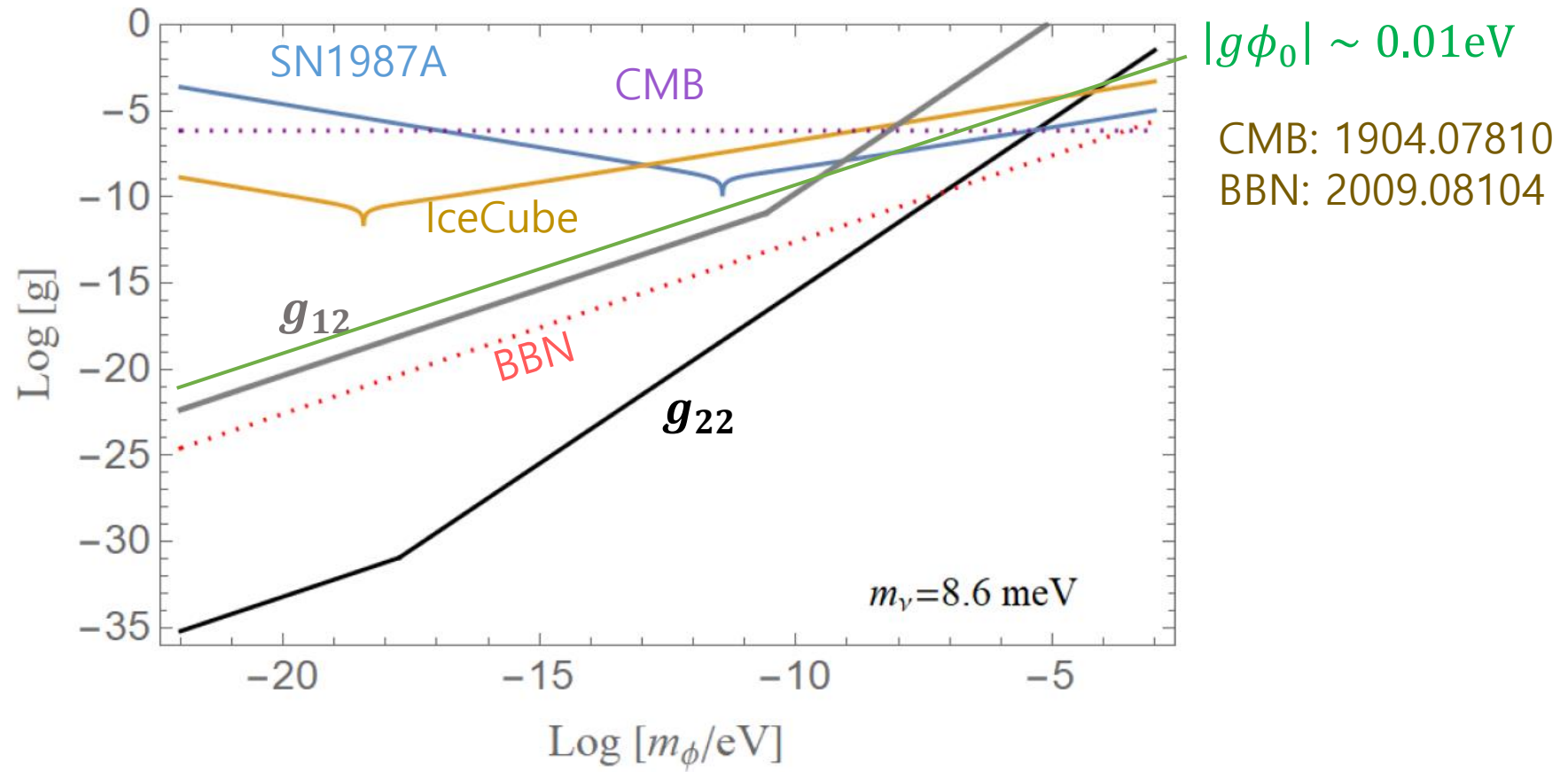
$$\delta P_{\mu e} \lesssim 0.1 P_{\mu e}^{\text{obs.}}$$

$$P_{\mu e} = \sin^2(2\theta) \sin^2\left(\frac{\Delta m_{21}^2 L}{4E_\nu}\right) \quad \Delta_{21}^\pm \equiv \frac{\Delta m_{21}^2}{2E_\nu} \pm m_\phi$$

$$+ \cos^2(2\theta) \frac{|g_{12}\phi_0|^2 m_{\nu 2}^2}{(\Delta_{21}^\pm E_\nu)^2} \sin^2\left(\frac{\Delta_{21}^\pm L}{2}\right)$$

$$+ \sin^2(2\theta) \frac{|g_{22}\phi_0|^2 m_{\nu 2}^2}{m_\phi^2 E_\nu^2} \sin^2\left(\frac{m_\phi L}{2}\right)$$

Constraints from neutrino oscillation data



Summary

- Interesting medium effects in neutrino propagation – medium potential, induced mass-squared, CPV, etc..
- Discussed yet another transition process $\nu_i(p) \rightarrow \nu_j(p \pm k)$ that occurs in the classical background of $\phi_k/\phi_{k'}^*$, and is highly constrained by the neutrino oscillation data.
- This process is suppressed in the limit of $m_\nu \rightarrow 0$, where neutrino oscillations can be solely due to the medium effect.