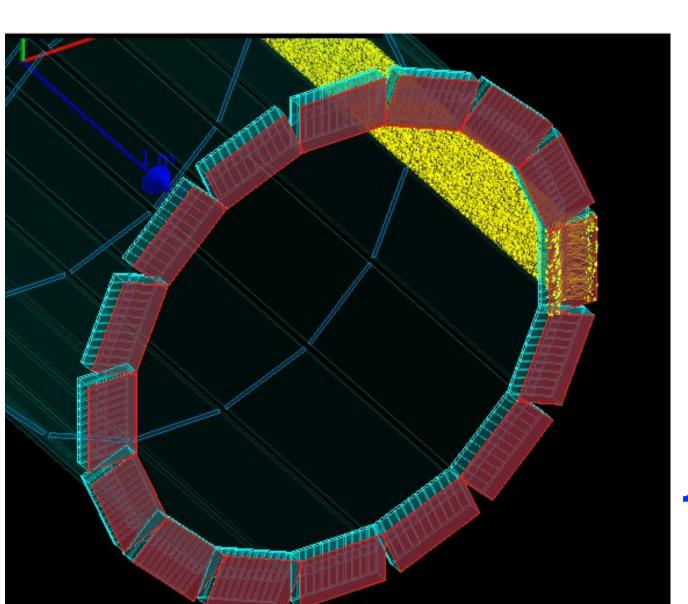
#### Plans for novel Cherenkov detectors at the Super Charm-Tau Factory at Novosibirsk CREMLIN



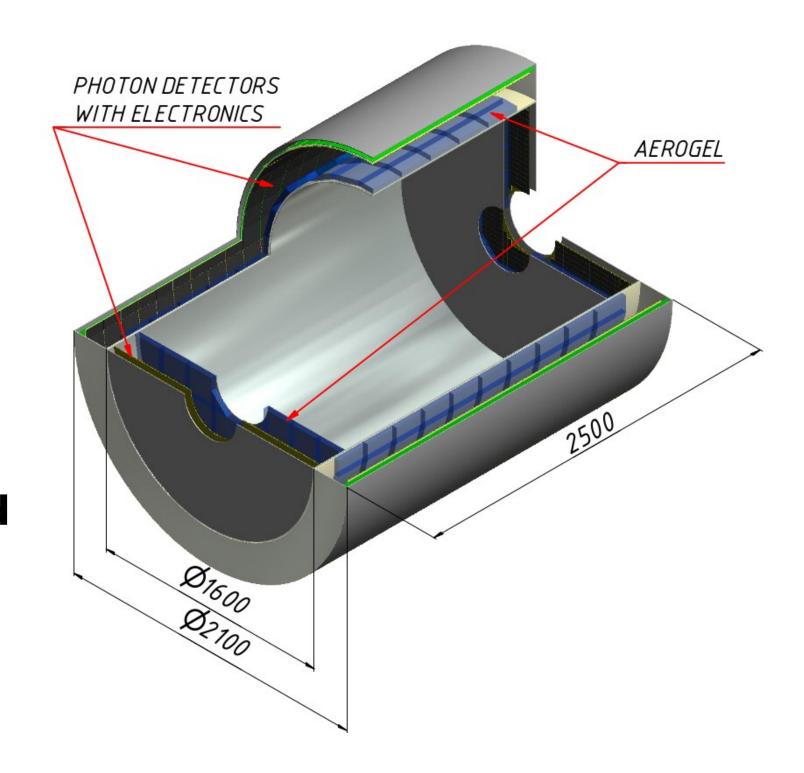


https://ctd.inp.nsk.su/c-tau/

Avetik Hayrapetyan, M. Dueren, S. Kegel, M. Schmidt,

A. Yu. Barnyakov, V.S. Bobrovnikov, S.A. Kononov

Justus-Liebig-University, Giessen, Germany and **Budker Institute of Nuclear Physics, Russian Federation** 

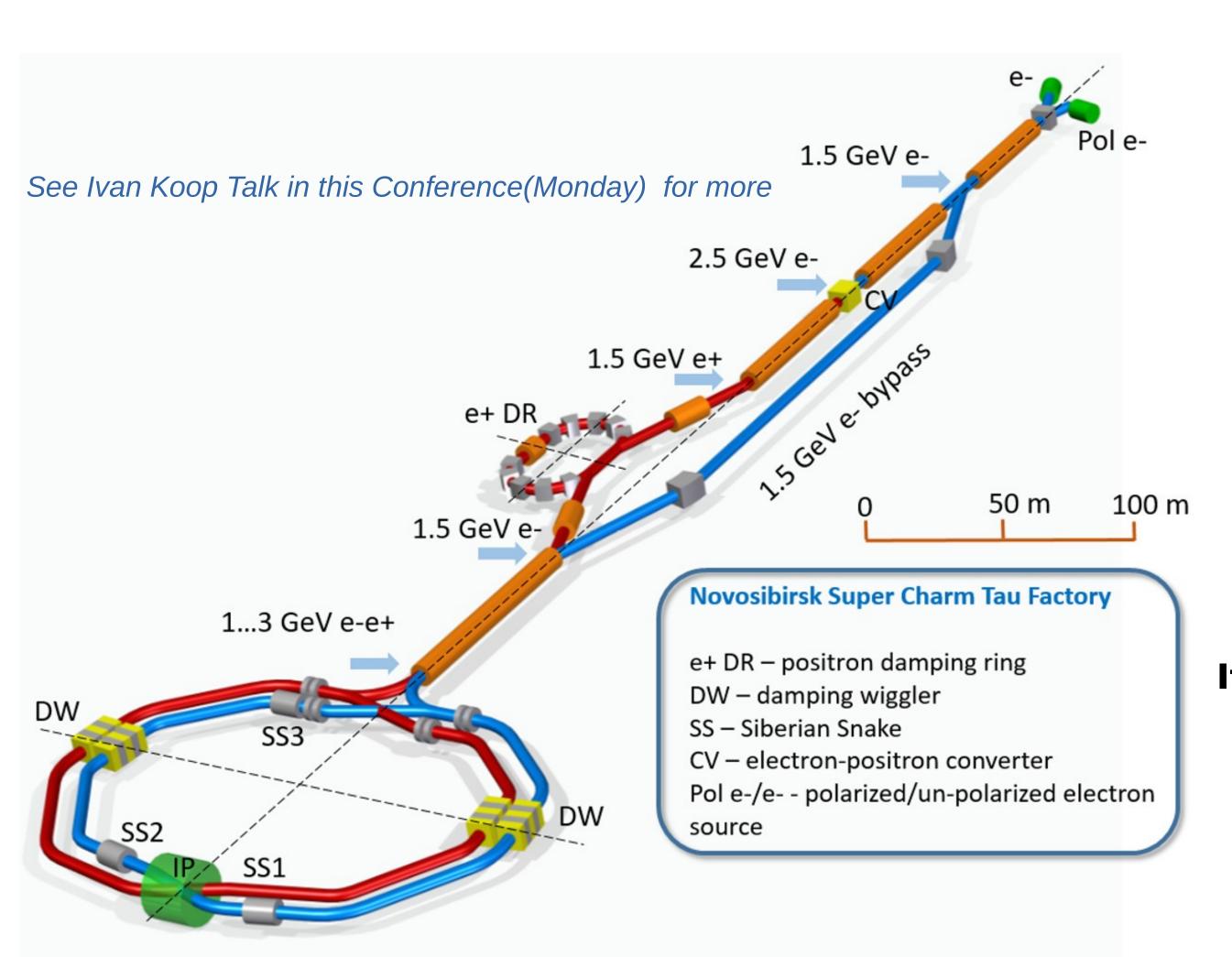






#### The new SCTF Collider





<b>A</b>				1 4		
Α.	Bog	omy	/agi	KOV(	RI	NP.

The 2020 joint international workshop for Super-Charm Tau Facility 16-18 November 2020

E(MeV)	1500	2000	3000
	2	2	2
I(A)			
	8	6	10
N_(e/bunch)×10^(-10)			
	323	431	259
N_bunches			
	0.4	0.5	1
Lumi 10^(-35) (cm^(-2) s^(-1))			
	2800	3200	1731
τ_Lifetime(s)			

It offers also a polarised e- beam, high (90%) at 1500 and moderate (40%) at 3000 MeV

Having such a hightec device one can "attack" a big goals, like a

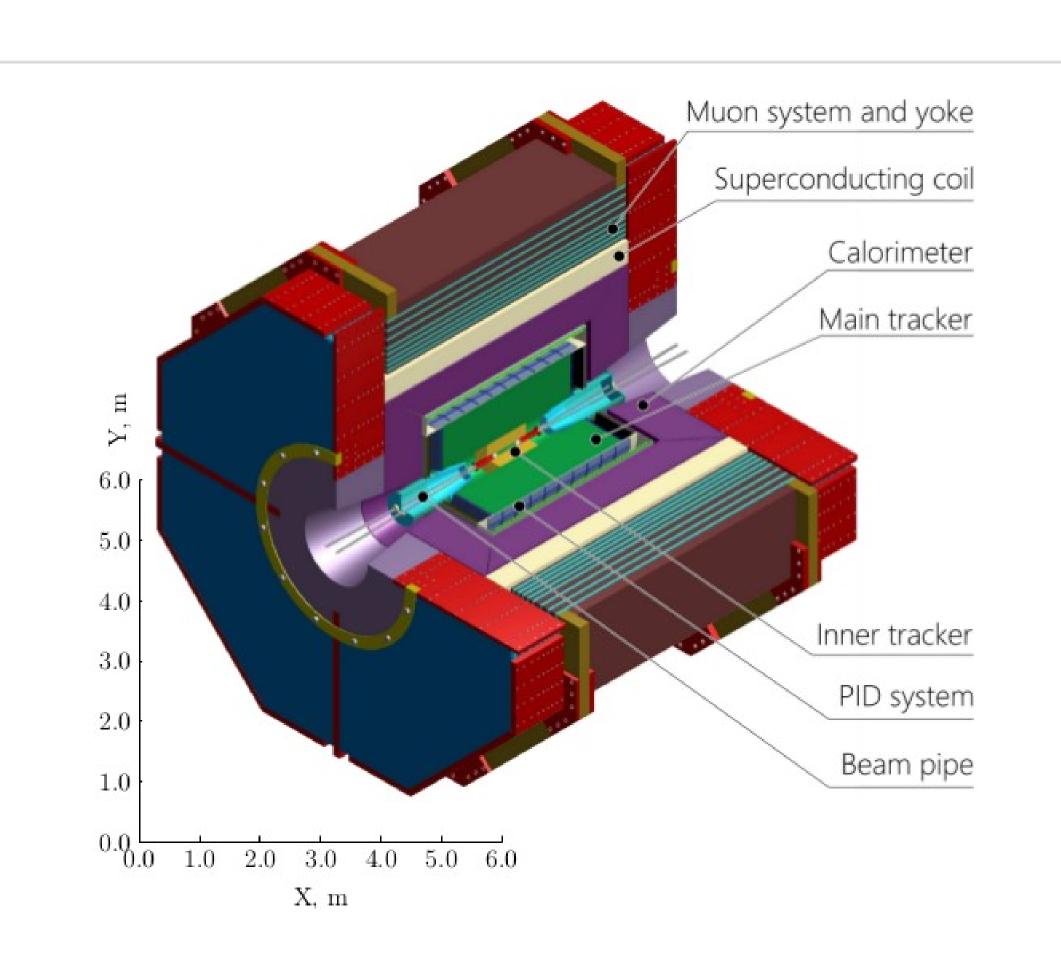
LFV

Decay of tau leptons into mu gamma

 $\tau \rightarrow \mu \gamma$ 

#### The Detector and requirement on it





There are very tight requirements on different detector parts:

Here we stress the part direct touching PID for LFV search, Because it is the strongest one is:

-Very good momentum resolution, below 1%

-Good μ/π/k separation till 1.5GeV/c

-special tight requirement on muon/pion acceptance and separation

-good  $\pi^0/\gamma$  separation in wide range of energies 10-3000 MeV

For detailed physics program and all requirements on detector see here:

В. Воробьев, ИЯФ СО РАН, "Проект «Супер Чарм-Тау Фабрики"

Seminar ,15.04.2021, ΦИΑΗ

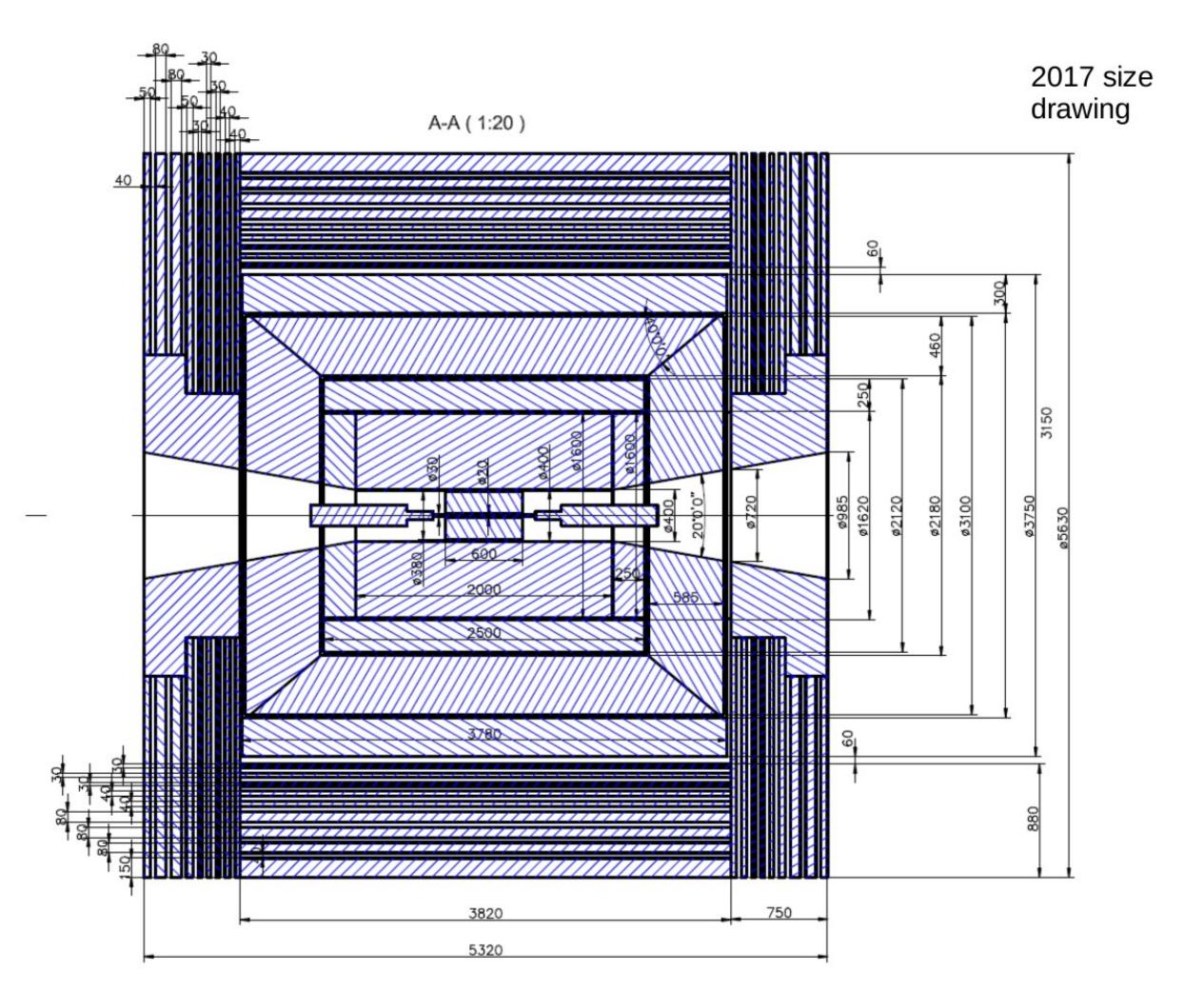
A. Sokolov "R&D activities on SCT detector"

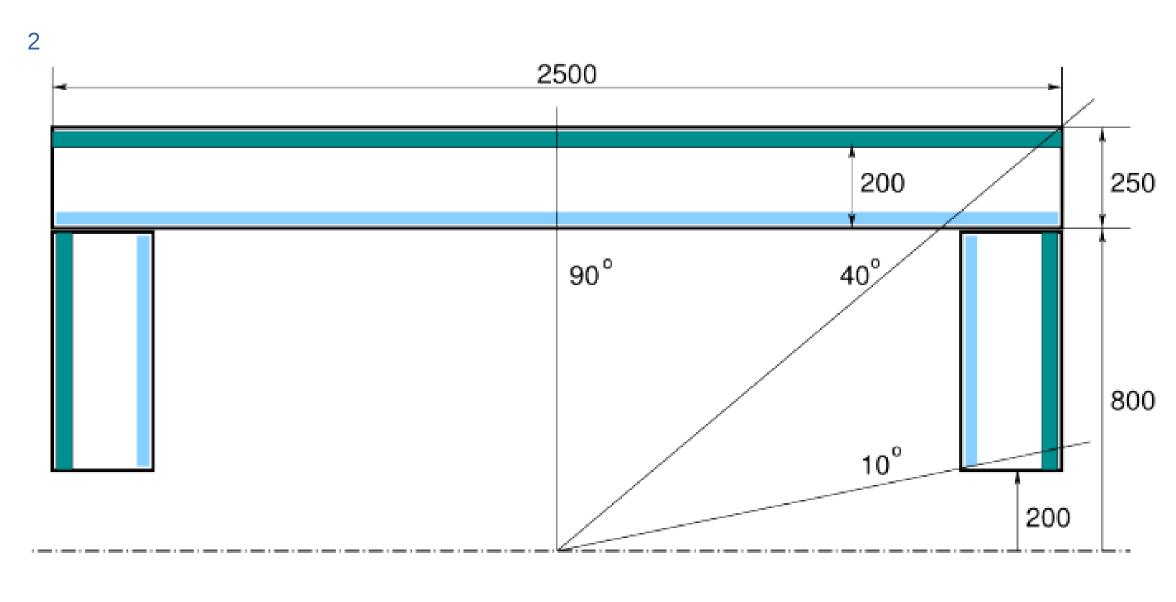
The 2020 joint international workshop for Super-Charm Tau Facility 16-18 November 2020

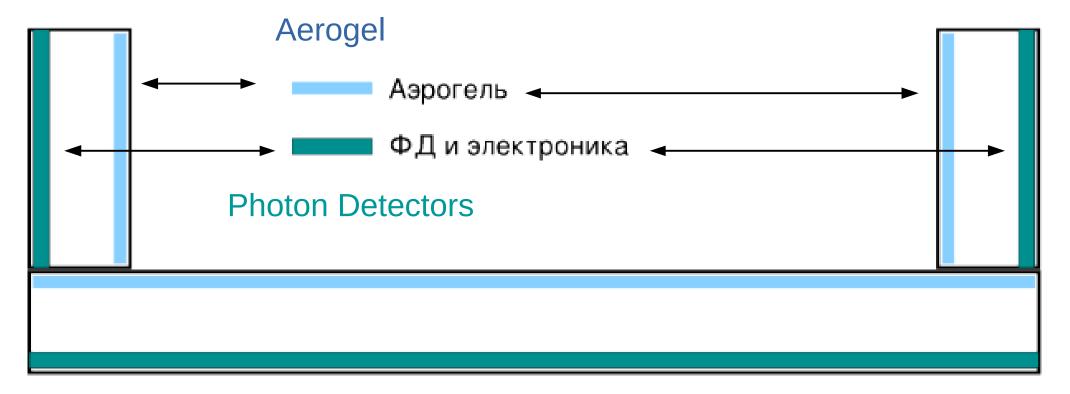
#### The SCTF detector volumes and PID



#### part (here the FARICH version)





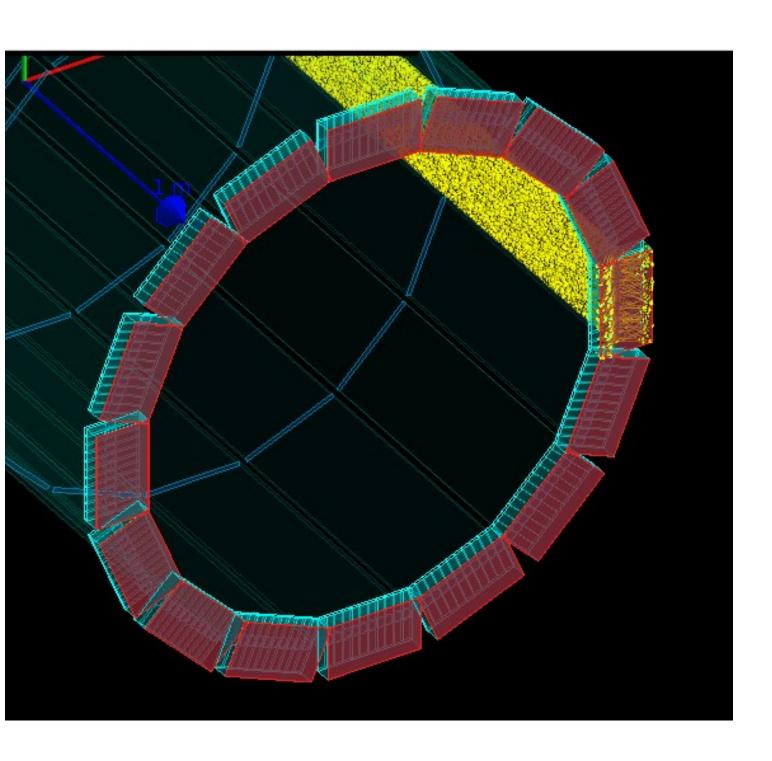


The fDIRC models we implement in Geant MC should reside inside this volumes



#### Two PID concepts are in consideration





The Outline of this talk

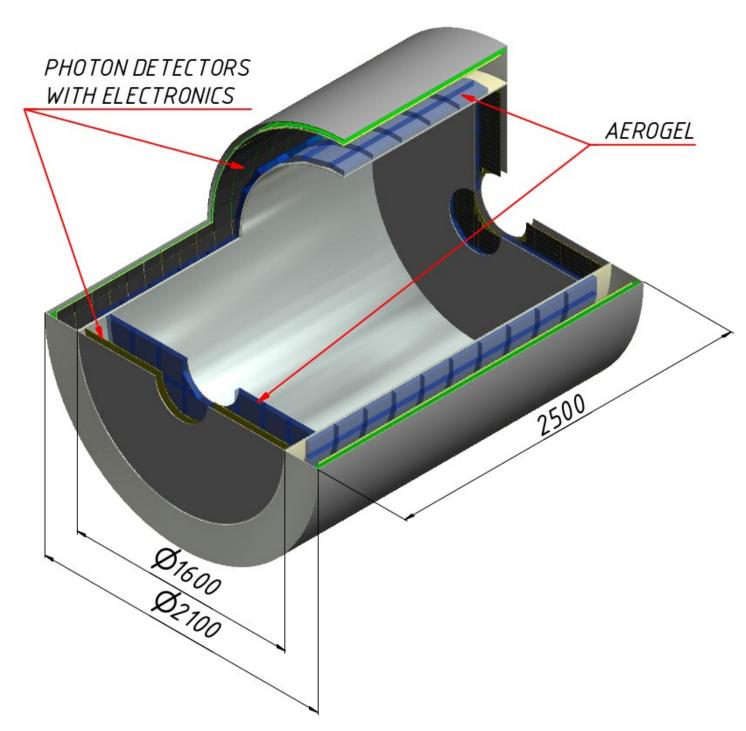
Barrel fDIRC in GEANT MC

Addressing the main "contributor" of the resolution

Candidate Prototype Photon detector tests at GCS

RD activities of FARICH part

Conclusions



fDIRC Option, two "PANDA like" endcap disc DIRC and one Barrel, both with focusing optics

Will project Cherenkov Photons on ~1.024 m² (Barrel) and ~0.6912 m² (Endcaps) Detection Surface with granularity to be optimized

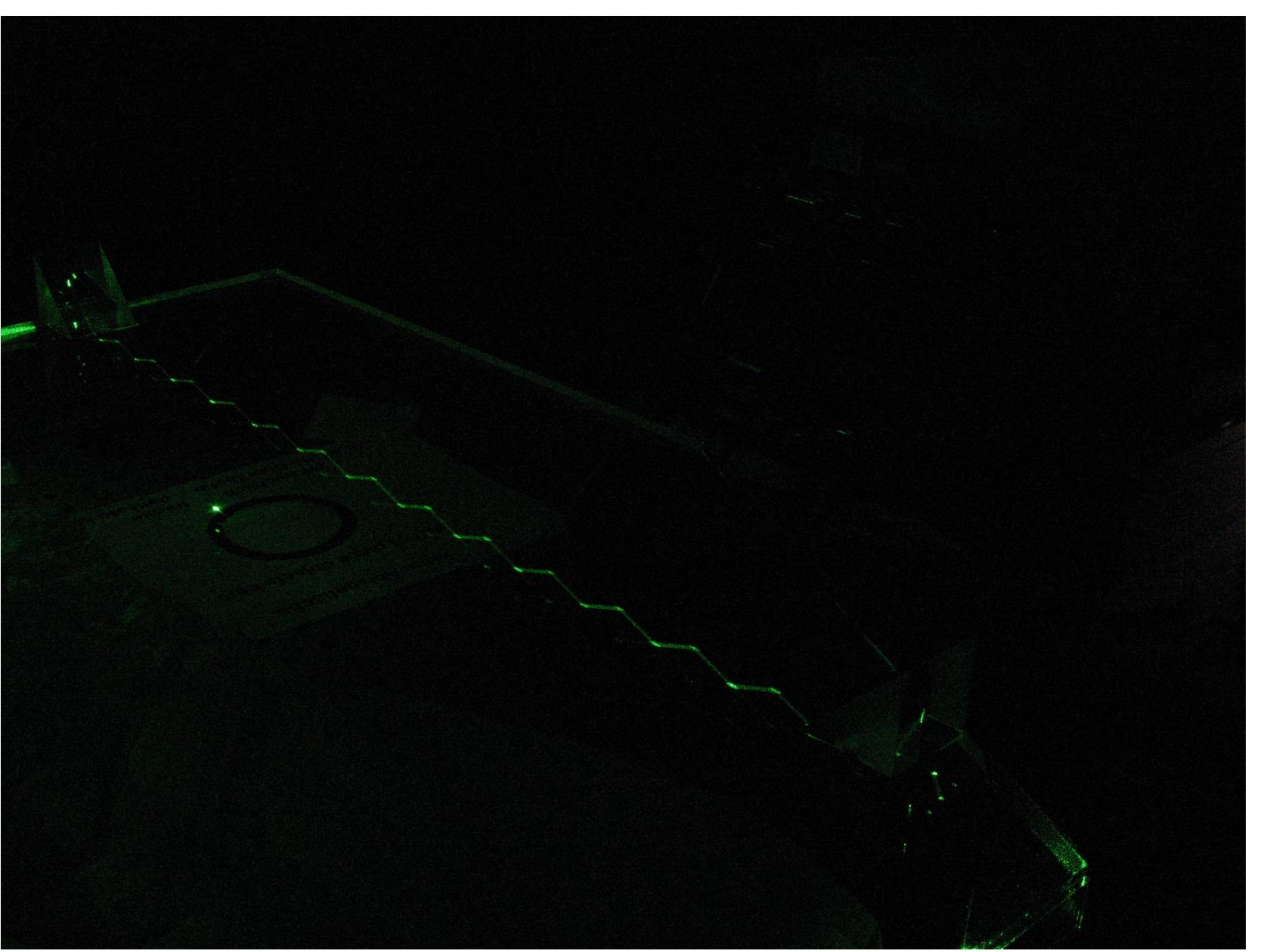
FocusingAerogelRICH, FARICH Option,

two BEII-2 ARICH geometry like endcap FARICH and one novel Barrel

Will project Cherenkov Photons on ~16.0 m² (Barrel) and ~5.0 m² (Endcaps) Detection Surface with granularity to be optimized

#### The Cherenkov radiation and the DIRC principle





nβ>1

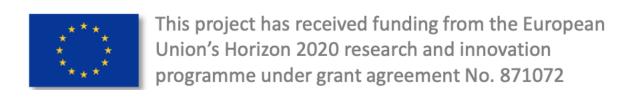
 $\theta_{c}$ =arccos(1.0/n(λ)β)



Such a stationary setup with possibility to measure radiator Quality parameters should be a good test bed for students to learn DIRC optics

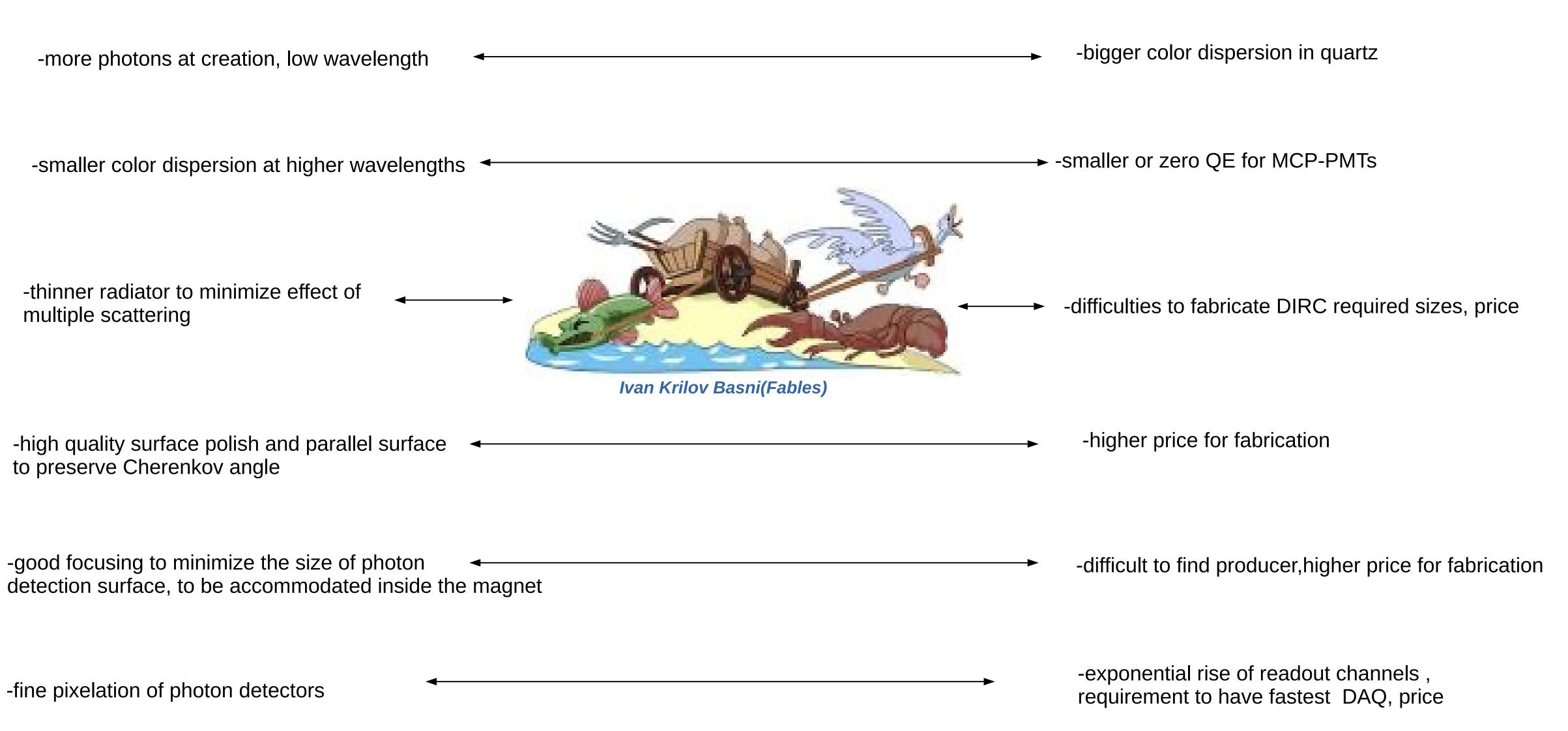
For the recent review of past, current, future DIRCs see here: J. Schwenning "DIRC technology requirements"

https://indico.cern.ch/event/999817/contributions/4253047/ attachments/2239796/3797410/20210506-ECFA-TF4-DIRC-schwiening.pdf



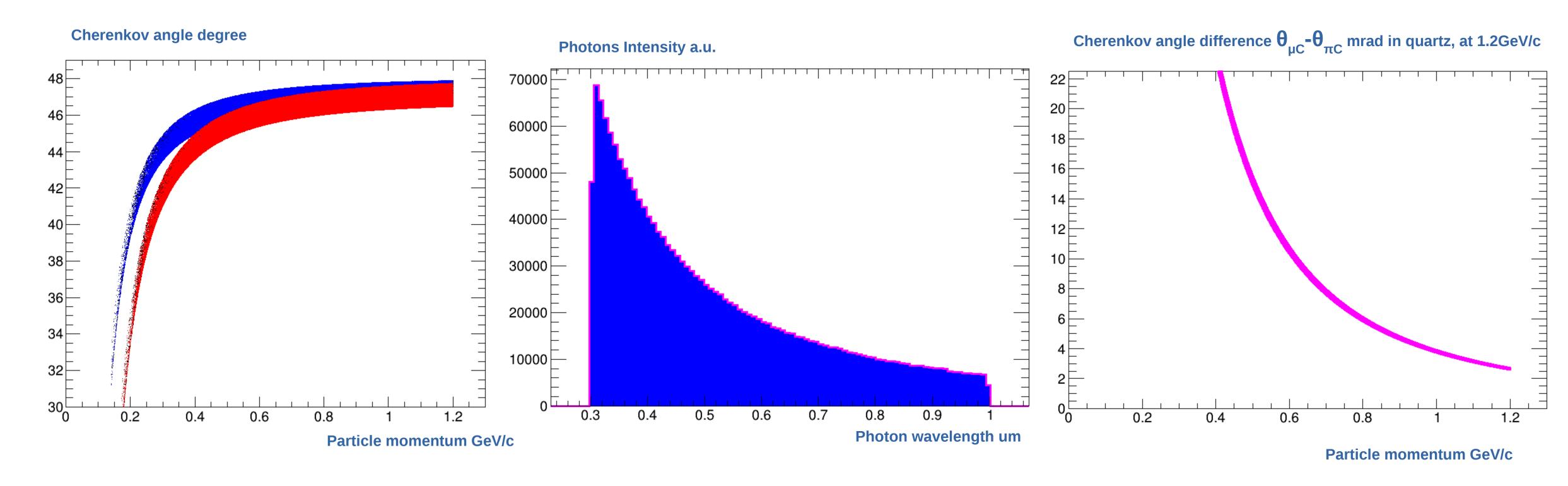
## The many opposite requirements to build a good fDIRC







#### The characteristic spectra of Cherenkov Photons in Quartz



From now we will discuss the case of  $\pi \mu$  in quartz radiator, because of as one see from plot at the right side

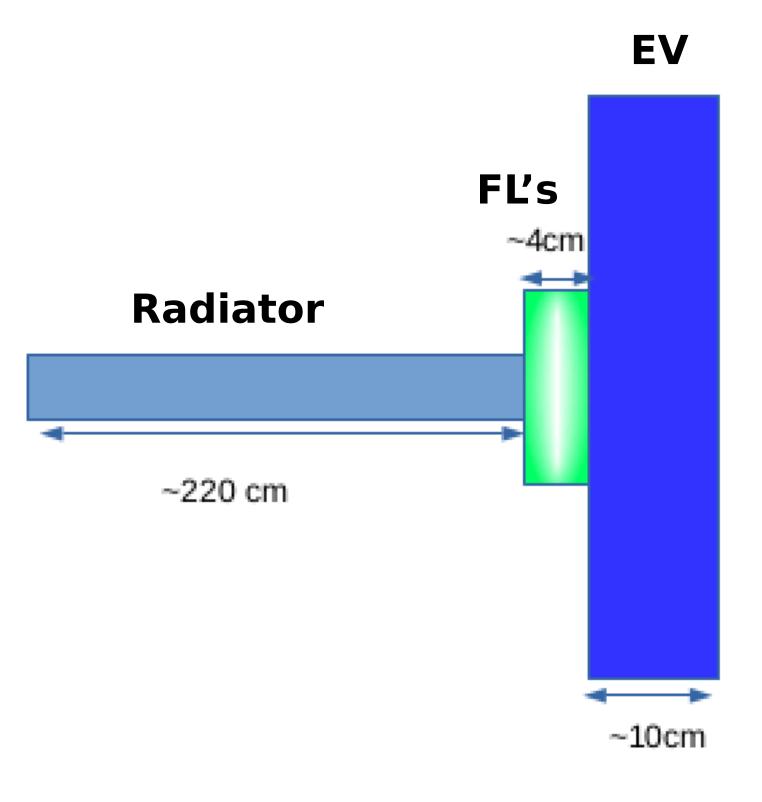
The separation of  $\pi/\mu$  is most challenging

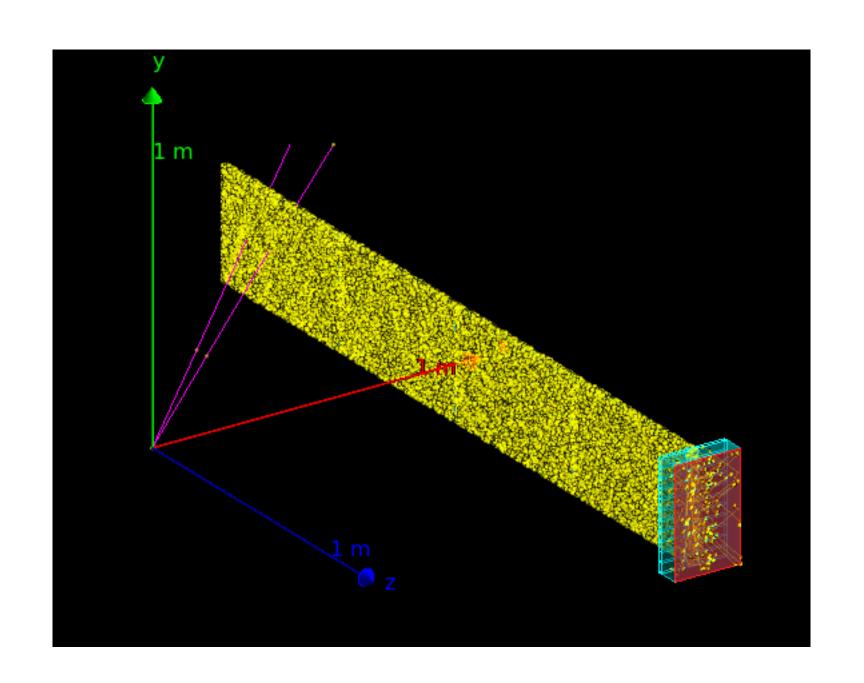
We start at 300 nm because of usage of only radiation hard glue Epotek-2 to connect optics, is transparent above this limits

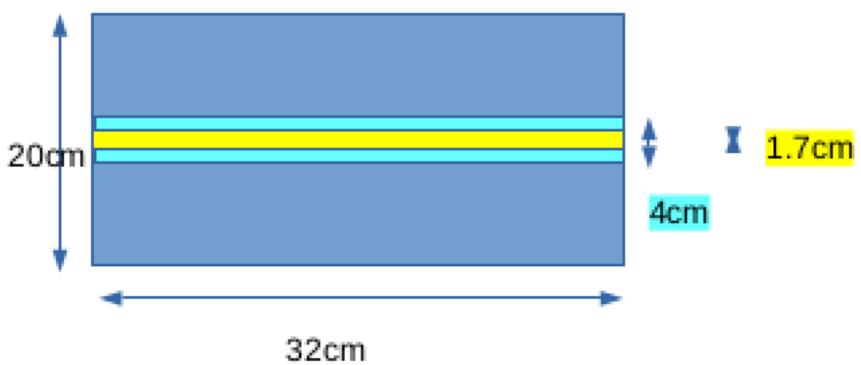
The upper limit is defined by quantum efficiency of common photon detectors

## Possible Module for barrel fDIRC







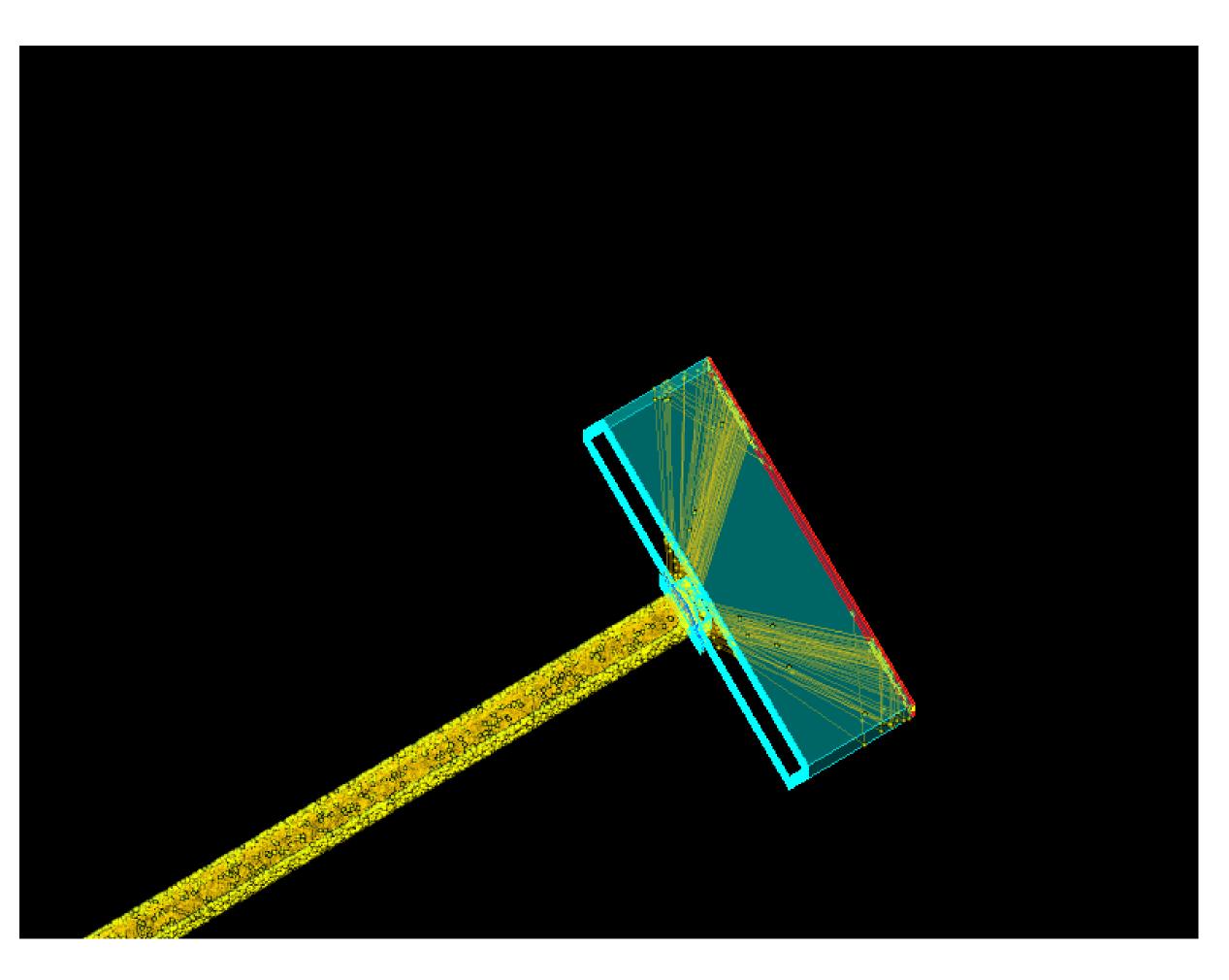


The Geant model is based on R. Dzhygadlo(GSI) EIC fDIRC, with modification of focusing lens(FL) and expansion volume (EV)

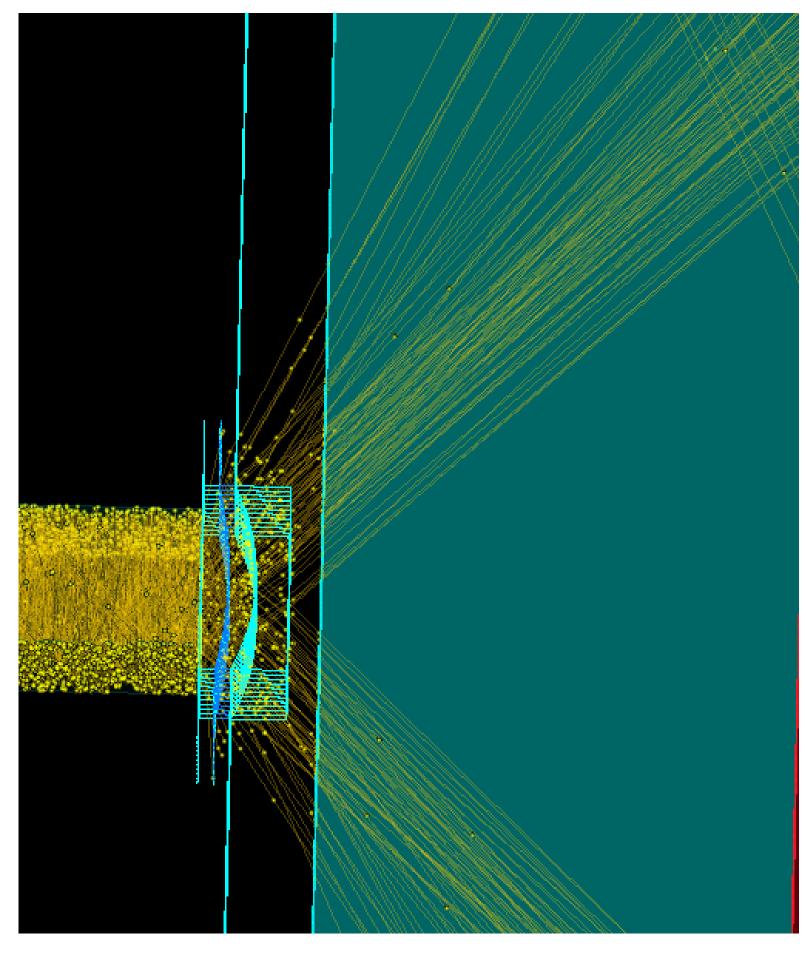


## The new Expansion Volume and 3 layer focusing Lens





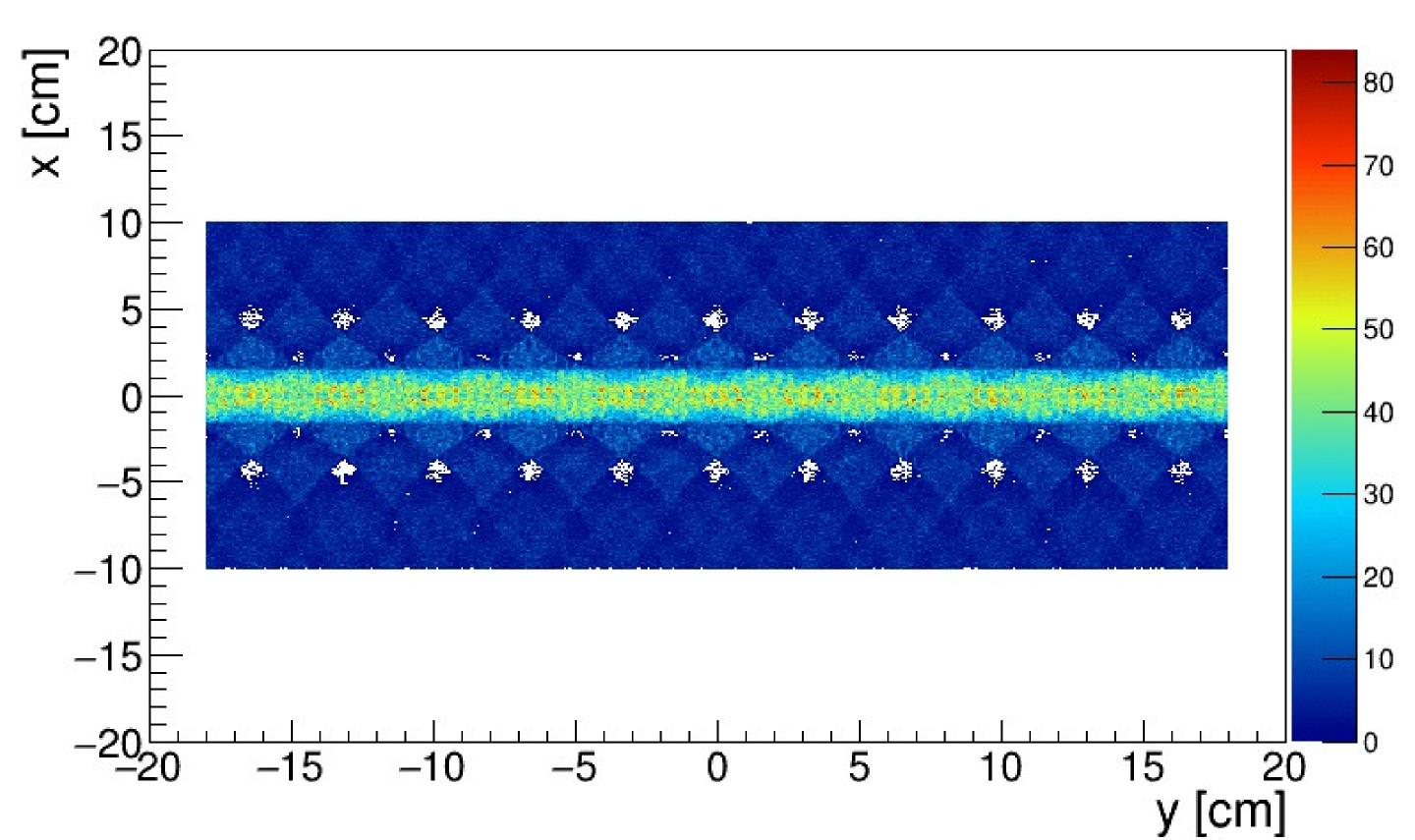
The new Expansion Volume is 320X200X100mm Quartz Module in MC, here zoomed with Cherenkov Photons produced by 1 GeV/C Muons



Here zoomed area shows the 3 Layer focusing Lens with circa 20mm thickness all together, again with MC photons There are ongoing MC studies for the Lens, 2 or 3 layer?? Focusing length, materials



## The angle coverage of Barrel DIRC



The hit map of Cherenkov photons on surface of EV where the Photon Detectors supposed to be coupled Starting from 40 till 140 degree polar angle

It is assumed that Barrel PID(DIRC) Should cover from 40 to 140 degree of polar angle and 360 in azimuthal

The magnet/detector geometry allows
To install 2500mm length radiator at
circa 930mm position in radial direction

The fDIRC option requires space for Expansion Volume(EV), focusing Lens (fL) And for Photon detector+Readout Electronic

Taking all this into account we suggest to have such a configuration at first in MC

Radiator 17\*320\*2200 mm EV 30\*200\*100 mm fL 8X20\*40\*40 mm

The remaining circa 160 mm place in Z direction One can use for a place for Photon detector and Readout electronic

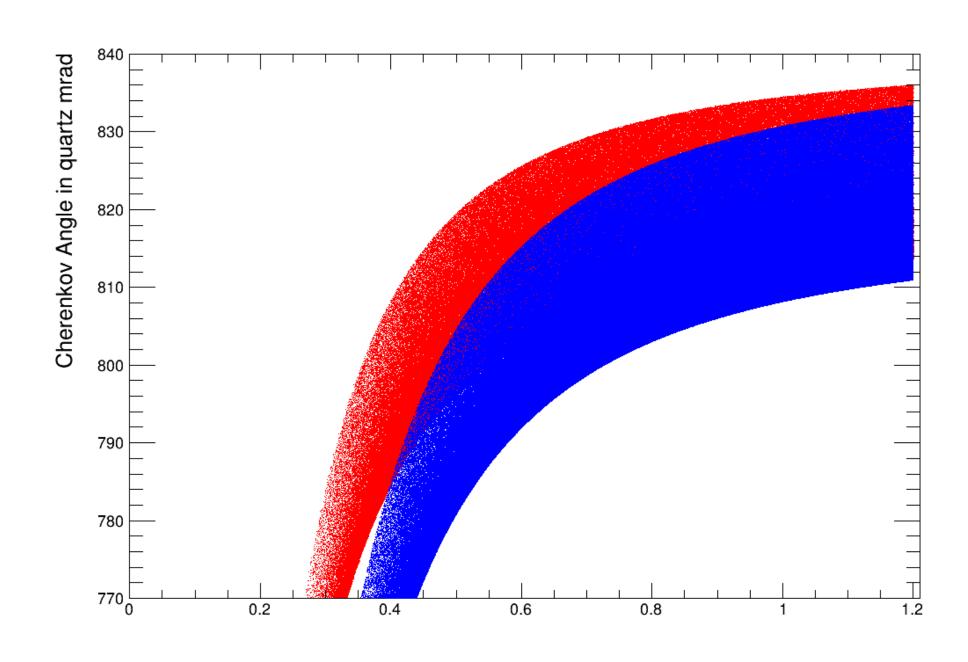
#### The importance of the wavelength restriction

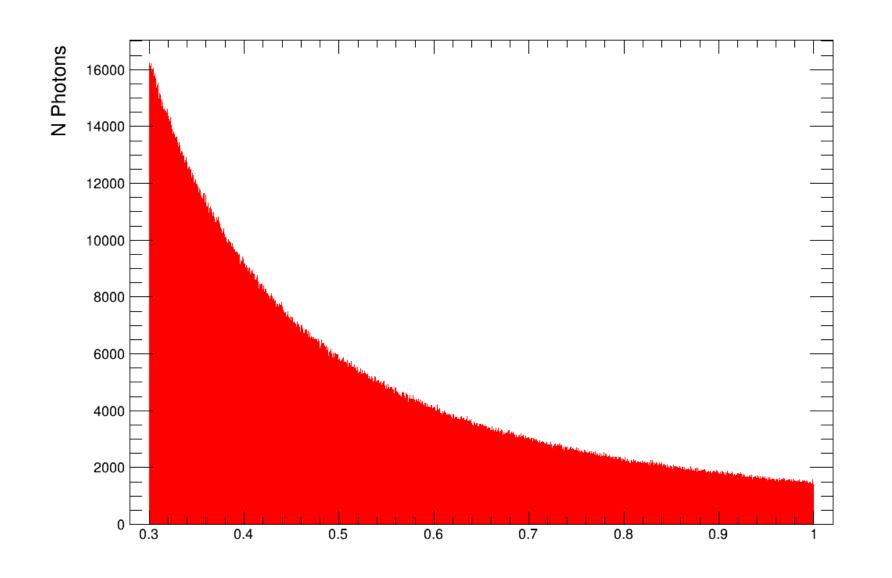
## In Silica produced Cherenkov Photons Number Against wavelength ->

#### **Pions**

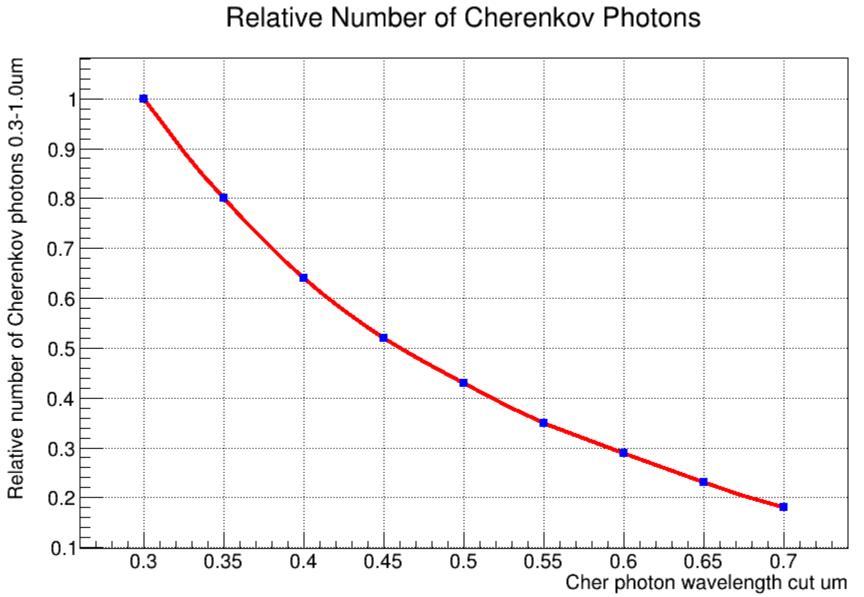
#### Muons

#### Cherenkov angle against momentum







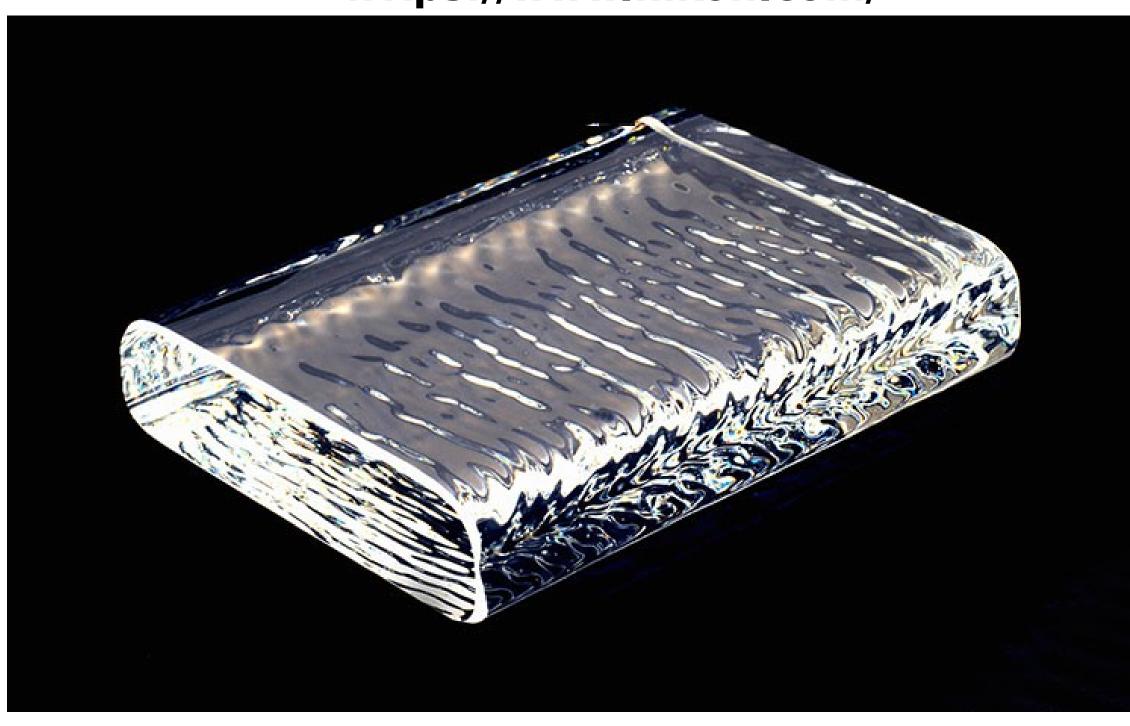


## Radiator with less dispersion



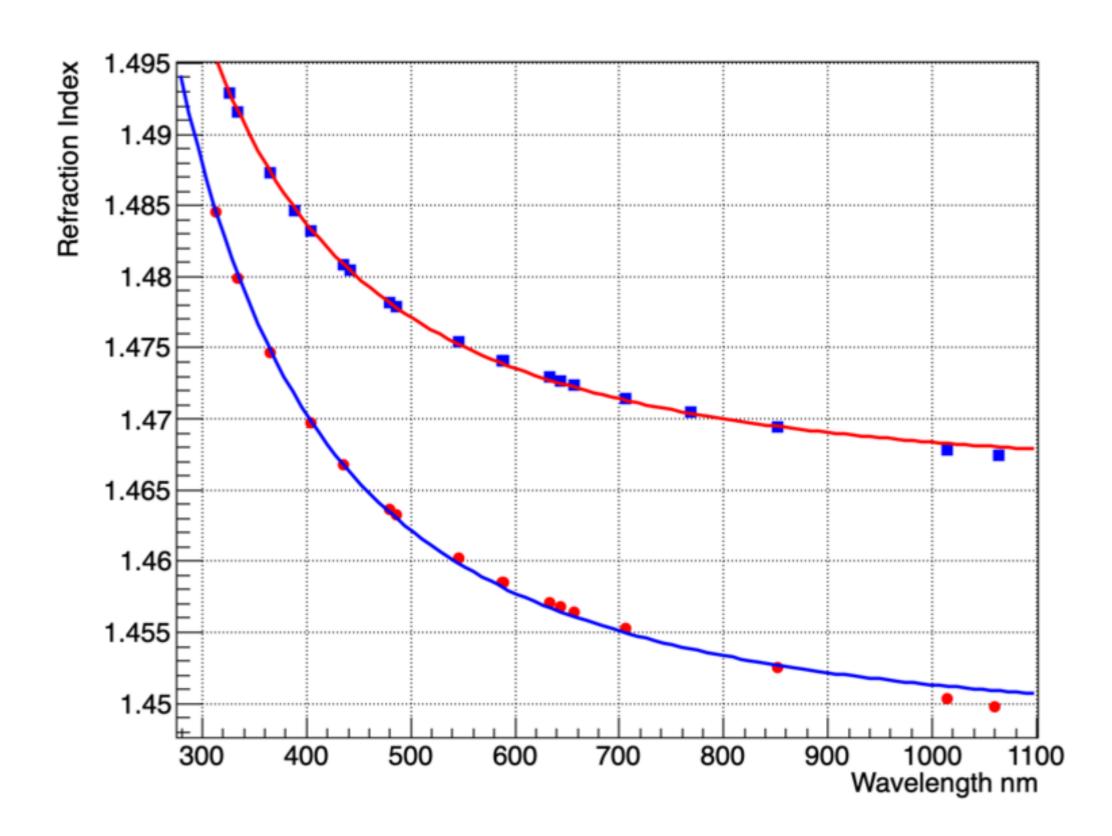
#### Nikon i-line Glass

https://www.nikon.com/



One of main components that restricts
DIRC angular resolution is radiator dispersive medium
while in case of disc DIRC one can go into LiF bars to correct
the Cherenkov angle, here we might try to use a material with
small dispersion

In both cases the radiation hardness could come in game

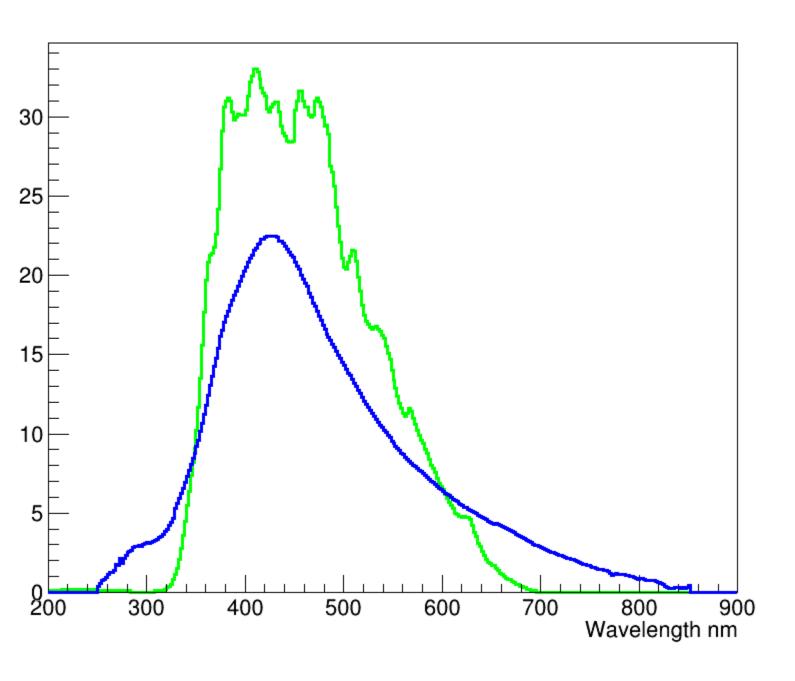


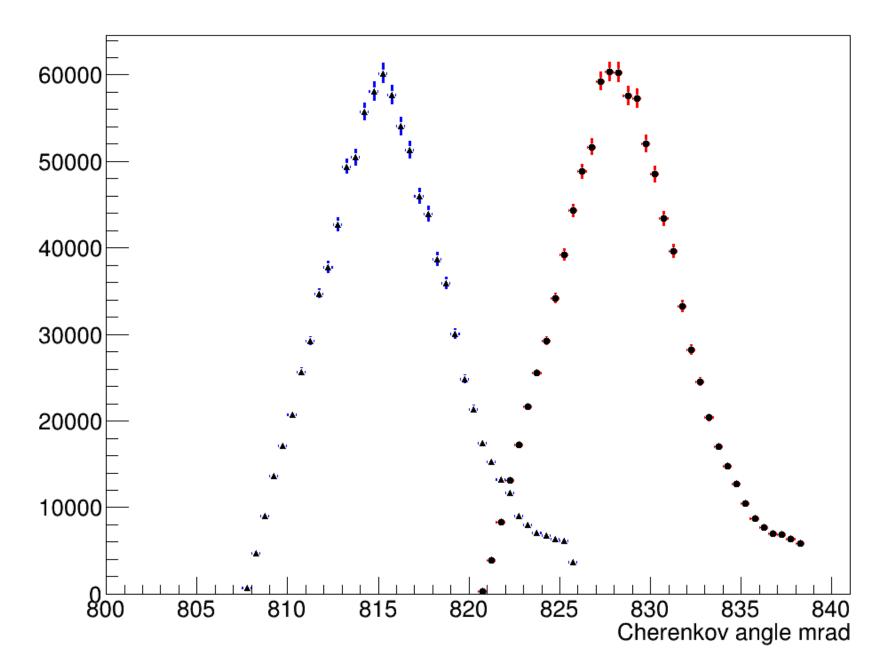
Nikon iGlass(blue points) and SiO2(red points) Refractive Index against photon wavelength, such a radiator could already minimize color dispersion 1-2 mrad at creation compared to quartz one

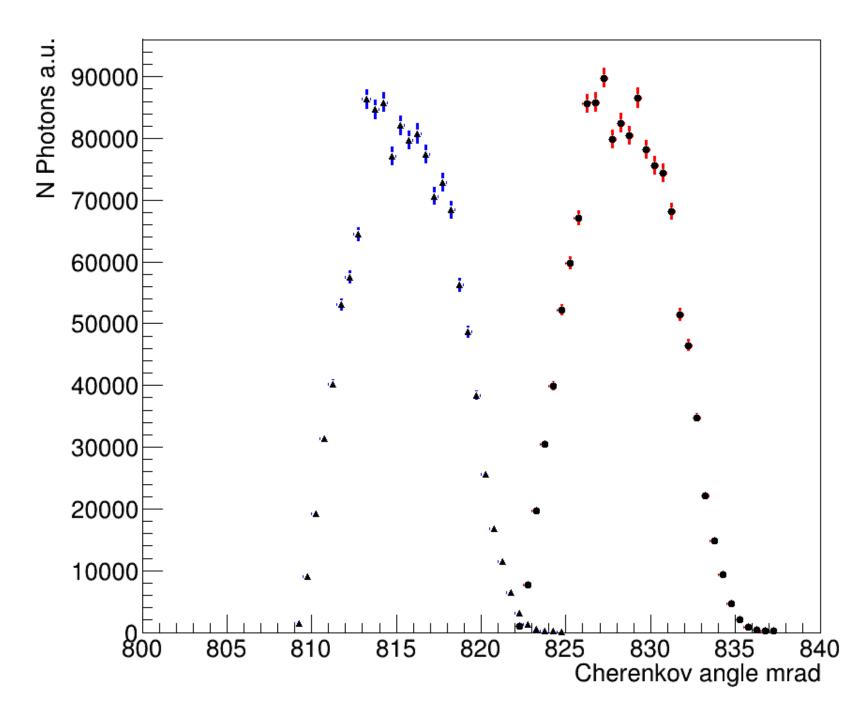


## The color dispersion, ways to overcome









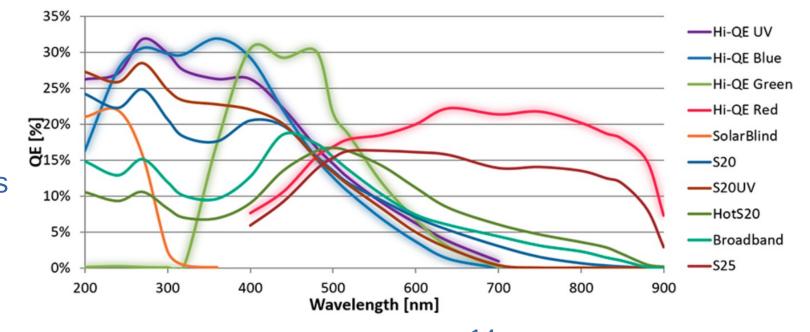
Even small restriction of wavelength interval, here our both MCP-PMTs QE against wavelength helps to reach a bit better separation

But if one can get Planacon MCP-PMT or SiPM like red cathode MCP's from Photonis this will ease the job significantly

See here for Photonis new development

https://www.photonis.com/products/hi-qe-photocathodes

Cherenkov Photons emission angle spectra from 4 GeV/C Pions(right) and Kaons(left) weighted with Photonis Aqua(blue curve in left) QE



Cherenkov Photons emission angle spectra from 4 GeV/C Pions(right) and Kaons(left) weighted with Photonis Green(green curve in far left) QE

For better separation then it will be a must:

-go into red side(High-QE Red is a candidate)

-use less disperse materials

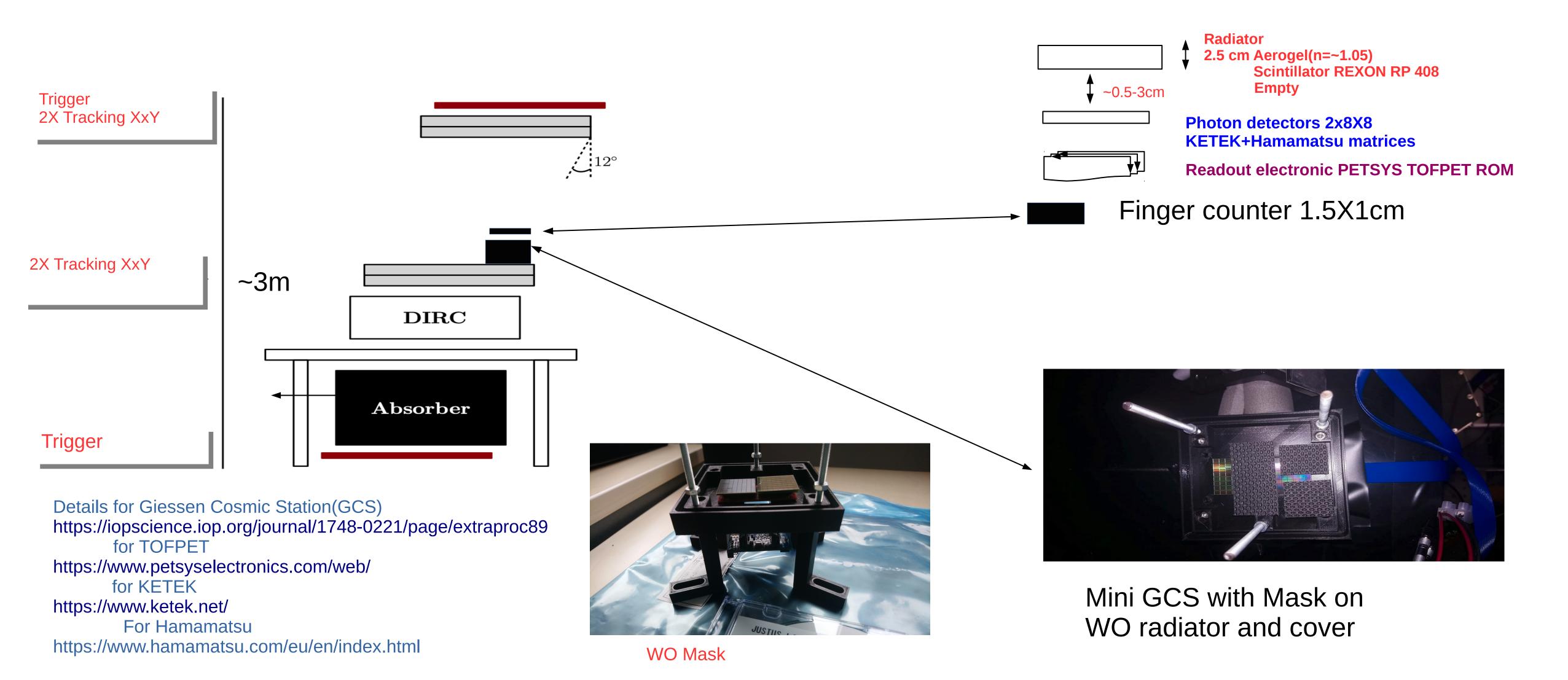
-use fast photon timing to perform a correction



This project has received funding from the European Union's Horizon 2020 research and innovation programme under grant agreement No. 871072

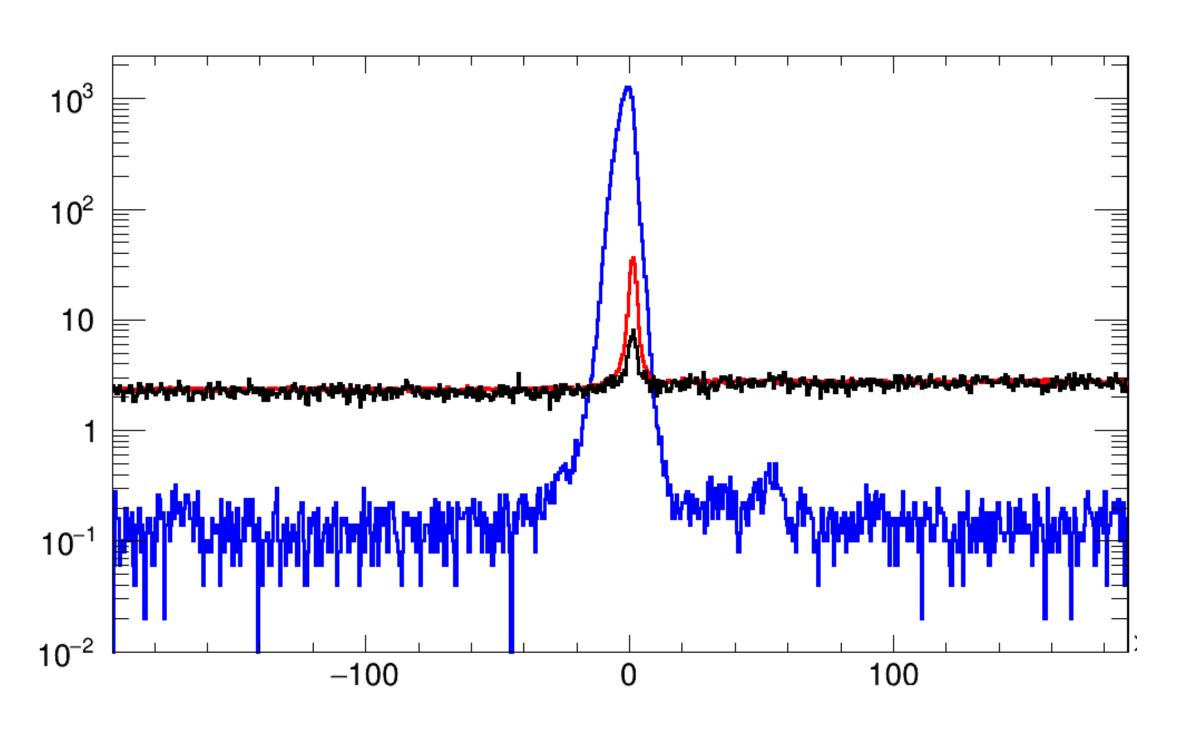


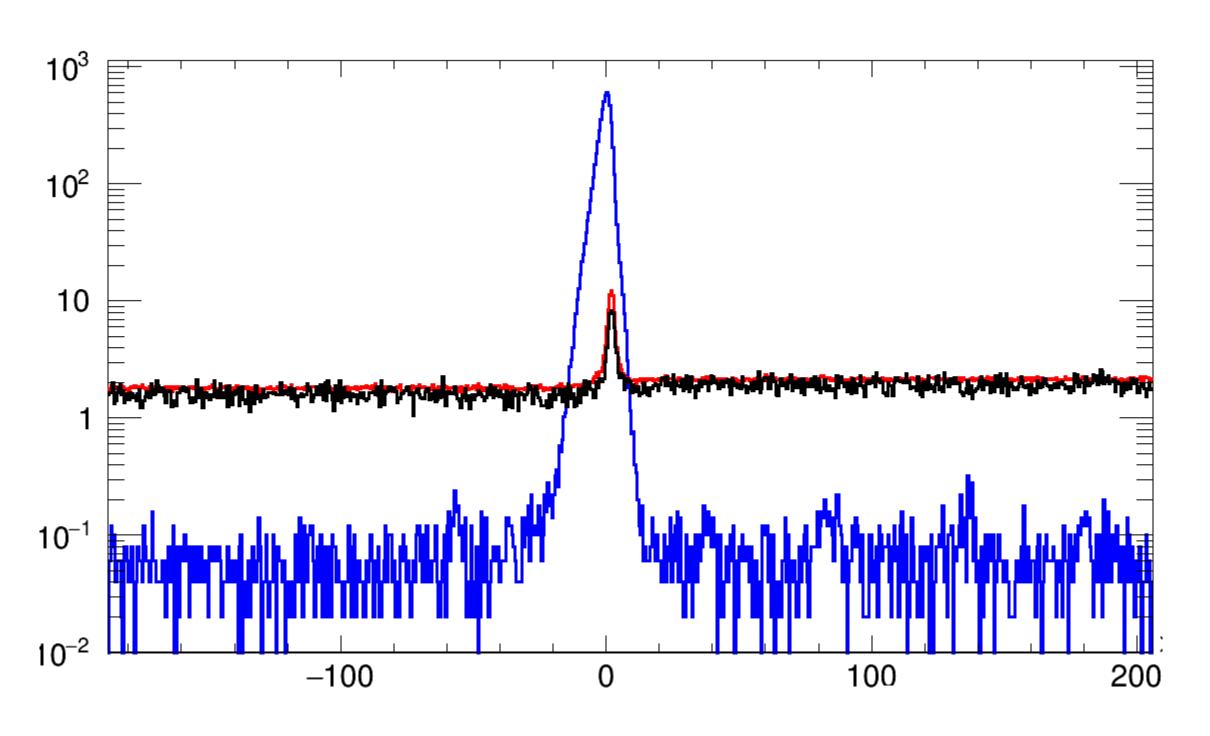
## The Giessen Cosmic Station(GCS), miniGCS Configuration for photon detector tests



#### Radiator comparison







Hodoscope Box time – miniGCS Time ns

Left side is for Hamamatsu, right is for KETEK

**Blue is from Scintillator** 

**Red is from Aerogel** 

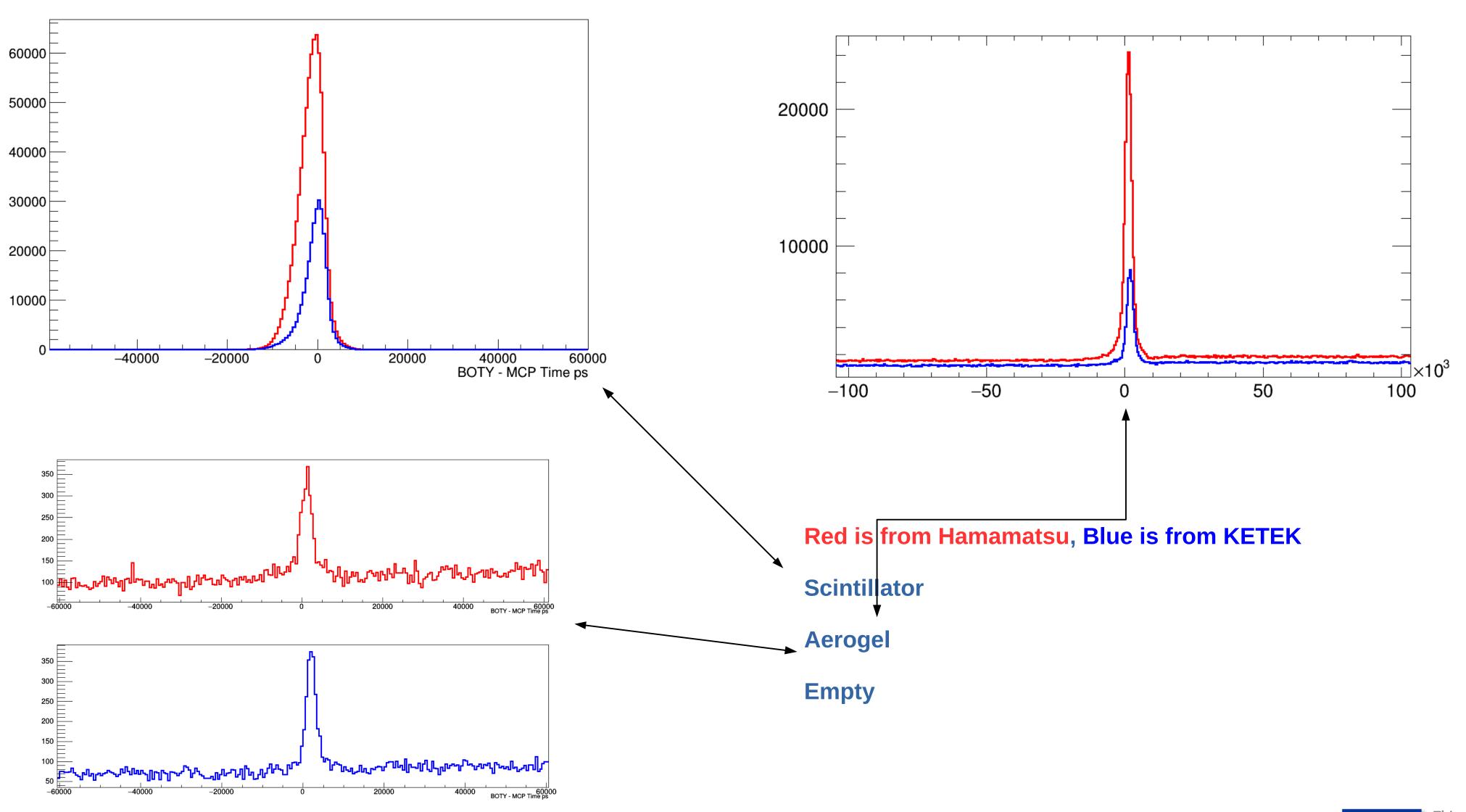
**Black is for Empty Runs** 

**Normalized for 30 min RUNS** 



## Detector comparison

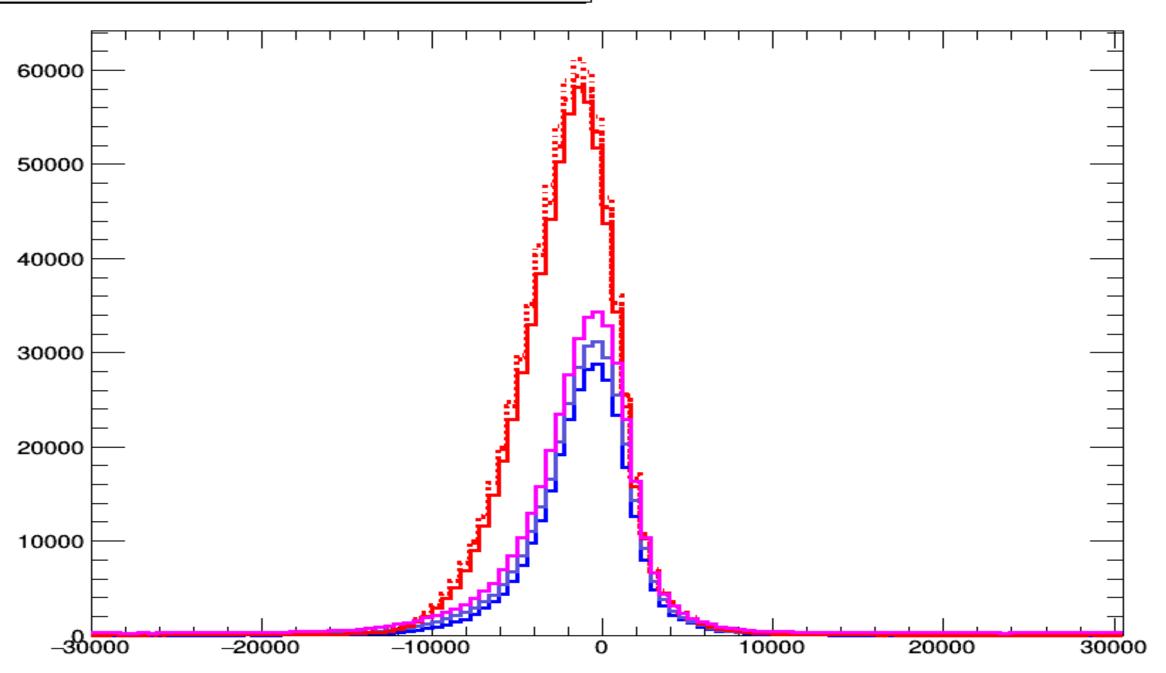




#### Comparisons



#### miniGCS TOPX TopXTrigger



Hamamatsu 57.6 Volt 3 Runs Overvoltage 3Volt Thodoscope-Tmini ps 31.5, overvoltage 3Volt 32.0, 3.5 32.5, 4 Volt

The difference between KETEK and Hamamatsu Cannot be explained by overvoltage

And here are the numbers for Radiator, Signal/Noise, Hamamatsu/KETEK

#### Aerogel

Hamamatsu Signal/Noise = 4.58 KETEK Signal/Noise = 1.83

HamamatsuSignal/KETEKSignal = 3.25

#### Scintillator

Hamamatsu Signal/Noise = 1468.7 KETEK Signal/Noise = 1440.4 HamamatsuSignal/KETEKSignal = 2.43

"Empty" Target

Hamamatsu Signal/Noise = 0.886 KETEK Signal/Noise = 1.26 HamamatsuSignal/KETEKSignal = 0.986

The good conclusion is that both matrices are seeing Cherenkov photons at room temperature (20°C)

Based on this numbers and the prices one can say in football language: after halftime the result is:

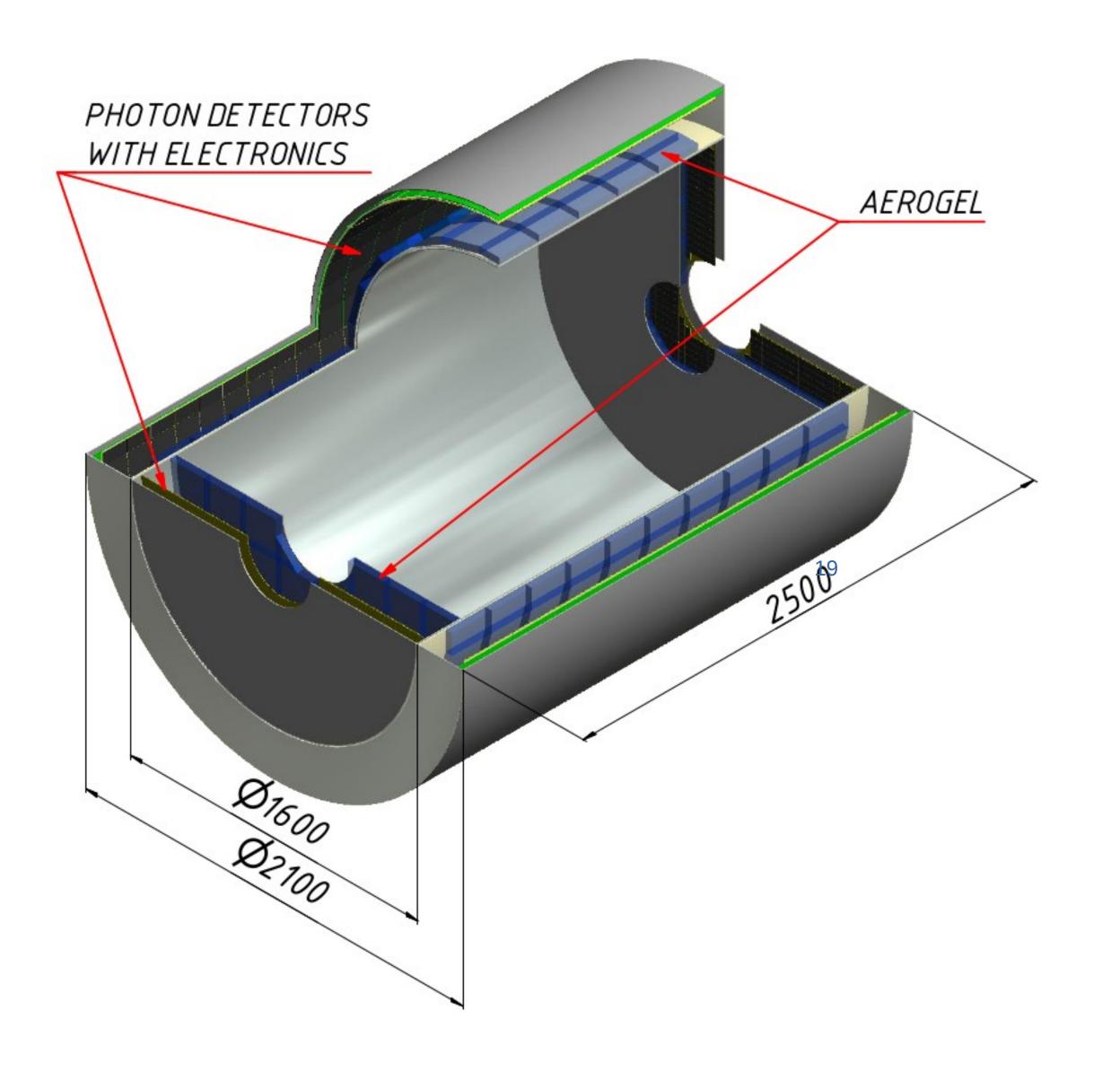
Hamamatsu KETEK 1



<sup>\*</sup>The stat error is very small

#### The FARICH option





Compact, hermetic design

Endcaps covering 10-40 and 140-180 polar angles

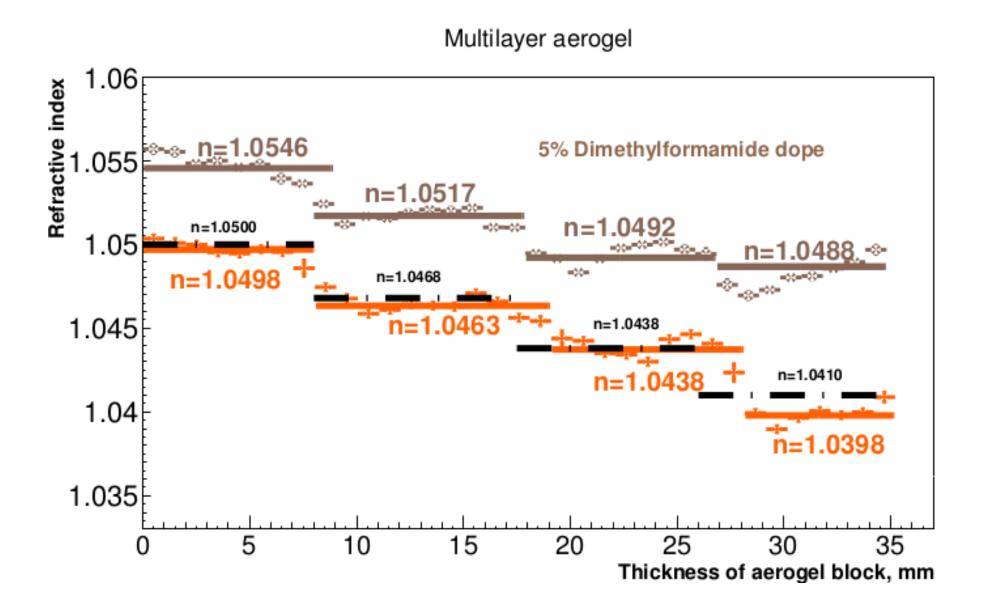
Barrel should cover 40-140 degree

4 layer of aerogel with different n for focusing

Fine pixelation for photon detection

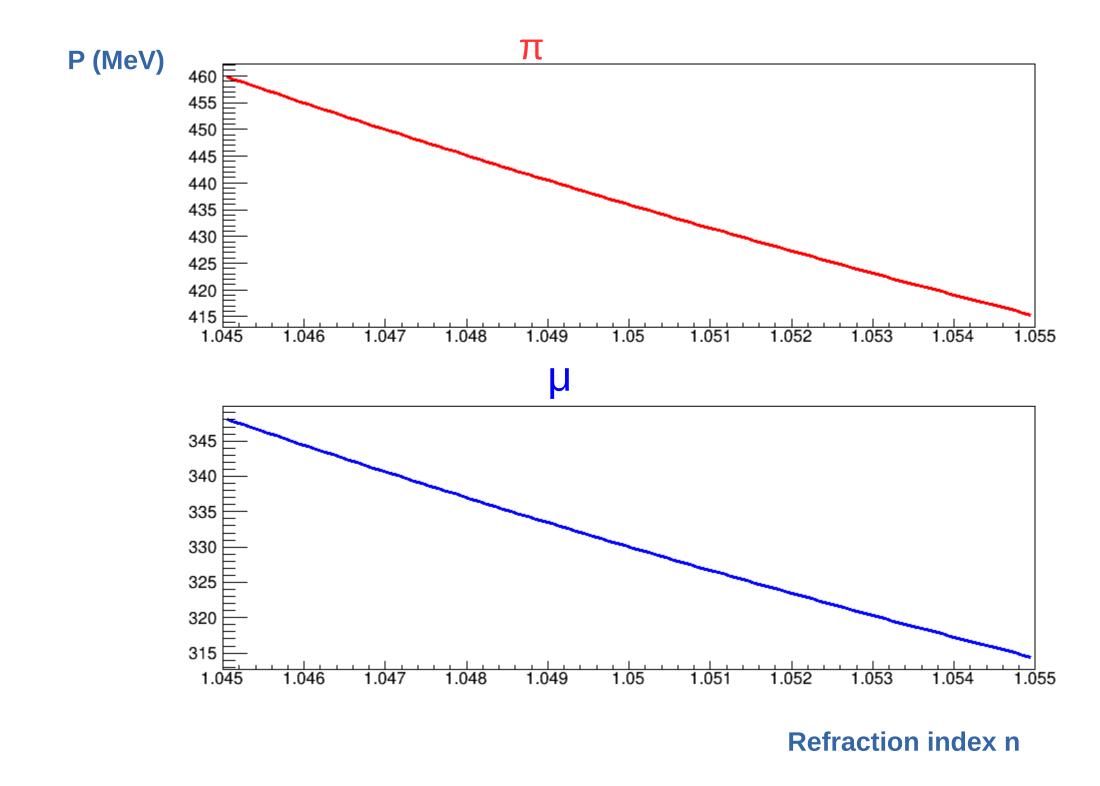
Ability to operate inside magnet

## The "entry point" of FARICH, the 4layer aerogel radiator



Most likely this configuration will be built in a new prototype

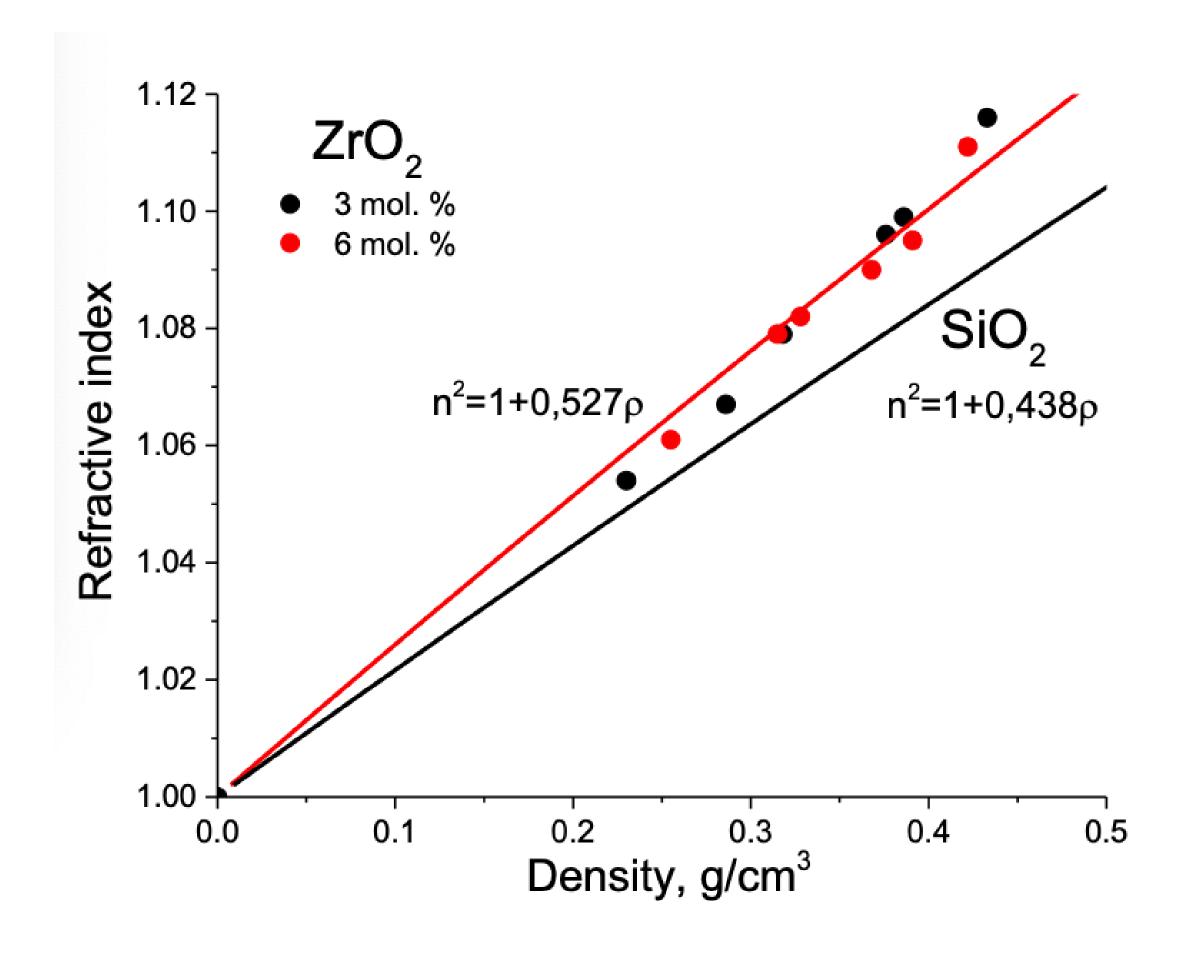
But the aerogel with n= $\sim$ 1.05 makes the threshold for  $\mu \pi$  higher



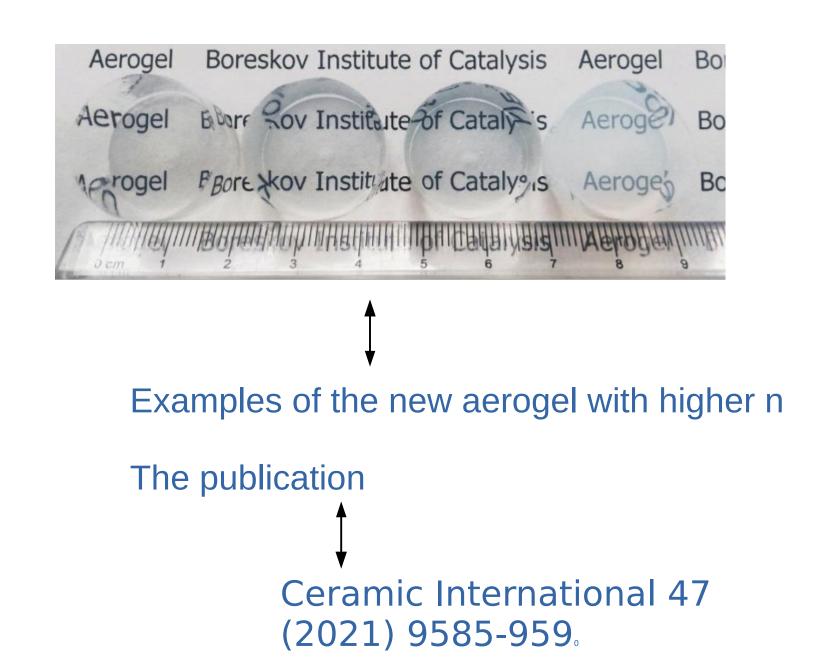
For the final detector RD for aerogel is also underway

#### The "entry point" of FARICH, the 4layer aerogel radiator





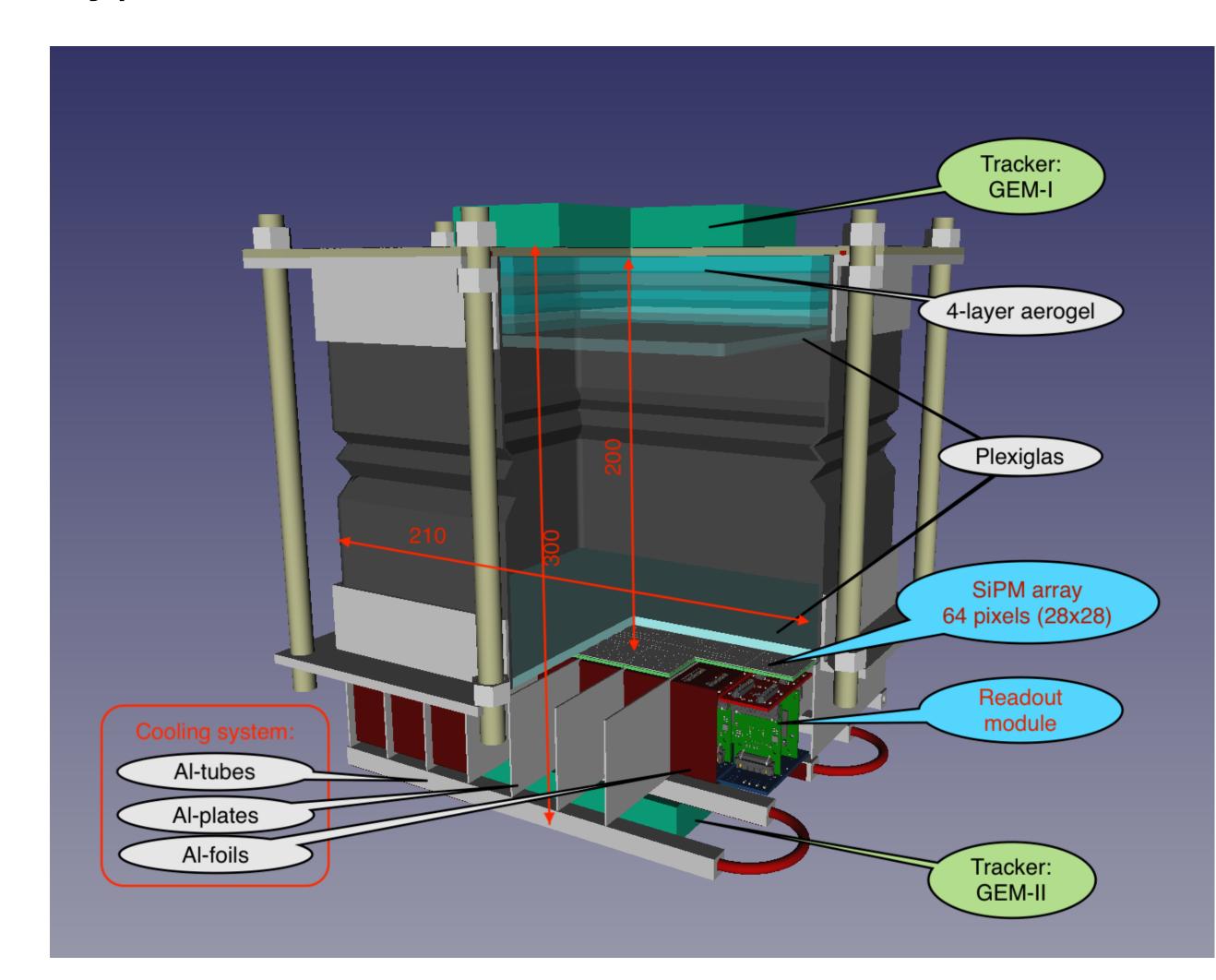
the aerogel with n= $\sim$ 1.12 will make threshold for  $\mu$   $\pi$  significantly lower



## The FARICH complete prototype under constructtion



- photon detector area ~21X21cm, almost the same size as in "proof of principle" setup at CERN
- 4 layer of aerogel 10X10cm with focusing
- Simultaneous tests with different photon detectors(KETEK, Hamamatsu,SensL)
- This version assumes PETSYS ToFPET readout
- Compact isolated environment allowing cool-down of sensitive elements



#### The FARICH complete prototype under constructtion



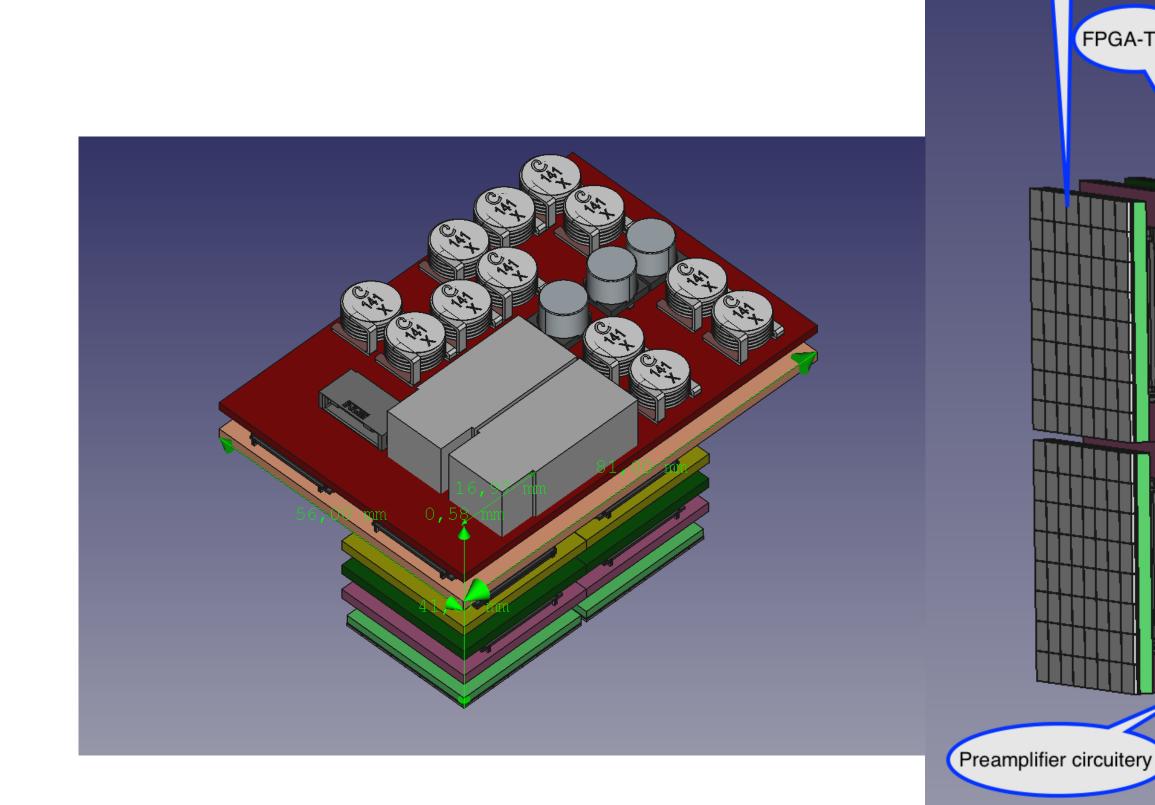
Sensor (KETEK:SiPM)

FPGA-TDC

PWR distr. board

DC/DC board

- photon detector area ~21X21cm, almost the same size as in "proof of principle" setup at CERN
- 4 layer of aerogel 10X10cm with focusing
- Simultaneous tests with different photon detectors(KETEK, Hamamatsu, SensL)
- In this version the FEE is GSI development with data transfer via optical link
- Compact isolated environment allowing cool-down of sensitive elements

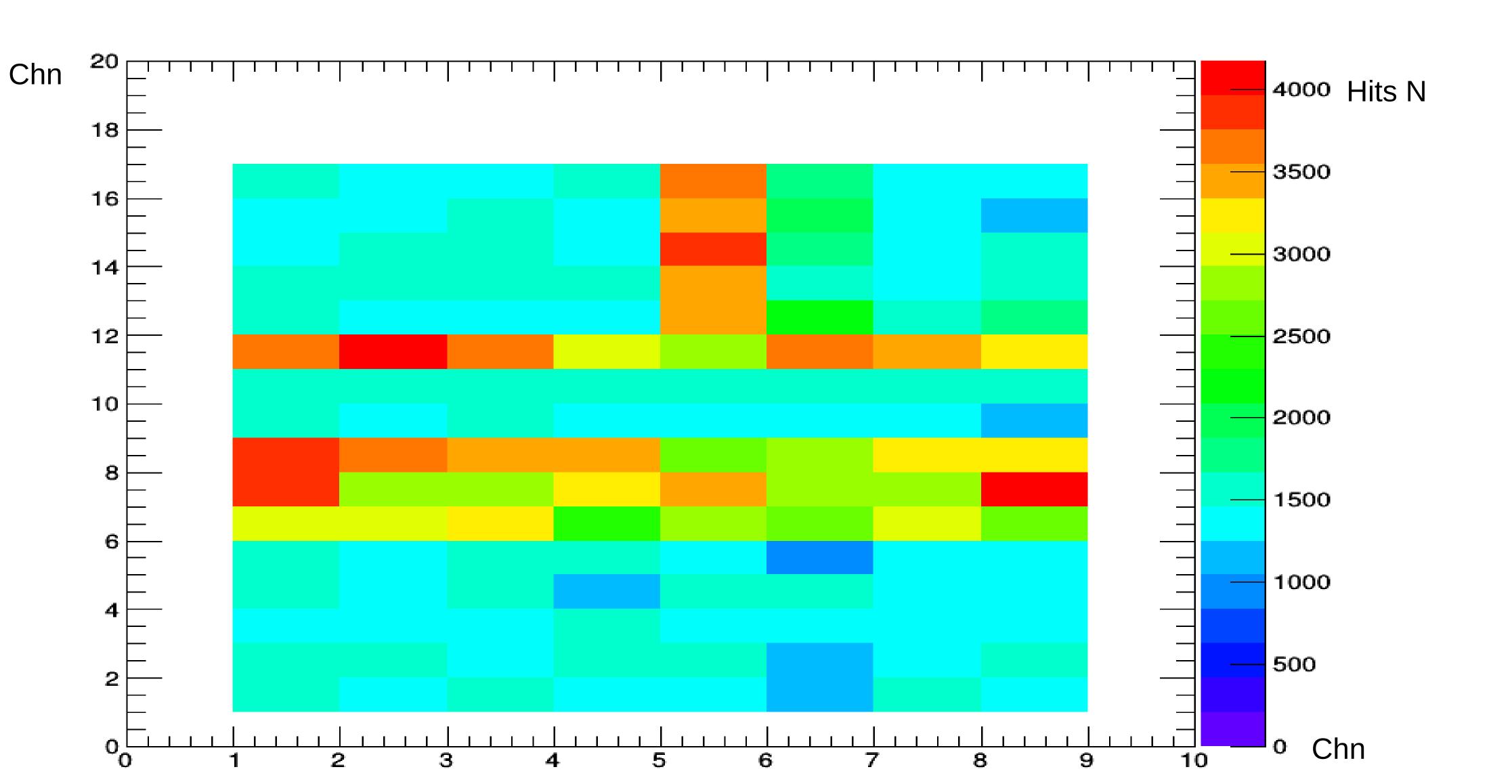




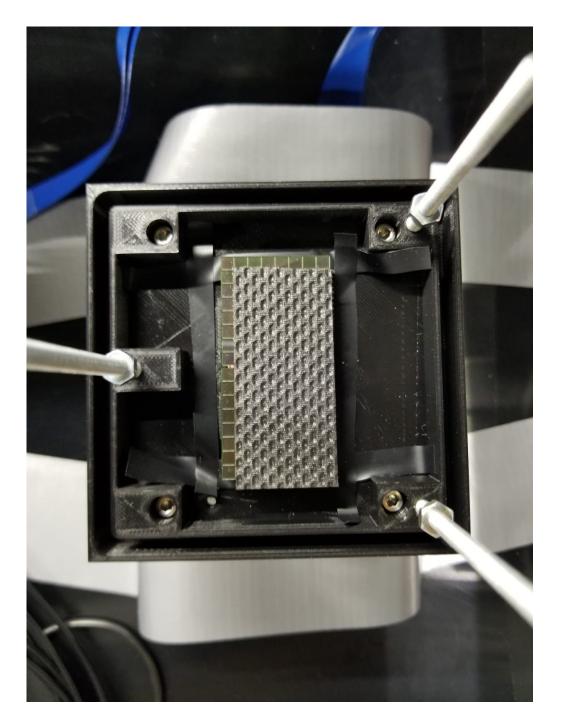
## Conclusions

- The models of FARICH and fDIRC with certain novelties are build in GEANT system and optimization are ongoing, focusing optics, photon detector pixelation and many more
- Small prototypes already exist and are under tests in beam-line (FARICH) and using cosmic muons (fDIRC)
- The GCS first runs allowed to get first comparison points for a candidate photon detectors, analysis ongoing to get time resolutions
- There are ongoing efforts on both sides to get a full prototype to be built and test
- Stay healthy

# The "Mask" is reproducible mapping is OK



## Photon detector candidate



KETEK 3X3 mm 8X8 two SiPM matrices are a good candidate for Photon detector
Already first checks was done to check their mapping

