

Concept of KamLAND2 DAQ system

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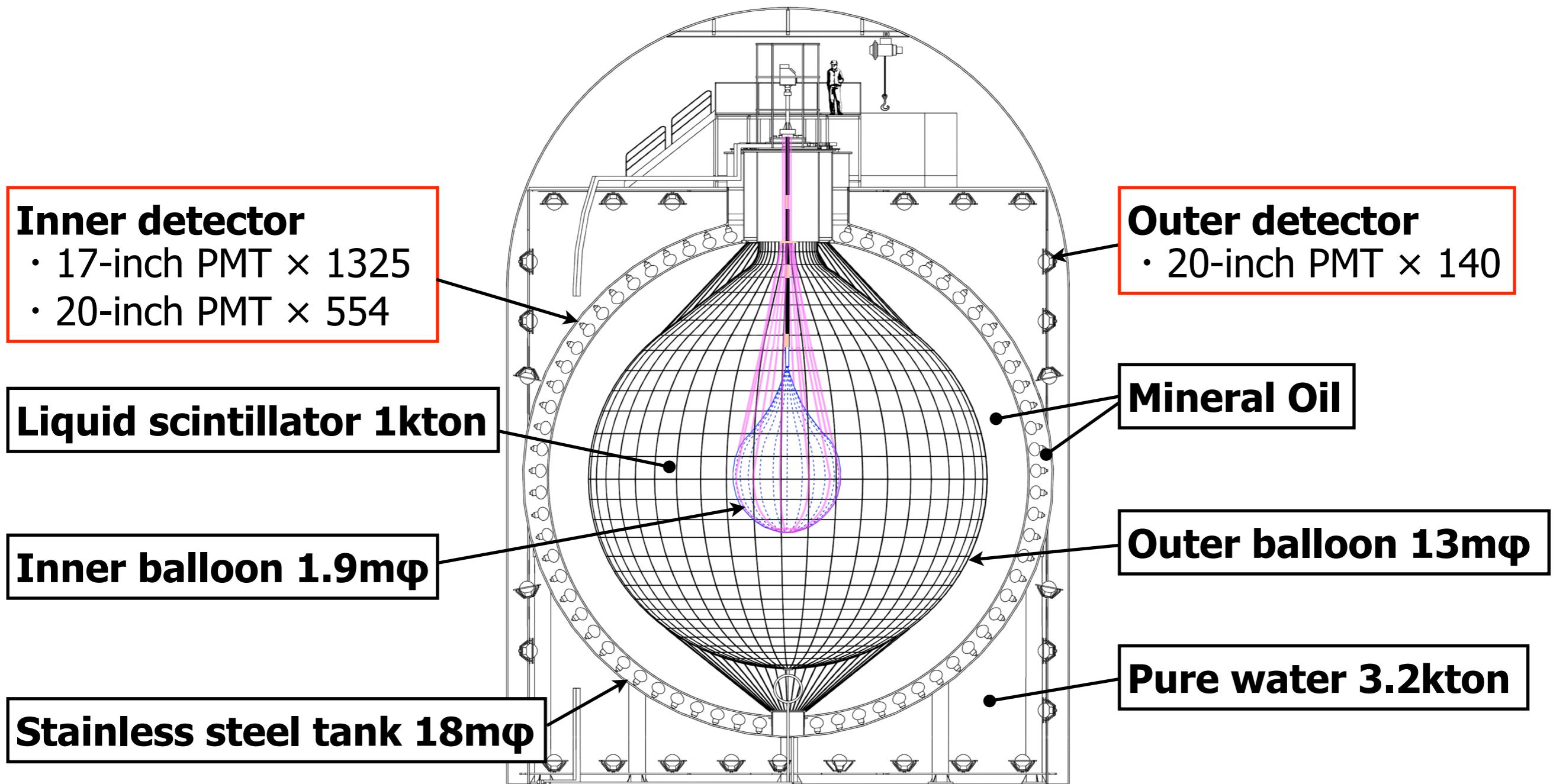
K. Ishidoshiro, T. Nakahata, K. Nakamura

May 25, 2021 TIPP2021 (online)

KamLAND

Kamioka Liquid scintillator Anti-Neutrino Detector

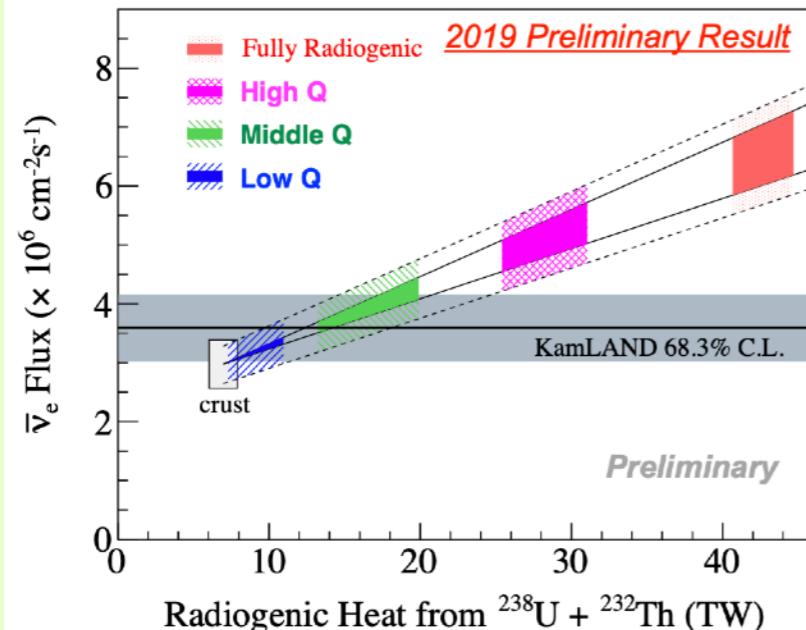
- KamLAND Zen 800 phase is running since Jan. 2019.



Neutrino physics @ KamLAND

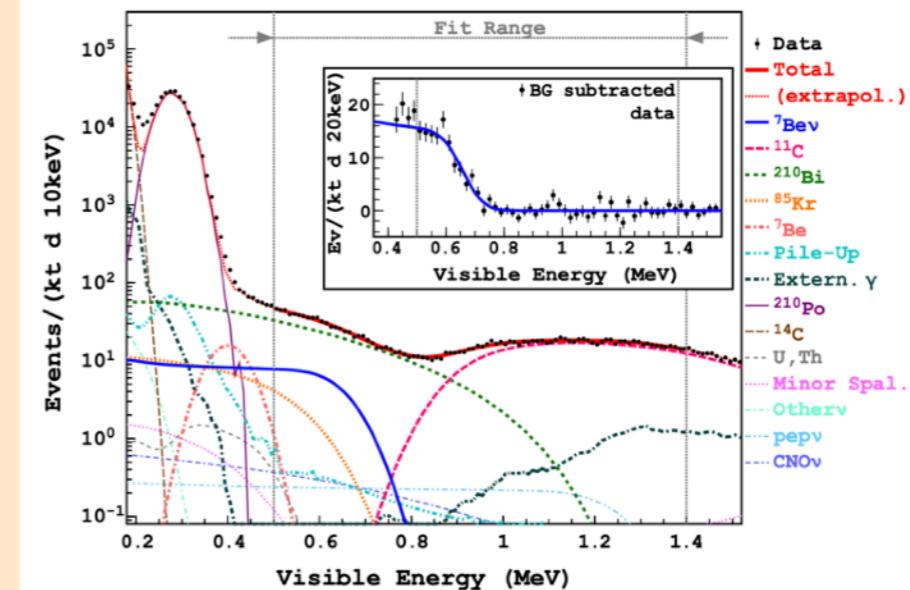
Geo neutrino

(Neutrino Geoscience 2019, H. Watanabe)



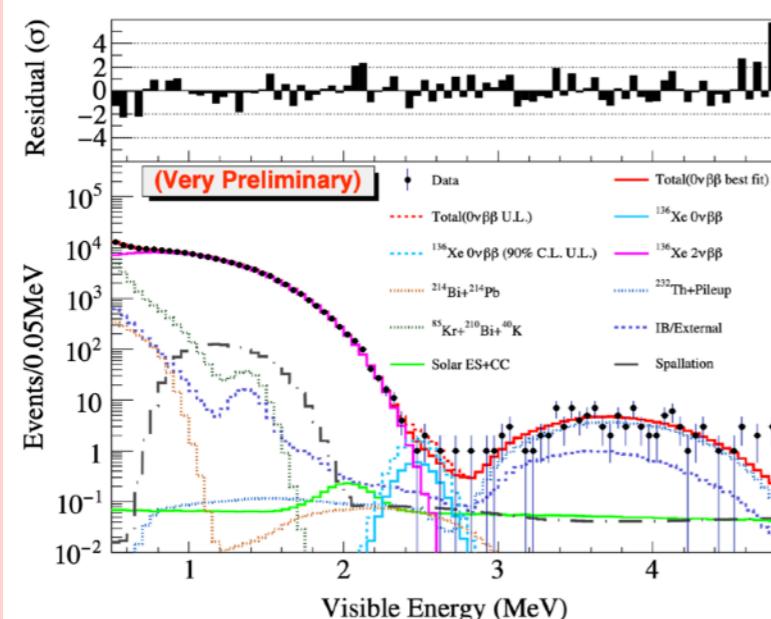
${}^7\text{Be}$ solar neutrino

(PR C 92, 055808 (2015))



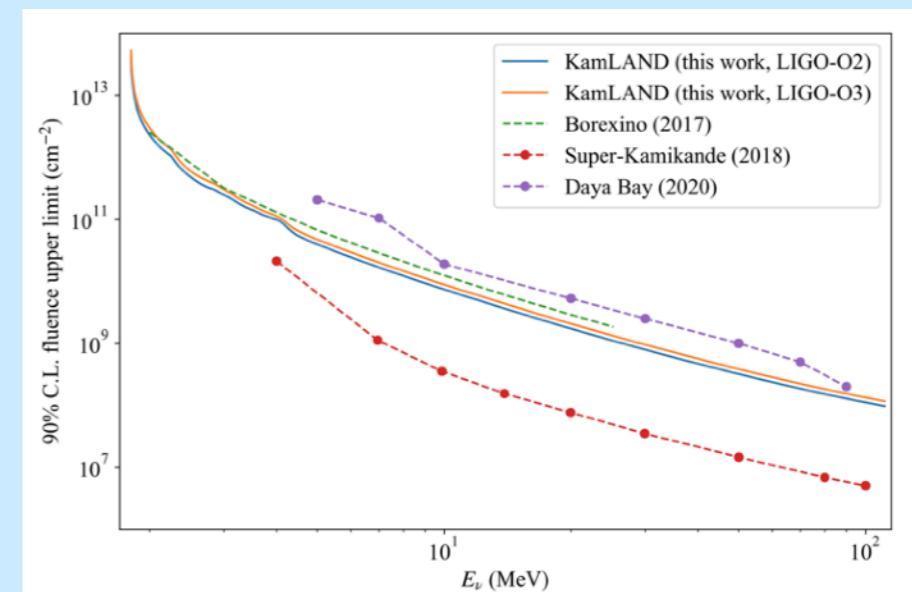
Neutrino-less double β decay

(TAUP2019, Y. Gando)



Gravitational Wave

(Astro. Phys. Journal, 909;116 (2021))



Neutrino physics @ KamLAND

SNEWS: SuperNova Early Warning System

International collaboration of experiments

- Super-K (Japan)
- LVD (Italy)
- Ice Cube (South Pole)
- KamLAND (Japan)
- Borexino (Italy)
- Daya Bay (China)
- HALO (Canada)

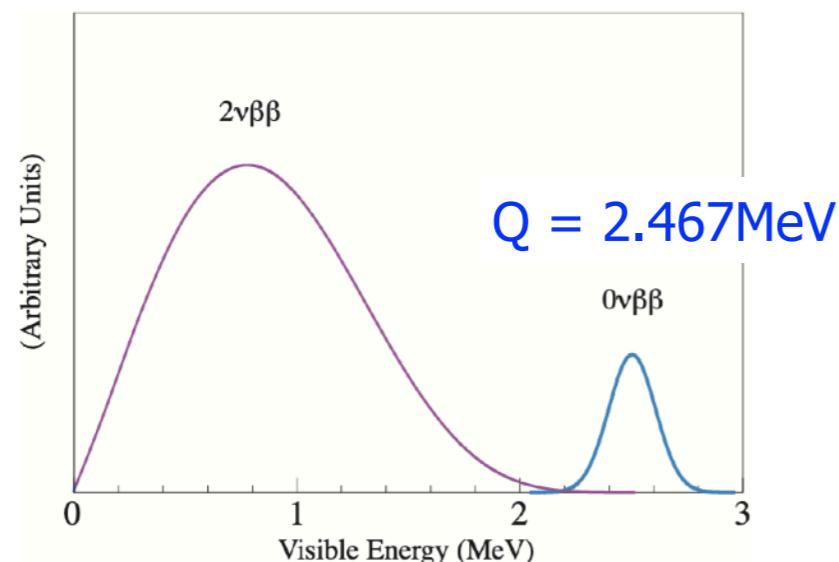


KamLAND → KamLAND2

Major backgrounds for $0\nu 2\beta$ decay

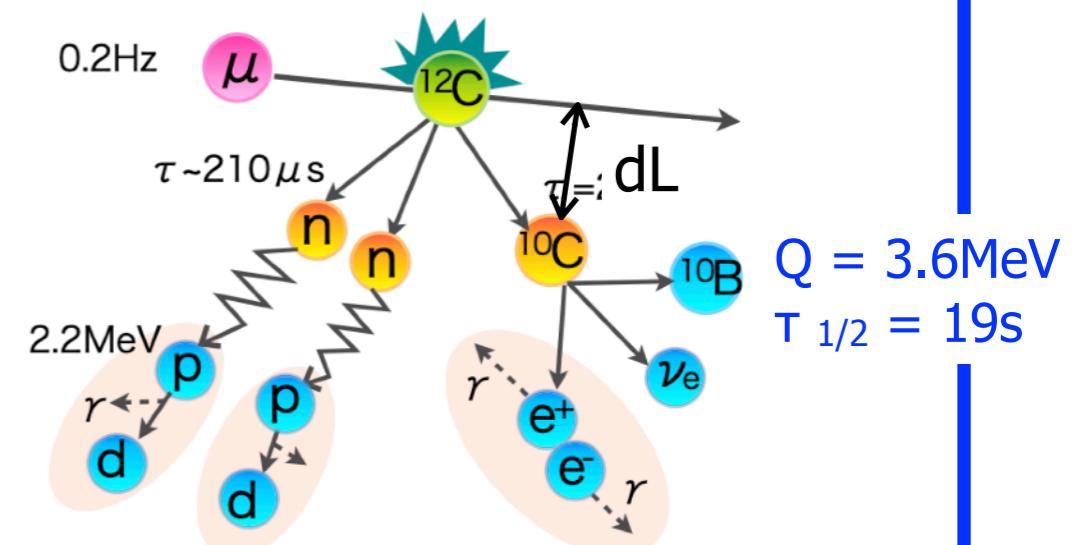
- $2\nu 2\beta$ decay
- Cosmic μ spallation backgrounds

2v2 β decay



To improve energy resolution is the only one solution!

Cosmic μ spallation backgrounds

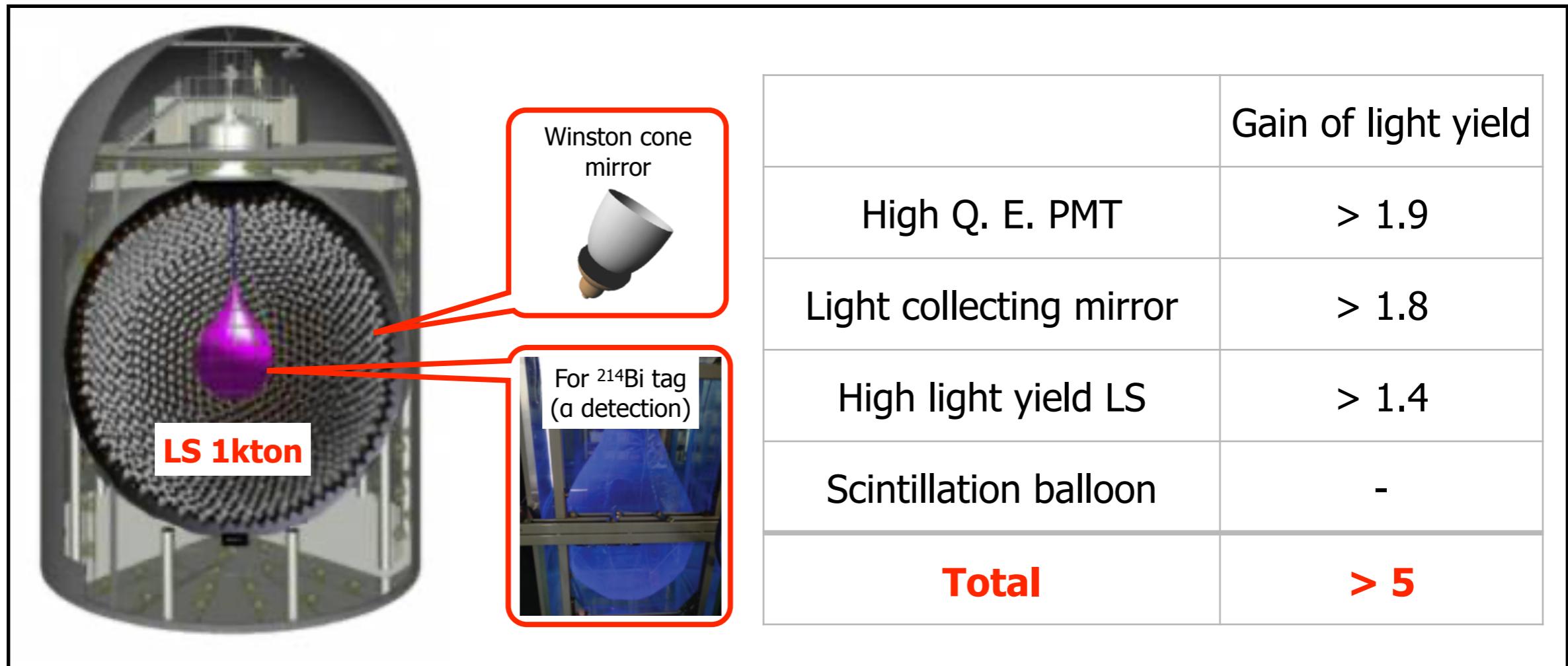


- Likelihood (dE/dX , dL and dT)
- 3-fold tagging (μ , ^{10}C and **neutron**).

KamLAND → KamLAND2

Major backgrounds for $0\nu 2\beta$ decay

- $2\nu 2\beta$ decay
- Cosmic μ spallation backgrounds



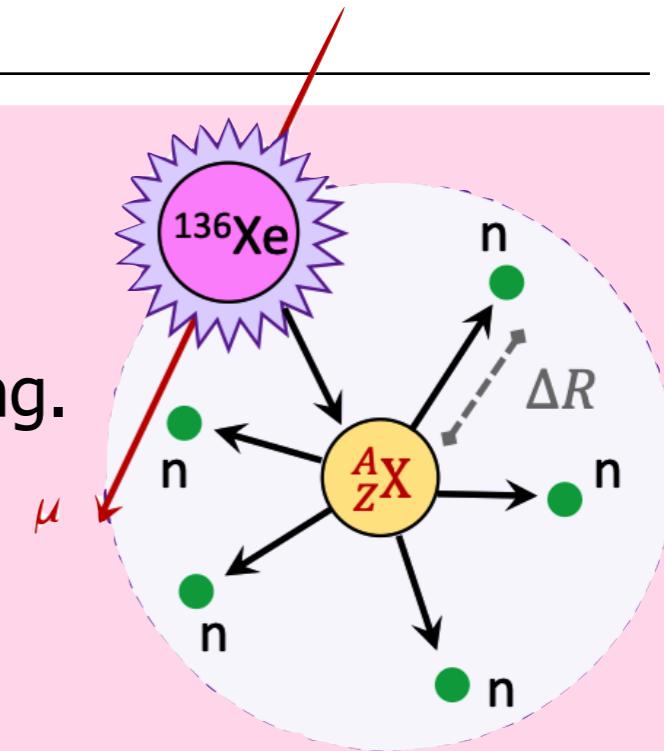
- Increase light yield & improve energy resolution!!
- Target $\langle m_{\beta\beta} \rangle$ sensitivity ~ 20 meV / 5 years

Challenges for electronics

- **Cosmic μ spallation backgrounds**

- Spallation on ^{136}Xe induce serious backgrounds.
- Number of surrounding neutrons is a key for tagging.

"High neutron detection efficiency"



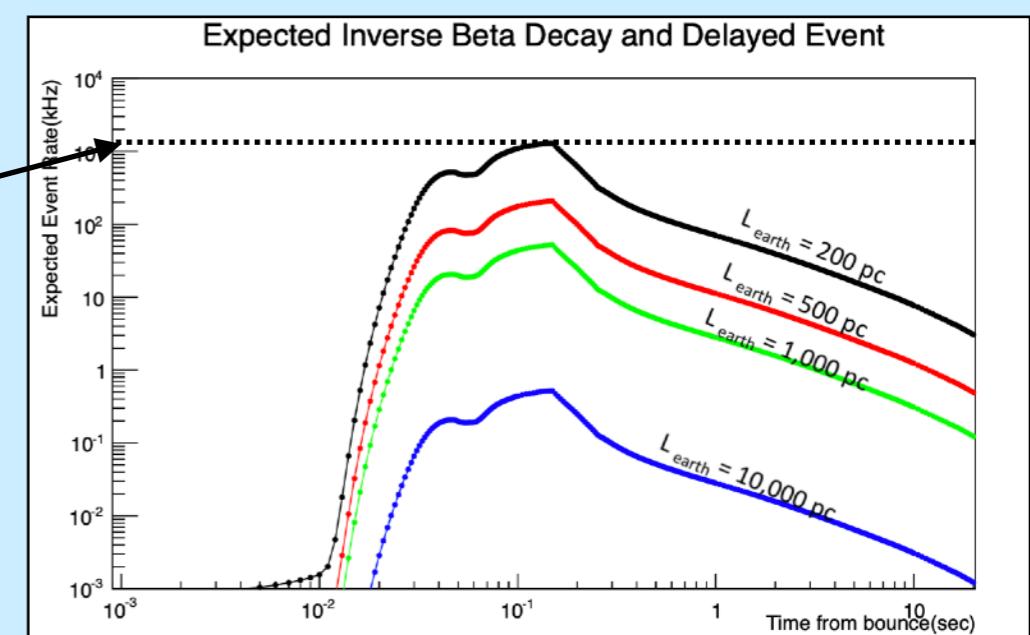
- **Supernova neutrinos**

- Time profile of the number of neutrinos is a probe of supernova explosion mechanism.

"High event rate tolerance"

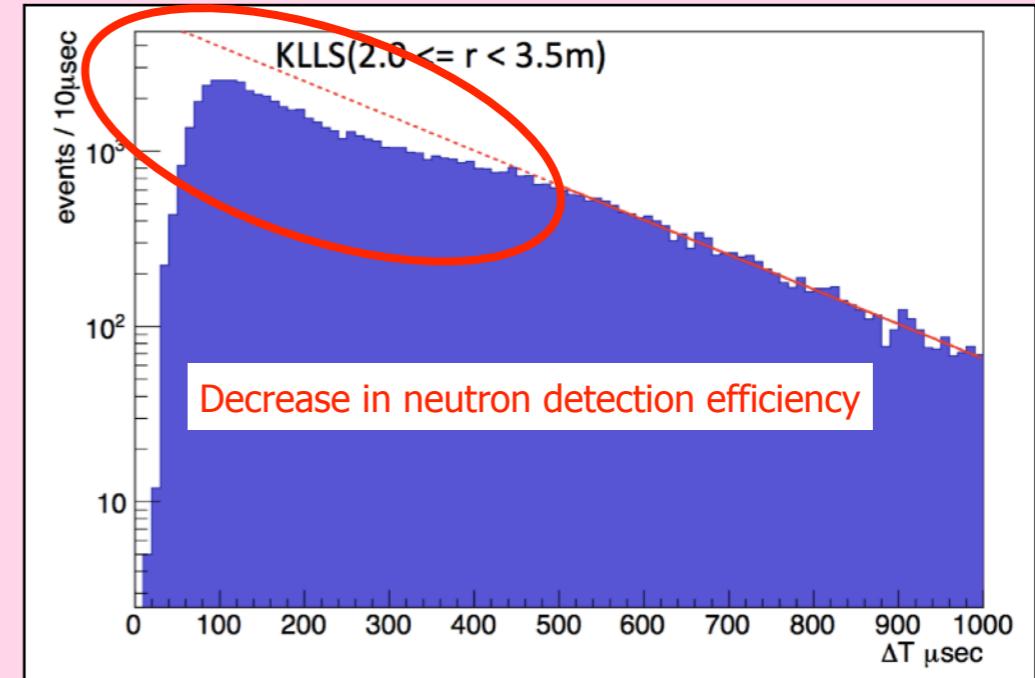
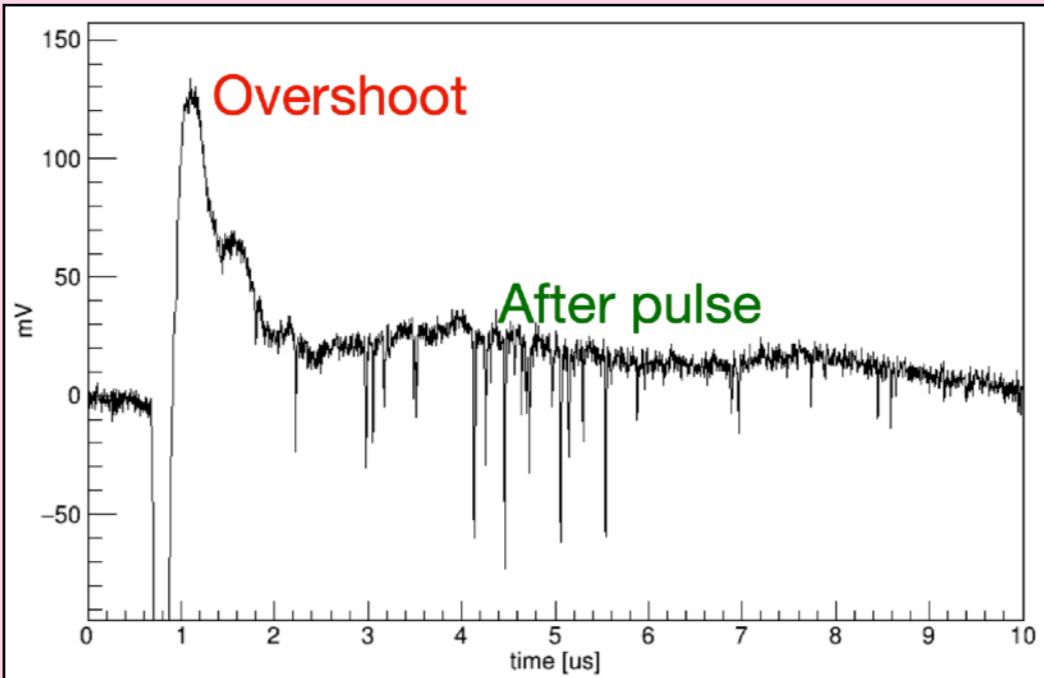
Betelgeuse (200pc)

Maximum instantaneous event rate $\sim 1\text{MHz}$



Current limitation

- **Cosmic μ spallation backgrounds**
 - Trigger decision become difficult under overshoot.



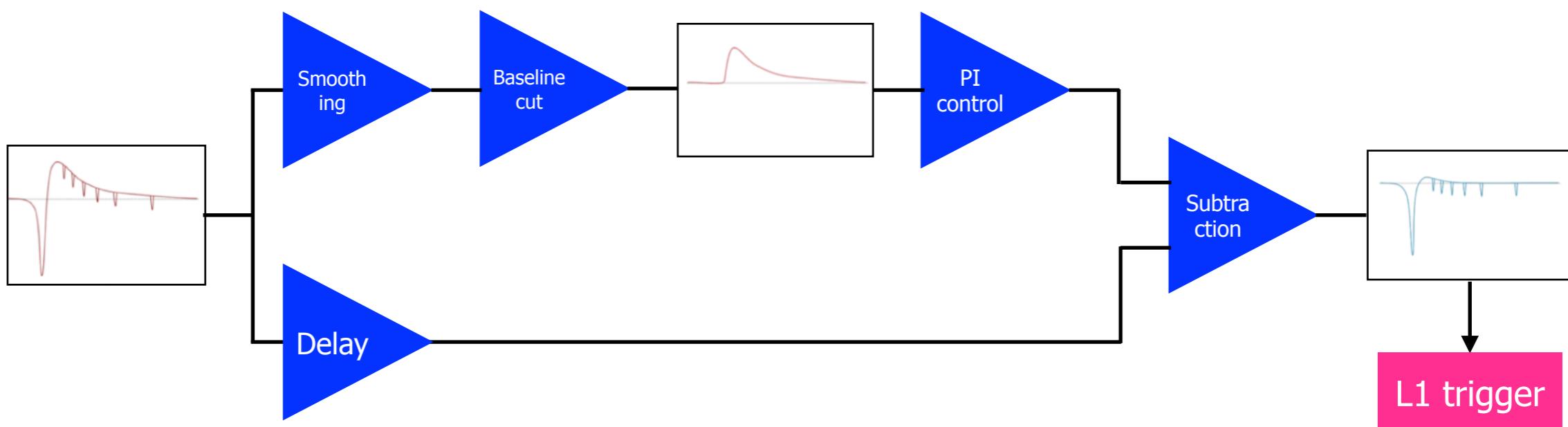
- **Supernova neutrinos**
 - Buffer will be full at around 300 μ s with event rate of 1MHz.

How to overcome

- **Cosmic μ spallation backgrounds**

- Trigger decision become difficult under overshoot.

"Digital BaseLine Restorer (BLR)"



- **Supernova neutrinos**

- Buffer will be full at around 300 μ s with event rate of 1MHz.

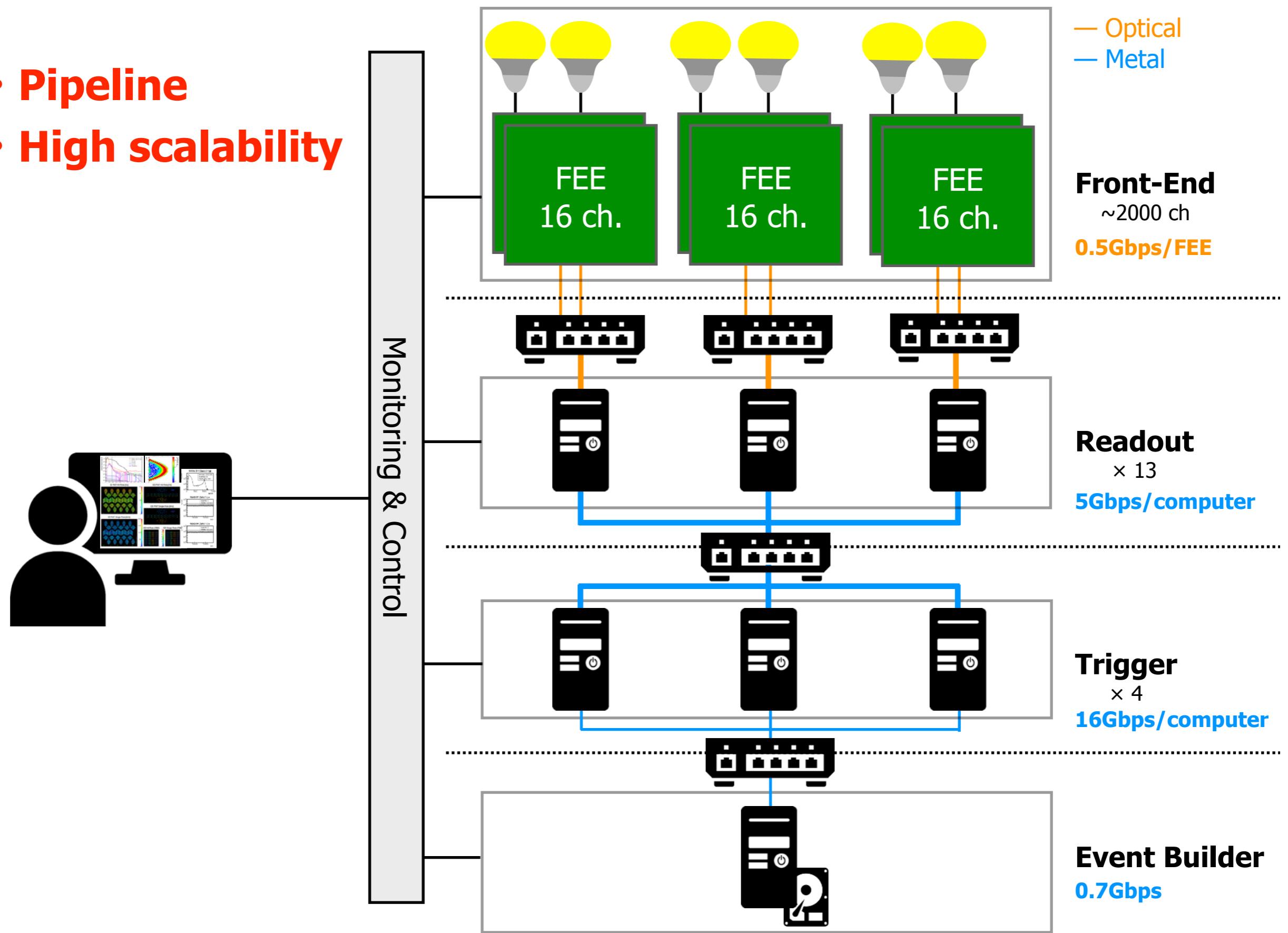
"RFSoC" transfer speed 76.8 Gbps

"DDR4" on-board buffer 4 GB



DAQ overview

- Pipeline
- High scalability



Event rate & Data rate

	Event rate	Data rate
Dark noise	20kHz	32 Mbps/PMT
Cosmic μ	0.3Hz	48 kbps/PMT
Supernovae	1MHz ($\sim 1s$)	1.6 Gbps/PMT

- **Dark noise**

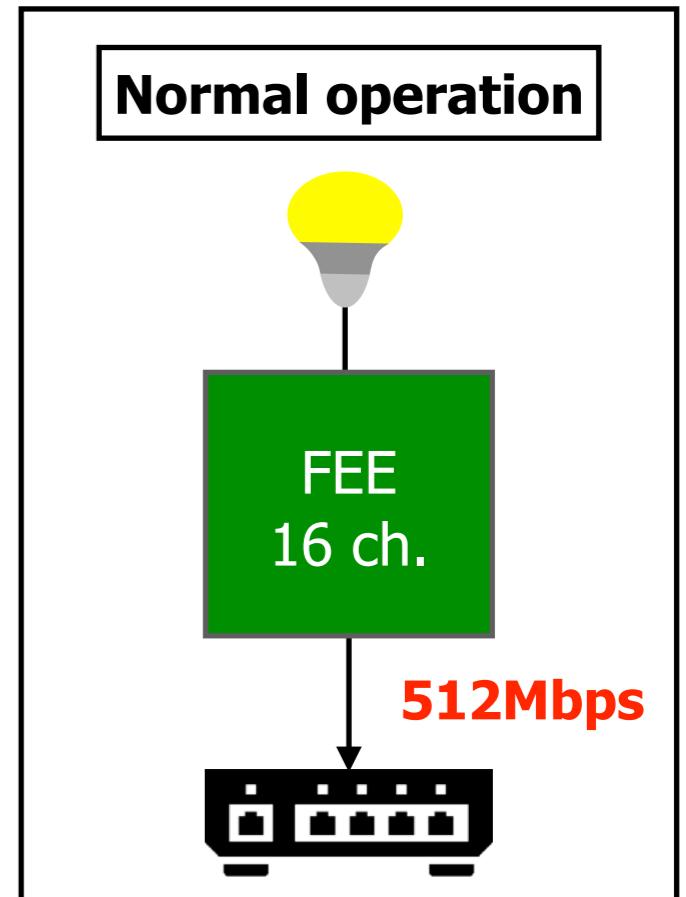
$$1\text{GSPS} \times 16\text{bit} \times 100\text{ns} \times 20\text{kHz} = 32 \text{ Mbps}$$

- **Cosmic μ**

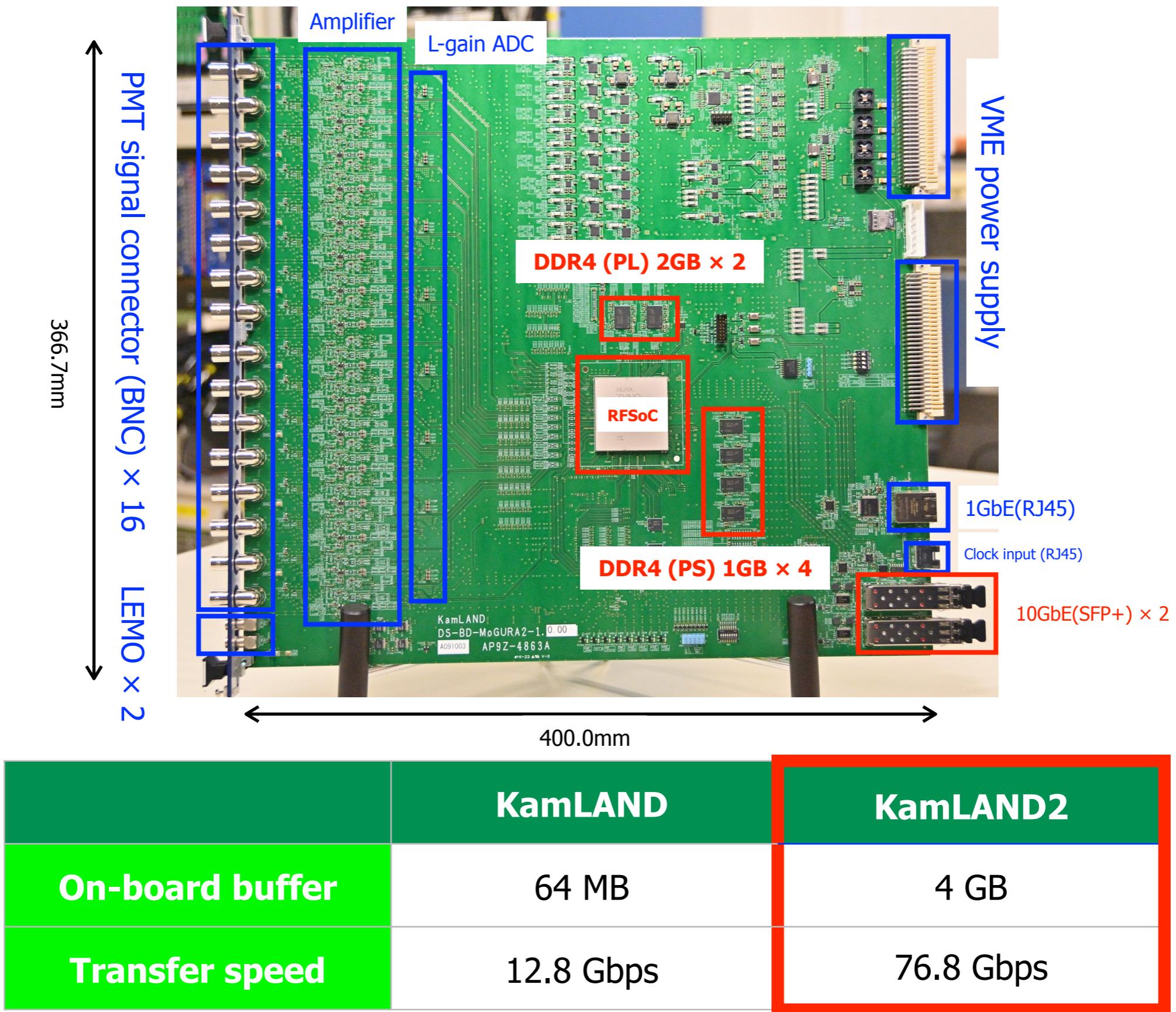
$$1\text{GSPS} \times 16\text{bit} \times 10\mu\text{s} \times 0.3\text{Hz} = 48 \text{ kbps}$$

- **Nearby Super novae**

$$1\text{GSPS} \times 16\text{bit} \times 100\text{ns} \times 1\text{MHz} = 1.6 \text{ Gbps}$$



Front-End Electronics

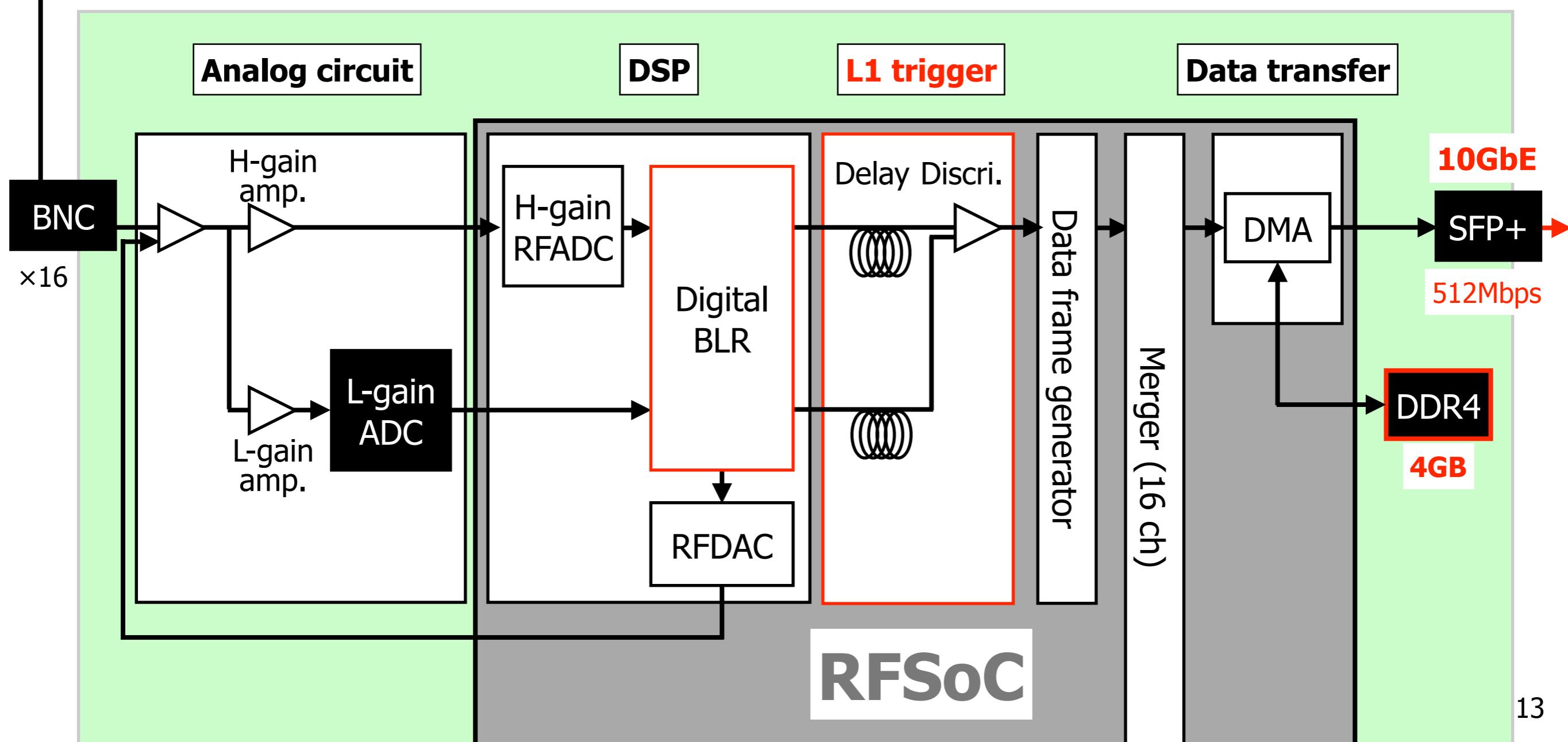


Front-End Electronics

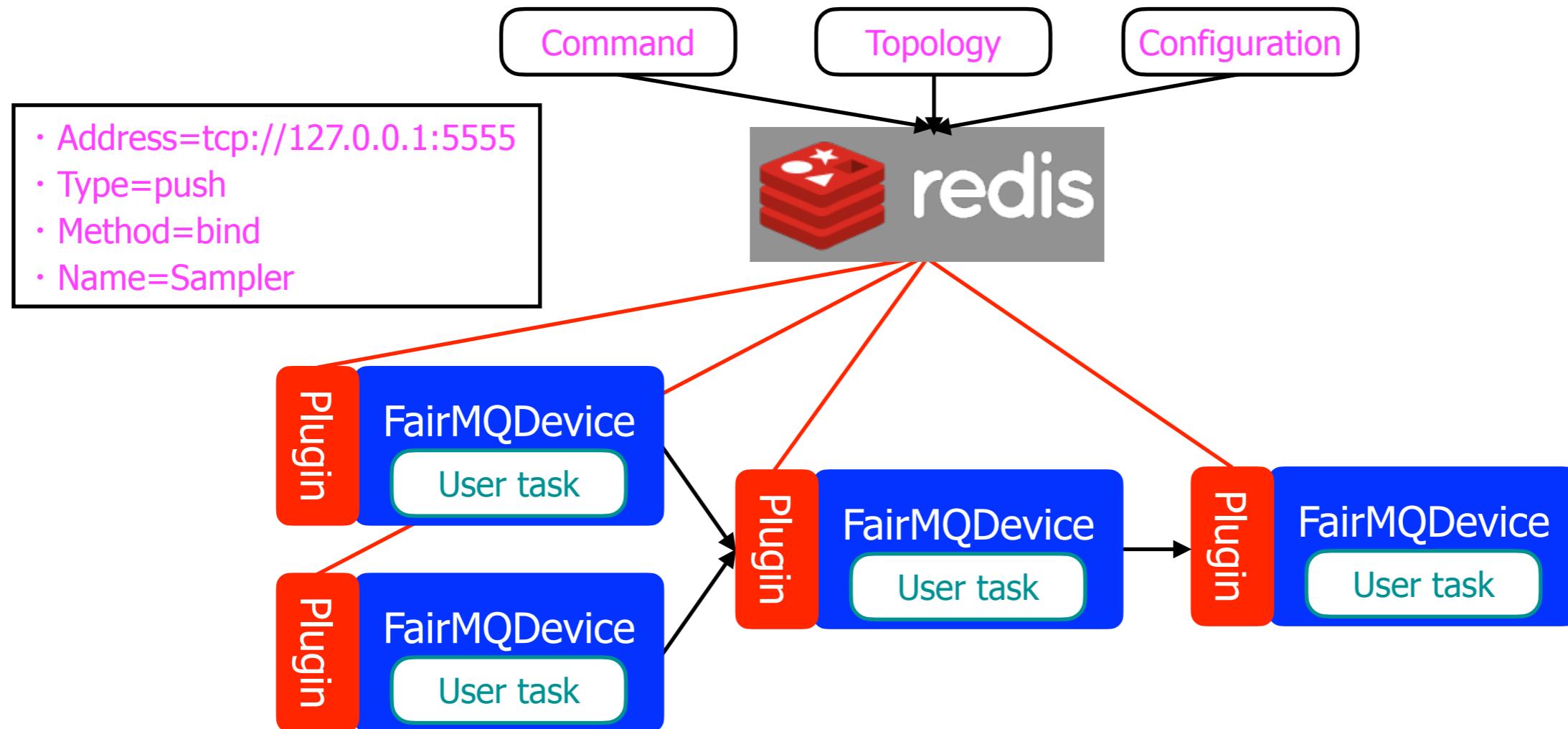
RFSoC (XCZU29DR-1FFVF1760E)

- High speed ADC and DAC
- Large number of Digital Signal Processors (DSP)
- Fine tunable digital BLR

ADC	16 ch
Resolution	12 bit
Max sample speed	2.058 GSPS
Logic cell	930300
DSP slice	4272



DAQ control



Plugin

- Development based on J-PARC E16 experiment
- Redis API
- Key-Value store

<https://github.com/redis/redis>

FairMQ

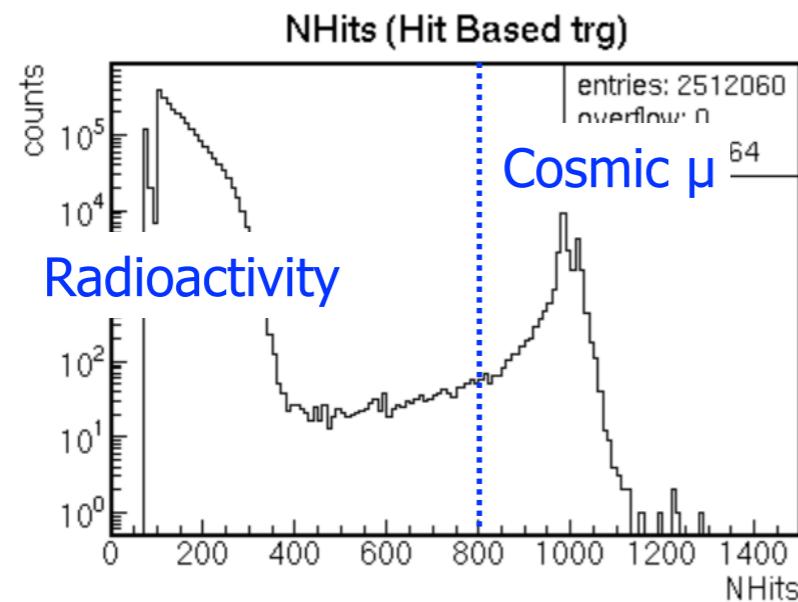
- Asynchronous message processing
- Various transports (TCP/IP, in-process, inter-process)
- Independent process (FairMQDevice)
- StateMachine

<https://github.com/FairRootGroup/FairMQ>

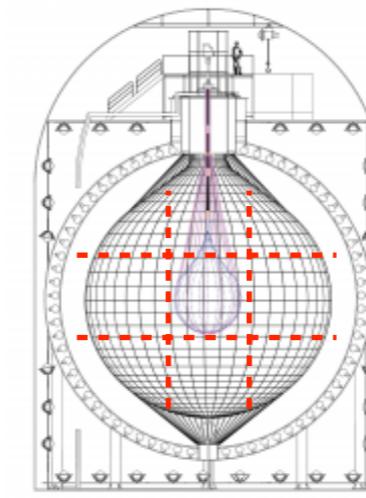
Trigger

Conceptual study is ongoing based on MC

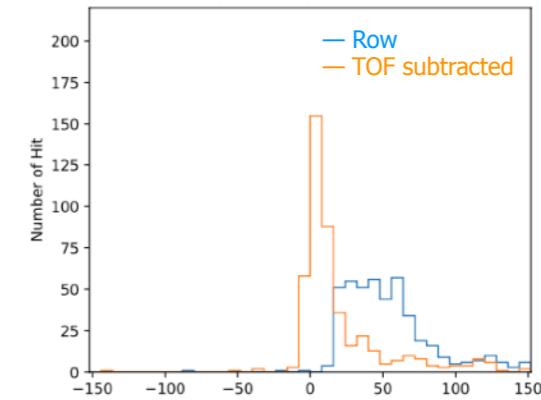
KamLAND



KamLAND2



Hit timing



L1 : 1 p.e. detection (hardware)

L2 : Hit based hardware trigger

Special trigger to detect neutrons after cosmic μ .

L1 : 1 p.e. detection (hardware)

L2 : Time-charge based software trigger

- TOF subtraction
- Vertex depend threshold

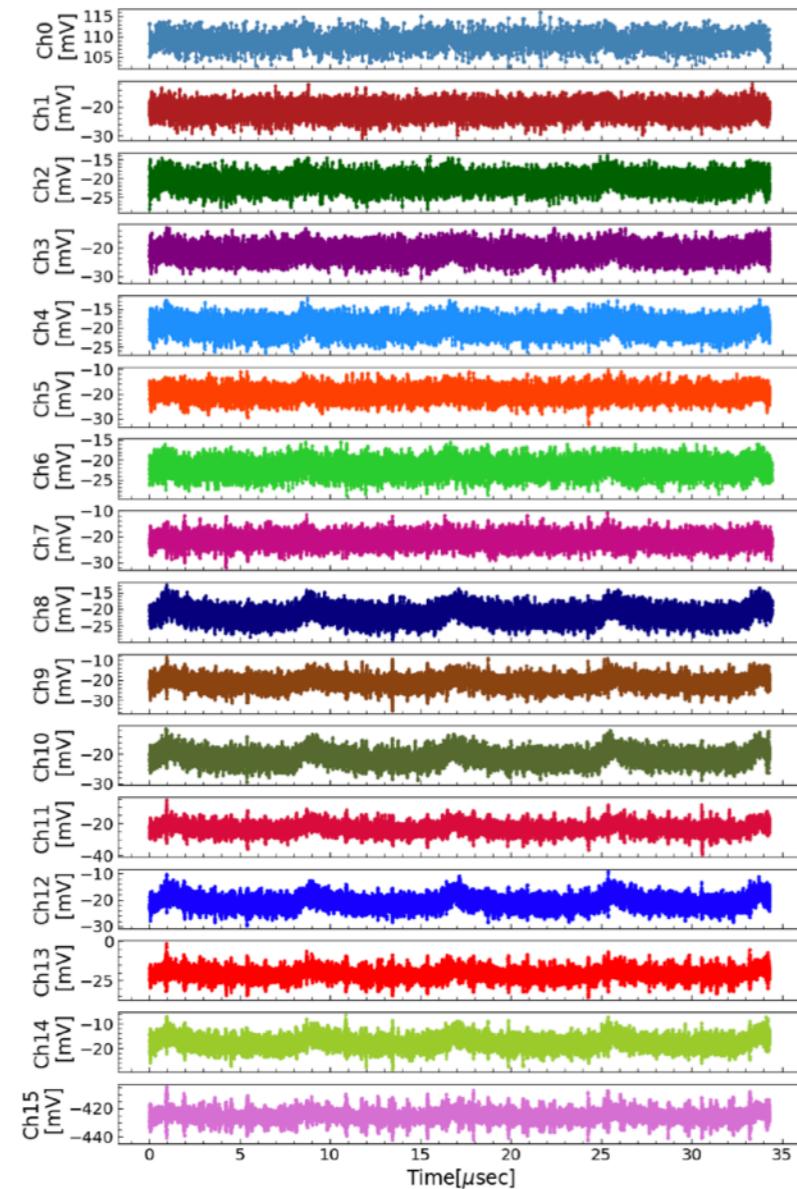
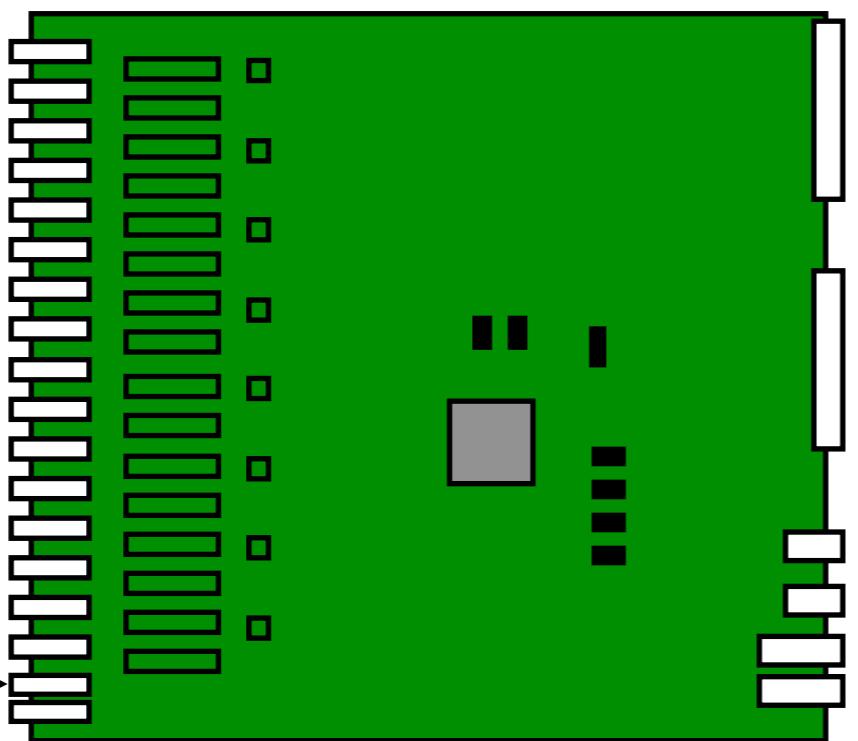
Waveform from all of hit channels to be recorded.

Development status

FEE

Recording waveform with a time window of 34 μ s.

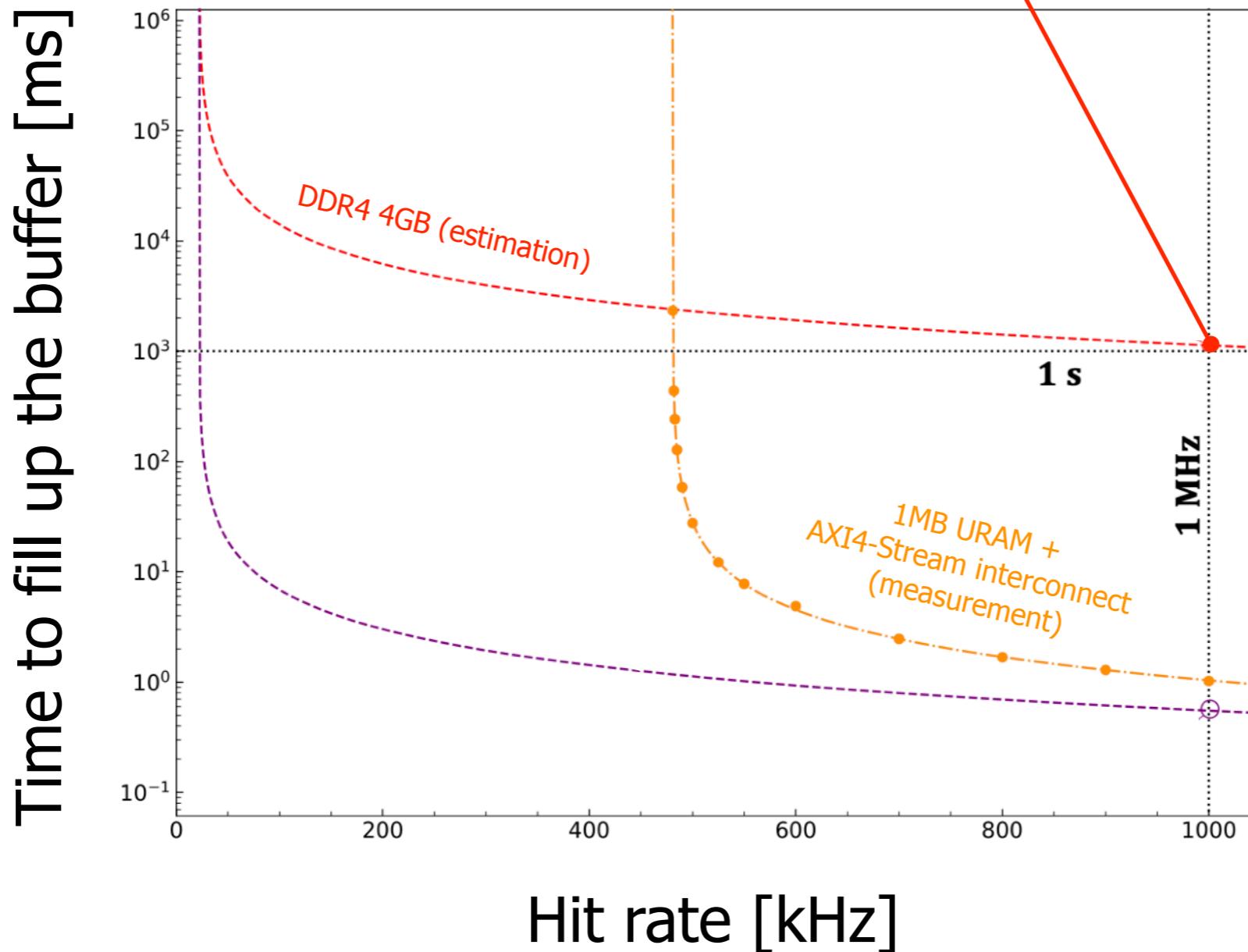
Long enough for cosmic μ events!



Development status

FEE

- Recording waveform with a time window of 34 μ s.
- Data acquisition of 1MHz events for 1s.



Summary

- KamLAND2 is planned to improve the energy resolution.
- Challenges for electronics are,
 - to achieve higher rejection efficiency of spallation backgrounds
 - to acquire data of supernova events.
- High speed processing is one of the important subjects.

Key for implementation

- RFSoC, DDR4 and 10GbE on FEE
- Network design and data processing with FairMQ
- Time-charge based software trigger

Next step

- Operation of KamLAND2 prototype detector around end of this year.