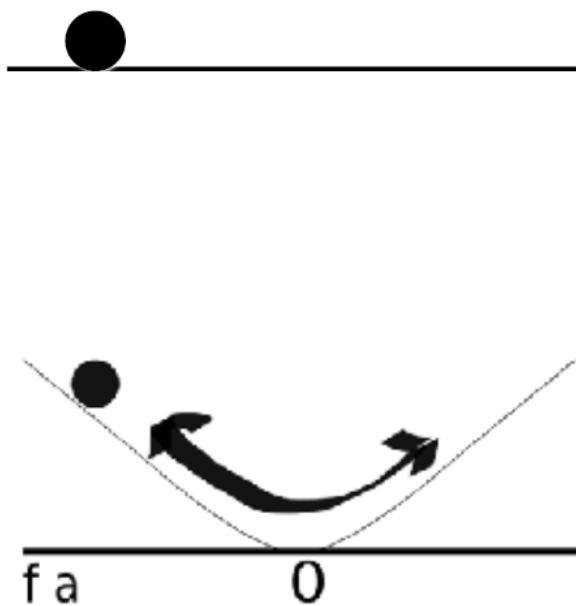


# Axion/ALP search program introduction

APPEC Direct Dark Matter Detection Report  
Community Feedback Meeting  
February 2, 2021

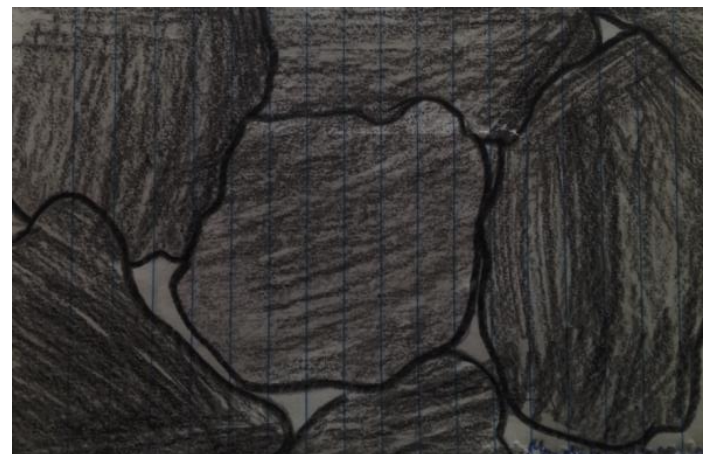
Béla Majorovits  
Max-Planck-Institut für Physik, München



Laura, physicist



Anna, 13



Ella, 11

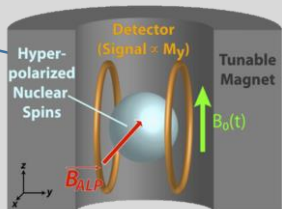


Emöke, 6

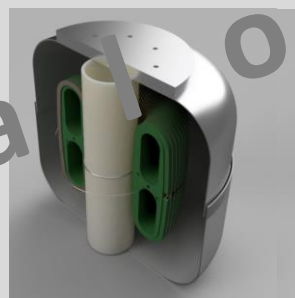
# Projects with (expected) axion sensitivity

Exciting last ~5 – 10 years:  
plethora of approaches emerging  
**VERY COMPLEMENTARY!**

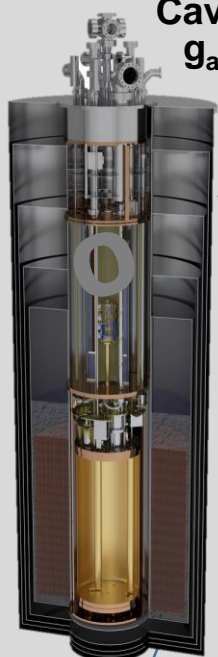
NMR / Spin-precession  
 $g_{aN}, g_{aEDM}$



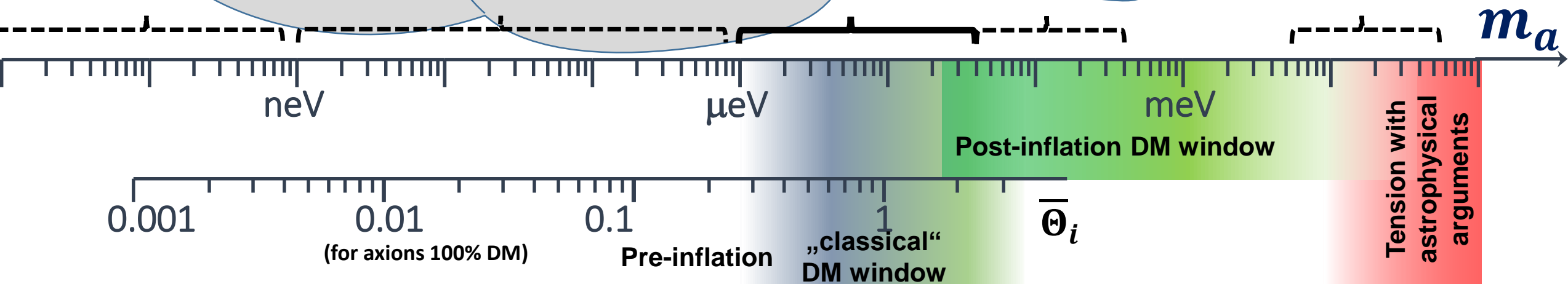
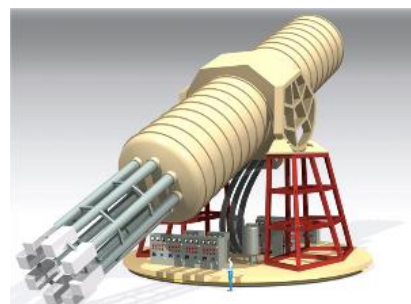
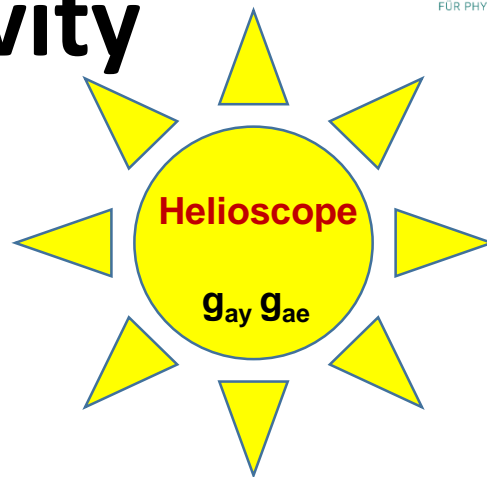
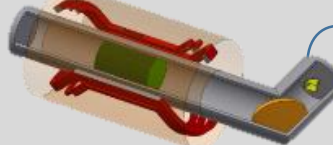
LC circuit  $g_{ay}$



Cavity  $g_{ay}$



Di-electric haloscope  $g_{ay}$



# Axion experiments: next ~decade prospect:

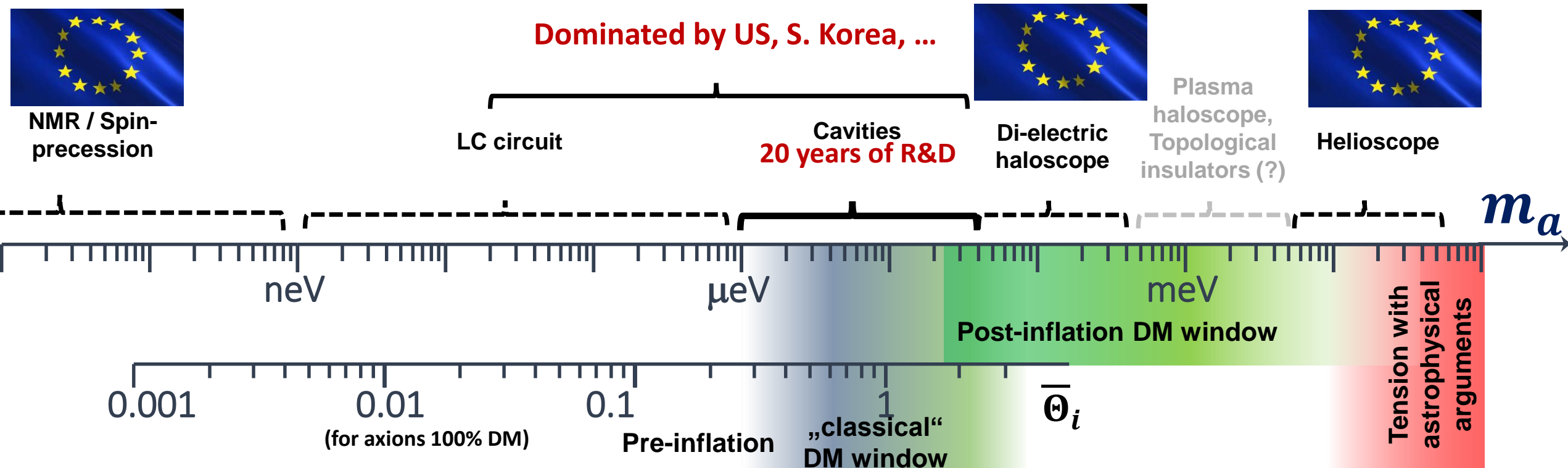
**Close the gap: haloscopes + helioscopes**

cover whole axion mass range compatible with dark matter!

**~ 10 orders of magnitude!**

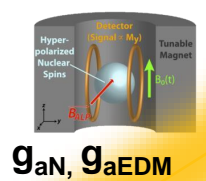
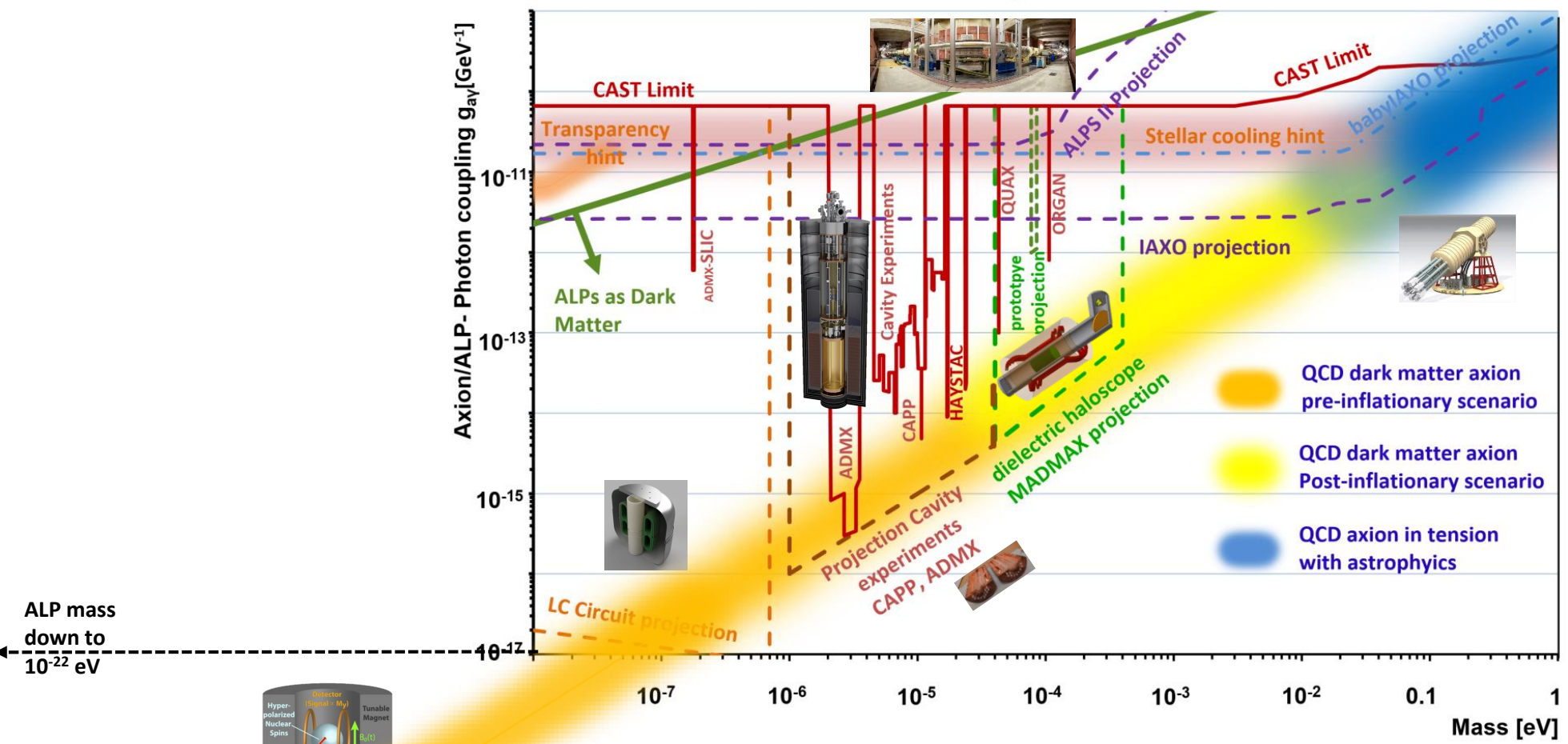


**Find the DM axion!**



# Axion & ALP experiments

## Axion search prototypes: very good ALP searches

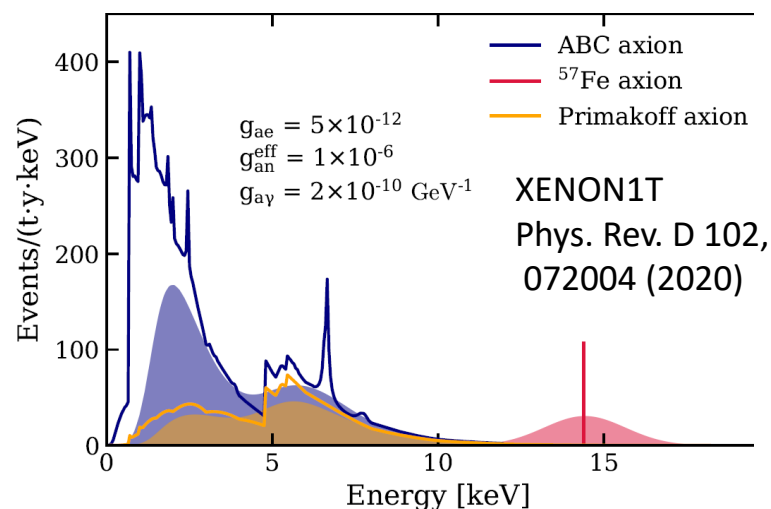


**Vast parameter range compatible with ALP DM**

**Some possibilities for "smaller scale" experiments**

# Sensitivity of "low background" experiments

## Solar axions - ALPs



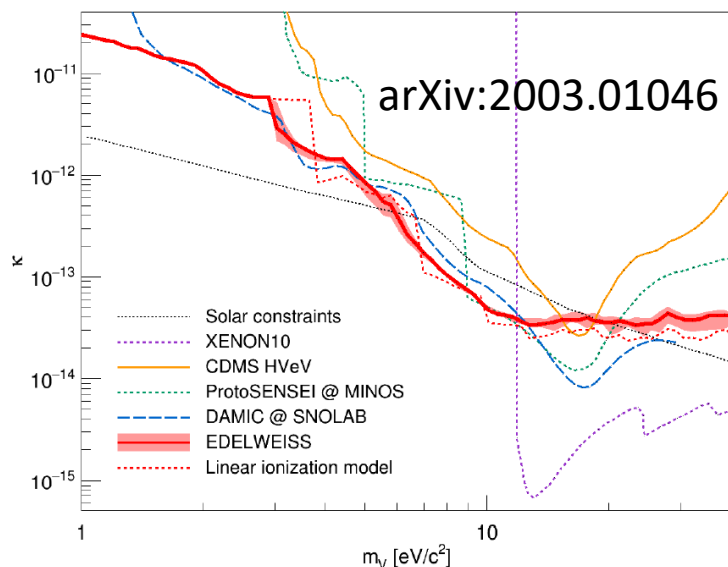
**Sensitivity "for free" to (solar) ALPs, Hidden Photon or Vector Boson dark matter with "high masses"**

**No QCD axion sensitivity**

compatible with astrophysical observations

But: **Independent limits for axions from sun!**

## Hidden Photons



**Some ALP sensitivity:**

Some "fine tuning needed" for  $g_{ae}$  vs  $g_{ay}$

**Hidden Photons/Vector bosons:**

sensitive to parameter range not yet excluded

DM & neutrino experiments: **sub eV thresholds!**

# S W O T table axion, ALP and HP experiments

	<b>S</b> trengths	<b>W</b> eaknesses	<b>O</b> pportunities	<b>T</b> hreats
<b>haloscopes</b>	can cover "classical" mass region in next decade(s)	DM abundance relies on cosmological models	profit from R&D on quantum computing	what if axions sub dominant DM contribution?
	partly already sensitive to QCD axion	high field magnets needed	straight forward to adapt to precise frequency	potentially suppressed signal due to axion miniclusters
	precise mass measurement		possibility of transients	
	easy to switch off signal			
<b>helioscopes</b>	all required technologies available	signal not related to DM halo density	could carry payload experiment	non DM related signal could deviate efforts
	easy to switch off signal			
	sensitivity to $g_{ae}$ and $g_{aN}$			
<b>LSW</b>	no model dependence of ALP source	QCD axion not accessible	can soon cover astrophysical hints	
		not related to DM halo densities	R&D synergies with other fields	
<b>Low background</b>	comes for free	not sensitive to QCD axion and large parts of ALP parameter range	exploitation of some new parameter range	
	can probe $g_{ae}$	not necessarily related to DM halo densities		

Axions ALP HPS

Experiment	Type	Techn.	$g_-$	Mass range	Status	Limits	Location	Timescale
<b>Experiments with expected sensitivity to DM axion benchmark models</b>								
CASPER-e <sup>a</sup>	∅	NMR	$a_N$	$10^{-13}$ eV – 1 neV	R&D	ALP	BU	
DM Radio <sup>b</sup>	∅	LC	$a_\gamma$	20 neV – 0.8 $\mu$ eV	R&D	HP	Stanford	2025-30
ADMX <sup>c</sup>	∅	C	$a_\gamma$	2 $\mu$ eV – 40 $\mu$ eV	running	<b>axion</b> <sup>†</sup>	UW	2017-30
HAYSTAC	∅	CS	$a_\gamma$	15 $\mu$ eV – 35 $\mu$ eV	running	axion <sup>‡</sup>	Yale	2015-25
CULTASK	∅	SC/MC	$a_\gamma$	3 $\mu$ eV – 70 $\mu$ eV	running	axion <sup>*</sup>	CAPP	2021-30
QUAX <sup>d</sup>	∅	SC/DC	$a_\gamma$	30 $\mu$ eV – 50 $\mu$ eV	in prep.	ALP <sup>*</sup>	INFN	2021-25
MADMAX <sup>e</sup>	∅	DH	$a_\gamma$	40 $\mu$ eV – 400 $\mu$ eV	prototype		DESY	2025-35 <sup>f</sup>
ORGAN <sup>d</sup>	∅	DC/CS	$a_\gamma$	60 $\mu$ eV – 210 $\mu$ eV	prototype	ALP	UWA	2025-35 <sup>f</sup>
IAXO <sup>g</sup>	⊙	XR	$a_{\gamma,ae}$	1 meV – 10 eV	in prep.		DESY	2023-35

ALP HPS

<b>ALP experiments</b>								
CASPER-w <sup>a</sup>	∅	NMR	ALPN	$10^{-22}$ eV – 1 $\mu$ eV	running	ALP	HIM/UCB	
GNOME	∅	NMR	ALPN	$10^{-21}$ eV – $10^{-10}$ eV	running	ALP	global	2017-24
DANCE	∅	OC	ALP $\gamma$	$\lesssim 10^{-10}$ eV	R&D	ALP	Tokyo	
Up/Download	∅	MO	ALP $\gamma$	$10^{-10}$ eV – $10^{-7}$ eV	prototype	ALP	UWA	
ABRA <sup>b</sup>	∅	LC	ALP $\gamma$	1 neV – $\mu$ eV	in prep.	ALP	MIT	
SHAFT	∅	LC	ALP $\gamma$	$\lesssim 10$ neV	R&D	ALP	BU	
ADMX-SLIC	∅	LC	ALP $\gamma$	$\lesssim 0.2$ $\mu$ eV	R&D	ALP	UFL	
ALPS II	$\mathcal{L}$	LSW	ALP $\gamma$	$\lesssim 0.1$ meV	constr.		DESY	2021
RADES	∅	MC	ALP $\gamma$	$\sim 30 - 50$ $\mu$ eV	R&D		CERN	
QUAX	∅	$e^-S$	ALP $e$	30 $\mu$ eV – 80 $\mu$ eV	R&D	ALP	INFN	2021-25
BRASS	∅	DA	ALP $\gamma$	1 $\mu$ eV – 1000 $\mu$ eV	in prep.		UH	2022-23
IAXO <sup>g</sup>	⊙	XR	ALP $\gamma$	$\lesssim 1$ eV	in prep.		DESY	2025-35

HPS

<b>Hidden photon experiments (no axion or ALP coupling)</b>								
SHUKET	∅	DA	$\epsilon$	20 $\mu$ eV – 30 $\mu$ eV	in prep.	HP	CEA	2024
FUNK	∅	DA	$\epsilon$	2 eV – 8 eV	upgrade	HP	KIT	

+ low background experiments

## Recommendation 6: Axion/ALP experiments



**European-led** efforts should

**focus on axion and ALPs mass ranges** that are **complementary to the established cavity approach**

and this is where **European teams** have

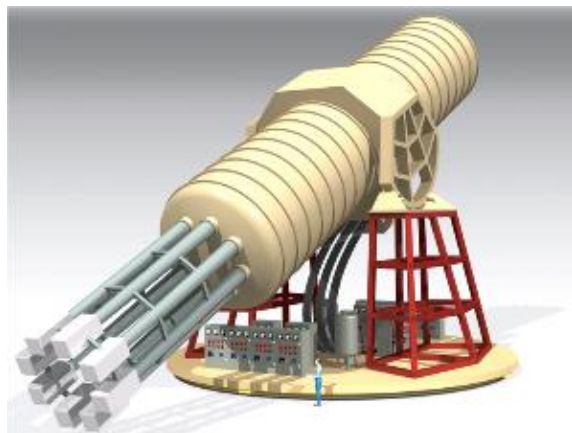


a **unique opportunity to secure the pioneering role**

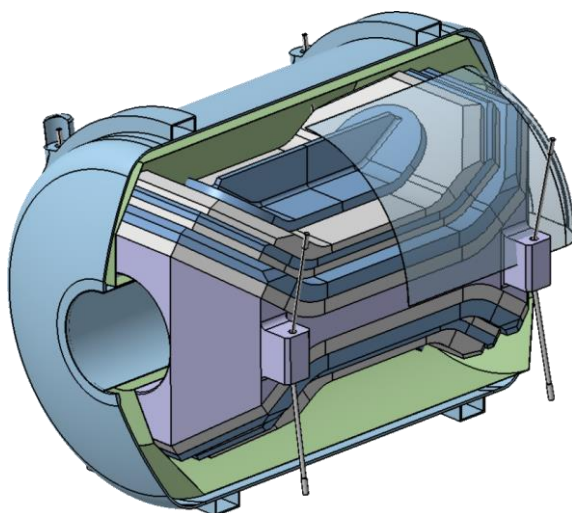
in achieving sensitivities **in axion/ALP mass ranges** not yet explored by experiments conducted elsewhere.

In parallel, **R&D efforts to improve experimental sensitivity** and to extend the accessible mass ranges should be supported.





IAXO magnet: based on existing CERN technology from accelerator magnets



MADMAX dipole magnet: being developed in terms of EU-innovation partnership between MPP, Bilfinger-NOELL and CEA-Irfu Saclay



# Technologies

## Magnets

Development of **strong (dipole) magnets** paramount for reaching goal of **covering axion DM mass range**

### Main challenges:

- **Huge superconducting dipole magnets**  
**large stored energy**  
(helioscopes, dielectric haloscopes, dish antenna)  
→ **Infrastructure**
- **Solenoids (cavity experiments)**

### Potential synergies:

- **accelerator magnets,**
- **medical physics,**
- **fusion reactors,**
- **some aspects of solid state physics,**
- ...

# Technologies

## Photon detection

Improve detector sensitivity

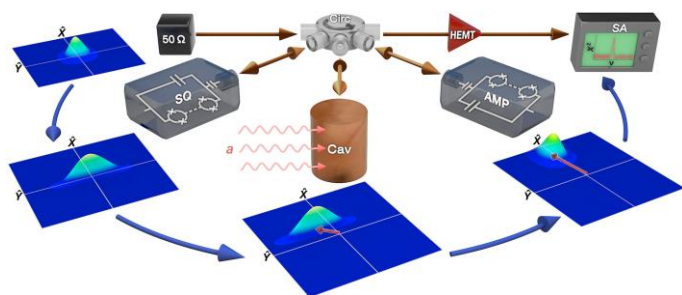
→ increases mass range that can be covered in given time

Main challenges:

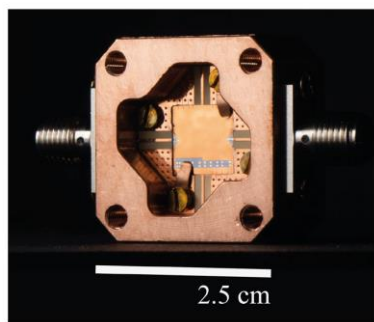
- **quantum detectors** (frequency < 100GHz):  
with broad bandwidth (JPA, TWPA)  
**or beat quantum limit by squeezing**
- **Single photon detectors** (frequency > 10 GHz):  
extremely low threshold, low background: (QMONs, TES)

Potential synergies:

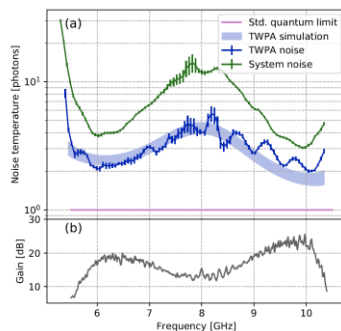
- **Quantum computing**
- **neutrino experiments & other light DM searches**
- ...



HAYSTAC: squeezed states arXiv:2008.01853



MADMAX TWPA arXiv:2101.05815



# Technologies

## Low loss dielectrics and cavities

Minimize RF losses in cavities/booster, etc.

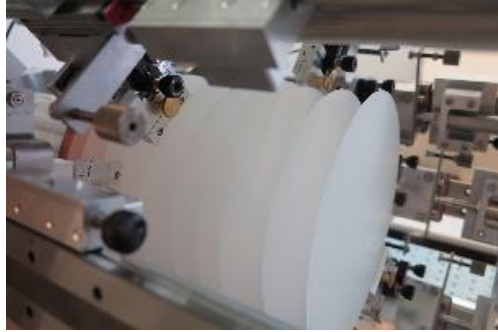
→ maximize signal

**Main challenges:**

- Grow large **single crystals** or produce **amorphous materials** low enough dielectric loss (tan delta)
- Develop suitable **superconducting cavities**,
- Develop **meta-materials** (plasma haloscope – topological insulators)

**Potential synergies:**

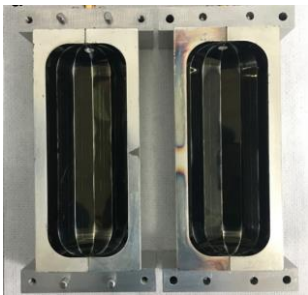
- **Radio astronomy**
- **Quantum computing**
- **RF engineering – telecommunication**
- ...



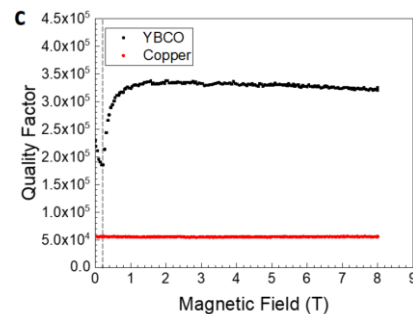
Low loss dielectrics for  
**MADMAX** arXiv:1901.07401



Photonic crystal **QUAX**  
arXiv:2002.01816



Superconducting cavities: **CAPP** arXiv:2002.08769



# Technologies

## Cryogenic precision engineering in B-field



Piezo linear stage developed for **MADMAX** by  
<https://www.jpe-innovations.com/>

**Precision displacement with large stroke at low temperature & strong B-field surrounding tune frequency , minimizing noise, maximize signal**

### Main challenges:

- Precision machining of large objects
- Precision displacement at extreme conditions
- Macroscopic stroke at cryogenic temperature
- Reliability

### Potential synergies

- **Cryo-electron microscopy**
- **Gravitational wave detectors**
- **Precision displacement technology in general**
- ...

Contact made by:



# Infrastructure

## for experiments with axion sensitivity

### Infrastructure for projects with QCD axion reach:

similar to particle physics – accelerator R&D needs:  
space in **well-equipped large (enough)**  
**experimental halls**  
with good **cryogenic infrastructure**

### Necessary boundary conditions:

- operation of **large aperture superconducting magnets**,
- operating **ultra-sensitive quantum detectors**,
- **minimize (thermal) RF noise & stable ground** in detector surrounding.



MORPURGO magnet at CERN



ALPSII at DESY

# Infrastructure for axion & ALP experiments

## Projects benefit from lab vicinity with world-leading knowledge:

- accelerator technology,
- superconducting magnet R&D,
- detector research,
- low loss RF-technologies,
- cryogenic engineering.

Very strong players in



....



+ top universities

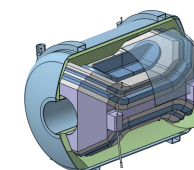
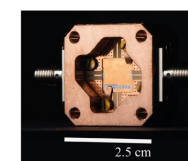
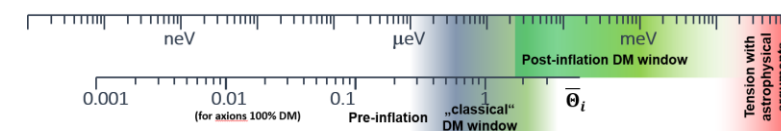
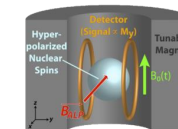
**Benefit of support** of large European and national labs demonstrated:  
e.g. by CAST & OSQAR experiments at CERN and ALPS at DESY

**Potential of fundamental discoveries with small(er) scale experiments**

Opportunities for research program complementary  
to large scale - for example accelerator - projects

# Concluding words:

- **Axions** naturally emerge as **excellent DM candidate** from natural solution to standard model strong CP problem
- Last decade: **new innovative initiatives - many within EU:**
  - possibility to **cover full mass range** by complementary efforts!
  - tentatively **high probability of success!**
- **Unique possibility for European community** to take leadership in covering significant part of well motivated mass range!
- Challenging **technologies to be developed** within axion community but also outside!
-  becoming **"axion hub"**?
-  has all prerequisites to find axion in **next ~ decade** and **solve the DM puzzle!**
- **Cross fertilization** between EU INDUSTRY & Scientific research!



**Lets take the opportunity, acquire the resources and DO IT ...**