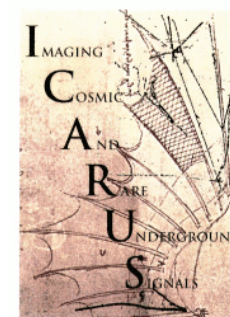


Commissioning, operation, and early result from the light collection system of the ICARUS T600 detector at the Short Baseline Neutrino project

Andrea Scarpelli, on behalf of ICARUS collaboration

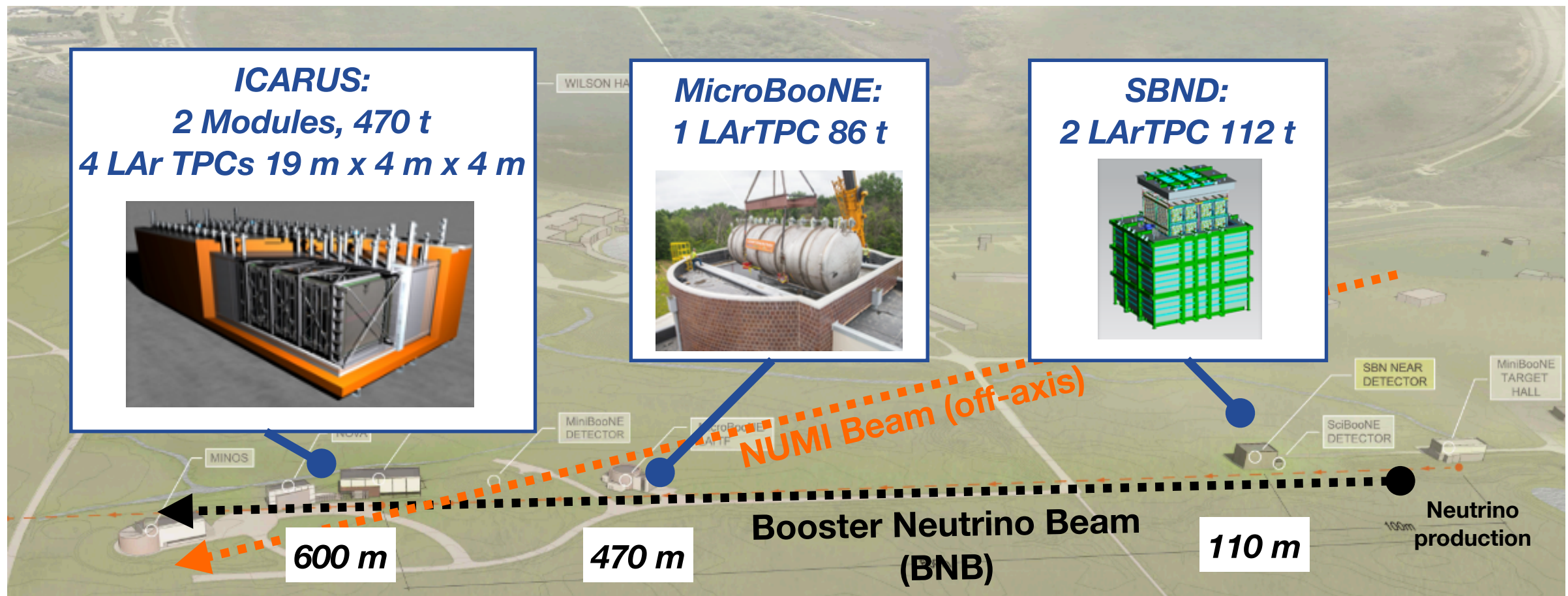
Brookhaven National Laboratory

TIPP2021 26 May 2021



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THE SBN PROJECT @FNAL



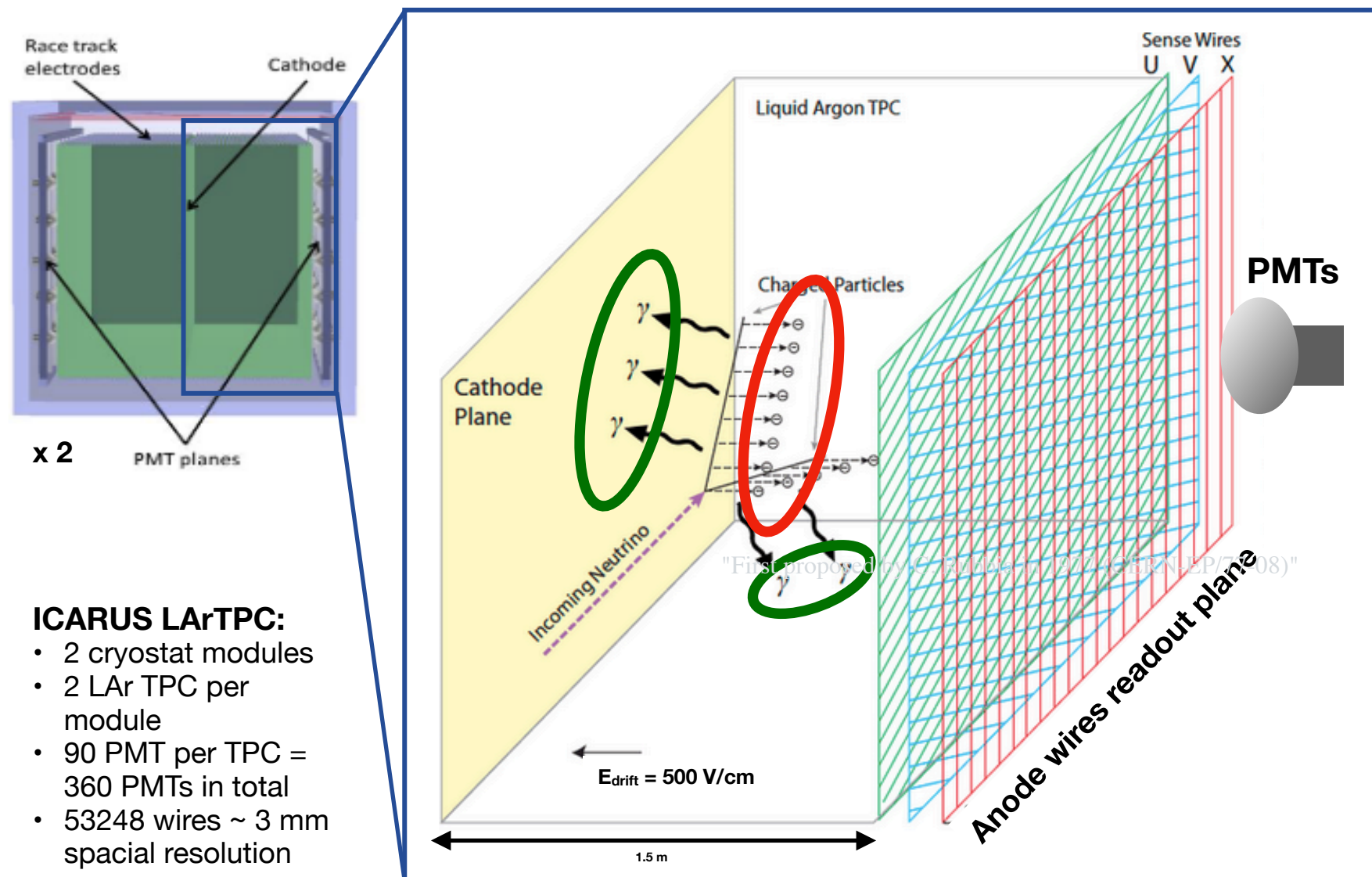
- **Physics goals**

- ▶ Search for anomalies in the Booster Neutrino Beam (BNB) ν_μ disappearance and ν_e appearance
- ▶ ν -Ar cross-section measurements with NUMI off-axis beam, dark matter searches, Neutrino-4 anomaly, ...

- **Detectors: ICARUS + MicroBooNE + SBND along the Booster Neutrino Beamline**

- ▶ All three detectors use the **Liquid Argon Time Projection Chamber technology (LArTPC)**

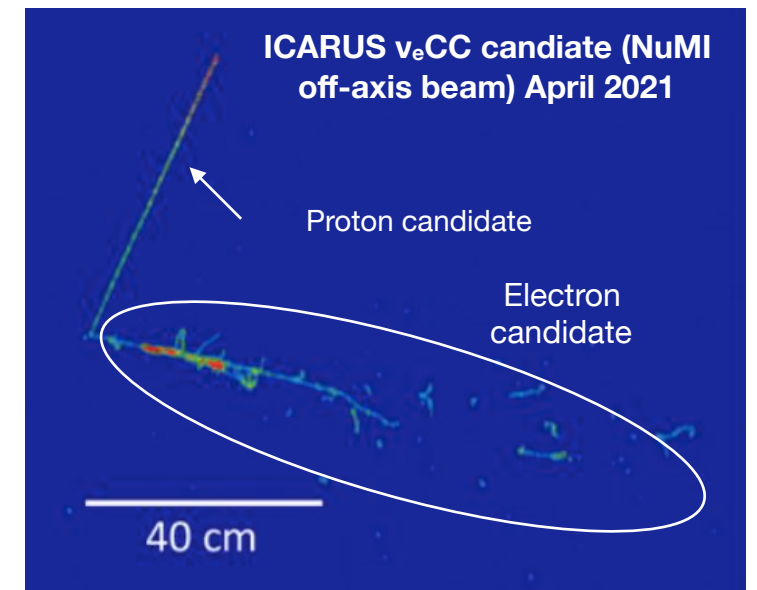
THE LArTPC TECHNOLOGY



ICARUS LArTPC:

- 2 cryostat modules
- 2 LAr TPC per module
- 90 PMT per TPC = 360 PMTs in total
- 53248 wires ~ 3 mm spacial resolution

Electrons from Ar ionization drift (**time ~1m/mm**) to the anode wire readout. Image of the event interaction is created



Scintillation light from de-excitations of Argon eximers

- Singlet state : **fast light** component ($\sim 6 \text{ ns}$)
- Triplet state: **slow light** component ($\sim 1.6 \mu\text{s}$)
- Light yield **$\sim 20\text{k photons /MeV}$** (@500 V/cm drift)
- Wavelength: 125 nm

• SBN Program uses the LArTPC technology for its imaging capability:

- ▶ Easy topological selection of the different flavor of neutrinos and identification of the sub-products of the interactions with particular attention to the separation between electrons and photons
- ▶ Early light signal used for timing the charge deposits, online event selection

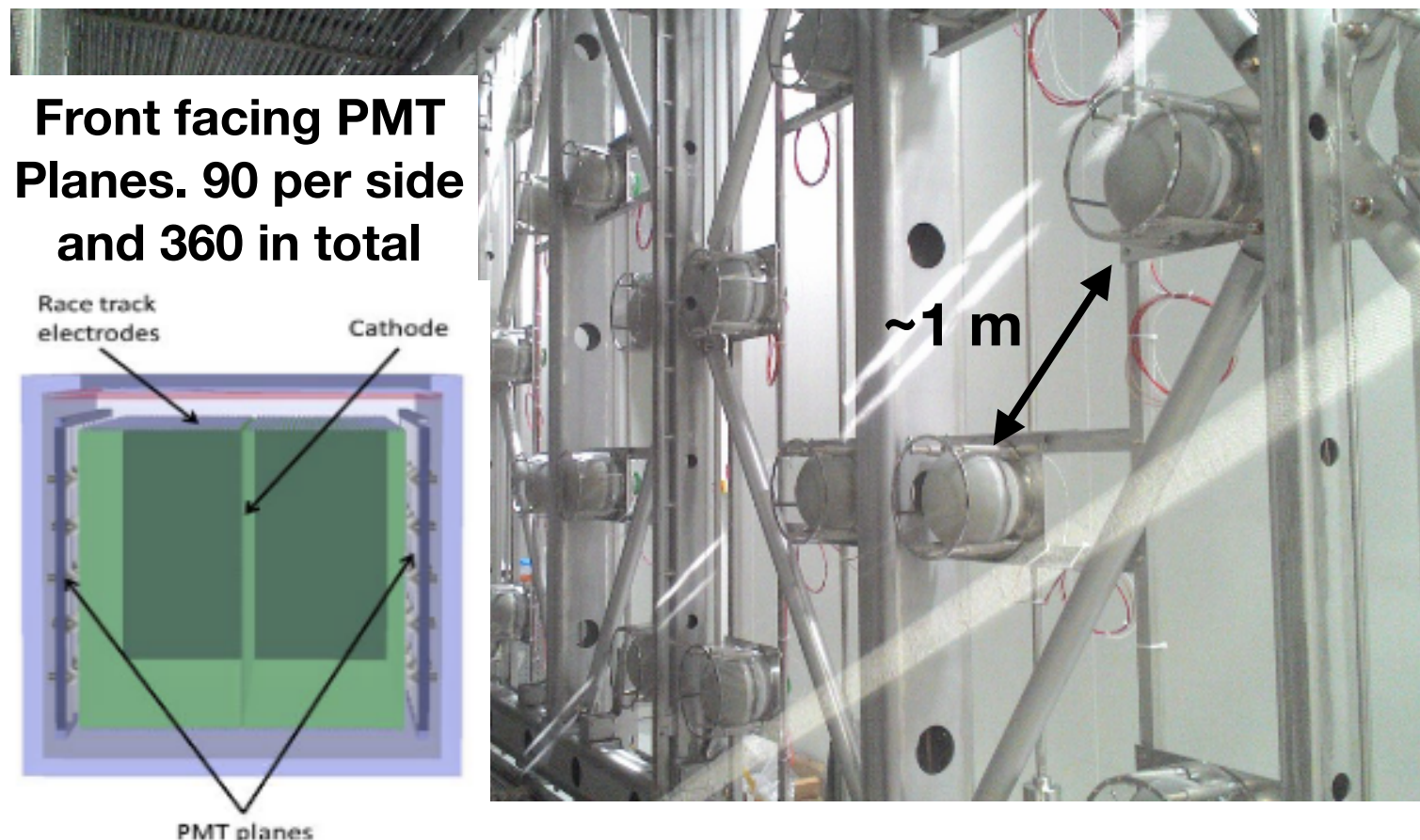
THE PHOTODETECTION SYSTEM

- **Main requirements for the experiment goals**

- ▶ Expected neutrino interaction event threshold ~ 15 Photoelectrons / MeV - Event localization better than 50 cm
- ▶ Excellent timing resolution ($\sim \text{ns}$) to properly identify the neutrino events within the beam arrival window

- **360 PMTs 8" HAMAMATSU R5912-MOD Fully characterized at CERN test-stand**

- ▶ Coated by evaporation of $250 \mu\text{g}/\text{cm}^2$ of Tera-Phenyl-Butadiene (wavelength shifter) *M. Bonesini, et al. JINST 12 P12020 (2018)*
- ▶ Transit Time resolution $\sim 1 \text{ ns}$, Dark rate $< 5 \text{ kHz}$, 12% uniform Quantum Efficiency *B. Ali-Mohammadzadeh et al 2020 JINST 15 T10007*
- ▶ Stable Gain of 10^7 at 87 K to detect down to Single Photoelectrons (SP) light *M. Babicz et al, 2018 JINST 13 P10030*



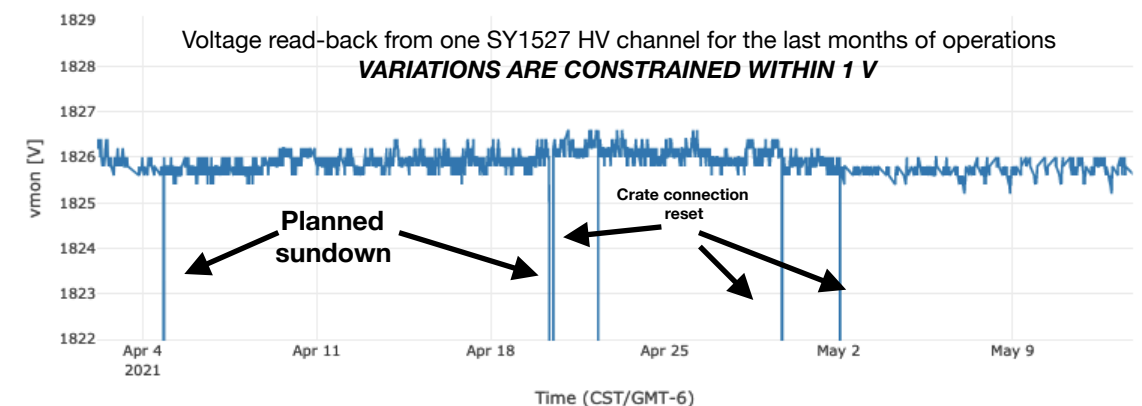
The 8" HAMAMATSU R5912-MOD

POWER SUPPLY AND DATA ACQUISITION



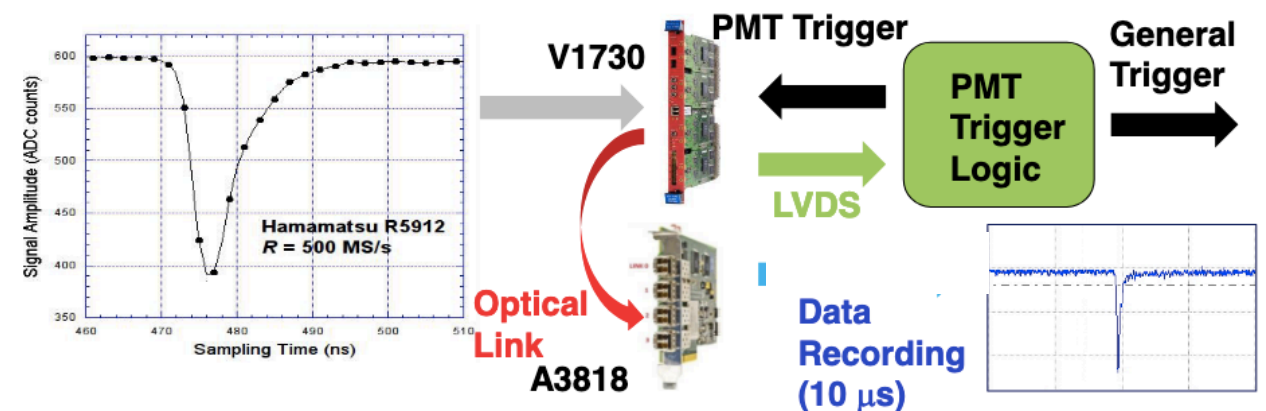
- The Power supply system

- ▶ Primary PS (BERTAN 210-02) 2000 V max (100 mA)
- ▶ 4 High density distributors (CAEN A1932AN): 45 independent channel (**1 V precision**) distributed the voltage to the 360 PMTs
- ▶ The CAEN SY1527LC crate allows control and monitoring. Included in the ICARUS Slow Control system



- The data acquisition system

- ▶ 24 CAENV1730 digitizers x 15 channels each
- ▶ Waveform digitization and read by a Flash ADC, 2Vpp dynamic range, 14-bit resolution and 500 MHz sampling rate (2ns /sample)
- ▶ Discriminated output pulses are presented in 8 programmable outputs (1 every 2 channels) using LVDS (Low Voltage Differential Signaling) standard used for the ICARUS trigger logic



TIMELINE OF COMMISSIONING @FNAL

First lockdown



BNB
Activated

**FY
2018-2019**

**Winter
2020**

**Spring
2020**

**Summer
2020**

**Fall
2020**

**Winter
2021**

**Spring
2021**

Test stand
@CERN and
installation
@ FNAL

Detector
Filling and
cooling
down

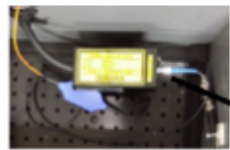
Noise stability

Begin of operation at cold
TPC and PMT High Voltage
commissioning

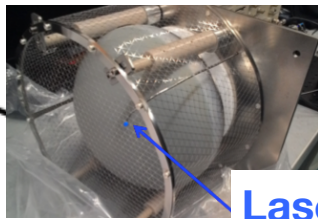
PMT
calibration
at cold

Trigger
commissioning

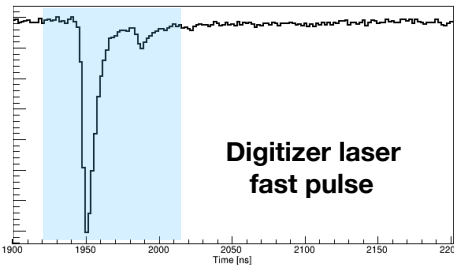
CALIBRATION AT COLD



Laser source:
Hamamatsu
PLP10, $\lambda=405$
nm, FWHM
<100 ps, peak
power ~130
mW.



Laser spot



- ▶ Calibration in situ done using an external fast (~ns) laser pulser flashed on each pmt by dedicated optical fibers
- ▶ Charge integrated in a 100 ns window timed with the laser pulse. Expected 0.005 dark rate pulses per event
- ▶ The PMT Response function is fitted simultaneously for 5 different voltage values to constrain μ (mean number of photoelectron, fixed per PMT). *Laser fluctuations: 5:1000 for > 4 hours*

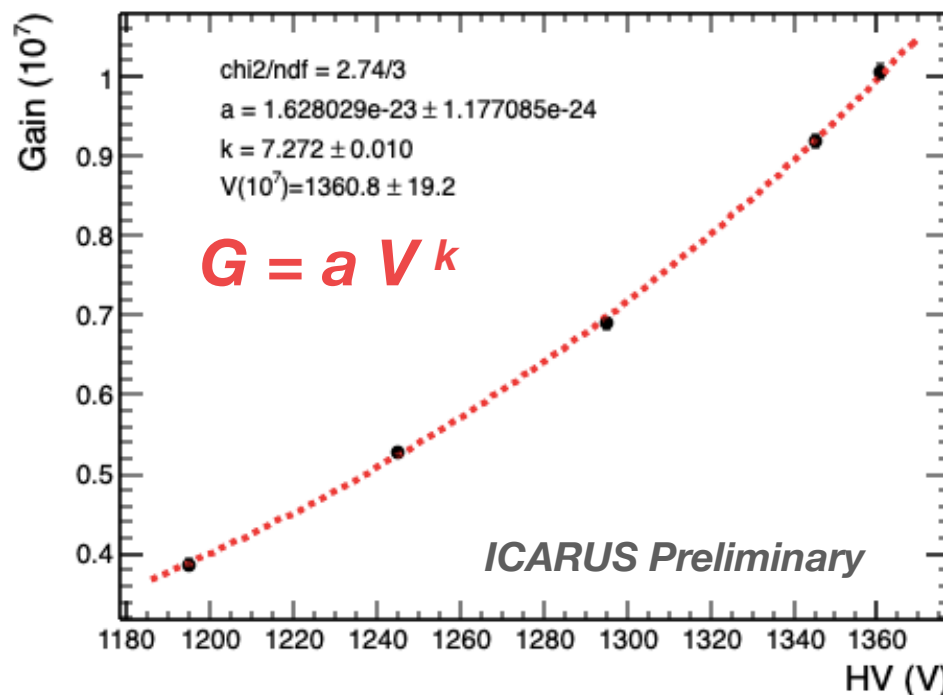
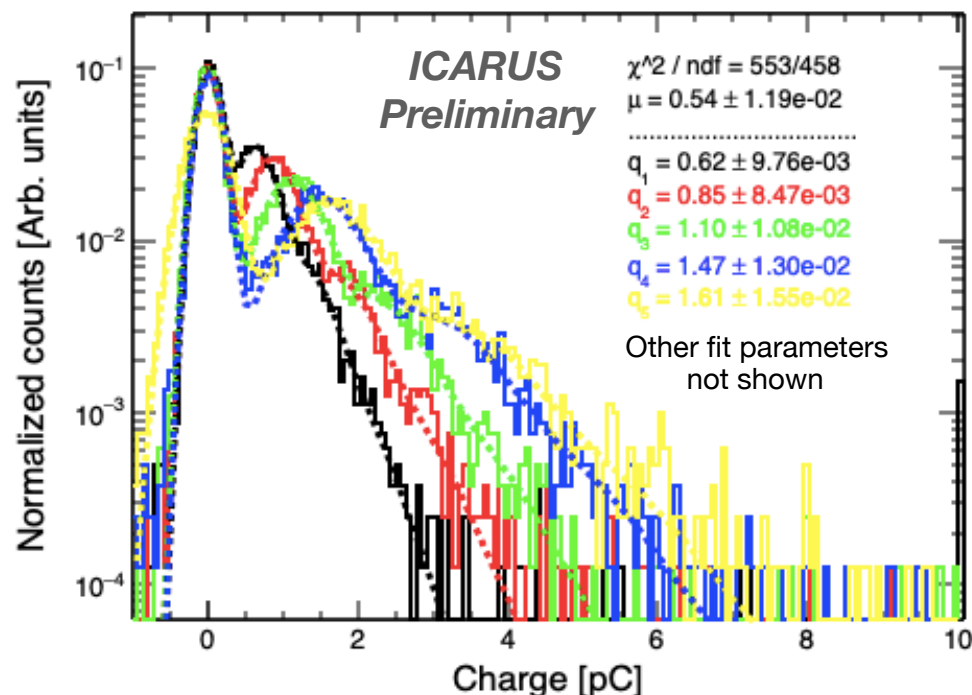
Pedestal modeled using an ExpNormal function:

Due to the low dark rate expected the noise is considered decoupled from the "signal" part

$$S_i(x) = a_0 E_N(x, q_0, \sigma_0, \tau_0) + a_i \sum_{n=1}^{100} \frac{1}{\sqrt{2\pi n \sigma_i}} \frac{\mu^n e^{-\mu}}{n!} \exp\left(\frac{-(x - q_i n)^2}{2n \sigma_i^2}\right)$$

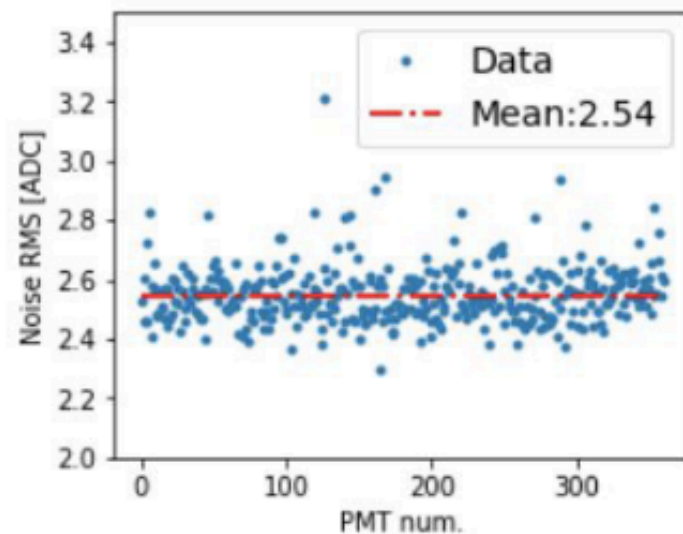
▶ E. H. Bellamy NIM A399 (1994) 468-476

▶ M.V. Diwan 2020 JINST 15 PO2001

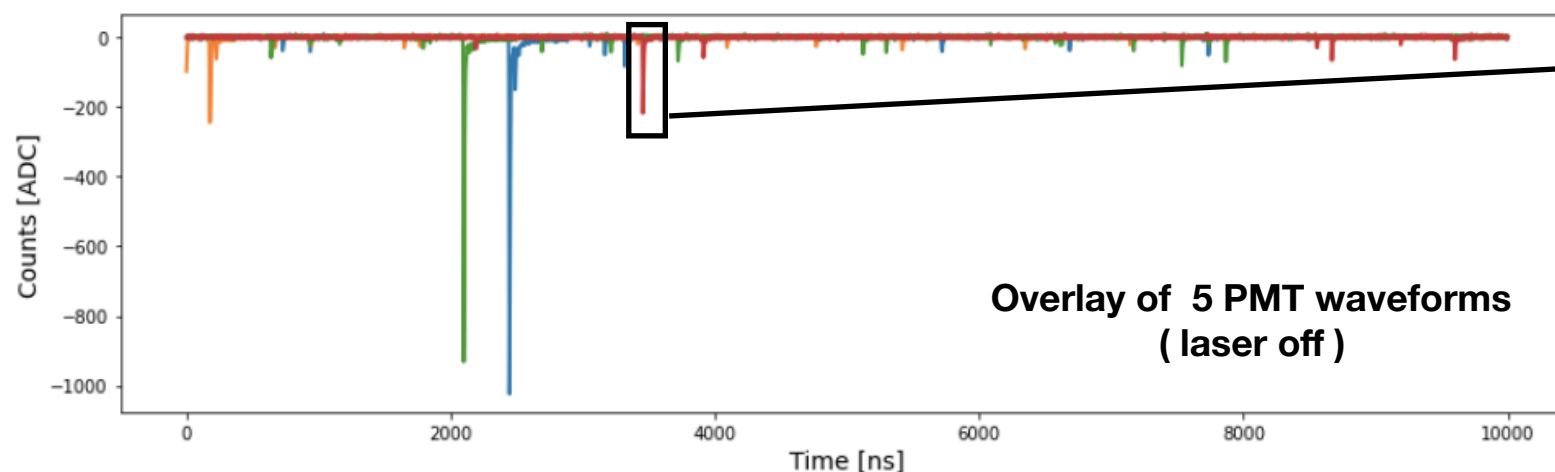


- ▶ Phototubes equalized to the value of $G=10^7$ using the **Power Law** fitted from the 5 gains as function of the applied voltage

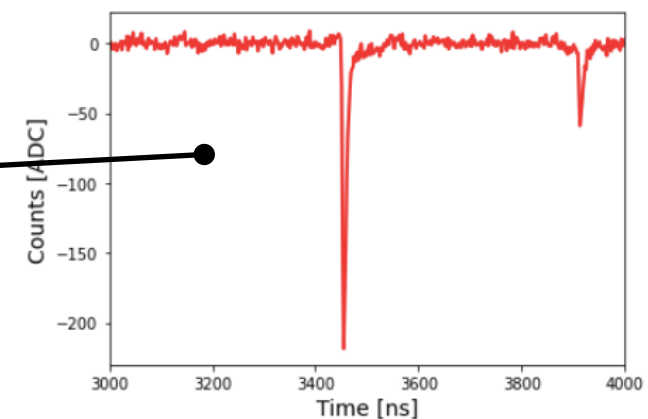
NOISE AND EQUALIZATION OF THE PMTs



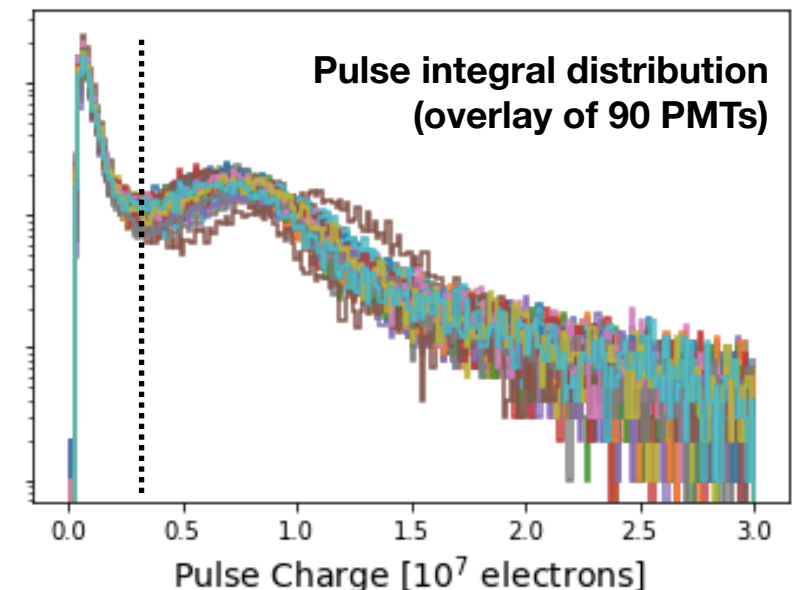
- ▶ Electronic noise has been measured at cold and room temperature. Uniform on all channels. Expected S/N with PMT at $G=10^7$ is **20 for a single photoelectron (SP)**
- ▶ *Isolated* photons activity is present normally in large quantity



Isolated Photons from reflection, dark counts, or other sources under investigation



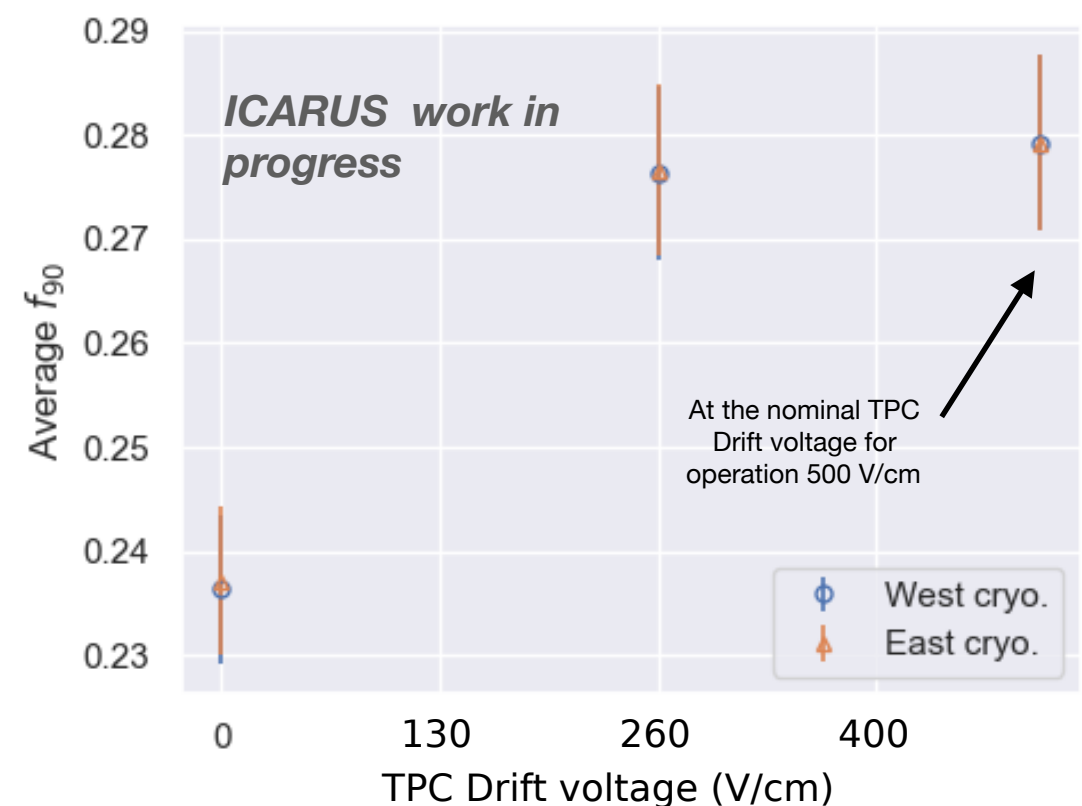
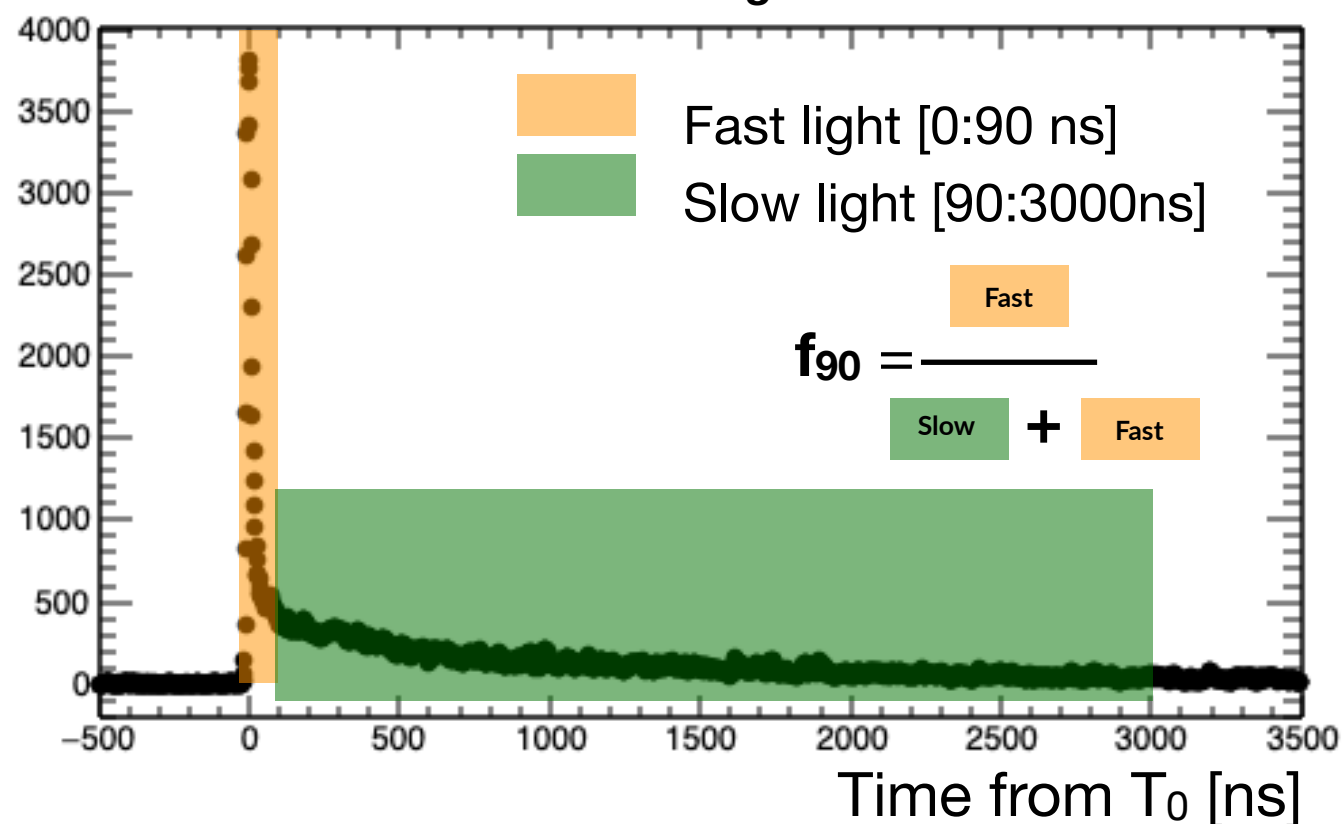
- ▶ Use isolated photoelectron pulses to probe the PMT system status and confirm the relative PMT response equalization
- ▶ May be done without the necessity of dedicated downtime (as happens for the laser)



EARLY STUDIES OF SCINTILLATION LIGHT

- Argon scintillation profile has been observed
 - Both **fast** ($\tau \sim 6 \text{ ns}$) and **slow** ($\tau \sim 1.6 \mu\text{s}$) light components are visible from the two de-excitation times of argon excimers
 - Study of a sample of cosmics at reduced drift voltage collected during commissioning seems to indicate a variation of the light profile similar to what observed by other experiment (B. Aimard *et al* 2021 *JINST* **16** P03007), which is not yet entirely understood. This requires more work and in particular a better control of the level of impurity of Argon.

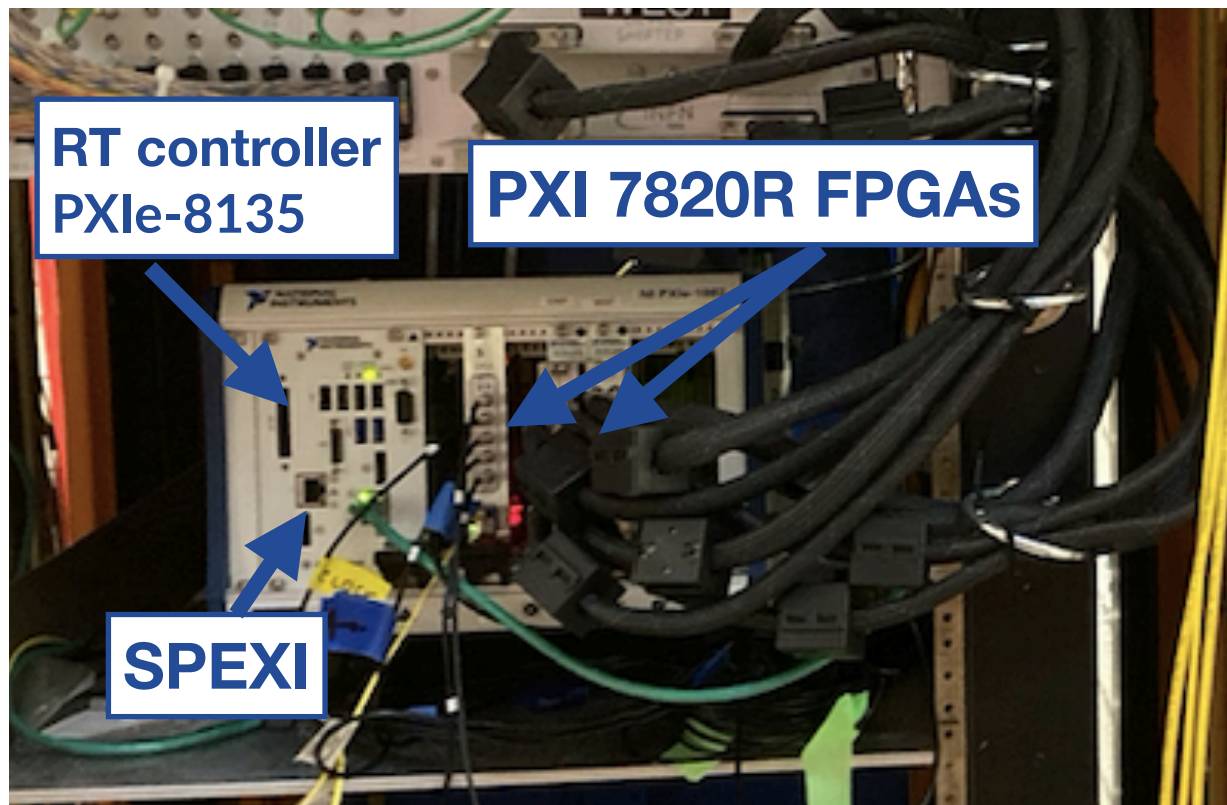
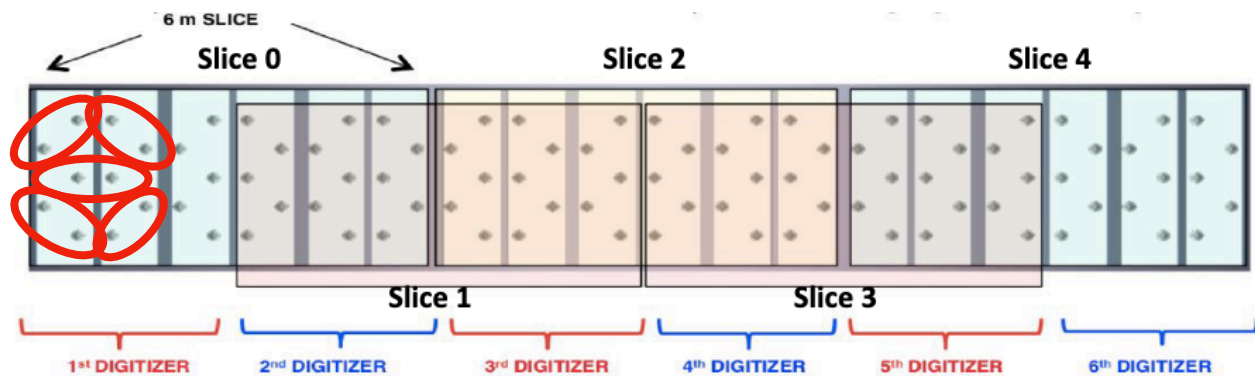
Average scintillation waveform



- Quite limited sample at reduced TPC drift voltage due to the commissioning schedule, studies may be enriched during the detector summer shutdown

THE TRIGGER SYSTEM

- The prompt scintillation light from physical events is used for the trigger system of ICARUS
 - ▶ Beam extraction signals at 5 Hz (BNB) and 0.7 Hz (NUMI), but **total physical events rate ~0.27 Hz**
 - ▶ **0.17 Hz** are in-time cosmic rays and **0.08 Hz** from beam neutrino

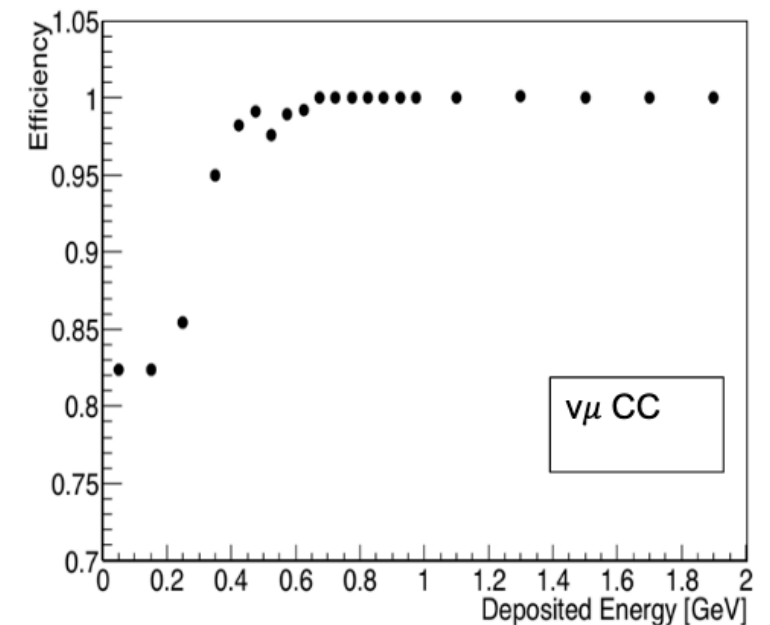
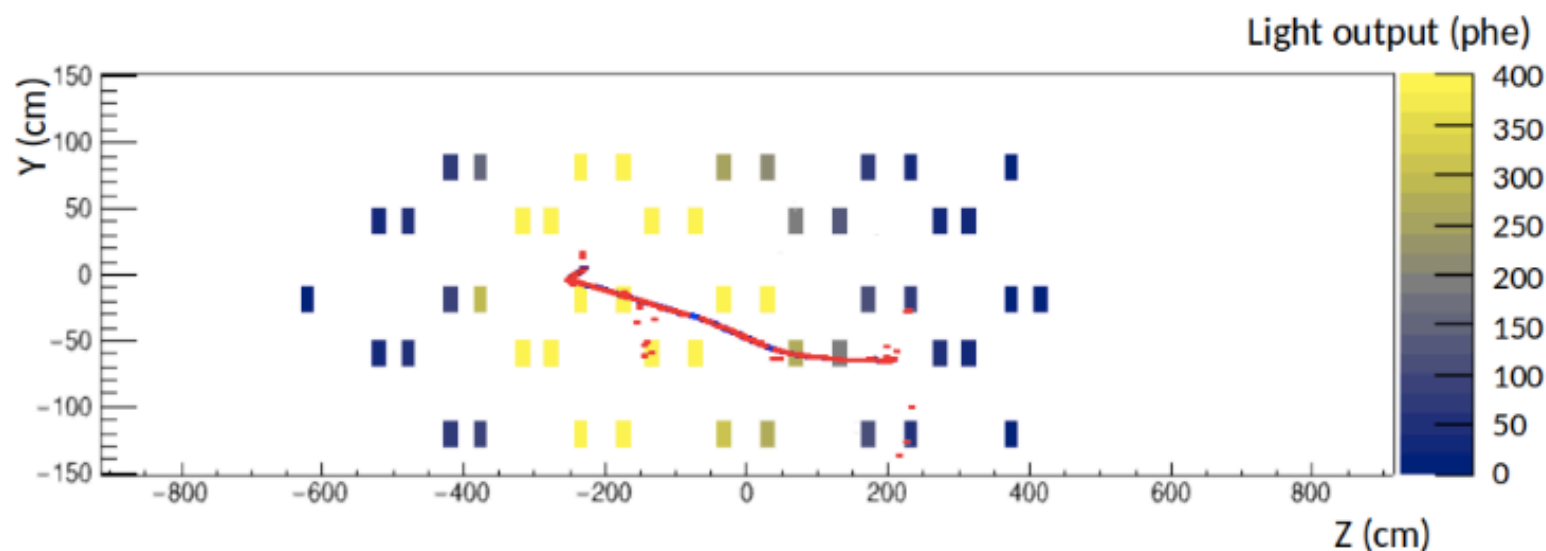


- The trigger is designed to maximize the capture of localized light activity in coincidence with the beam spill window time
 - ▶ Trigger logic based on combinations of PMT **LVDS pairs** over a certain threshold
 - ▶ The trigger system use the NI-PXIe instrumentation with FPGAs to customize different settings and logics
 - ▶ Precision information of the beam extraction timing are transmitted to the system through a *White Rabbit network* with sub-ns accuracy synchronization

TRIGGER: MONTECARLO STUDIES

- Efficiency studies implementing different combination of threshold discriminations and local PMT trigger logics are carried out with a detailed Monte Carlo simulation

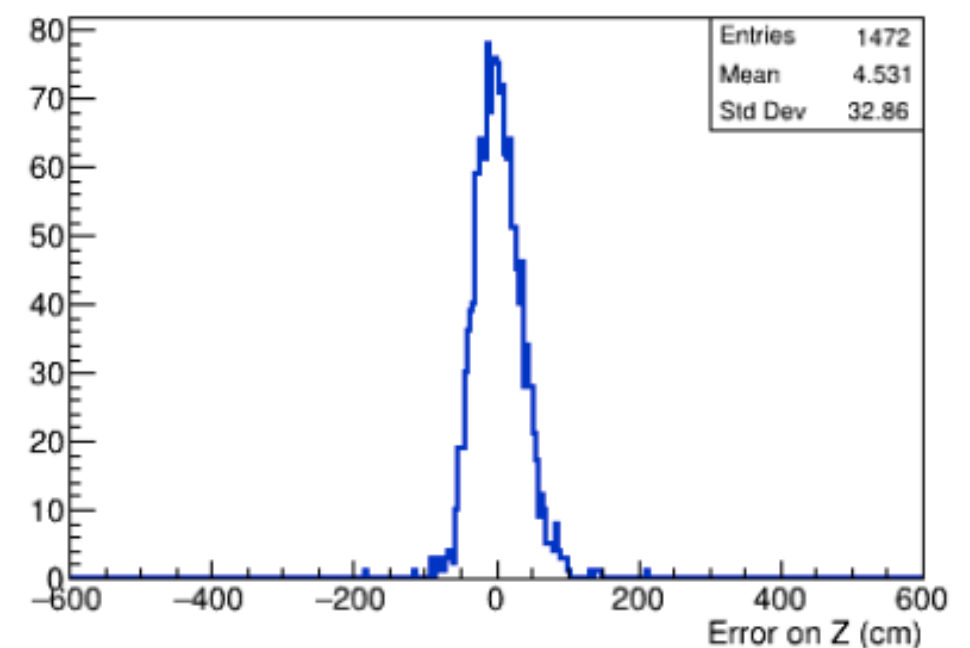
- ▶ Studies to be refined with data-driven inputs



- Optimization of the light reconstruction and identification of the vertex of the neutrino candidates is also in progress

- ▶ Preliminary simulation studies suggest vertex resolution ~ 30 cm to be confirmed also with data

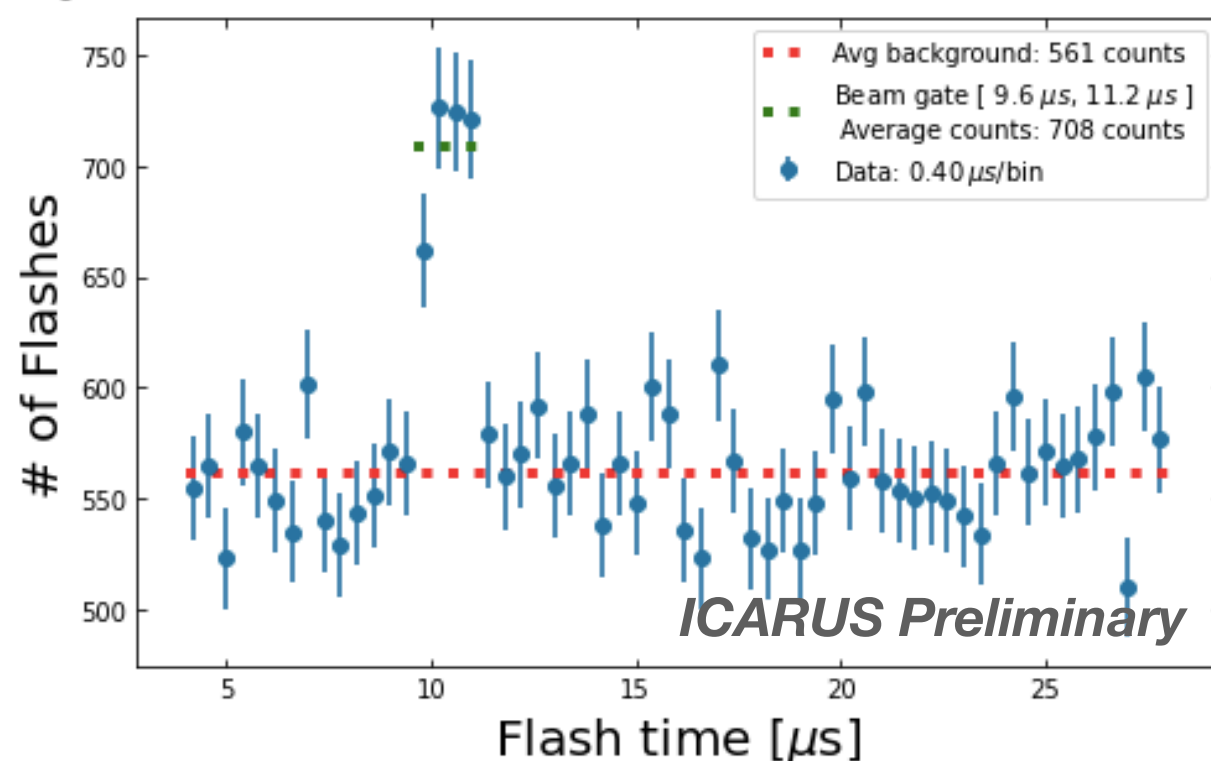
- ▶ B. Ali-Mohammadzadeh et al JINST 15 (2020) 10, T10007



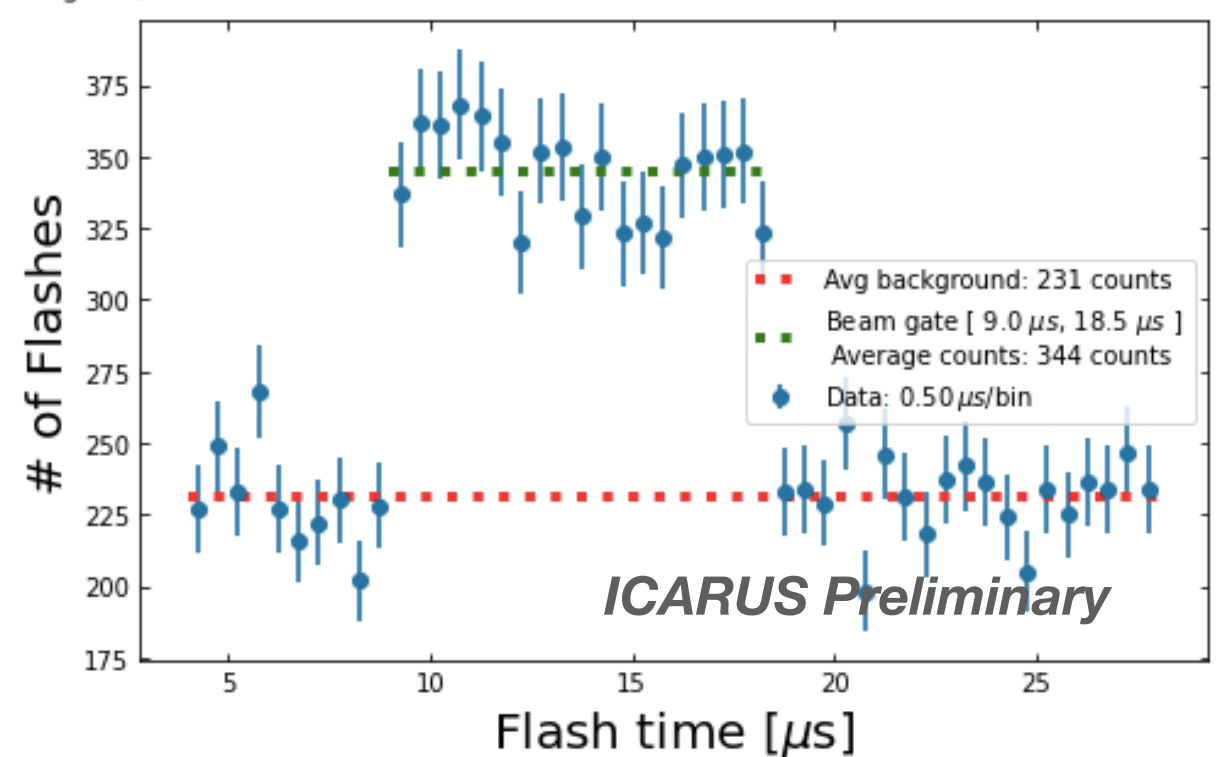
BEAM SPILLS IDENTIFICATION

- Example of a trigger commissioning exercise: identify timed light activity in coincidence with the beam extraction signals
 - ▶ Offline selection: Flashes are defined as the 100 ns coincidence between pulses on 3 PMTs with more than 10 PE. Only events with flashes with 10 PMTs for two front-facing walls are kept

BNB SPILL WINDOW 1.6 μ s



NuMI SPILL WINDOW 9.5 μ s



- Background represented by cosmic activity. The excess of flashes is timed and compatible with the presence of neutrinos from the beam spill window

SUMMARY AND OUTLOOK

- The photodetector system is a fundamental part of the LArTPC technology allowing the precise timing and interpretation of a TPC event
- The ICARUS photodetection system is fully commissioned and in good shape:
 - ▶ The characterization of the PMTs in a test and in-situ has terminated
 - ▶ The calibration of the system is completed and stability proven
 - ▶ Supports the Trigger commissioning activities
- Outlook
 - ▶ The final timing resolution of the system is targeting a *ns* resolution. This aspect is currently under study
 - ▶ Analysis of the scintillation light from both cosmic and neutrino interaction is in progress to provide data-driven inputs to the detailed Monte Carlo simulation and trigger efficiency studies
 - ▶ ***ICARUS is planning to start physics data collection as early as September!***

Thank you!