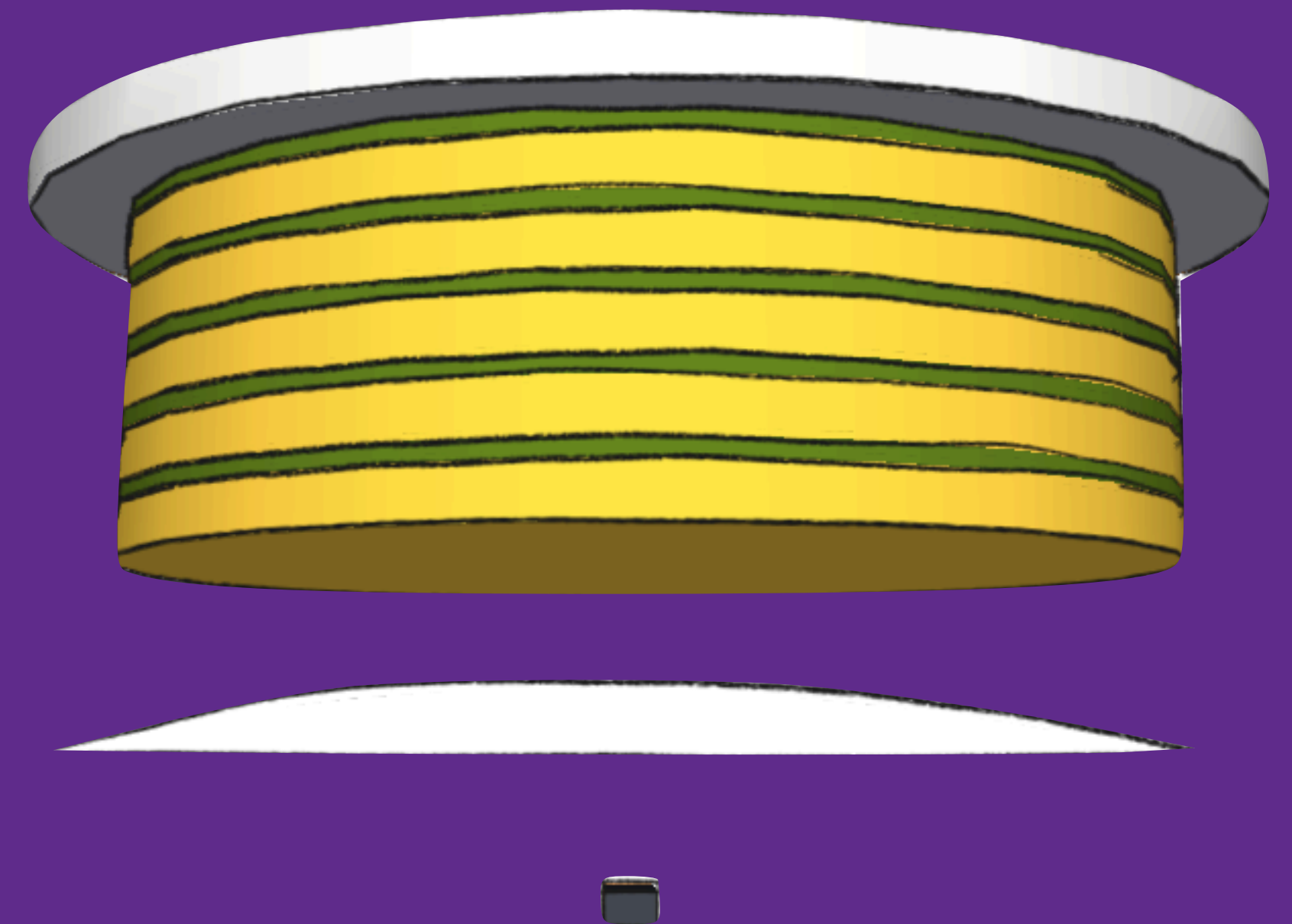


# DARK PHOTON SEARCH WITH OPTICAL HALOSCOPES

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NYUAD and WIS Collaboration

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## Dark Photons

# Dark Photons

Among the **most favourable candidates** for DM are WIMPs.

But what about new “species” of light dark matter  $< \text{GeV}$ ?

Many of the existing experiments that look for light DM are not sensitive to DM masses far above the **microwave range**.

We extend the search to higher masses, from 0.1 to 10 eV.

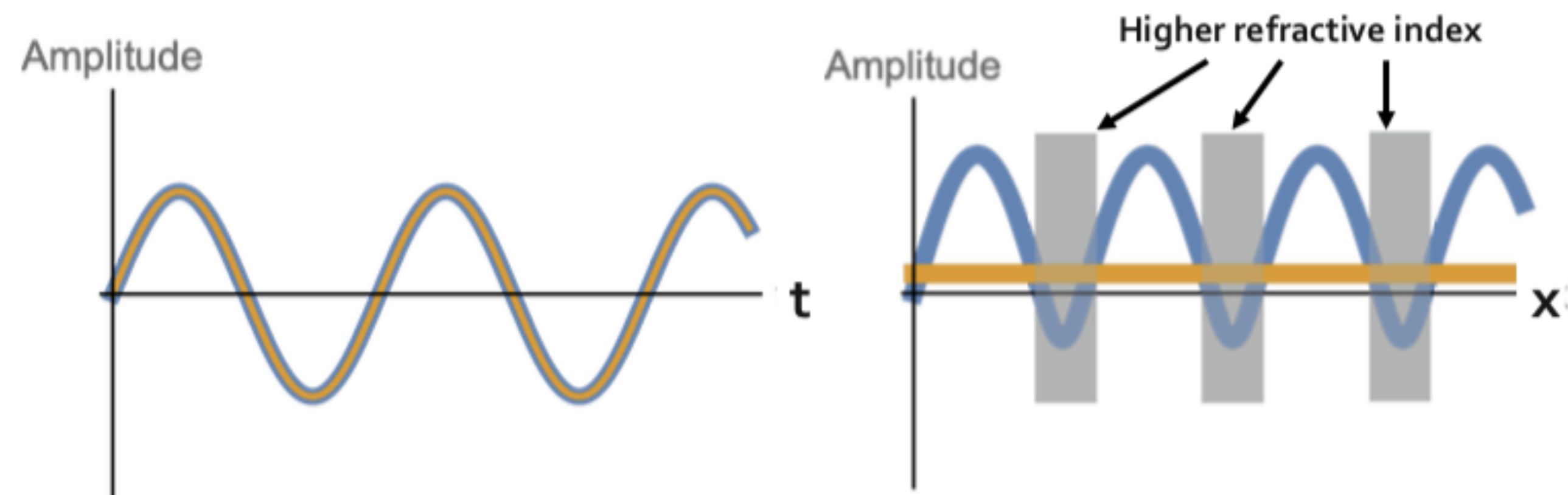
In particular, we look for dark photons around 1.5 eV.

How?

# Dark Photon conversion

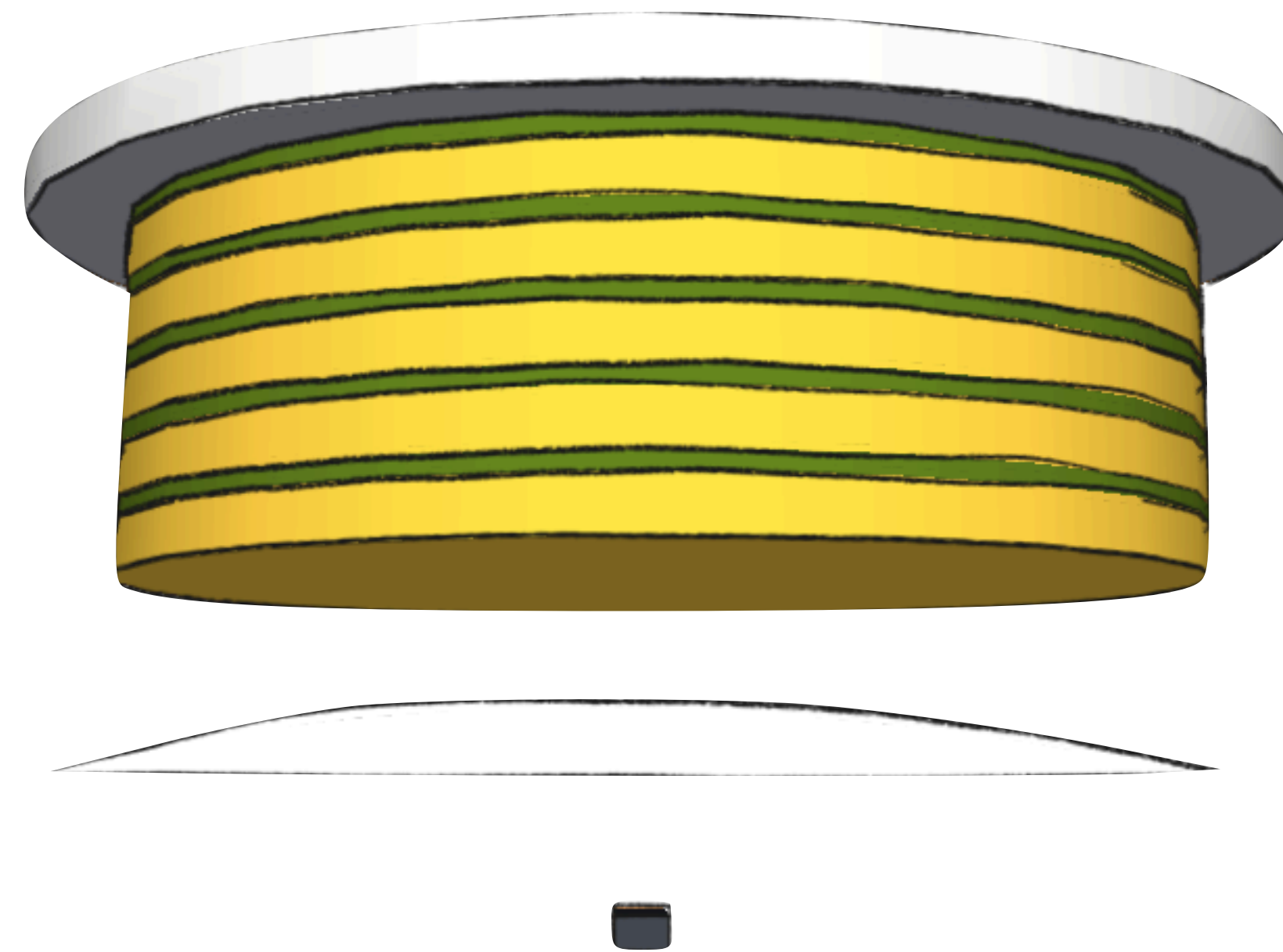
Dark Photons are converted to SM (detectable) photons.

A periodic index of refraction corrects the momentum mismatch between the DP and the SM photon, such that the electric field does not integrate to zero.



Dark Photon  
conversion

# Multilayer optical haloscopes

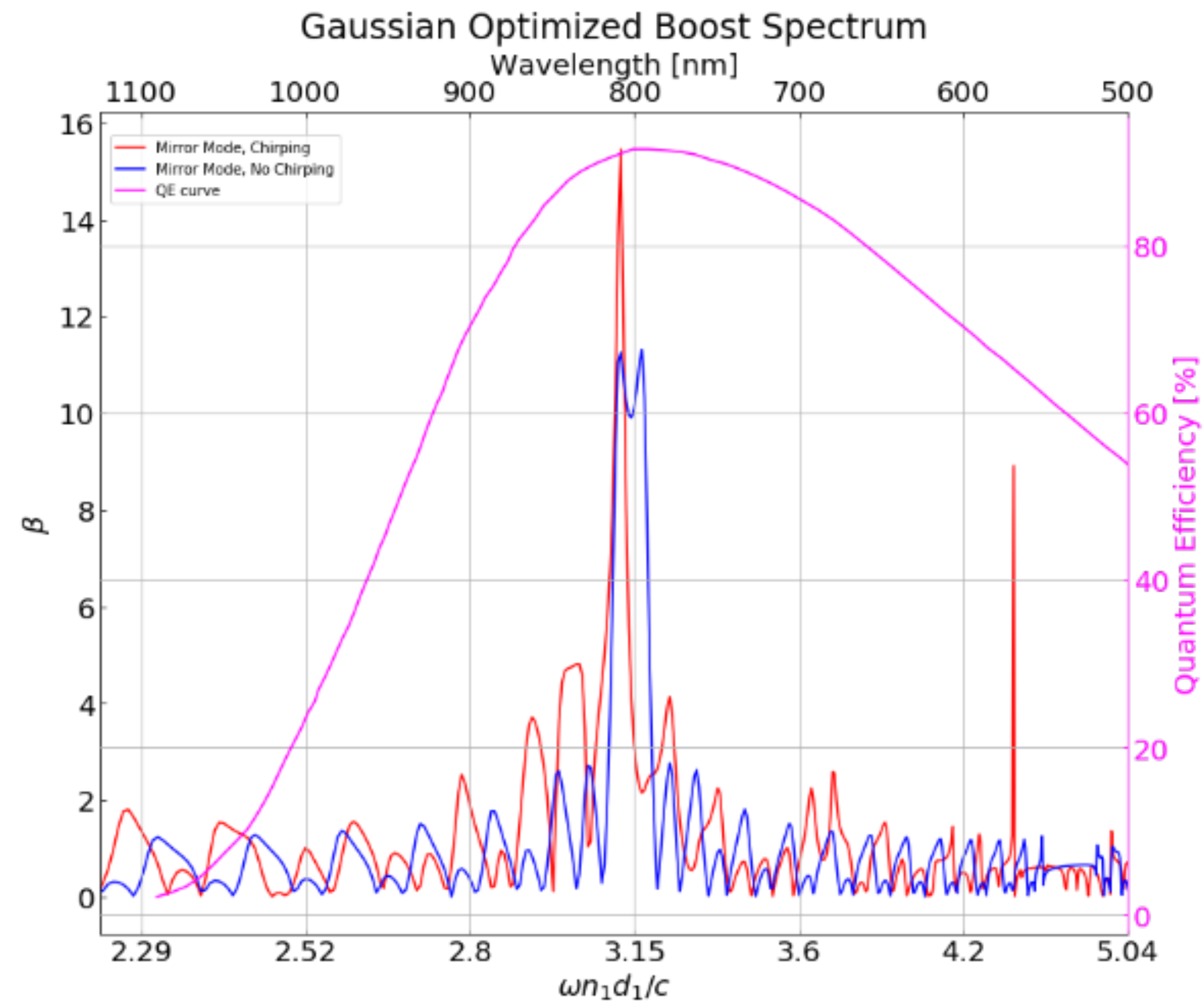
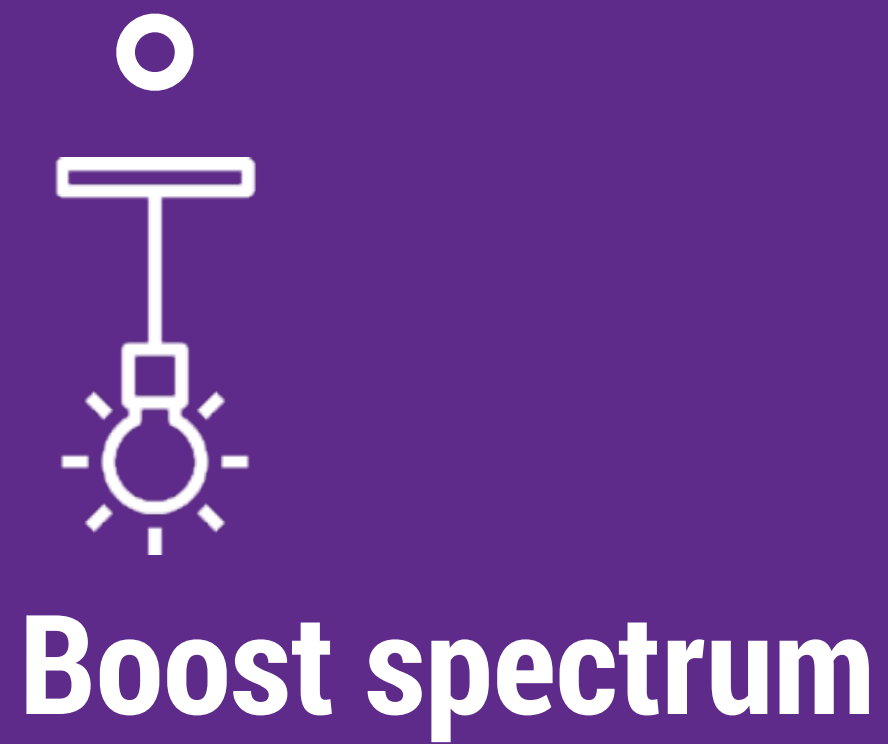


- Mirror
- Dielectric bilayers, with high contrast indices of refraction  $n_1$  and  $n_2$  (e.g. SiO<sub>2</sub> and Si<sub>3</sub>N<sub>4</sub>, with indices of 1.46 and 2.05 respectively)
- Lens
- Photon detector (QE peaked at the DP mass)



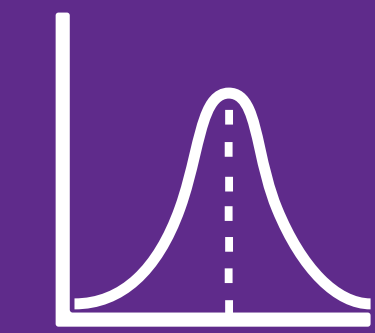
Multilayer  
dielectric  
haloscope

# Boost factor (beta) spectrum



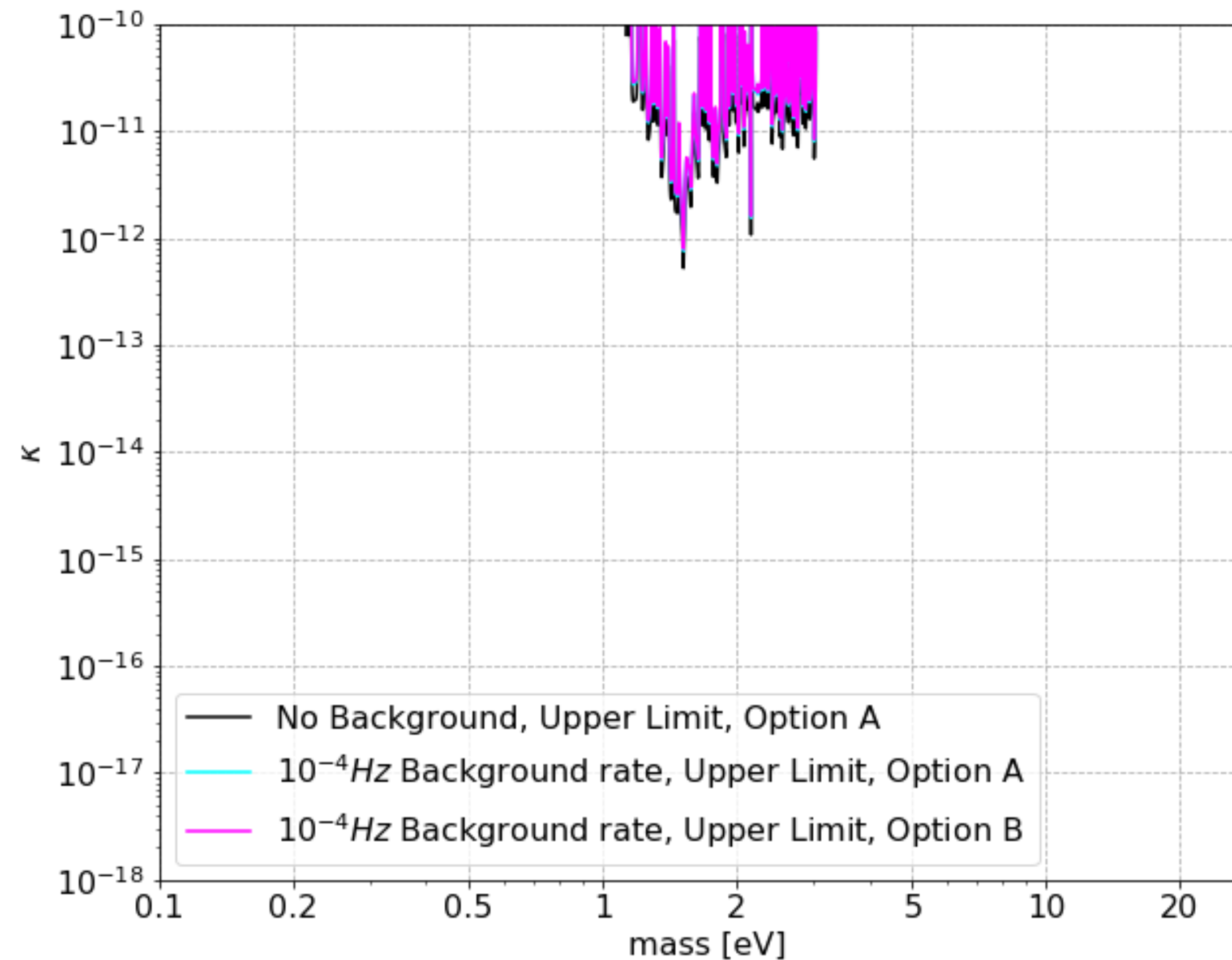
$$\Phi = 2.6 \times f_{DM}\eta \left( \frac{1}{m} \right) (\beta\kappa)^2 [(eV \times \text{days})^{-1} (\text{cm})^{-2}]$$





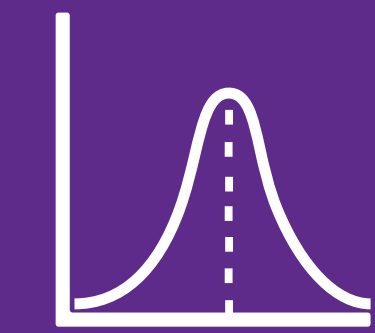
## Statistical analysis

# Upper limits



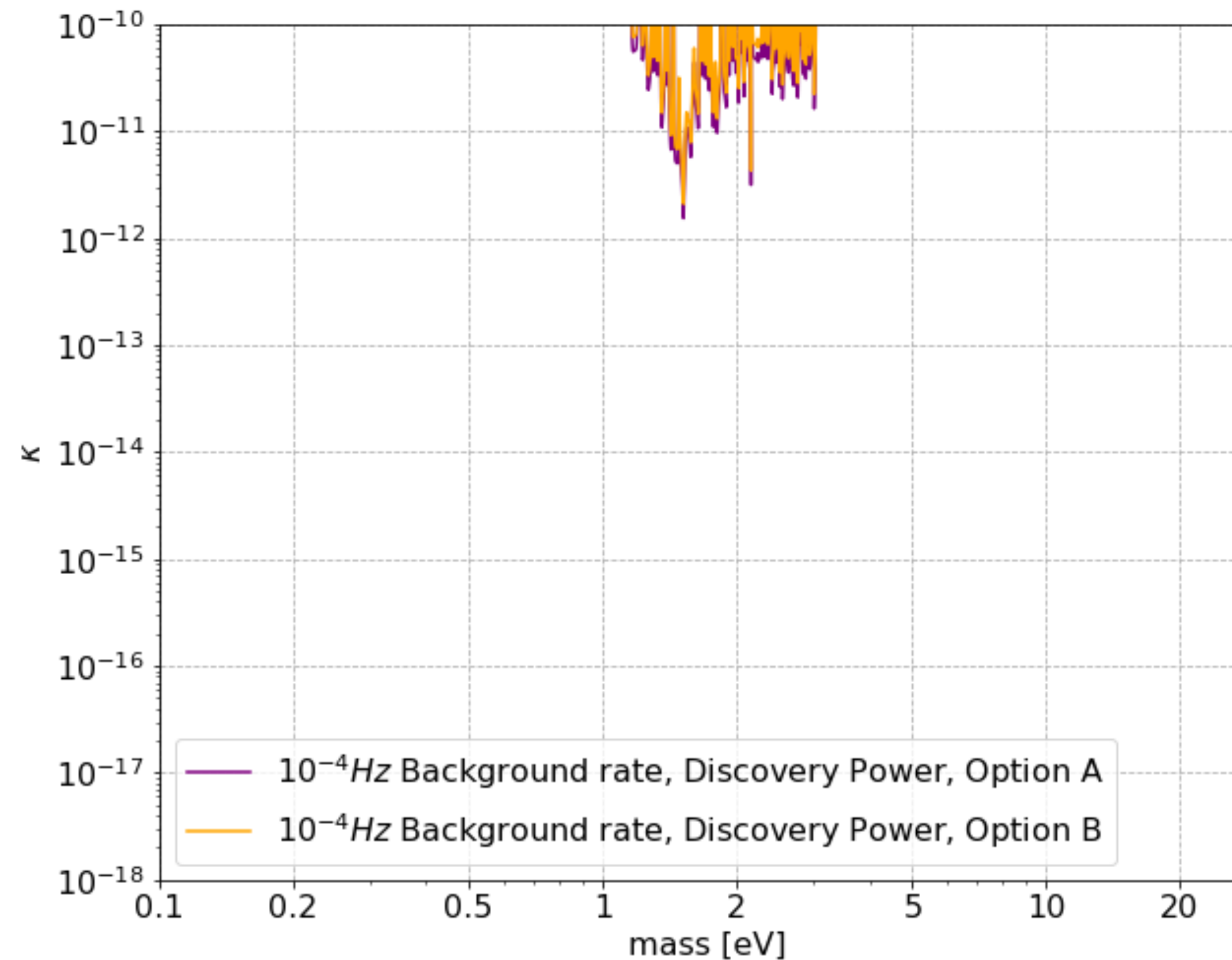
Option A: Poisson background with no uncertainty

Option B: Poisson background with uncertainty



## Statistical analysis

# Discovery limits



Option A: Poisson background with no uncertainty

Option B: Poisson background with uncertainty

# The stack



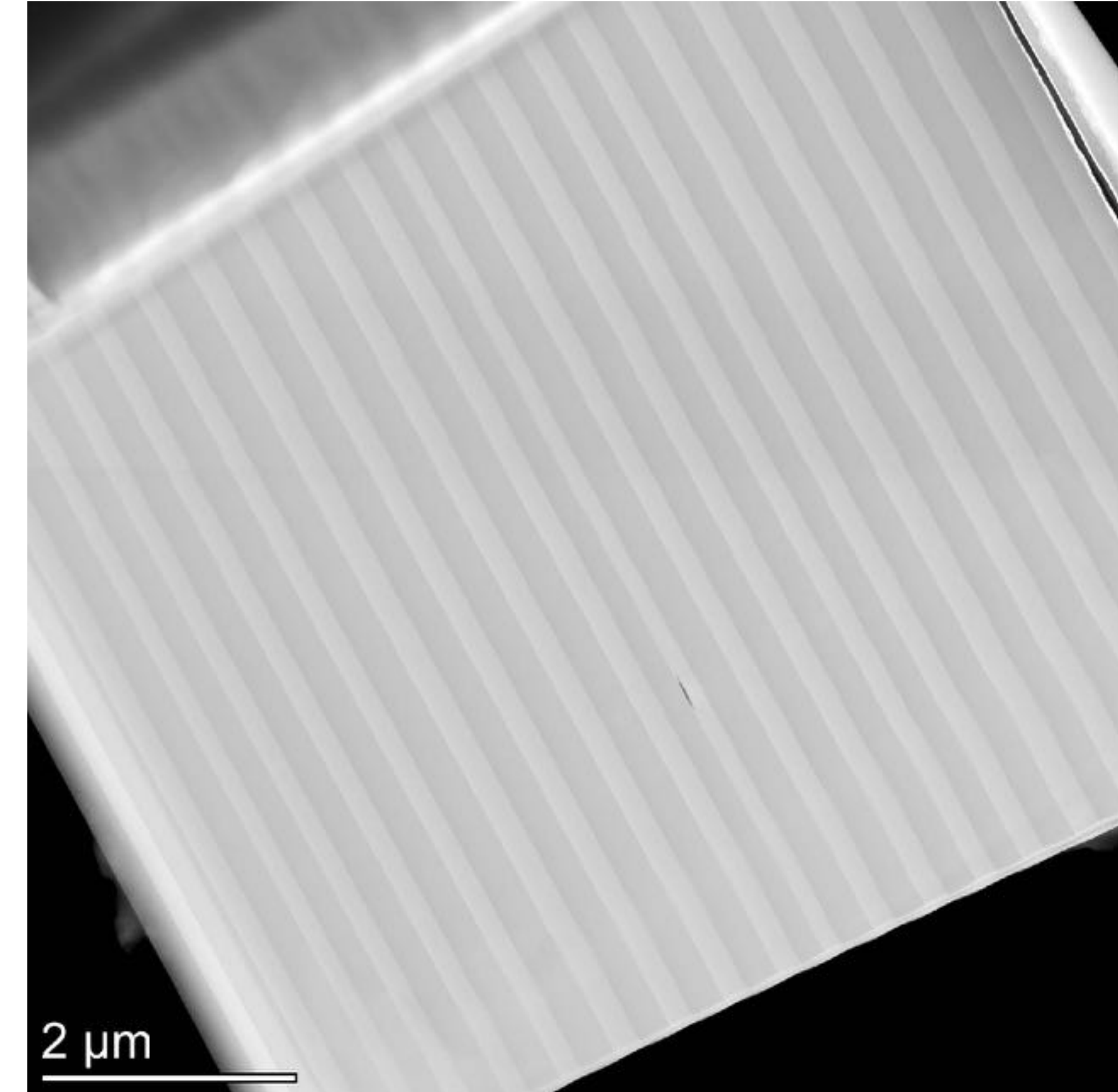
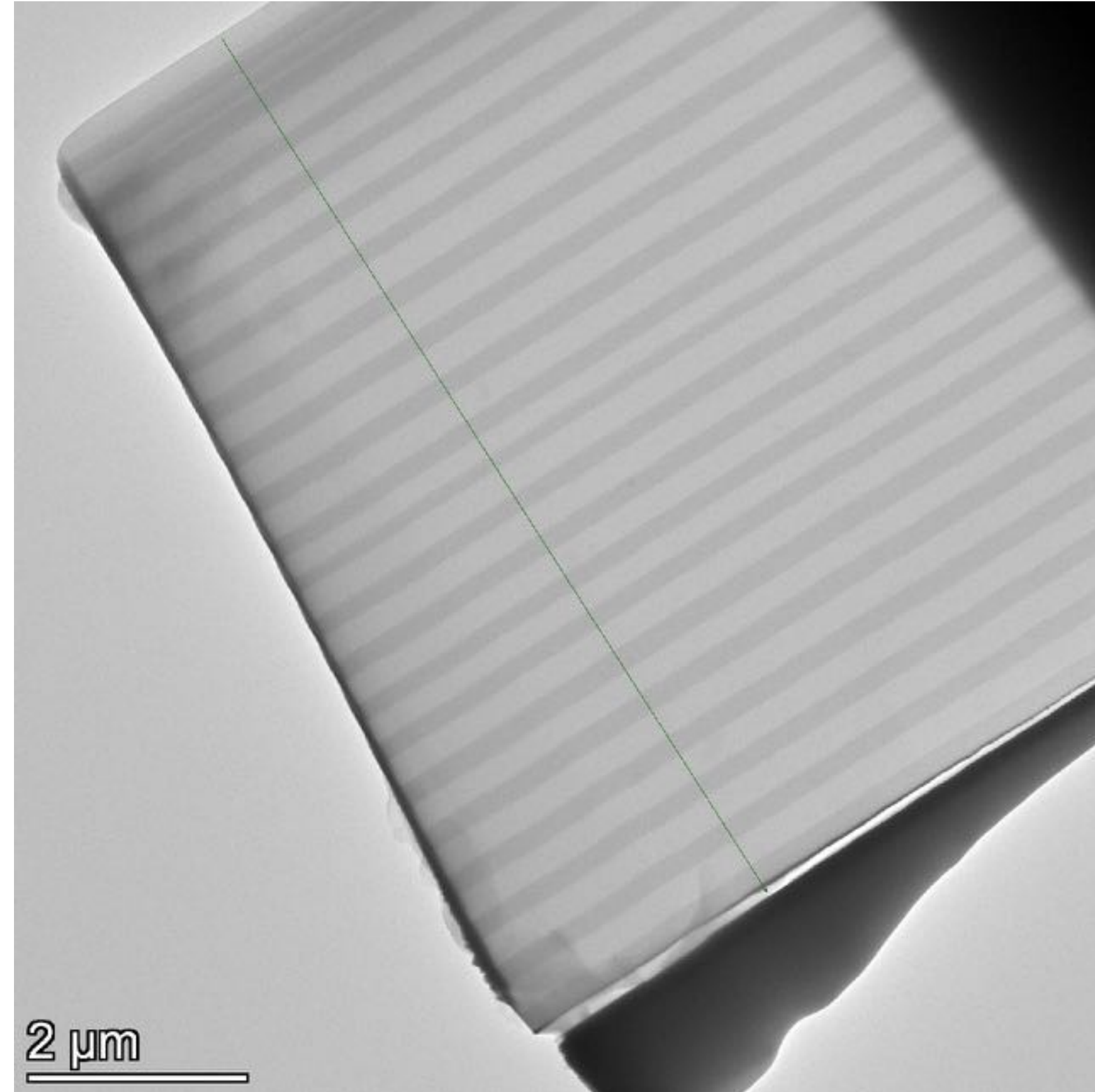
- The stack (actually 5 of them) was produced at the PoliFAB, the micro and nano technology center of the Politecnico di Milano.
- Technique used: plasma-enhanced chemical vapor deposition
- 46 layers (23  $\text{SiO}_2$  and 23  $\text{Si}_3\text{N}_4$ )



The stack



# TEM analysis

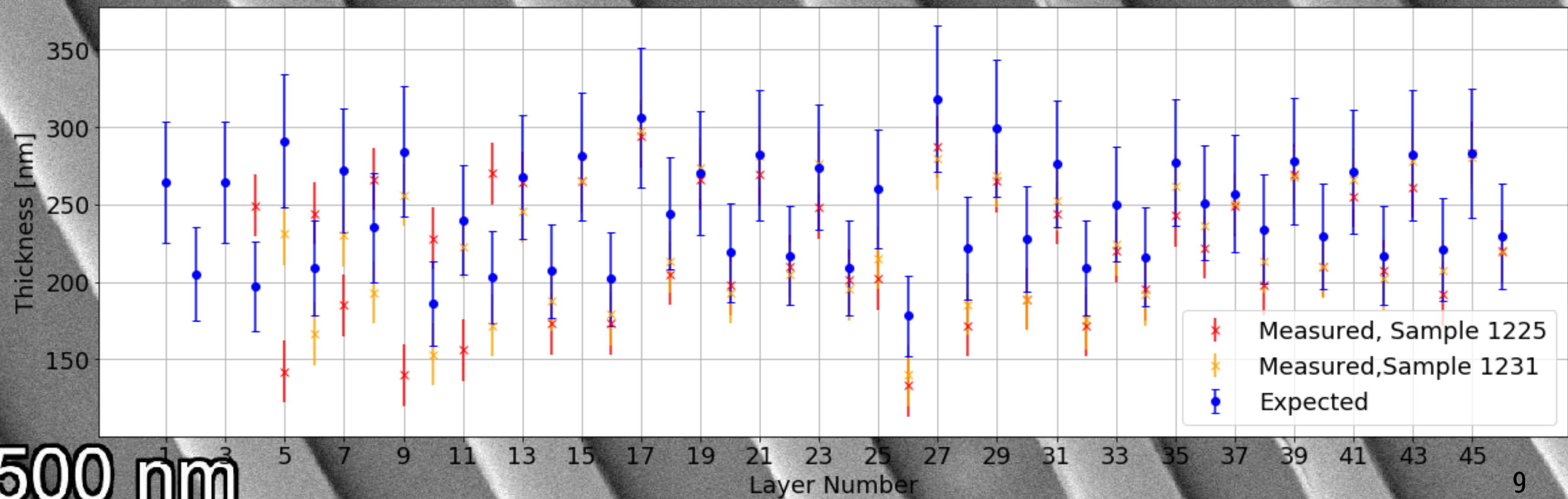
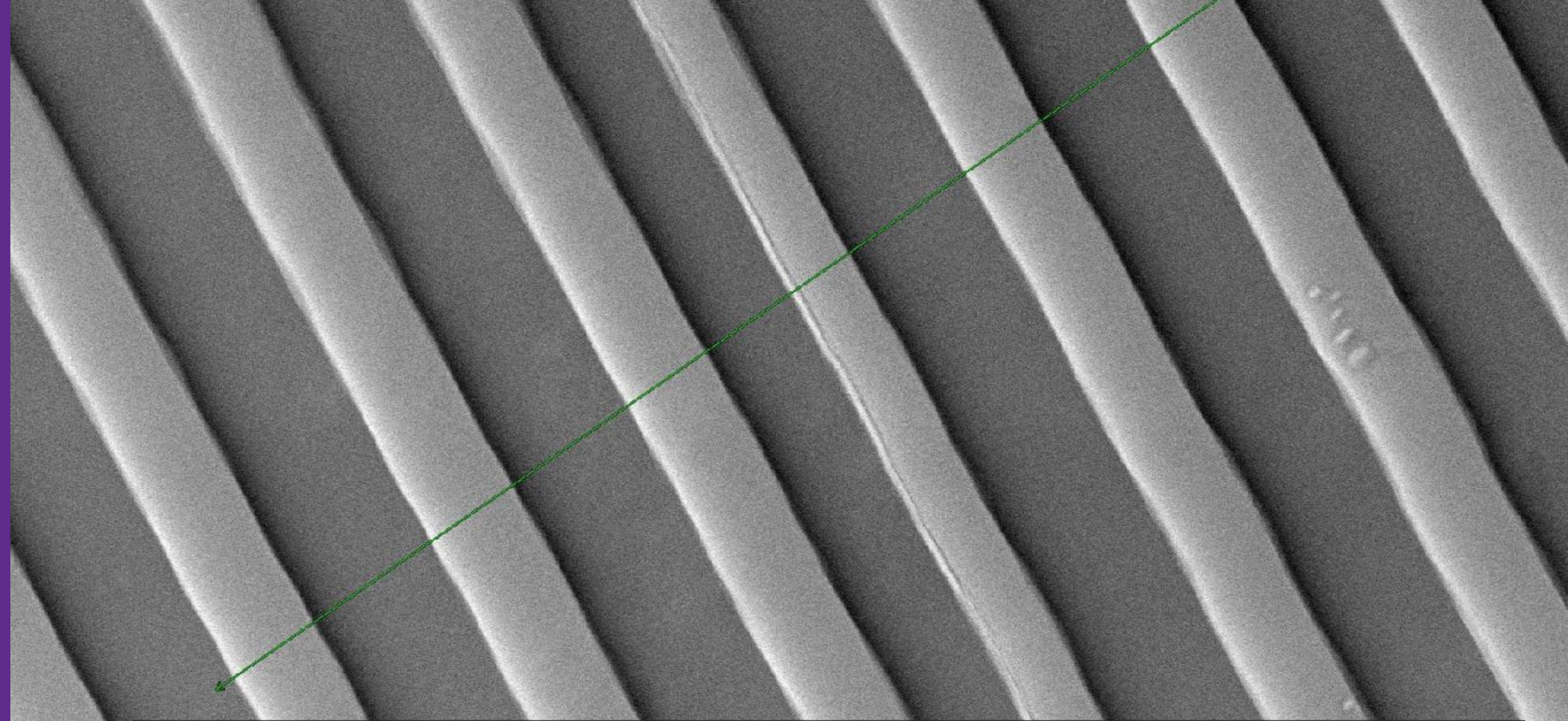


TEM analysis





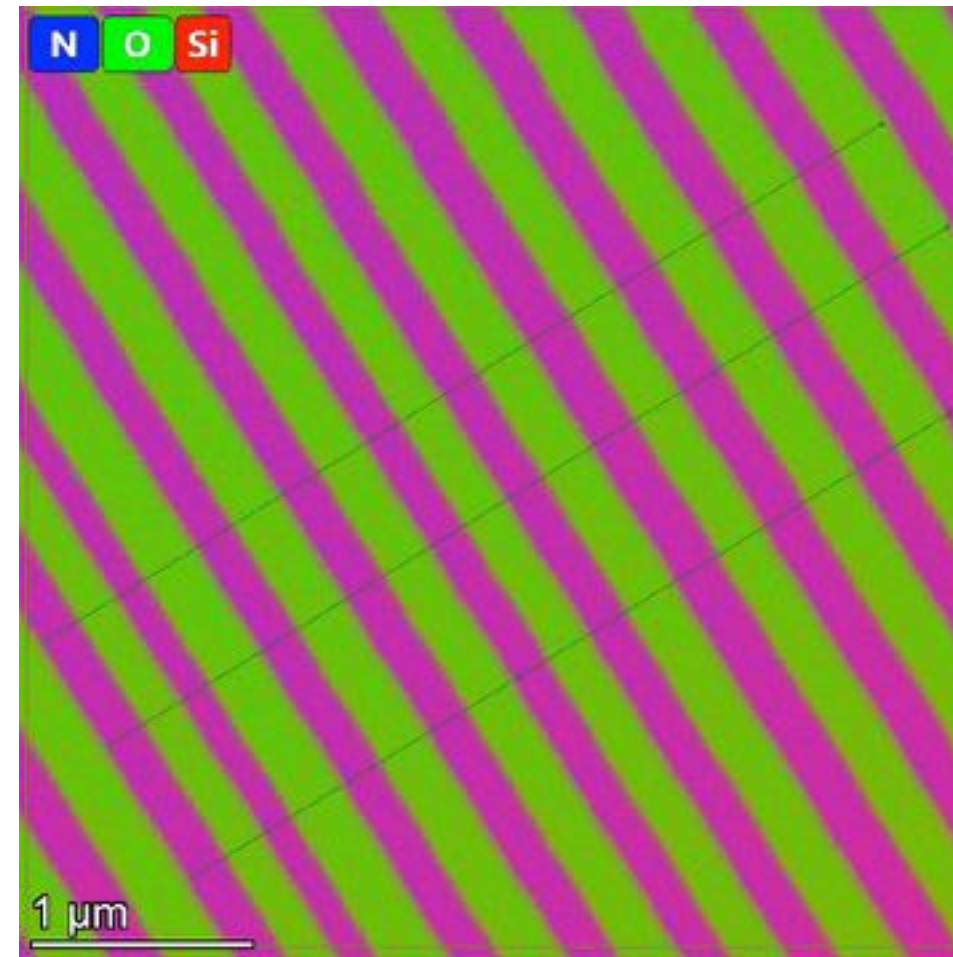
# TEM analysis



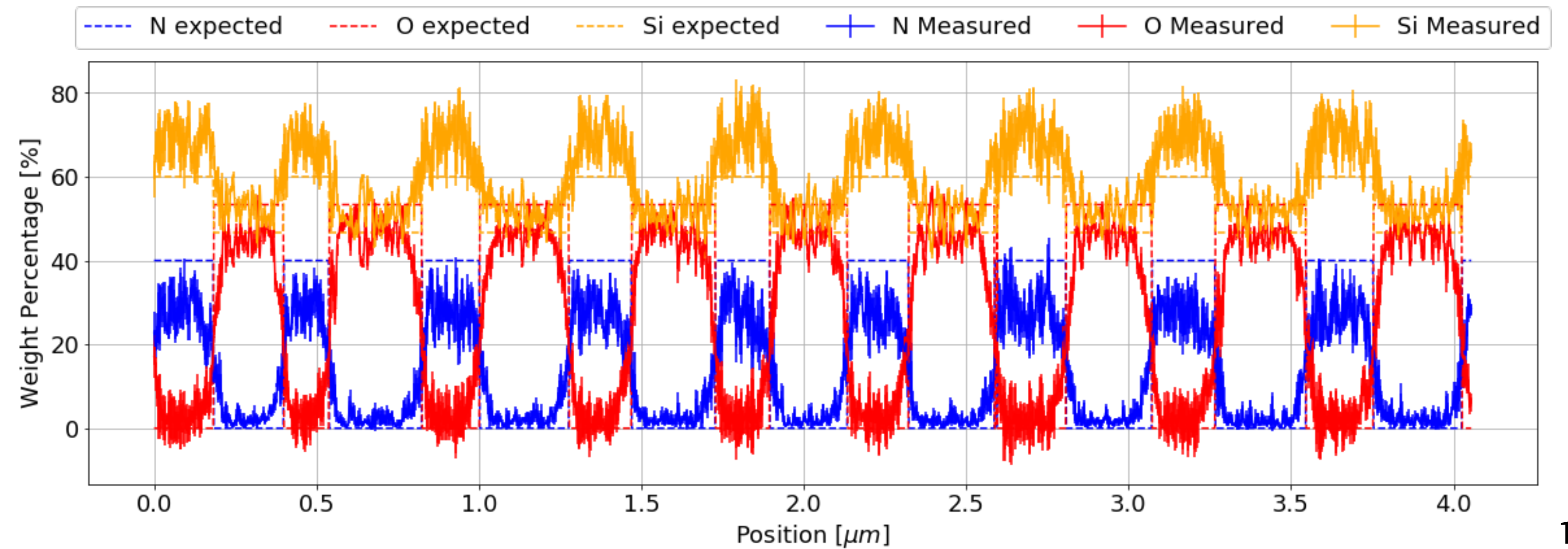
500 nm



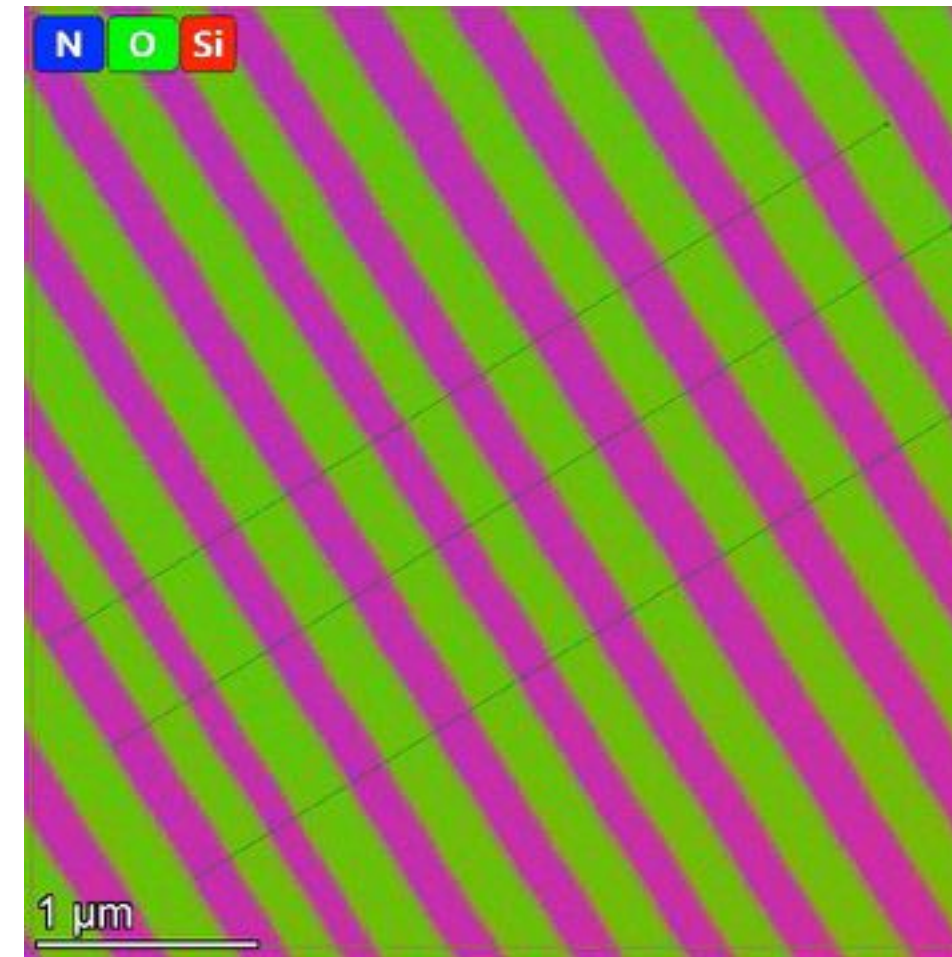
# EDAX analysis



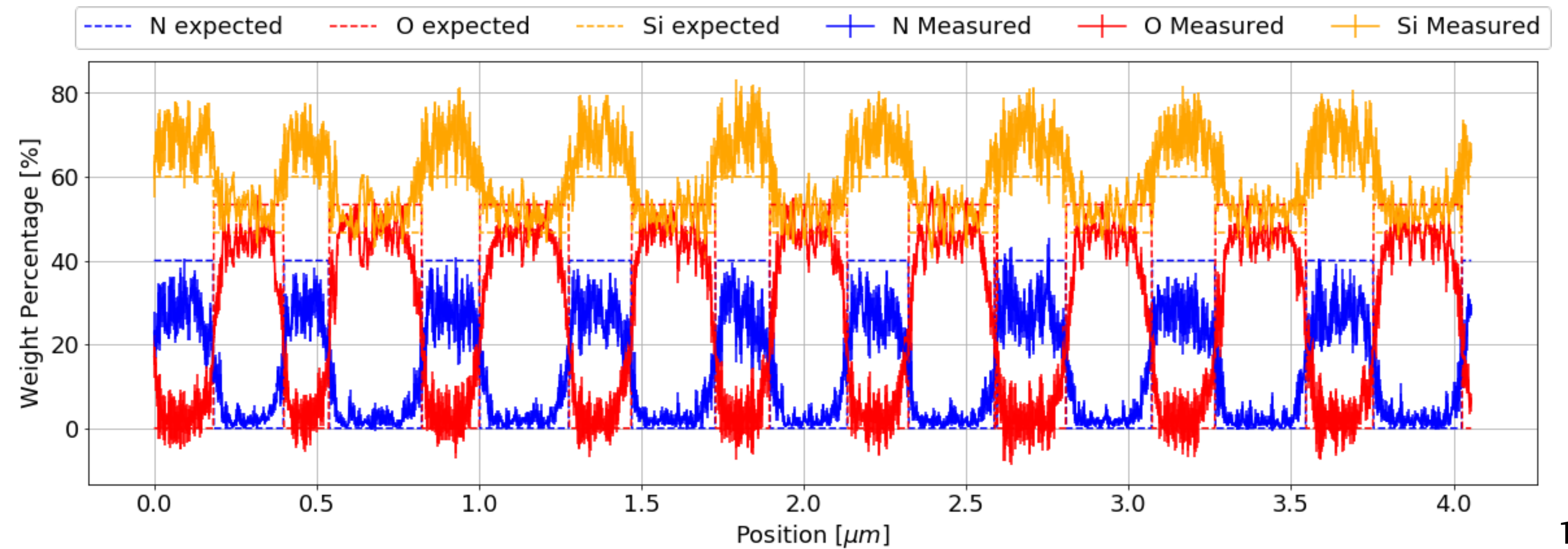
EDAX analysis



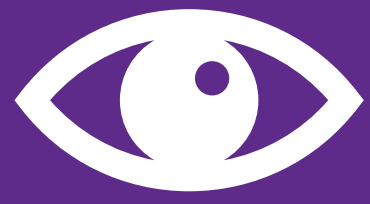
# EDAX analysis



EDAX analysis

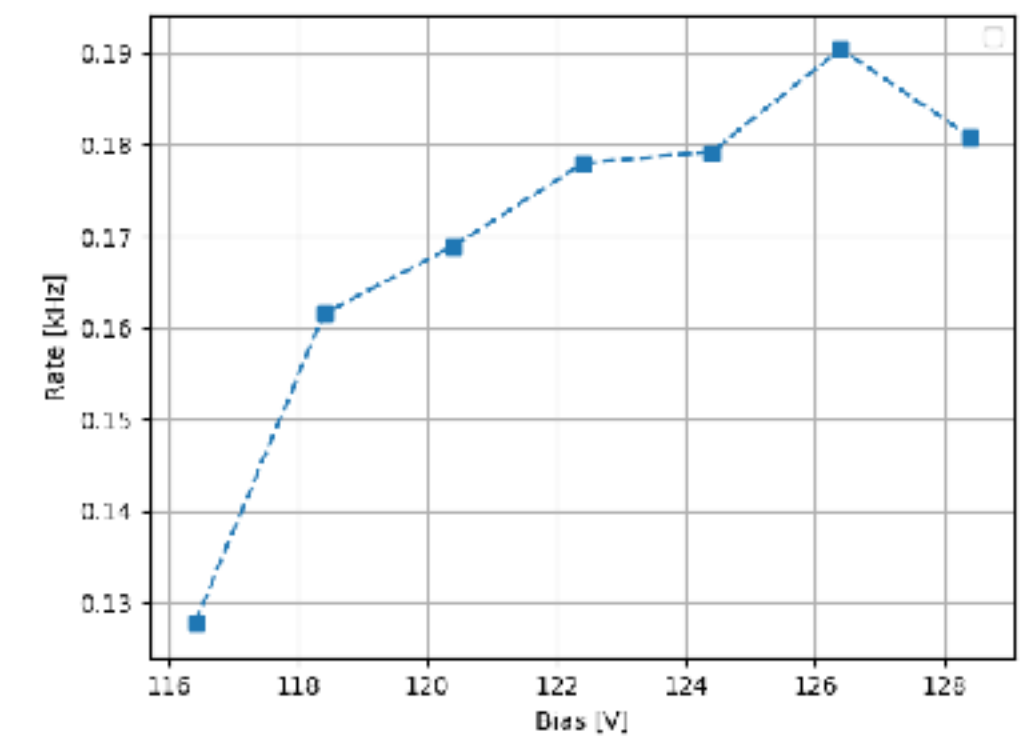
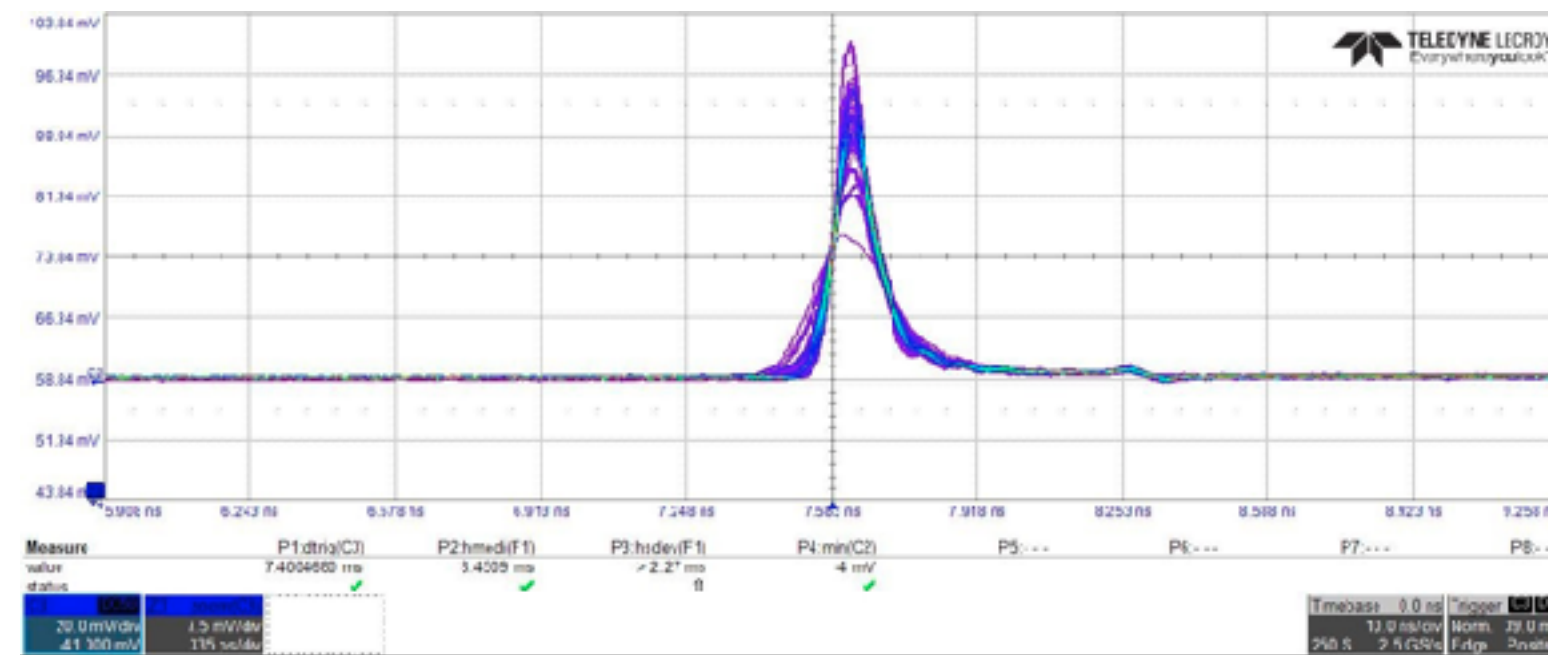
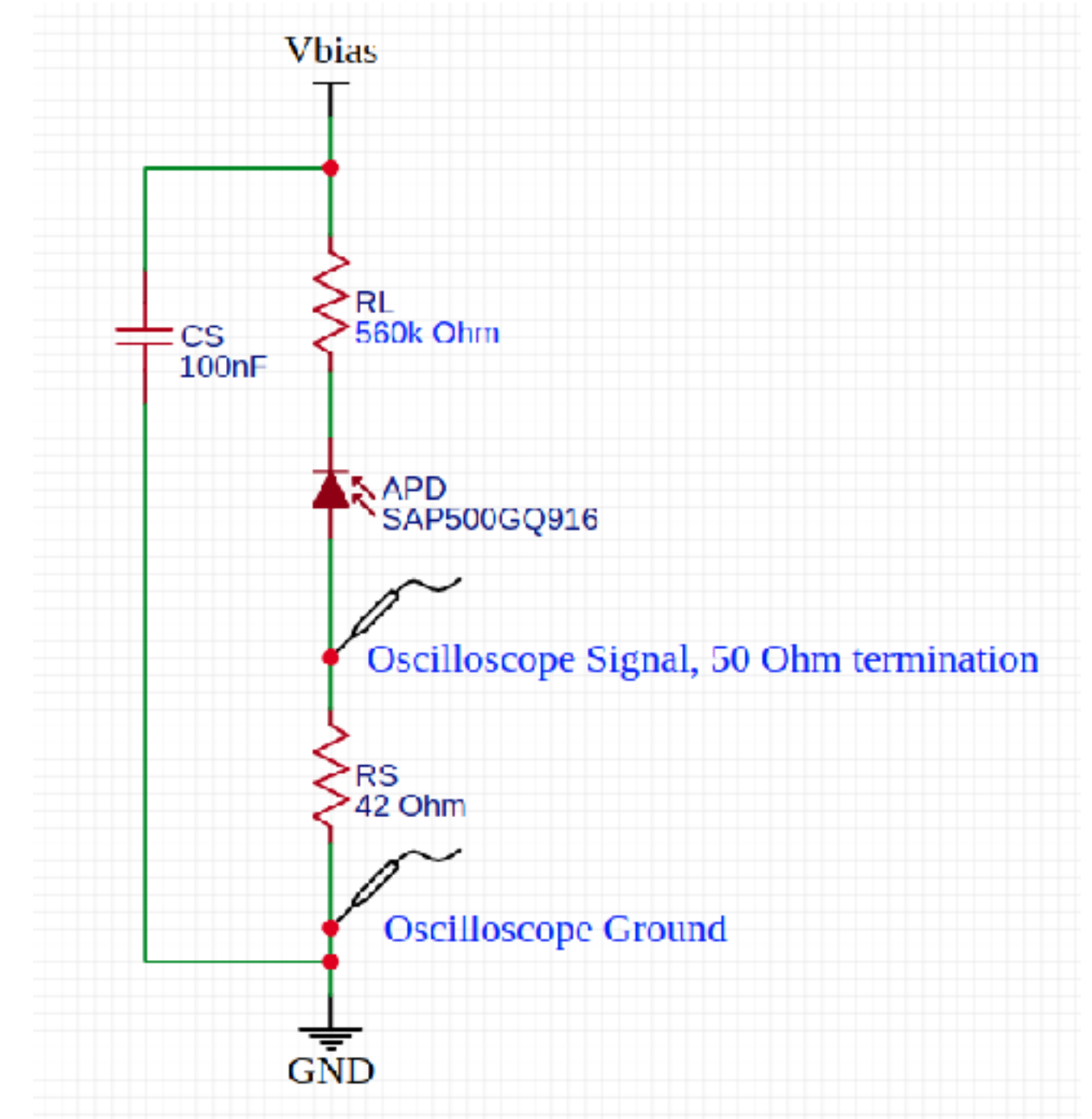
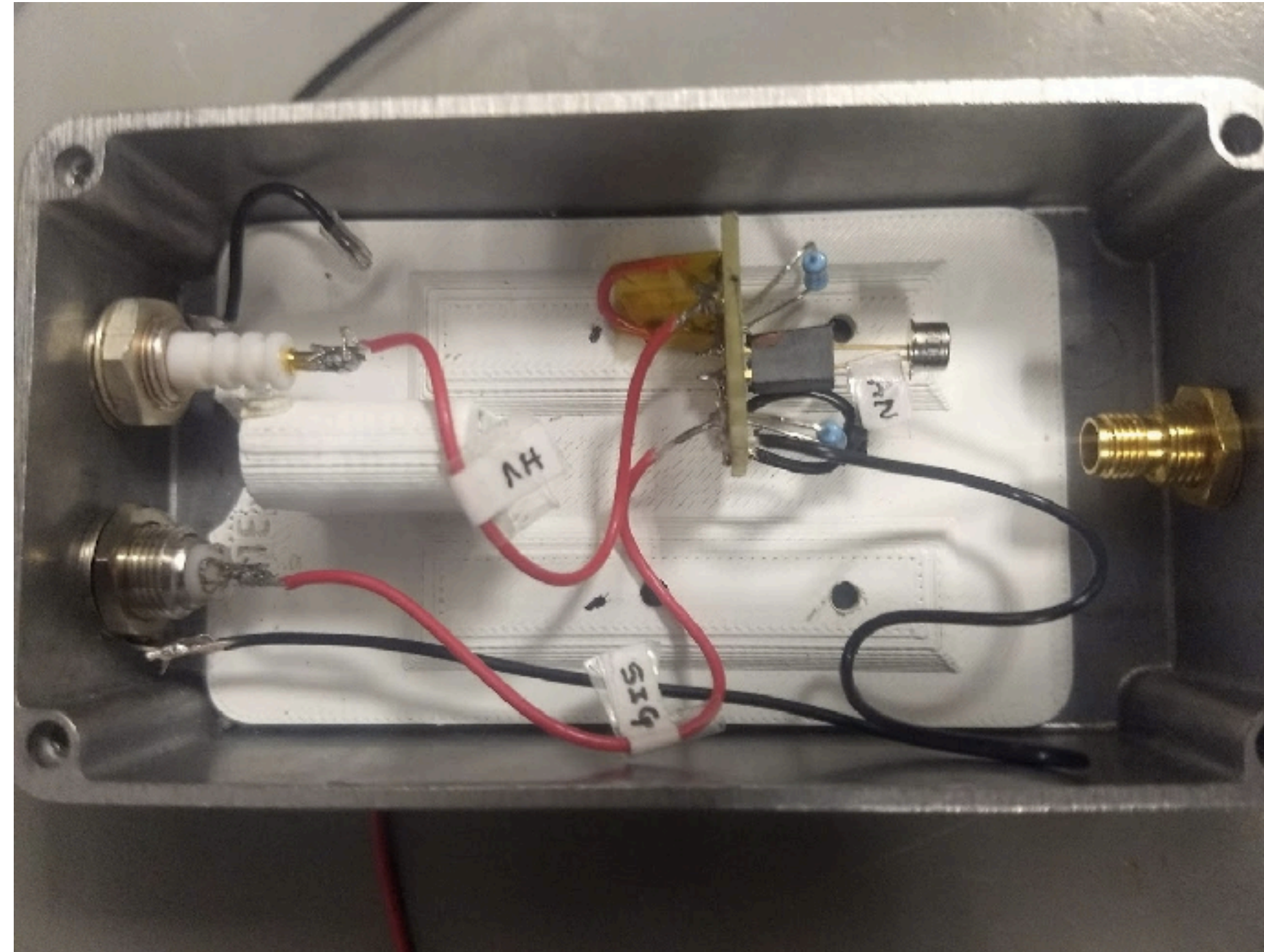






# The photosensor

# SPAD





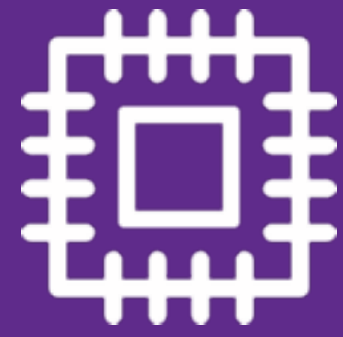
# SPAD cryo test

The photosensor





# Transition Edge Sensor



TES





## Next steps

# Next steps

- Run the full experiment using the SPAD
- Construction and characterisation of the TES
- Final experiment using the TES