



Quasi-Dirac neutrinos and the baryon asymmetry of the universe

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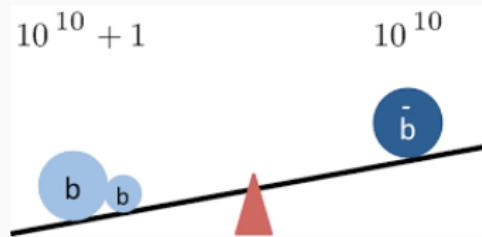
May 26, 2021

based on [C.S. Fong, T. Gregoire, AT, 2007.09158]

Motivations

Motivations

- Tiny neutrino masses and baryon asymmetry are two of the biggest puzzles in particle physics and both require the presence of Beyond the Standard Model physics



- Leptogenesis is an elegant mechanism that explains the baryon asymmetry in connection with neutrino masses
- In this work we propose a novel leptogenesis model with quasi-Dirac neutrinos that can be tested in oscillation experiments

Neutrino masses

Nature of neutrinos

Neutrinos can be:

- Majorana (m_L and/or $m_R \neq 0$)
- Dirac ($m_L = m_R = 0$)
- Quasi-Dirac ($m_L, m_R \ll m_D$)

$$\mathcal{L}_{\text{mass}} = -\frac{1}{2} \begin{pmatrix} \bar{\nu}_L & \bar{\nu}_R^c \end{pmatrix} \begin{pmatrix} m_L & m_D \\ m_D & m_R \end{pmatrix} \begin{pmatrix} \nu_L^c \\ \nu_R \end{pmatrix} + \text{h.c.}$$

Quasi-Dirac states are almost degenerate Majorana fermions

$$m_L, m_R \ll m_D \quad \Rightarrow \quad \begin{cases} m_{1,2} \simeq m_D(1 \pm \epsilon) \\ \nu_{1,2} \simeq \frac{1(i)}{\sqrt{2}} [(\pm 1 + \theta)\nu_L + (1 \mp \theta)\nu_R^c] \end{cases}$$

with $\epsilon = \frac{m_L + m_R}{2m_D} \ll 1$ and $\theta = \frac{m_L - m_R}{4m_D} \ll 1$ (in the limit $m_L, m_R \rightarrow 0$ one recovers the Dirac case)

Seesaw mechanisms with HNL

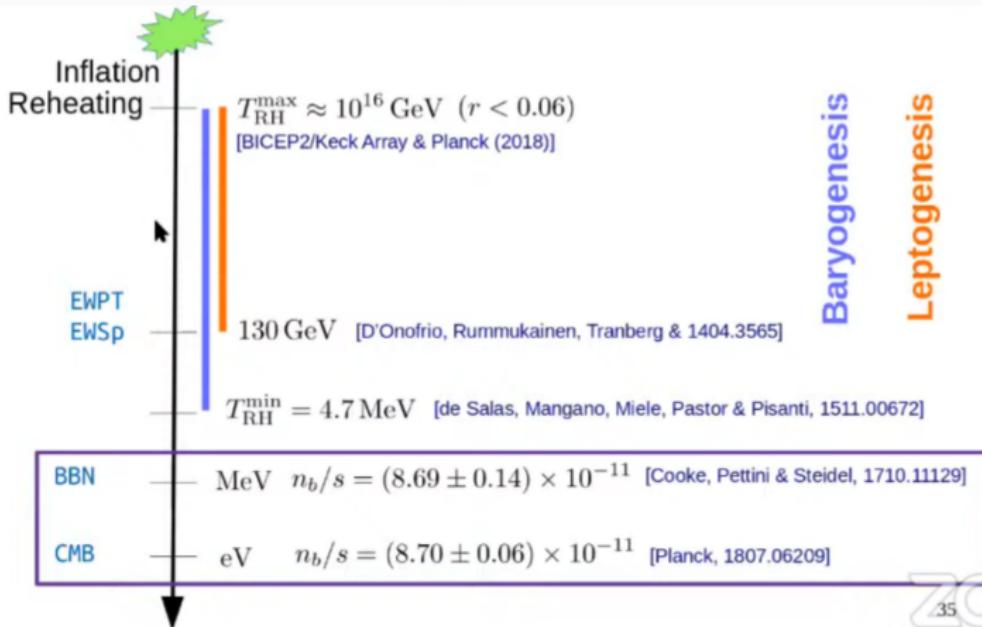
Elegant way to explain light neutrino (ν) masses due to the presence of Heavy Neutral Leptons (N)

Type	light ν	Heavy N	Content
Type I	Majorana	Majorana	ν_L, N_R
Inverse/Linear	Majorana	quasi-Dirac	ν_L, N_R, N'_R
Extended	quasi-Dirac	Majorana	ν_L, ν_R, N_R
Dirac	Dirac	Dirac	ν_L, ν_R, N_R, N'_R
quasi-Dirac	quasi-Dirac	quasi-Dirac	ν_L, ν_R, N_R, N'_R



Leptogenesis

Leptogenesis



The lepton asymmetry generated by the decay of Heavy Neutral Leptons (N) is converted into a baryon asymmetry by $B + L$ -violating sphalerons [M. Fukugita and T. Yanagida Phys. Lett. B 174 (1986) 45]

Quasi-Dirac leptogenesis

The model

[CS Fong, T. Grgoire, AT, 2007.09158]

Inspired by mirror world models (copy of SM)

- L-conserving lagrangian ($i, j = 1, \dots, N_f$)

$$\begin{aligned}\mathcal{L} = & i\bar{N}_{Ri}\not{\partial}N_{Ri} + i\bar{N}'_{Ri}\not{\partial}N'_{Ri} - \textcolor{blue}{M}_i\bar{N}_{Ri}N'^c_{Ri} + \text{h.c.} \\ & -\textcolor{blue}{y}_{\alpha i}\bar{l}_{L\alpha}\tilde{\Phi}N_{Ri} - \textcolor{blue}{y}'_{\alpha i}\bar{l}'_{L\alpha}\tilde{\Phi}'N'_{Ri} + \text{h.c.}\end{aligned}$$

- Small L-violating terms ($m, m' \ll M, \tilde{y} \ll y, \tilde{y}' \ll y'$)

$$\begin{aligned}\not{\mathcal{L}} = & -\frac{1}{2}\textcolor{red}{m}_{ij}\bar{N}^c_{Ri}N_{Rj} - \frac{1}{2}\textcolor{red}{m}'_{ij}\bar{N}'^c_{Ri}N'_{Rj} \\ & -\tilde{y}_{\alpha i}\bar{l}_{L\alpha}\tilde{\Phi}N'_{Ri} - \tilde{y}'_{\alpha i}\bar{l}'_{L\alpha}\tilde{\Phi}'N_{Ri} + \text{h.c.}\end{aligned}$$

In the limit $m, m', \tilde{y}, \tilde{y}' \rightarrow 0$ the total baryon minus lepton number

$$\Delta_{\text{tot}} = (B - L) - (B' - L')$$

is a good symmetry of the model

$N_f = 1$ case

We have a pair of Heavy Neutrinos N_R and N'_R . In the limit $m, m' \ll M$ we get the following mass eigenstate

$$\begin{pmatrix} N_1 \\ N_2 \end{pmatrix} = U^\dagger \begin{pmatrix} N_R \\ N'_R \end{pmatrix} \quad \begin{aligned} N_1 &\simeq (N_R + N'_R)/\sqrt{2} \\ N_2 &\simeq i(-N_R + N'_R)/\sqrt{2} \end{aligned}$$

with quasi-degenerate masses

$$M_{1,2}^2 \simeq M^2 \left(1 \mp \frac{|m' + m^*|}{M} \right)$$

The heavy singlet fermions split into quasi-Dirac pairs and CP violation in their decays can naturally be enhanced to realize resonant leptogenesis

N decays and CP violation

- Lagrangian ($Y_{\alpha i} = y_\alpha U_{1i} + \tilde{y}_\alpha U_{2i}$ and $Y'_{\alpha i} = \tilde{y}'_\alpha U_{1i} + y'_\alpha U_{2i}$)

$$-(\mathcal{L} + \mathcal{L}') \supset \frac{1}{2} M_i \bar{N}_i^c N_i + Y_{\alpha i} \bar{l}_{L\alpha} \tilde{\Phi} N_i + Y'_{\alpha i} \bar{l}'_{L\alpha} \tilde{\Phi}' N_i + \text{h.c.}$$

- Total decay width

$$\Gamma_i = \frac{M_i}{8\pi} [(Y^\dagger Y)_{ii} + (Y'^\dagger Y')_{ii}]$$

- CP violating parameters

$$\begin{aligned}\epsilon_i &\equiv \frac{\sum_\alpha \Gamma(N_i \rightarrow l_\alpha \Phi) - \Gamma(N_i \rightarrow \bar{l}_\alpha \bar{\Phi})}{\Gamma_i} \\ \epsilon'_i &\equiv \frac{\sum_\alpha \Gamma(N_i \rightarrow l'_\alpha \Phi') - \Gamma(N_i \rightarrow \bar{l}'_\alpha \bar{\Phi}')}{\Gamma_i}\end{aligned}$$

Resonant enhancement

- Resonant enhancement occurs when the mass splitting is of the order of the decay width [*Pilaftsis and Underwood 0309342*]

$$|M_1 - M_2| \simeq |m' + m^*| \simeq \frac{\Gamma}{2}$$

- We consider Z_{2D} symmetric limit

$$m = m' \quad y = y' \quad \tilde{y} = \tilde{y}'$$

- Maximum CP parameters ($y^2 \equiv \sum_{\alpha} y_{\alpha}^2$ and $y'^2 \equiv \sum_{\alpha} y'_{\alpha}^2$ and $w \equiv \sum_{\alpha} y_{\alpha} \tilde{y}_{\alpha}$)

$$|\epsilon_1^{\max}| = |\epsilon_1'^{\max}| \quad \simeq \quad \frac{|\text{Re } w|}{y^2} \quad \sim \frac{\tilde{y}}{y}$$

- Final baryon asymmetry

$$Y_B \propto \epsilon_1$$

Quasi-Dirac seesaw mechanism

- Integrating out the Heavy N s we get the following mass matrix in the basis (ν_L, ν'_L)

$$M_\nu = -m_D \hat{M}^{-1} m_D^T \quad m_D \equiv \begin{pmatrix} vY \\ fY' \end{pmatrix}$$

- Diagonalization

$$U_\nu^T M_\nu U_\nu = \text{diag}(m_-, m_+, 0, 0, 0, 0)$$

- Non-zero quasi-Dirac mass eigenstates

$$m_{\mp} = m_\nu \mp \delta m$$

with $\delta m \ll m_\nu$

$$m_\nu \equiv \frac{yy'vf}{M} \quad \delta m \simeq \left| \frac{w^*v^2 + w'f^2}{M} \right|$$

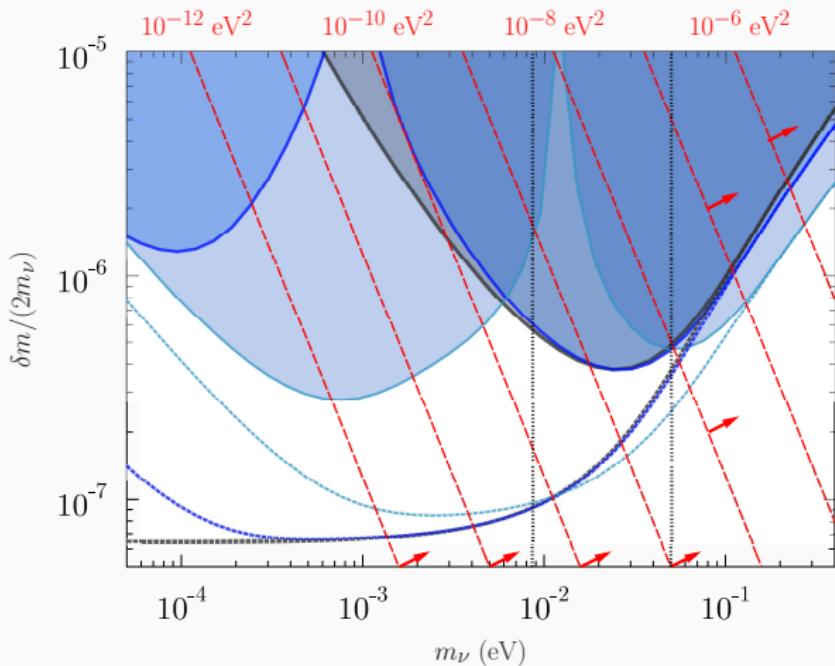
- In the Z_{2D} symmetric limit we have

$$\delta m = 2|\text{Re } w|v^2/M = 2m_\nu|\text{Re } w|/y^2$$

Light neutrinos are quasi-Dirac and there exist an intriguing relation between the CP parameter and light neutrino mass splitting

$$|\epsilon_1^{\max}| \simeq \frac{\delta m}{2m_\nu}$$

Results Z_{2D} symmetric limit



$M \gg 1$ TeV (gray), $M = 1$ TeV (blue) and $M = 500$ GeV (light blue).

Dashed red lines indicate isocurves of constant $m_+^2 - m_-^2 \equiv \epsilon^2 = 4m_\nu\delta m$

Quasi-Dirac oscillations

- Phenomenological study [Anamiati, Fonseca, Hirsch 1710.06249]

$$m_{i\mp}^2 = m_i^2 \mp \frac{\varepsilon^2}{2}$$

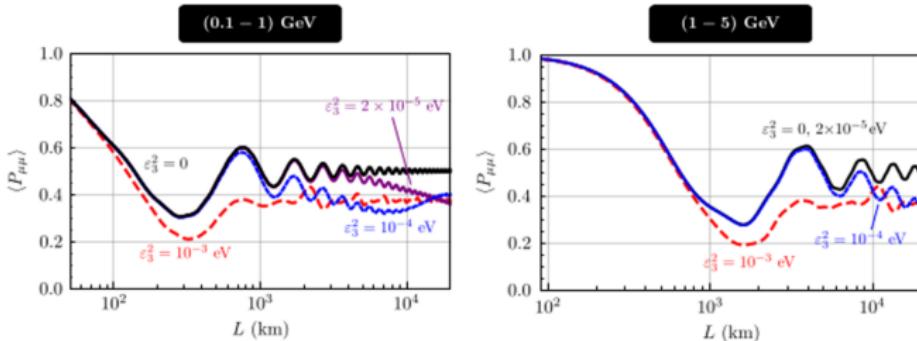


Figure 2: Averaged atmospheric muon neutrino survival probability for neutrinos with energies $E_\nu = (0.1 - 1)$ GeV (left) and $E_\nu = (1 - 5)$ GeV (right), as a function of distance (L), for different choices of ε_3^2 . Lower neutrino energies are more sensitive to small ε_3 values. This plot is calculated with the simplifying assumptions of $\sin^2 \theta_{23} = 1/2$, $\theta_{13} = 0$ and $\Delta m_{\odot}^2 = 0$.

Experiment	ε_1^2 [eV ²]	ε_2^2 [eV ²]	ε_3^2 [eV ²]
KamLAND	$7.7(3.4) \times 10^{-6}$	$1.7(1.0) \times 10^{-5}$	—
Solar + KamLAND	$1.7(1.3) \times 10^{-11}$	$1.7(1.5) \times 10^{-11}$	—
DayaBay + MINOS + T2K	—	$1.5(0.9) \times 10^{-4}$	$1.3(0.074) \times 10^{-3}$
Super-K + DayaBay + MINOS + T2K	—	$1.9(1.8) \times 10^{-5}$	$1.2(1.1) \times 10^{-5}$
JUNO	$1.7(0.07) \times 10^{-5}$	$2.3(0.09) \times 10^{-5}$	$6.0(2.2) \times 10^{-5}$

Conclusions

- We proposed a novel model with small $B - L$ breaking where Heavy Neutrinos as well as light neutrinos are split into quasi-Dirac pairs
- This implies a natural realization of resonant leptogenesis
- The parameter space for viable leptogenesis spans over the neutrino mass squared difference in the range $10^{-12} - 10^{-6}$ eV² which can be probed in solar and atmospheric neutrino oscillation experiments
- From minimality considerations, observation of small mass splitting would strongly suggest that neutrinos are indeed Majorana particles
- More realistic study with $N_f = 2, 3$ needed (in progress)



Thank you

BACK UP

Baryogenesis

Sakarov conditions to dynamically generate the baryon asymmetry

- Baryon number violation
- C and CP violation
- Out of equilibrium dynamics

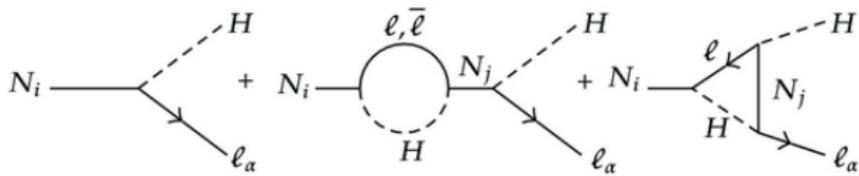
All three ingredients are present in the SM but no successful mechanism has been found within the SM \Rightarrow NP is needed

Minimal leptogenesis model I

- Type I seesaw lagrangian

$$\mathcal{L}_{\text{NP}} = -\frac{1}{2} M_i \bar{N}_{Ri}^c N_{Ri} - y_{\alpha i} \bar{l}_{L\alpha} \tilde{\Phi} N_{Ri} + \text{h.c.}$$

- L-violating and CP-violating decay of heavy N



- CP-violating parameters

$$\epsilon_i \equiv \frac{\sum_\alpha \Gamma(N_i \rightarrow l_\alpha \Phi) - \Gamma(N_i \rightarrow \bar{l}_\alpha \bar{\Phi})}{\Gamma_{N_i}} \propto \sum_j \text{Im}[(y^\dagger y)_{ij}^2]$$

- Baryon asymmetry (in the limit $M_1 \ll M_{2,3}$)

$$Y_B \simeq -\frac{1}{3} \epsilon_1 \eta Y_{N_1}^{eq}$$

Minimal leptogenesis model II

- Casas-Ibarra parametrization

$$y_\nu = \frac{1}{v} U_\nu^* \sqrt{\hat{m}} R \sqrt{\hat{M}} \quad \Rightarrow \quad m_\nu = v^2 y_\nu M^{-1} y_\nu^T = U_\nu^* \hat{m} U_\nu^\dagger$$

- y_ν has 15 independent parameters, but **6 moduli + 3 phases** cannot be probed at low energy
- Davidson-Ibarra bound

$$\epsilon_1 \gtrsim 10^{-6} \quad \Rightarrow \quad M_1 \gtrsim 10^{9-10} \text{GeV}$$

- We can have only circumstantial or supporting evidences through L-violation in $0\nu\beta\beta$ decay and CP-violation in leptonic sector δ_{CP}

Dirac seesaw general idea

- i) initial condition (no asymmetry)

$$B = 0 \quad L = 0 \quad B' = 0 \quad L' = 0$$

- ii) hierarchical spectrum (leptogenesis dominated by N_1 decay)

$$M_1 \ll M_2, M_3$$

- iii) CPV decay of heavy Dirac N_i generates asymmetry in L and L'

$$B = 0 \quad L = w \quad B' = 0 \quad L' = w$$

- iv) sphalerons redistribute the asymmetry in $B - L$ and $B' - L'$

$$B = a \quad L = w + a \quad B' = a' \quad L' = w + a'$$

- v) all N_i decay, the two sectors decouple and sphaleron freeze out
 $\rightarrow B, L, B', L'$ separately conserved

Dirac seesaw neutrino masses

- Integrating out the heavy N_i we get

$$\mathcal{L}_{\text{eff}} = (y M^{-1} y'^T)_{\alpha\beta} \bar{l}_{L\alpha} \tilde{\Phi} \tilde{\Phi}'^T (l'_{L\beta})^c + \text{h.c.}$$

- after EWSB in both sectors

$$\nu_R = (\nu'_L)^c$$

$$\mathcal{L}_{\text{mass}} = (\mathcal{M}_\nu)_{ij} \bar{\nu}_{Li} \nu_{Rj} + \text{h.c.}$$

where (f is the Φ' vev)

$$\mathcal{M}_\nu = v f y M^{-1} y'^T$$

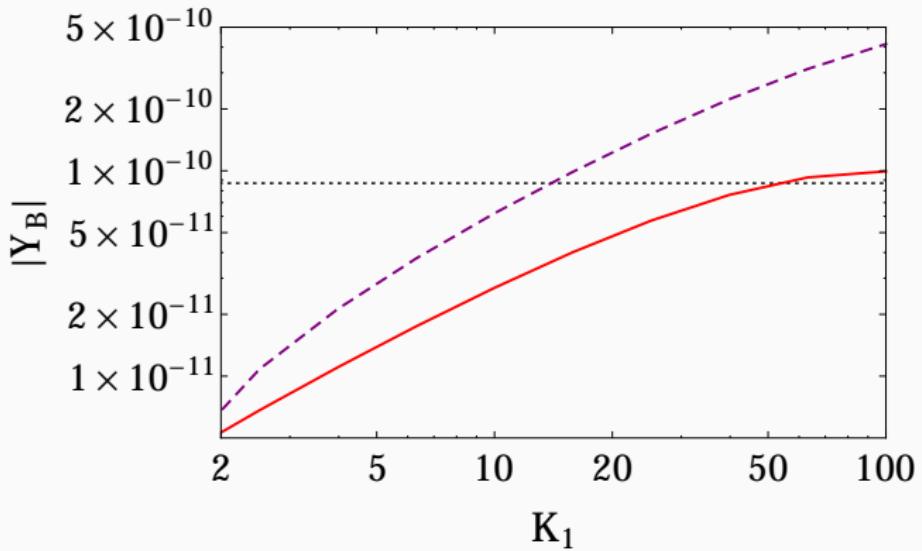
- DI bound on CP parameter

$$|\epsilon_1| \leq \frac{M_1(m_3 - m_1)}{16\pi} \frac{1}{vf} = \frac{M_1 |\Delta m_{\text{atm}}^2|}{16\pi(m_3 + m_1)} \frac{1}{vf} \equiv \epsilon_1^{\max}$$

Dirac leptogenesis

[K. Earl, CS Fong, T. Grgoire, AT, 1903.12192]

Baryon asymmetry as function of $K_1 \equiv \frac{\Gamma_{N_1}}{H(T=M_1)}$



zero initial abundance

$M_1 = 8 \times 10^8 \text{ GeV}$, $f = 500 \text{ GeV}$, $m_3 + m_1 = 0.1 \text{ eV}$, $|\epsilon'_{i\alpha}| \gg |\epsilon_{i\alpha}|$

thermal initial abundance

Other low energy tests

- LFV can be induced at one-loop level through loop of heavy quasi-Dirac fermions N , we have ($r \equiv y/y'$)

$$\text{Br}(\mu \rightarrow e\gamma) \approx 6 \times 10^{-26} r^2 \left(\frac{m_\nu}{0.1 \text{ eV}} \right)^2 \left(\frac{500 \text{ GeV}}{M} \right)^2$$

- Current experimental bound

$$\text{Br}(\mu \rightarrow e\gamma) < 4.2 \times 10^{-13}$$

- Neutrinoless beta decay rate is prop to

$$(M_\nu)_{ee} \simeq -\frac{2y_e \tilde{y}_e v^2}{M}$$

- $(M_\nu)_{ee}$ suppressed by $\tilde{y}_e \ll y_e$, not likely to be measured even in the next generation experiments which aim to probe $(M_\nu)_{ee} \sim 10 \text{ meV}$

CP parameters

Δ_{tot} -conserving terms $\epsilon_{i\alpha}^c$ and $\epsilon'_{i\alpha}^c$

$$\begin{aligned}\epsilon_{i\alpha}^c &\equiv \frac{M_i}{(8\pi)^2 \Gamma_i} \sum_{j \neq i} \left\{ \text{Im}[(Y'^{\dagger} Y')_{ij} Y_{\alpha i}^* Y_{\alpha j}] f_{ij} + \text{Im}[(Y^{\dagger} Y)_{ji} Y_{\alpha i}^* Y_{\alpha j}] g_{ij} \right\} \\ \epsilon'_{i\alpha}^c &\equiv \frac{M_i}{(8\pi)^2 \Gamma_i} \sum_{j \neq i} \left\{ \text{Im}[(Y^{\dagger} Y)_{ij} Y_{\alpha i}^* Y'_{\alpha j}] f_{ij} + \text{Im}[(Y'^{\dagger} Y')_{ji} Y_{\alpha i}^* Y'_{\alpha j}] g_{ij} \right\}\end{aligned}$$

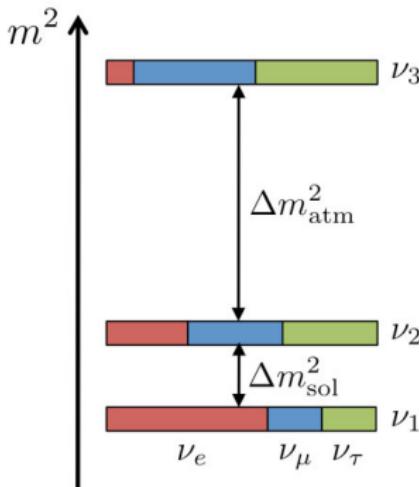
Δ_{tot} -violating terms $\epsilon_{i\alpha}^v$ and $\epsilon'_{i\alpha}^v$

$$\begin{aligned}\epsilon_{i\alpha}^v &\equiv \frac{M_i}{(8\pi)^2 \Gamma_i} \sum_{j \neq i} \left\{ \text{Im}[(Y^{\dagger} Y)_{ij} Y_{\alpha i}^* Y_{\alpha j}] f_{ij} + \text{Im}[(Y'^{\dagger} Y')_{ji} Y_{\alpha i}^* Y_{\alpha j}] g_{ij} \right\} \\ \epsilon'_{i\alpha}^v &\equiv \frac{M_i}{(8\pi)^2 \Gamma_i} \sum_{j \neq i} \left\{ \text{Im}[(Y'^{\dagger} Y')_{ij} Y_{\alpha i}^* Y'_{\alpha j}] f_{ij} + \text{Im}[(Y^{\dagger} Y)_{ji} Y_{\alpha i}^* Y'_{\alpha j}] g_{ij} \right\}\end{aligned}$$

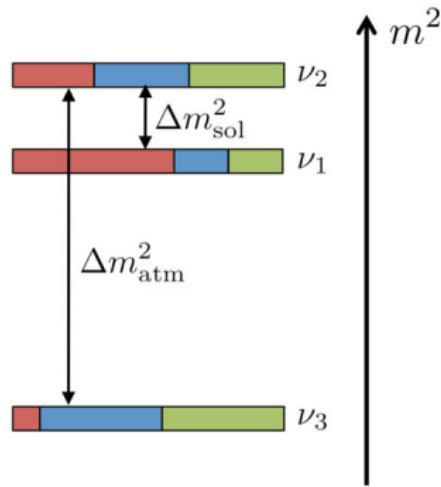
The regulated one-loop functions are given by $f_{ij} \equiv \frac{\sqrt{x_{ji}}(1-x_{ji})}{(1-x_{ji})^2+a_{ji}}$ and $g_{ij} \equiv \frac{1-x_{ji}}{(1-x_{ji})^2+a_{ji}}$, with $x_{ji} \equiv M_j^2/M_i^2$ and $a_{ji} \equiv \Gamma_j^2/M_i^2$.

Neutrino mass hierarchy

normal hierarchy (NH)



inverted hierarchy (IH)



General neutrino masses

Lagrangian

$$\begin{aligned}\mathcal{L} = & \mathcal{L}_{kin} - \frac{1}{2}(m_n)_{ij}n_i n_j - \frac{1}{2}(m'_n)_{ij}n'_i n'_j - (m_D)_{ij}n'_i n_j + \text{h.c.} \\ & - y_{\alpha i}(l\Phi)_\alpha n_i - y'_{\alpha i}(l'\Phi')_\alpha n'_i + \text{h.c.} \\ & - \tilde{y}_{\alpha i}(l\Phi)_\alpha n'_i - \tilde{y}'_{\alpha i}(l'\Phi')_\alpha n_i + \text{h.c.}\end{aligned}$$

Integrating out the heavy n, n' fields

$$\begin{aligned}\mathcal{L}_{eff} = & \left[\frac{-y^2 m'_n - \tilde{y}^2 m_n + 2y\tilde{y}m_D}{2(m_D^2 - m_n m'_n)} \right] (l\Phi)^2 \\ & + \left[\frac{-y'^2 m_n - \tilde{y}'^2 m_{n'} + 2y'\tilde{y}'m_D}{2(m_D^2 - m_n m'_n)} \right] (l'\Phi')^2 \\ & + \left[\frac{-y\tilde{y}' m'_n - y'\tilde{y} m_n + (yy' + \tilde{y}\tilde{y}')m_D}{(m_D^2 - m_n m'_n)} \right] (l\Phi)(l'\Phi')\end{aligned}$$

Comments

- With 3 families we have 30 total parameters (5 unobservables Majorana phases): 9 new angles, 12 new phases, 3 ε_k^2
- A two-parameter fit was performed (turning on one ε_k^2 and another new mixing angle at a time), leading to constraints in the range $\varepsilon_k^2 \lesssim 10^{-12} - 10^{-5} \text{ eV}^2$ for $k = 1, 2$. Larger values of ε_k^2 are also allowed for fine-tuned values of the mixing angle. For $k = 3$, the bound is in general much weaker: $\varepsilon_3^2 \lesssim 10^{-5} \text{ eV}^2$ [*Anamiati, Fonseca, Hirsch 1710.06249*]
- Solar neutrino experiments are not sensitive to values of $\varepsilon^2 \lesssim 10^{-12} \text{ eV}^2$, but this could be probed by measuring the flavor content of high-energy astrophysical neutrinos [*A. Esmaili, arXiv:hep-ph/0909.5410*]