

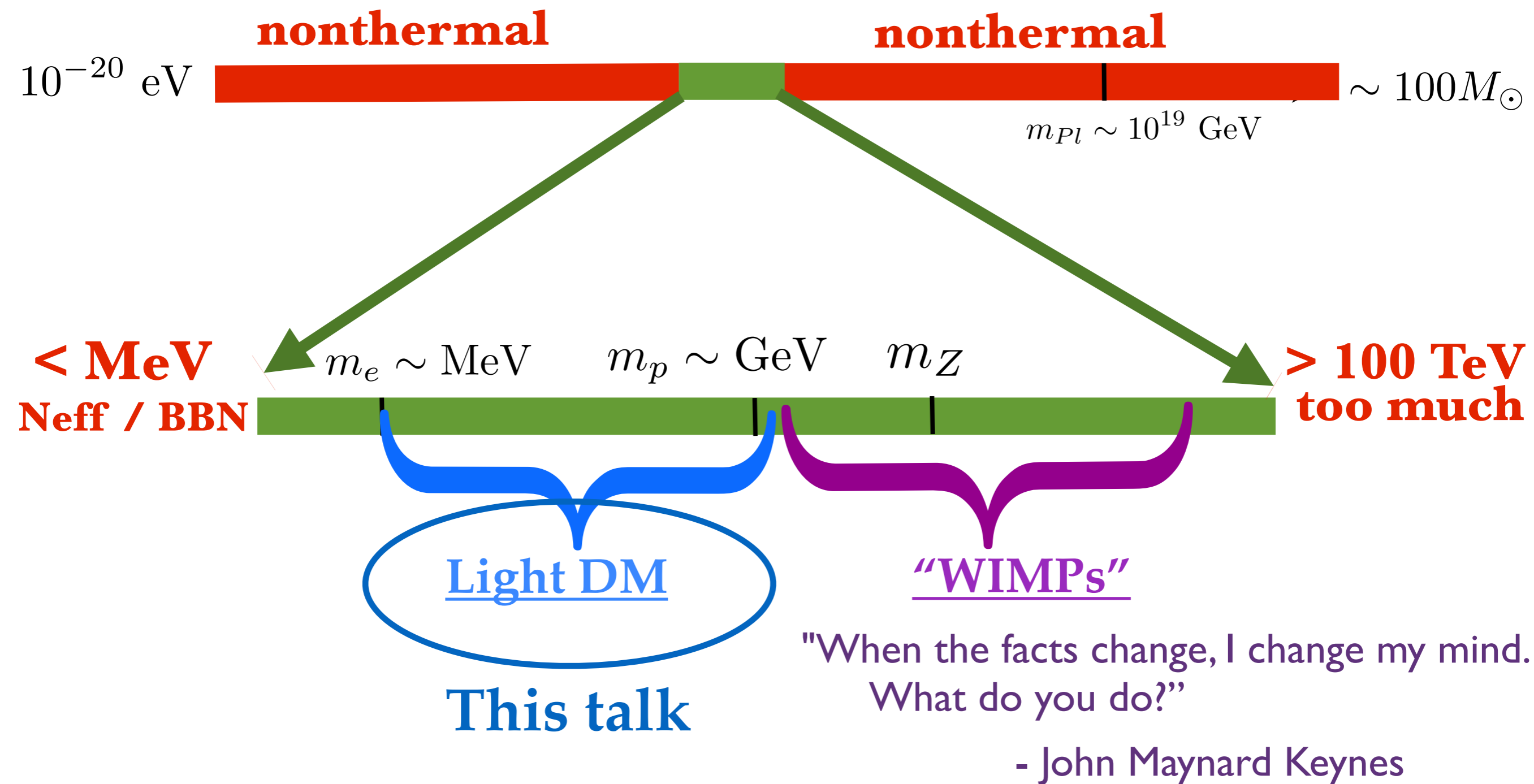


Secluded Dark Matter ($< \text{GeV}$)

Gordan Krnjaic

RF6 Meeting, Dec 4, 2020

Thermal Dark Matter



Light DM vs. WIMPs

LDM must be SM neutral

Else would have been discovered

LDM needs light new force carriers

Overproduced without light, neutral “mediators”

$$\sigma v \sim G_F^2 m_\chi^2 \sim 10^{-29} \text{ cm}^3 \text{ s}^{-1} \left(\frac{m_\chi}{\text{GeV}} \right)^2$$

Always too small if weak scale

Annihilation through *renormalizable* interactions

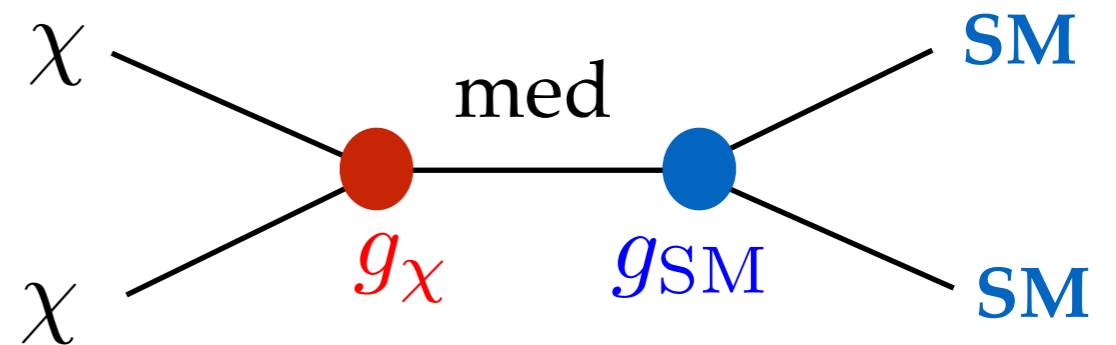
Easy to categorize options

Who's Heavier: DM or Mediator?

See Asher's talk
five minutes ago

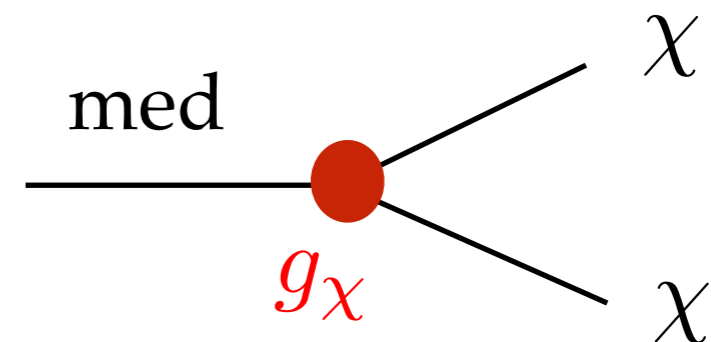
Direct Annihilation

$$m_\chi < m_{\text{med}}$$



Predictive thermal targets

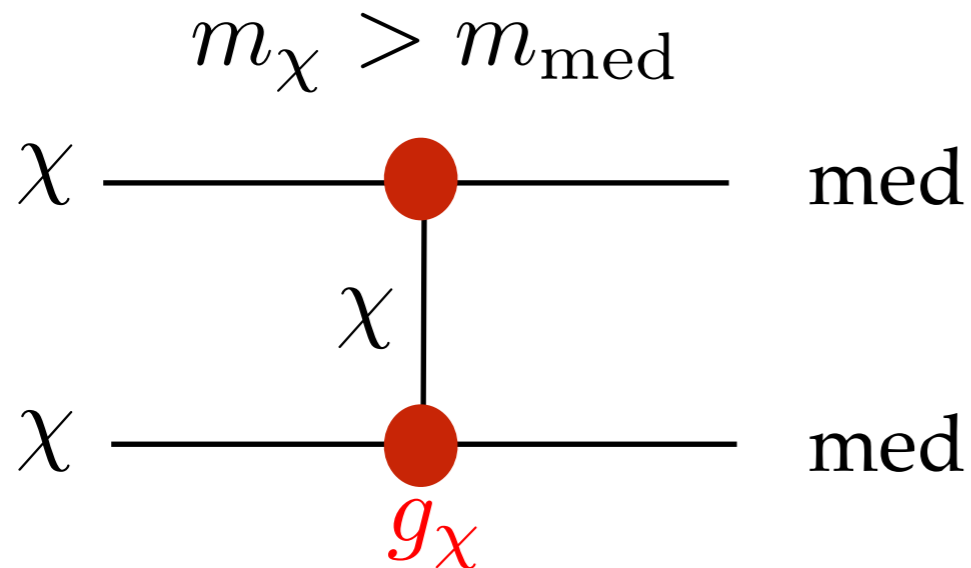
Abundance depends on g_{SM}



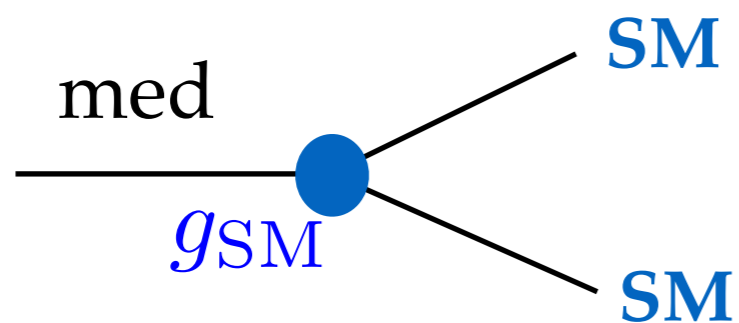
Mediator decays **to DM***

Who's Heavier: DM or Mediator?

Secluded Annihilation



No clear experimental target
Abundance set by g_χ



Mediator decays **to SM**

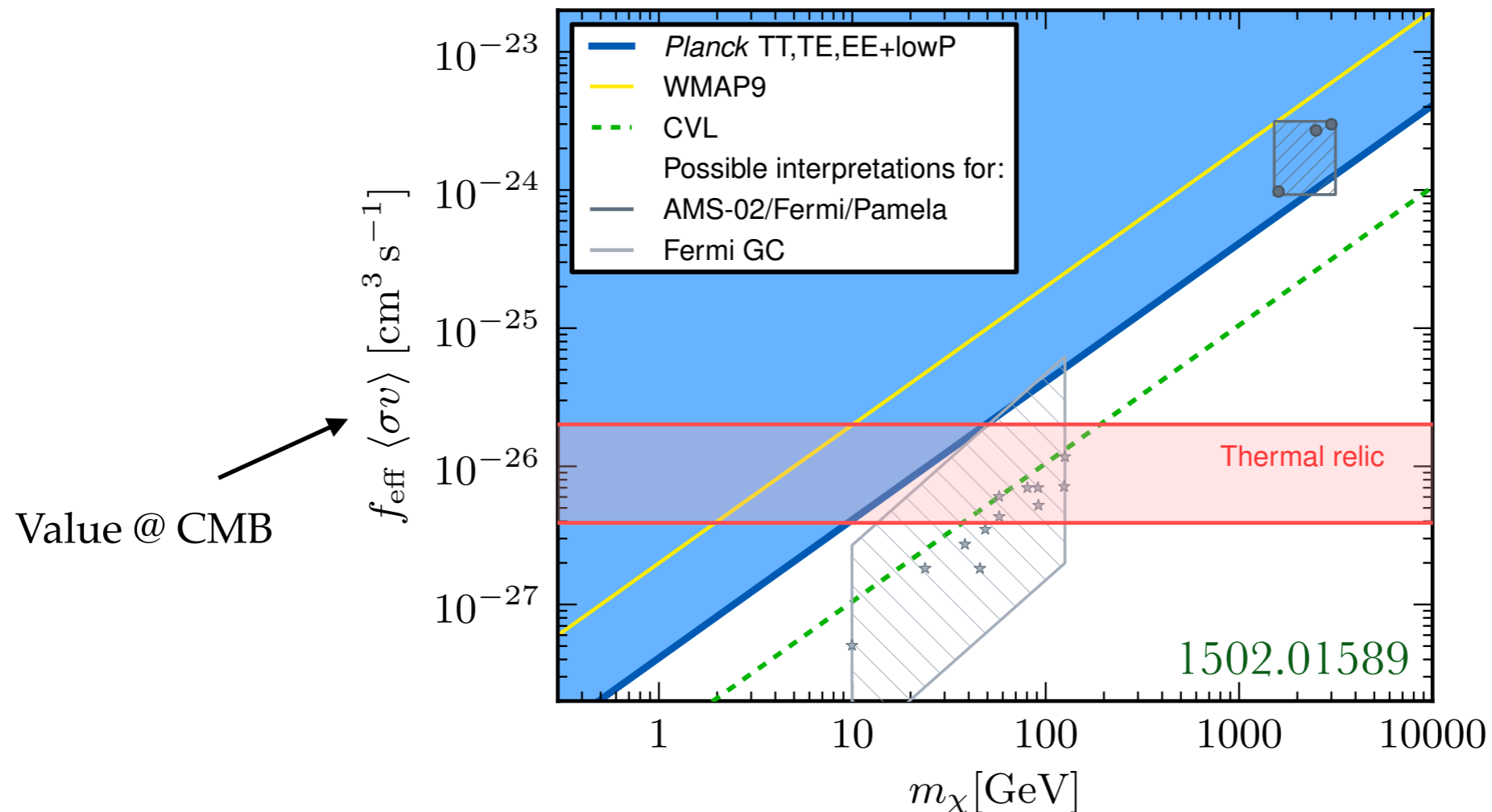
This talk

Early universe: DM+SM in equilibrium
due to g_{SM}

Freeze-out: independent of SM coupling

Secluded U(1) below the weak scale
Pospelov 0811.1030

CMB excludes many $< \text{GeV}$ thermal DM Models



Out-of-equilibrium annihilation injects energy during CMB era

Options for
safe models:

- 1) P-Wave Annihilation
- 2) Annihilation Stops pre-CMB
- 3) Mediator Decays to Neutrinos

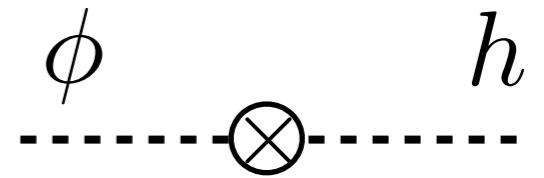
Overview

- 1) **P-Wave Annihilation**
- 2) **Annihilation Stops pre-CMB**
- 3) **Mediator Decays to Neutrinos**

Higgs Portal Scalar Mediator

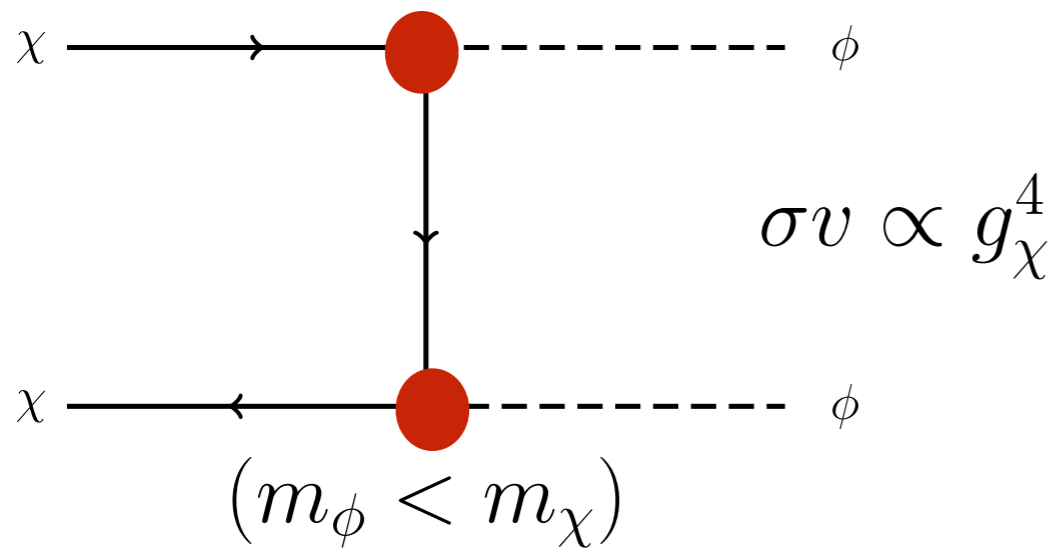
Scalar / Higgs mixing

$$\phi H^\dagger H \rightarrow \phi h$$

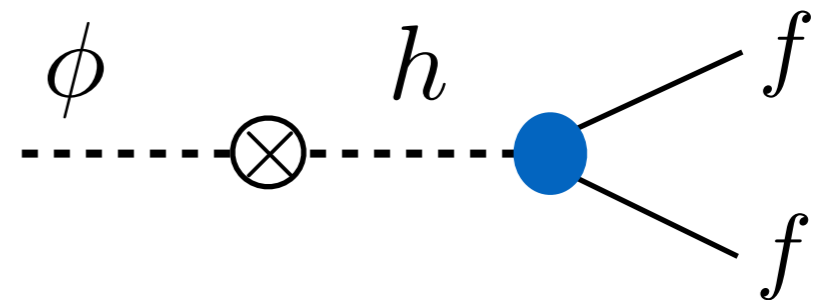


Coupled to DM/SM

$$\mathcal{L} \supset g_\chi \phi \bar{\chi} \chi + \phi \sin \theta \sum_f \frac{m_f}{v} \bar{f} f$$



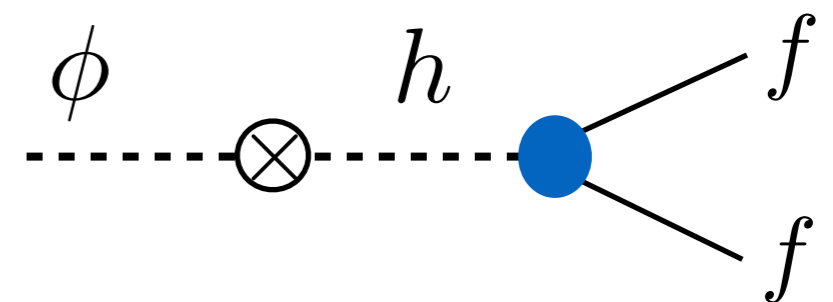
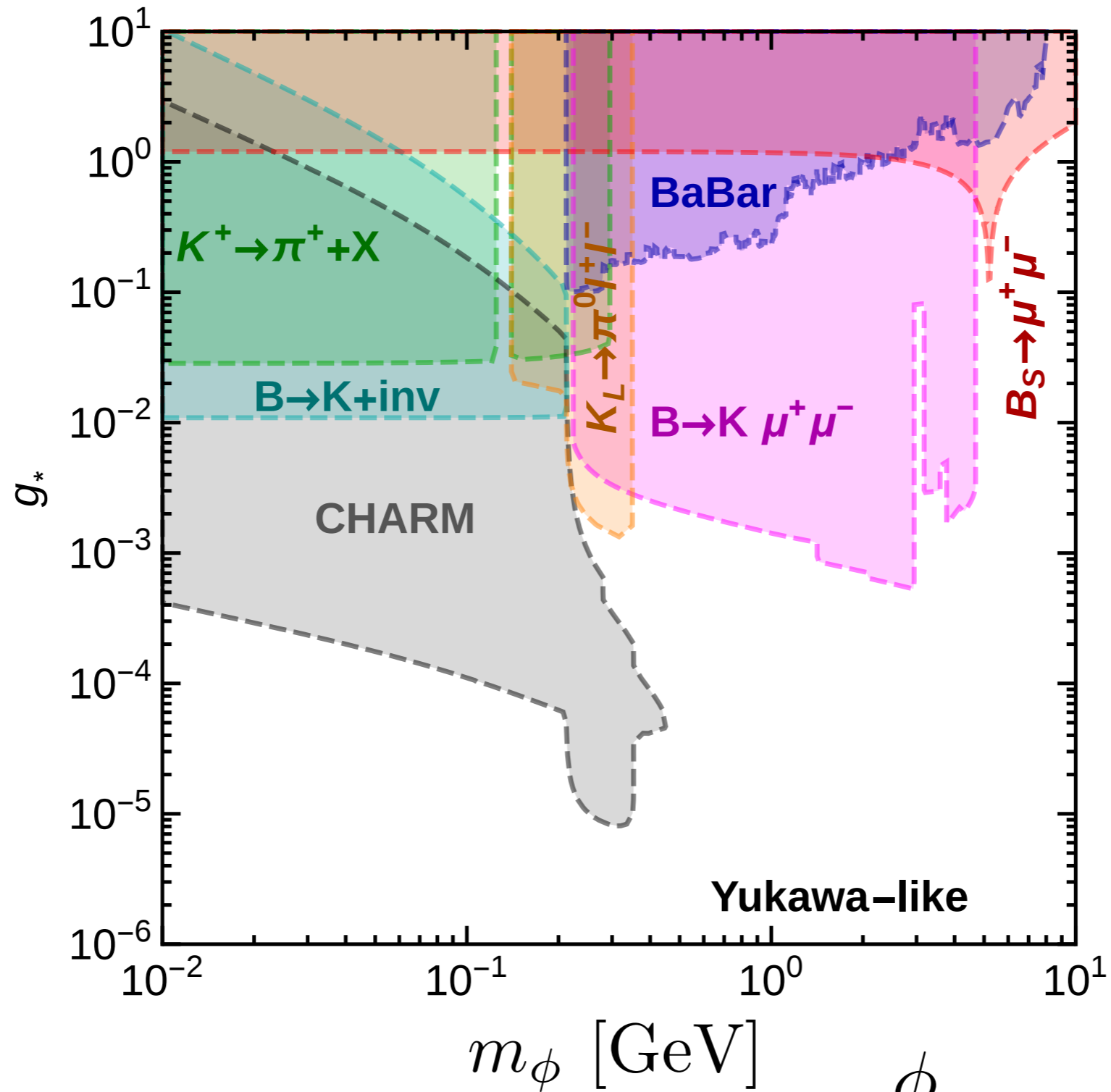
Decays like a light Higgs



Requires fermion DM for p-wave annihilation

$$\sigma v \propto v^2$$

Higgs Portal Scalar Mediator



Flavored Scalar Mediator

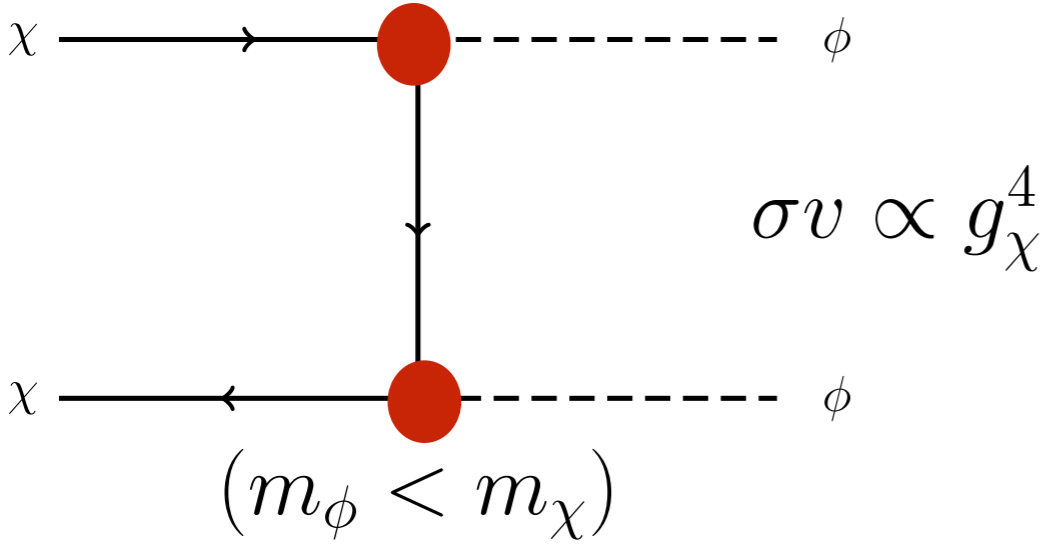
UV Theory gives flavor specific operator

Batell, Freitas, Ismail, McKeen 1712.10022

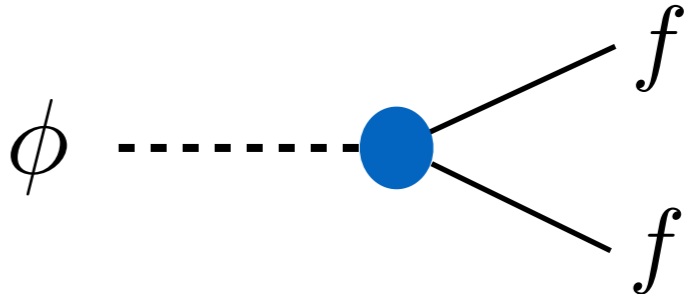
$$\mathcal{L}_{\text{eff}} = \frac{c_f}{\Lambda} \phi H \bar{f} f \rightarrow \frac{c_f v}{\Lambda} \phi \bar{f} f$$

Coupled to DM/SM

$$\mathcal{L} \supset g_\chi \phi \bar{\chi} \chi + g_f \phi \bar{f} f$$



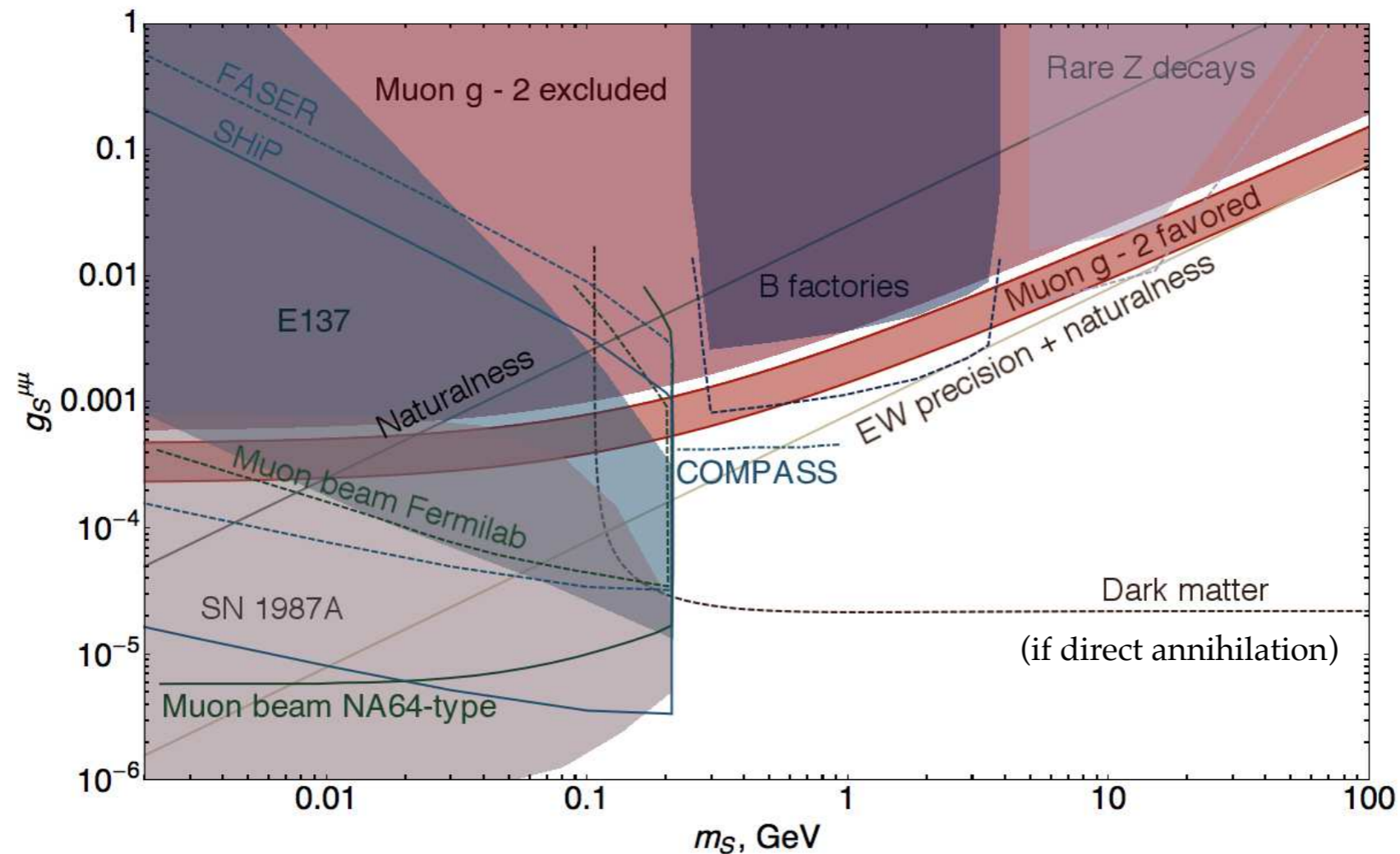
Decays to its favorite flavors



Requires fermion DM for p-wave

$$\sigma v \propto v^2$$

Flavored Scalar Mediator



Heavy vectorlike leptons mix with muon, to generate mediator coupling
Mediator enables DM $>$ muon annihilation

“Axion Portal” Pseudoscalar Mediator

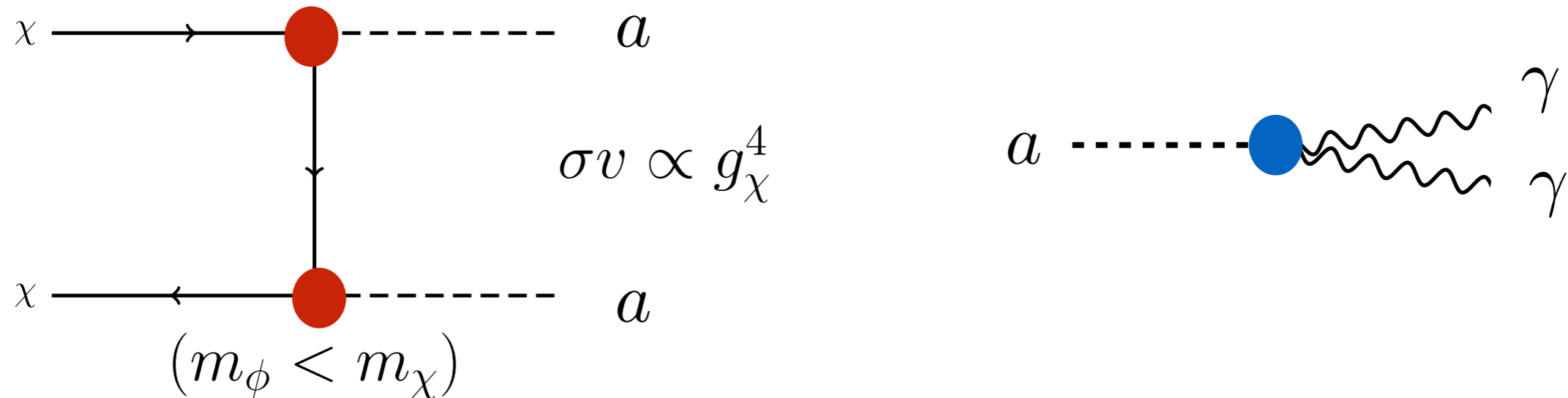
Axion like mediator

$$\mathcal{L}_{\text{eff}} \supset \frac{c_\gamma}{4\Lambda} a F^{\mu\nu} \tilde{F}_{\mu\nu} - \frac{4\pi\alpha_s c_g}{\Lambda} a G^{\mu\nu} \tilde{G}_{\mu\nu}$$

Add coupling to DM

$$\mathcal{L} \supset ig_\chi a \bar{\chi} \gamma^5 \chi$$

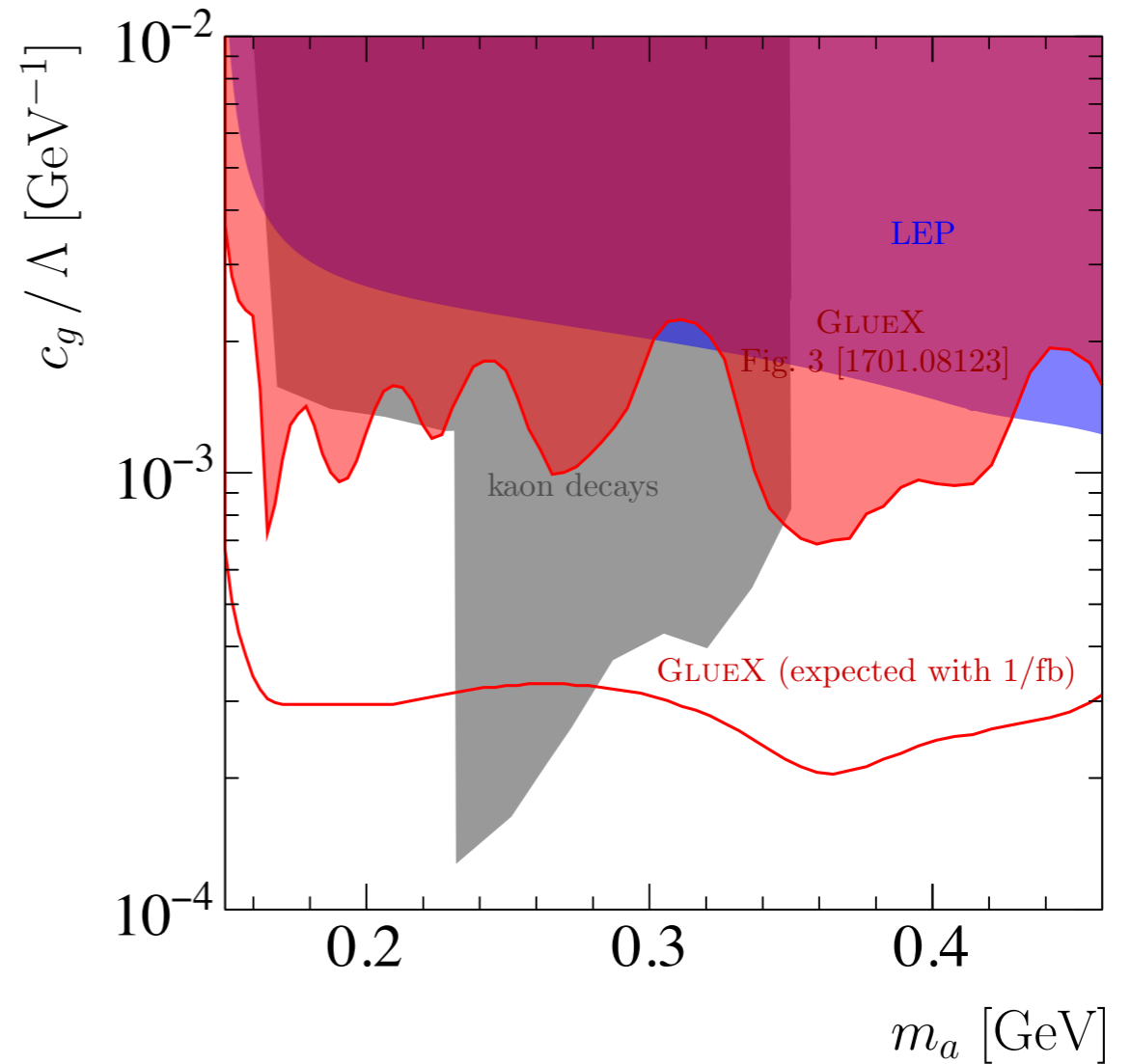
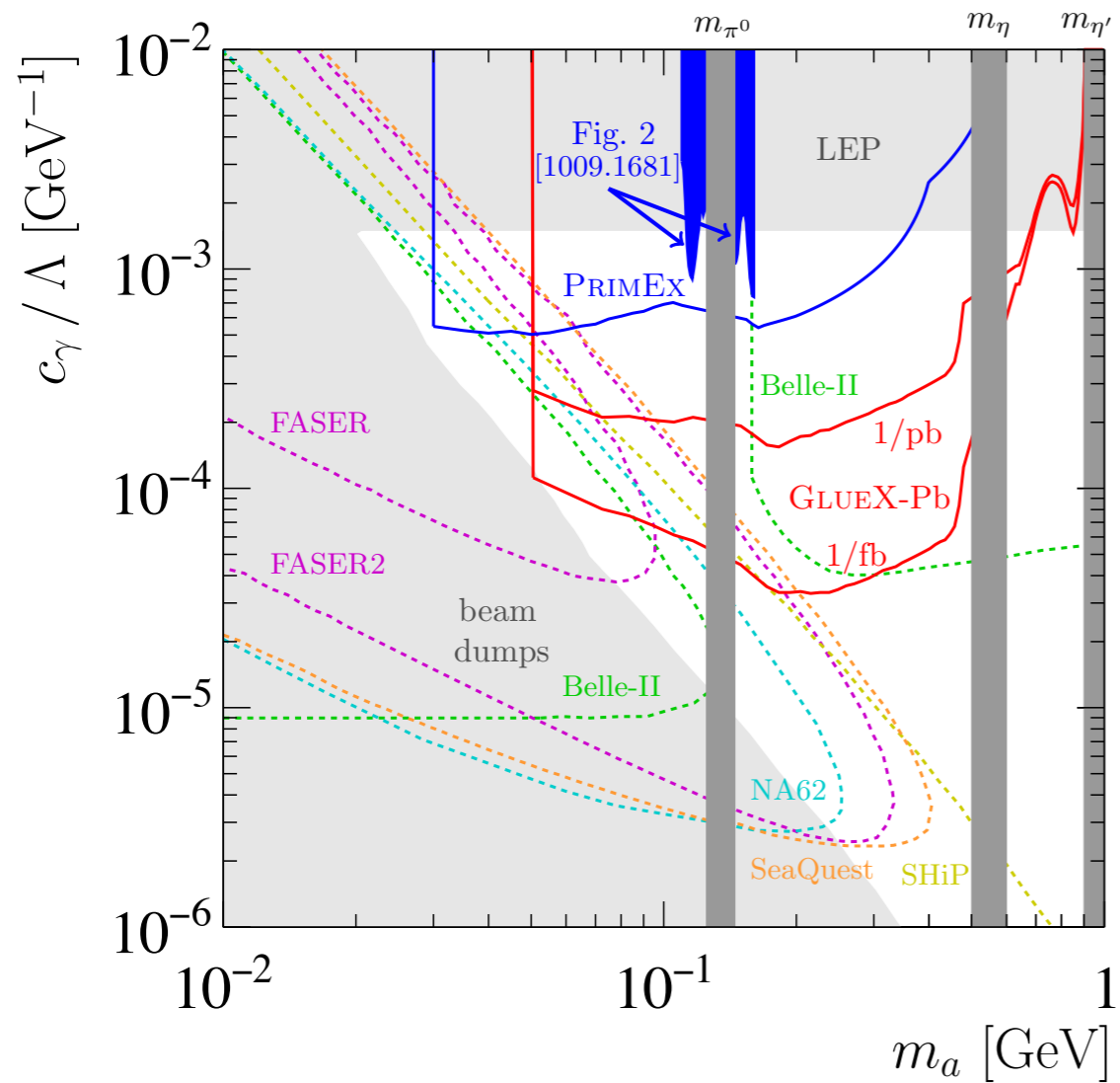
Nomura, Thaler 0810.5397



Requires fermion DM for p-wave $\sigma v \propto v^2$

Mixed annihilation bad $\chi\chi \rightarrow \phi$ (scalar) + a (pseudo) is s-wave

“Axion Portal” Pseudoscalar Mediator



Overview

- 1) P-Wave Annihilation
- 2) Annihilation Stops pre-CMB**
- 3) Mediator Decays to Neutrinos

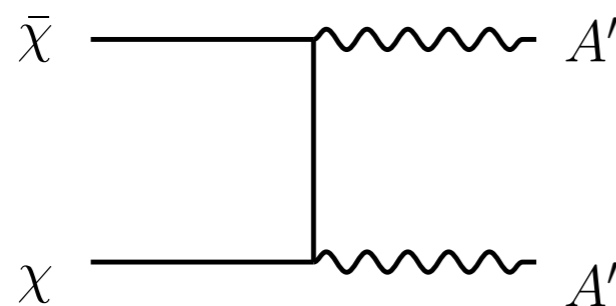
Asymmetric Dark Matter

Large particle/antiparticle asymmetry is safe from CMB

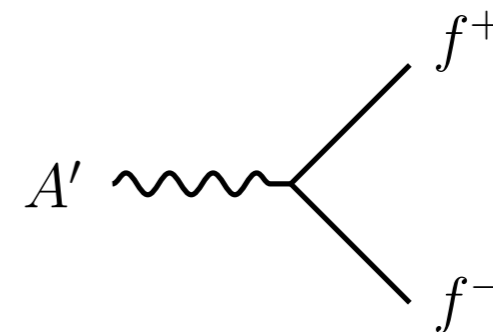
Lin, Yu, Zurek 1111.0293

$$n_{\bar{\chi}} \propto e^{-\sigma v}$$

Familiar vector mediators can now work for $< \text{GeV}$ asymmetric DM



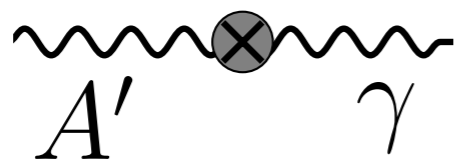
$$\sigma v \propto v^0 = \text{const.}$$



Works for kinetic mixing (dark photon A') or new gauged 5th force (V)

$$\epsilon e A'_\mu J_{\text{EM}}^\mu$$

$$g V_\mu J_{\text{SM}}^\mu$$

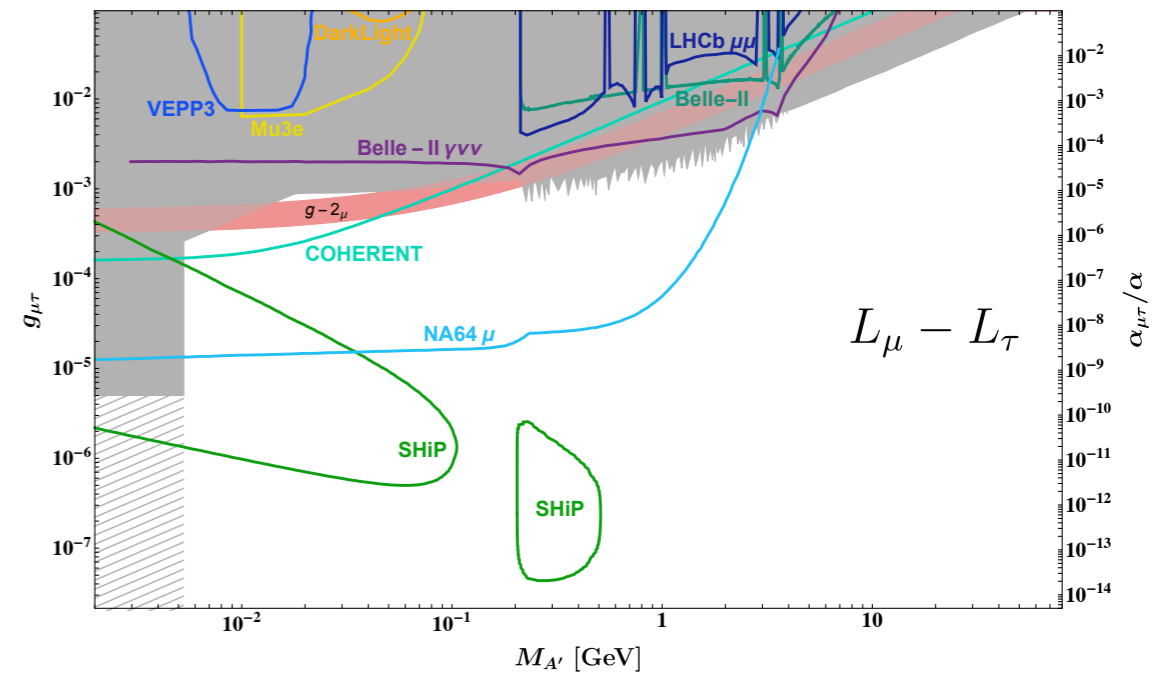
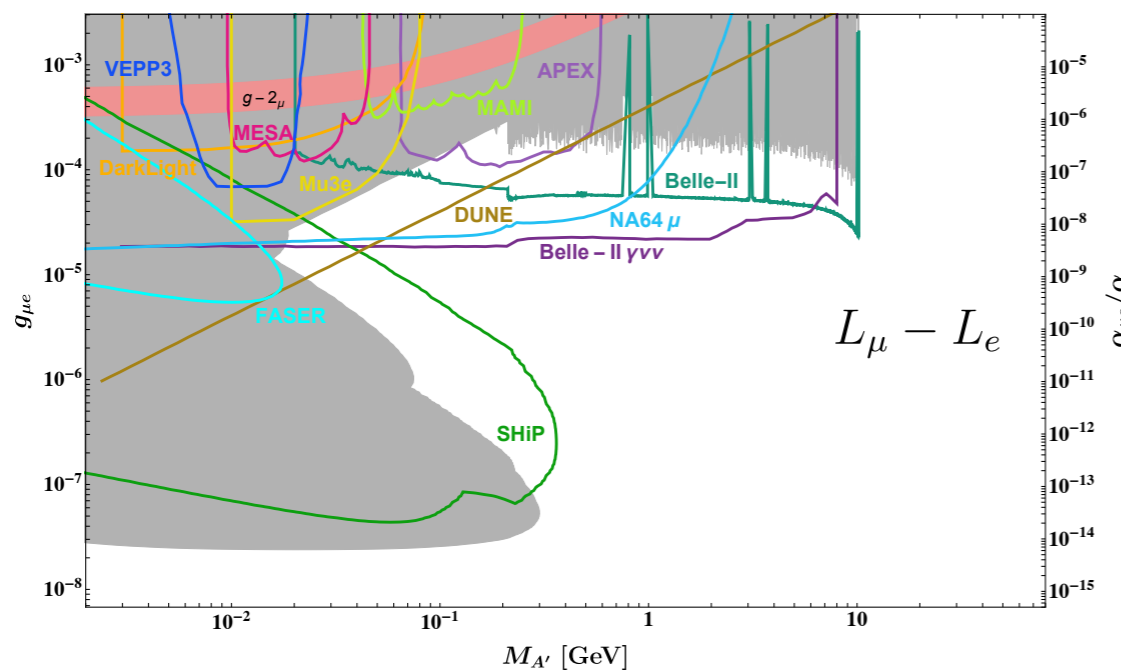
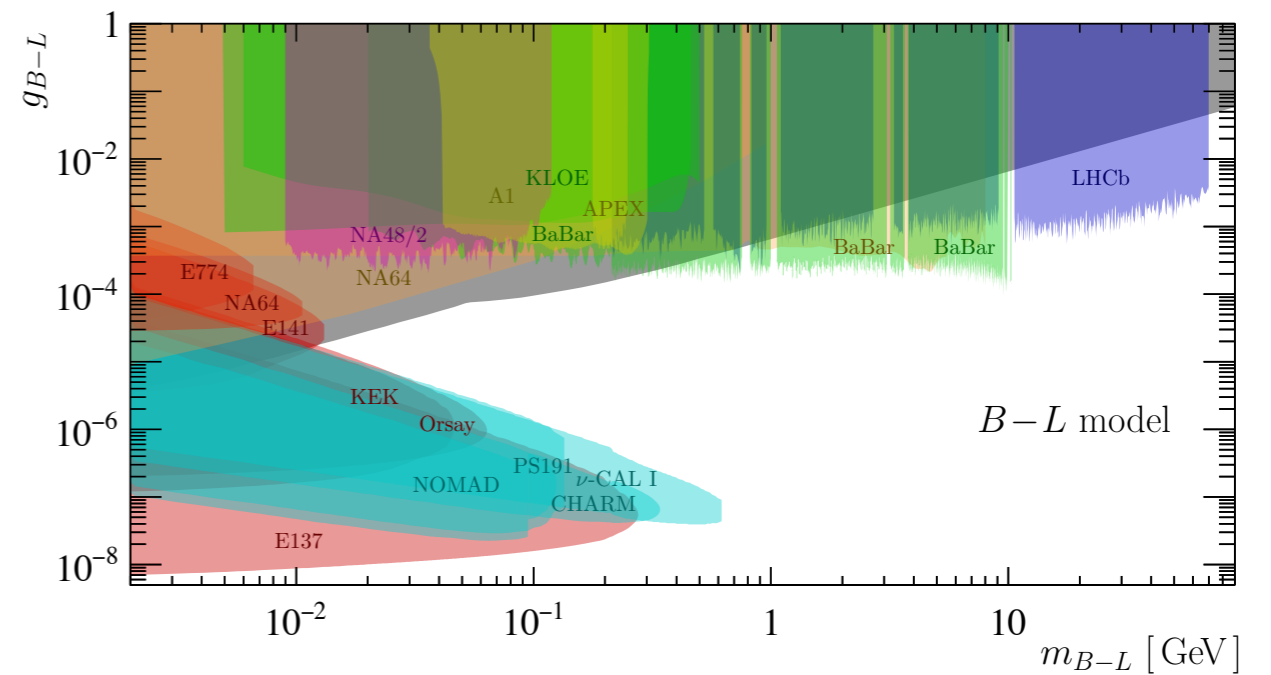
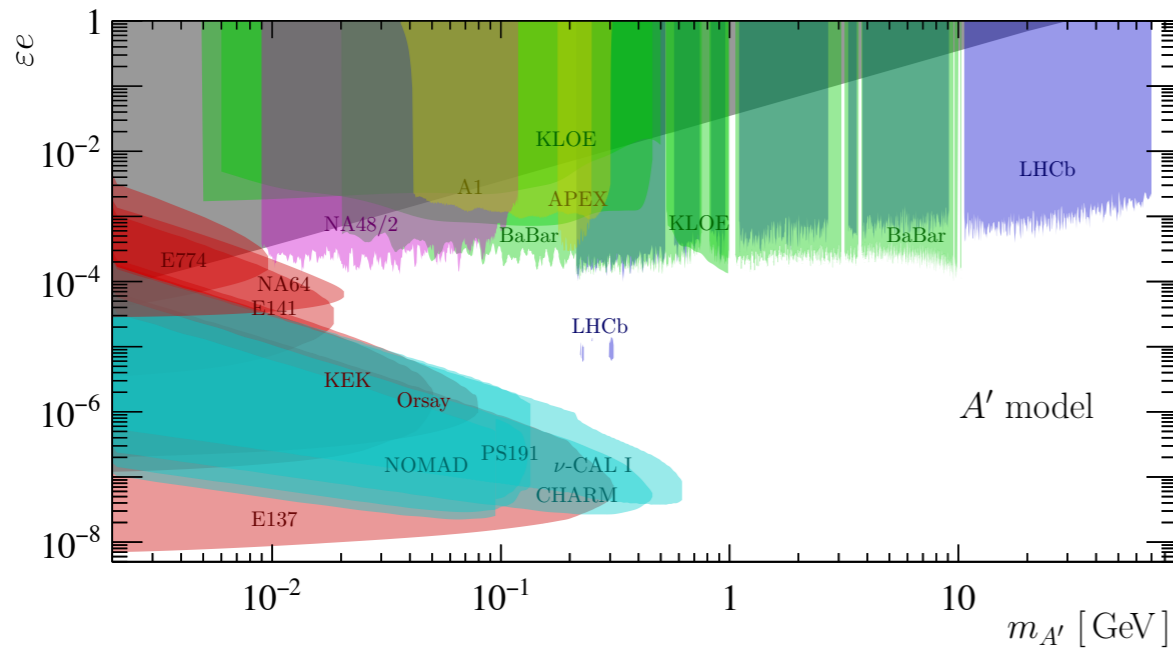


Anomaly free 5th force options

$$B - L, \quad L_i - L_j, \quad B - 3L_i$$

Asymmetric Dark Matter

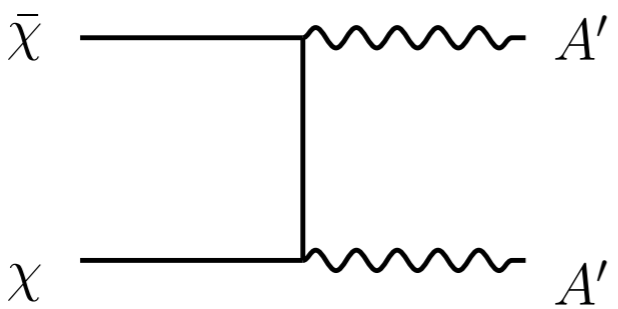
Iten Soreq Williams Xue 1801.04847



Bauer, Foldenauer, Jaeckel, 1803.05466

Forbidden Dark Matter

Mediator is *mildly heavier* than DM, but SM coupling is tiny



$$2m_\chi > m_{A'} > m_\chi$$

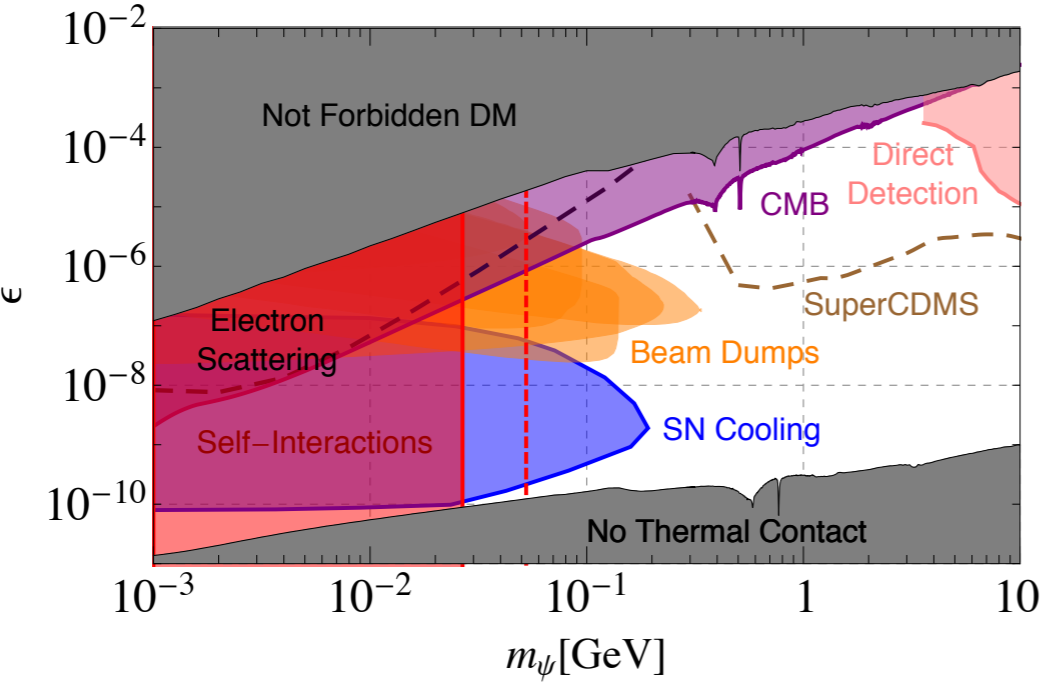
A' decays to SM

$$g_\chi \gg \epsilon$$

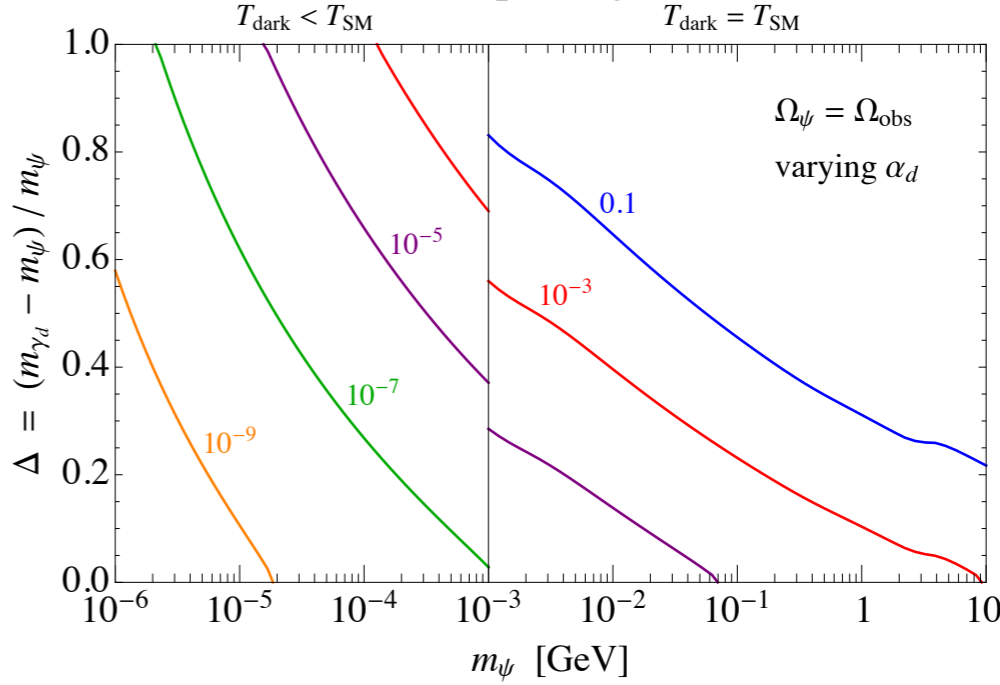
Suppress direct annihilation

Early universe DM has velocity tail, can annihilate to heavier A'

Signals from Kinetic Mixing ($\alpha_d=0.1$)

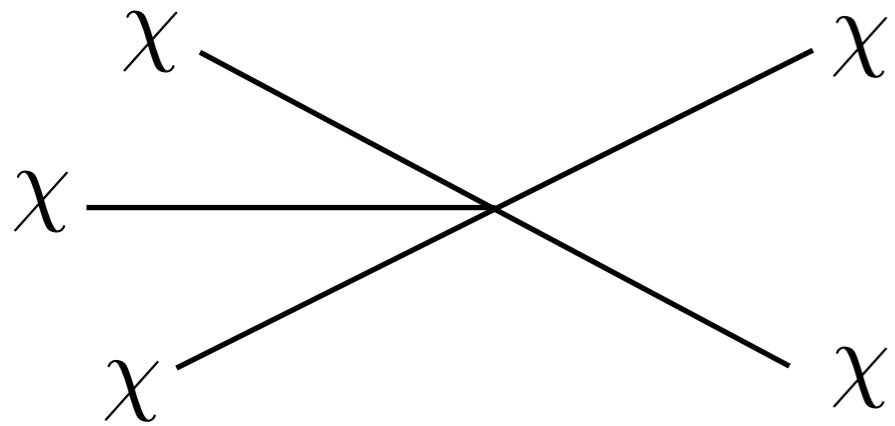


Splitting

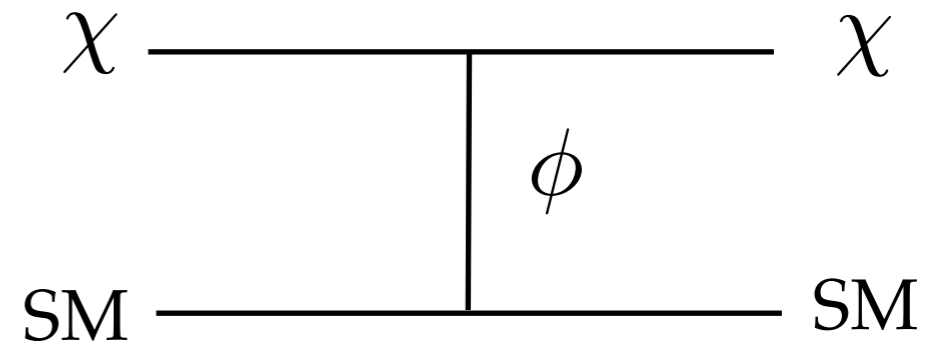


SIMPs and 3-2 Annihilation

Annihilation



Cooling w/ mediator



Cannibalization: 3-2 annihilation only (DM hot, ruled out)

Carlson, Machacek, Hall '92

SIMP: 3-2 freeze out, then SM scattering cools DM

Hochberg Kuflik Volansky Wacker 1402.5143

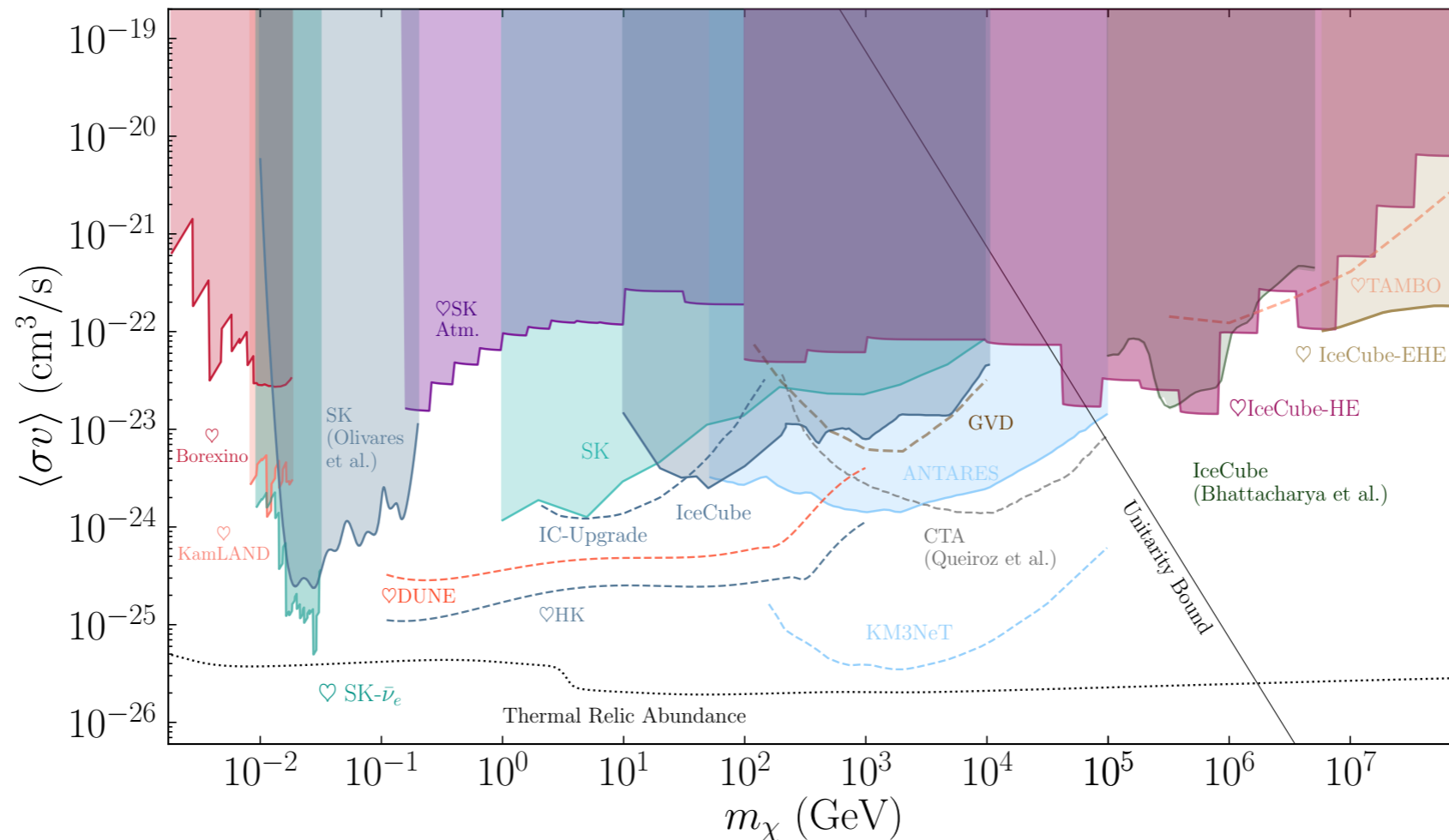
ELDER: SM-DM scattering decouples first, 3-2 freeze out later

Kuflik Prelstein Rey-Le Lorier, Tsai 1512.04545

Overview

- 1) P-Wave Annihilation
- 2) Annihilation Stops pre-CMB
- 3) Mediator Decays to Neutrinos**

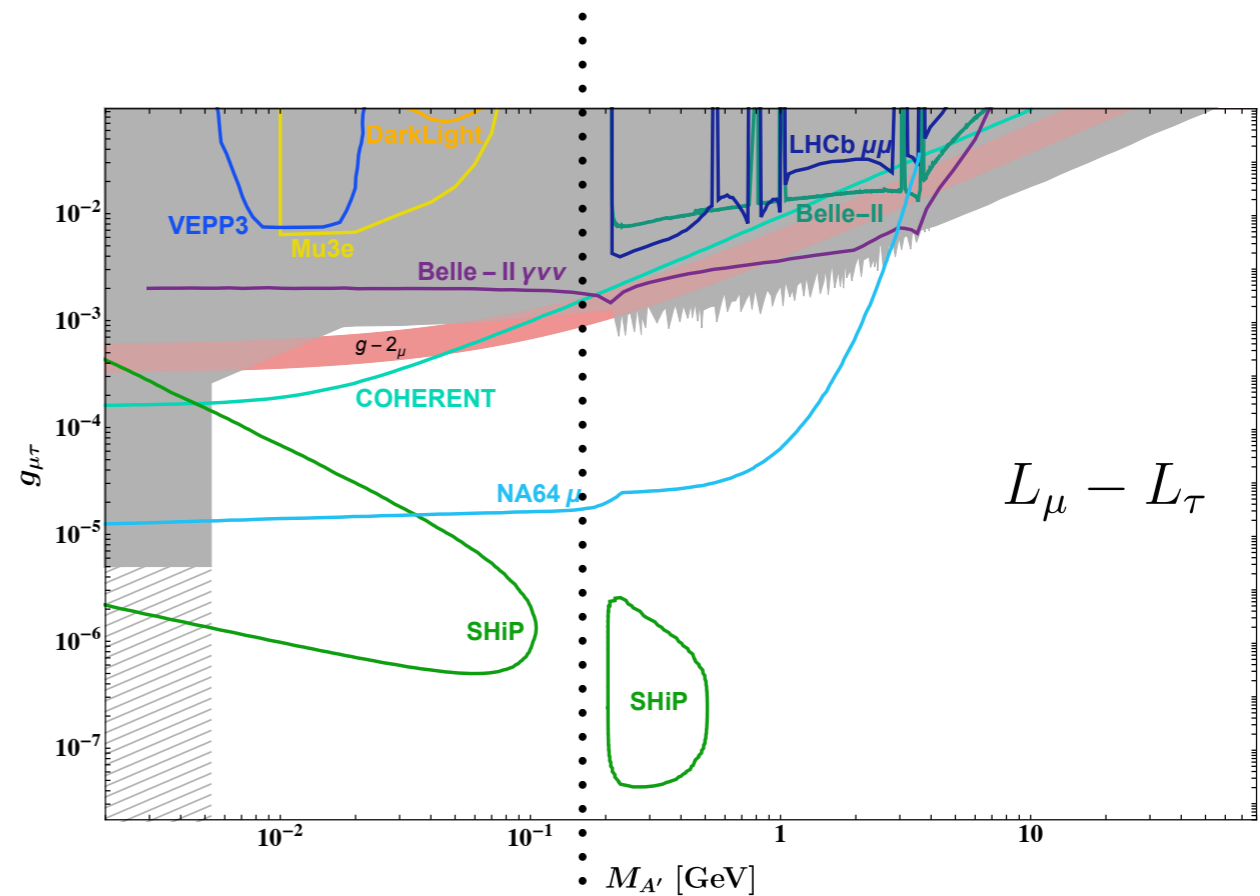
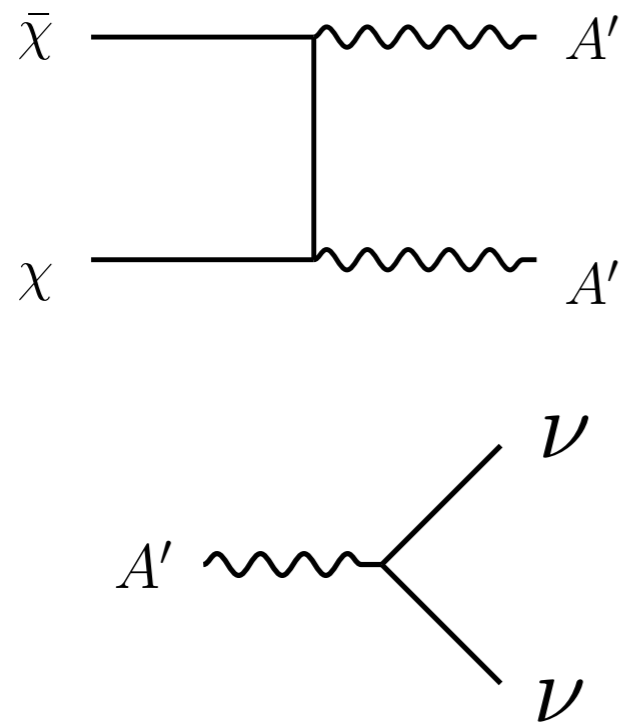
Agnostic Parameter Space



Bounds apply to direct and secluded
Easy to engineer scalar/pseudoscalar mediator

Vector Mediator

Leptophilic gauge boson mediator couples mainly 2nd & 3rd gen

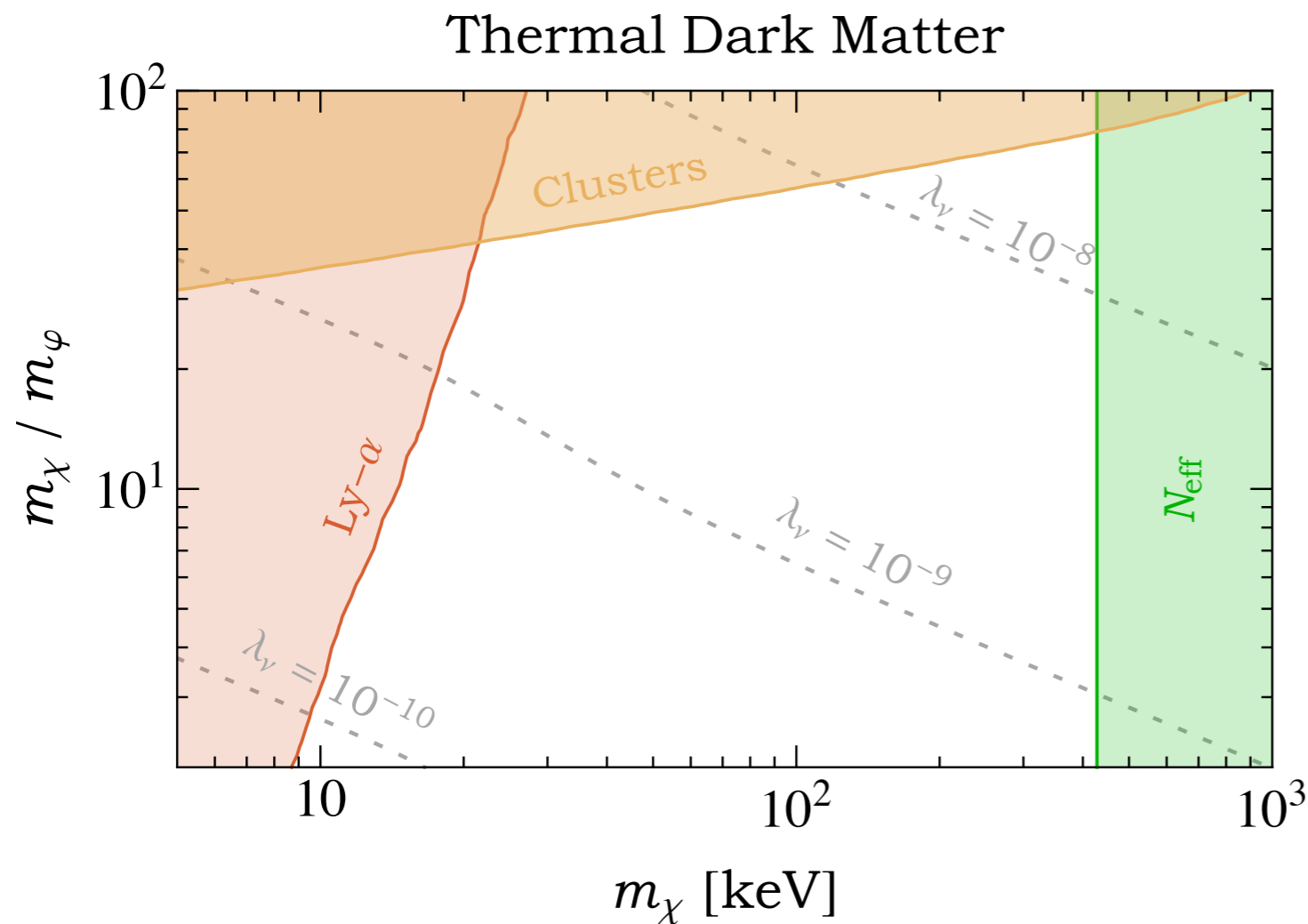


Bauer, Foldenauer, Jaeckel
1803.05466

Below muon mass, must decay to neutrinos $U(1)_{L_\mu - L_\tau}$
Only anomaly free vector option; others decay A' to electrons

Late Time Thermalization

Early times DM is not present, but produced when production rate exceeds Hubble after neutrinos decouple from photons



Light mediator $m \ll T$

$$\Gamma \sim n\sigma v g_\chi^4 T$$

$$H \sim T^2 / M_{\text{Pl}}$$

Equilibrate and freeze-out
below $T < \text{MeV}$

$$\chi\chi \rightarrow \phi\phi \rightarrow 4\nu$$

Berlin, Blinov 1706.07046

Chacko, Hall, Oliver, Perelstein 0405067

Lightest possible thermal DM ~ 10 keV

Concluding Remarks

Secluded Dark Matter

Equilibrium with SM but abundance independent of SM coupling

Need light new mediator that decays to SM particles

CMB Safety Organizes Options

1) P-Wave Annihilation

Motivates (pseudo)scalar mediators

DM annihilation to vectors almost always s-wave

2) Annihilation Stops pre-CMB

Asymmetric DM, SIMPs/EDLERS, forbidden etc.

3) Mediator Decays to Neutrinos

Neutrino-philic mediators (majorons mu-tau gauge bosons)

Late time thermalization (~ 10 keV scale thermal DM)