

Precision Gravitational Wave Spin Observables from EFT



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Outline of my talk:

- Motivation
- Overview of the EFT (Nonrelativistic GR) formalism
- Computing NLO Spin Effects
- Concluding remarks

Motivation

Compact Object Binaries and LIGO/Virgo/LISA

• First GW detection from black hole merger by LIGO in 2015

- Why?
 - Strong field G.R.
 - Structure of Neutron Stars
 - Multi-messenger physics

| GW150914 |
|---|
| LVT151012 |
| ~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~ |
| GW151226 |
| GW170104 |
| GW170814 |
| 0 sec. 1 sec. 2 sec. time observable by LIGO-Virgo |

LIGO/University of Oregon/Ben Farr

- Populations in stellar graveyard (spins, masses, etc)
- Others?

Big Picture: Producing precise templates for GW detectors

• Challenge: Solve Einstein's field equations



Solutions give waveforms to be used in templates and analysis



EFT approach takes advantage of separation of scales



LIGO/T. Pyle

- Post-Newtonian Approximation: $\mathcal{O}(v^n) = \left(\frac{n}{2}\right) PN$ •
- By Virial theorem: $\frac{G_N m}{r} = v^2 \rightarrow \frac{r_s}{r} \sim v^2, \ \frac{r}{\lambda} \sim v$

 $r_s \equiv$ Size of Compact Object $r \equiv$ Orbital Radius $\lambda \equiv$ Radiation Wavelength

• Start with full GR coupled to point particles:

$$S = S_{\text{matter}}(\dot{x}_A, g, S) + S_{\text{Einstein-Hilbert}}(g) + S_{\text{gauge fix}}$$

• Expand metric: $g^{\mu\nu} = \eta^{\mu\nu} + \frac{h^{\mu\nu}}{M_{\text{Pl}}} = \eta^{\mu\nu} + \begin{pmatrix} \text{Radiation} \\ \text{Modes} \\ \bar{h}^{\mu\nu} \\ & \end{pmatrix} + \begin{pmatrix} \text{Potential} \\ \text{Modes} \\ H^{\mu\nu} \\ & \end{pmatrix}$ •----•

• Scaling: $(k^0, \mathbf{k})_{\text{pot}} \sim (v/r, 1/r)$ and $(k^0, \mathbf{k})_{\text{rad}} \sim (v/r, v/r)$

$$\exp\left[iS\left(x_{A},\bar{h}\right)\right] = \int \mathcal{D}H_{\mu\nu} \exp\left[iS\left(x_{A},\bar{h}+H\right)\right]$$

(Goldberger and Rothstein 2005)

• Action reproduces GR spin equations of motion:

$$\frac{Dp^{\mu}}{D\tau} = -\frac{1}{2} R^{\mu}_{\nu\alpha\beta} u^{\nu} S^{\alpha\beta}, \quad \frac{DS^{\mu\nu}}{D\tau} = p^{\mu} u^{\nu} - u^{\mu} p^{\nu}$$

• (Porto 2006) Use Routhian mechanics + tetrads: $\eta_{IJ} = e_I^{\mu} e_J^{\nu} g_{\mu\nu}$

$$\mathscr{R} = -p^{\mu}u_{\mu} - \frac{1}{2}\omega_{\mu}^{ab}S_{ab}u^{\mu} \qquad (\omega_{\mu}^{ab} \equiv e_{\nu}^{b}\nabla_{\mu}e^{a\nu})$$

• Compute equations of motion from:

$$\frac{\delta}{\delta x^{\mu}} \int \mathcal{R} \, d\sigma = 0, \qquad \frac{dS^{ab}}{d\sigma} = \left\{ S^{ab}, \mathcal{R} \right\}$$

• Power counting for maximally rotating objects:

$$S \sim mv_{\rm rot}r_s < mr_s \sim Lv$$

Nonrelativistic General Relativity

EFT spin defined in the locally-flat frame

Need additional constraint to go from tensor to vector, known as spin-supplementary condition:

- Covariant: $p_{\mu}S^{\mu\nu} = 0$
- Newton-Wigner: $S^{\mu 0} S^{\mu j} \left(\frac{\tilde{p}^j}{\tilde{p}_0 + m} \right) = 0$



(Fleming 1964)

Appropriate spins in adiabatic approximation are *conserved norm spin vectors* (Will 2005) related to LFF vector by

$$\mathbf{S}_{A} = \left(1 + \frac{1}{2}\mathbf{v}_{A}^{2}\right)\mathbf{S}_{A}^{c} - \frac{1}{2}\mathbf{v}_{A}\left(\mathbf{v}_{A}\cdot\mathbf{S}_{A}^{c}\right) + \cdots$$

Spin evolution takes precession form, i.e.,



(Owen et al 1997, Faye et al 2006, Porto 2010)

Schematic of EFT



"Full" theory



1.5PN potential from Lagrangian:

$$V_{1.5PN}^{SO} = \frac{G_N m_2}{r^3} x^j \left(S_1^{j0} + S_1^{jk} \left(v_1^k - 2v_2^k \right) \right) + 1 \leftrightarrow 2$$

Euler-Lagrange
$$\mathbf{a}_{SO}^{(cov)} = \frac{1}{r^3} \left\{ 6\hat{\mathbf{n}} \left[(\hat{\mathbf{n}} \times \mathbf{v}) \cdot \left(2\mathbf{S} + \frac{\delta m}{m} \mathbf{\Delta} \right) \right] - \left[\mathbf{v} \times \left(7\mathbf{S} + 3\frac{\delta m}{m} \mathbf{\Delta} \right) \right] + 3\dot{r} \left[\hat{\mathbf{n}} \times \left(3\mathbf{S} + \frac{\delta m}{m} \mathbf{\Delta} \right) \right] \right\}$$

In general: $\mathbf{a} = \mathbf{a}_{N} + \mathbf{a}_{1PN} + \mathbf{a}_{1.5PN}^{SO} + \mathbf{a}_{2PN} + \mathbf{a}_{2PN}^{SS} + \mathbf{a}_{2.5PN}^{RR} + \cdots$

Effective action:

$$S_{\text{rad}}\left[\bar{h}, x_{a}\right] = -\int dt \sqrt{\bar{g}_{00}} \left[M(t) + \sum_{l=2}^{\infty} \left(\frac{1}{l!} I^{L} \nabla_{L-2} E_{i_{l-1}i_{l}} - \frac{2l}{(2l+1)!} J^{L} \nabla_{L-2} B_{i_{l-1}i_{l}} \right) \right]$$

Coupling of one radiation mode and the stress-energy pseudotensor:

$$\Gamma[\bar{h}] = -\frac{1}{2m_{Pl}} \int d^4x T^{\mu\nu}(x) \bar{h}_{\mu\nu}$$

Compute single radiation mode diagrams to extract $T_{\mu\nu}$



Use general expressions for moments as functions of $T_{\mu\nu}$ (Ross 2013)

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Next-to-leading order spin-orbit effects in the equations of motion, energy loss and phase evolution of binaries of compact bodies in the effective field theory approach

Brian Pardo¹ and Natália T. Maia¹ (arX

(arXiv:2009.05628)

Gravitational radiation from inspiralling compact objects: Spin-spin effects completed at the next-to-leading post-Newtonian order

Gihyuk Cho,¹ Brian Pardo,² and Rafael A. Porto¹ (arXiv:2103.14612)

Purpose: Complete NLO spin effects

- Equations of motion
- Adiabatic invariants
- Flux-Balance Laws
- Accumulated orbital phase and recoil velocity circular orbits
- Compare with literature (Blanchet et al. 2006, Bohe et al. 2015, Racine et al. 2009, Faye et al. 2006, Blanchet et al. 2006, Bohe et al 2013)

NLO Spin Effects

- Potentials at NLO already computed by Porto and Rothstein
 - Use to compute accelerations and spin evolution
 - (ArXiv 1005.5730, 0804.0260, 0802.0720)

- Multipole moments at NLO already computed by Porto, Rothstein, and Ross
 - Use to compute flux-balance laws
 - (ArXiv 1007.1312)
 - We showed these are equivalent to literature with transformation to PN spins/conserved norm spins

NLO Spin Effects

Example: NLO Spin-Orbit Potential (Porto 2010):

$$V_{\text{SO}}^{\text{NLO}} = \frac{Gm_2}{r^3} \left[\left\{ S_1^{i0} \left(2\mathbf{v}_2^2 - 2\mathbf{v}_1 \cdot \mathbf{v}_2 - \frac{3}{2r^2} \left(\mathbf{v}_2 \cdot \mathbf{r} \right)^2 - \frac{1}{2} \mathbf{a}_2 \cdot \mathbf{r} \right) + \left(2\mathbf{v}_1 \cdot \mathbf{v}_2 + \frac{3\left(\mathbf{v}_2 \cdot \mathbf{r} \right)^2}{r^2} - 2\mathbf{v}_2^2 + \mathbf{a}_2 \cdot \mathbf{r} \right) S_1^{ij} \mathbf{v}_2^{i} - \left(\frac{3}{2r^2} \left(\mathbf{v}_2 \cdot \mathbf{r} \right)^2 + \frac{1}{2} \mathbf{a}_2 \cdot \mathbf{r} \right) S_1^{ij} \mathbf{v}_1^{i} + 2S_1^{ij} \mathbf{a}_2^{i} \mathbf{v}_2 \cdot \mathbf{r} + r^2 S_1^{ij} \mathbf{a}_2^{j} \right] \mathbf{r}^{i} + S_1^{i0} \left(\left(\mathbf{v}_1 - \mathbf{v}_2 \right)^i \mathbf{v}_2 \cdot \mathbf{r} - \frac{3}{2} \mathbf{a}_2^{i} r^2 \right) + S_1^{ij} \mathbf{v}_2^{i} \mathbf{v}_1^{j} \mathbf{v}_2 \cdot \mathbf{r} - r^2 S_1^{ij} \mathbf{a}_2^{j} \mathbf{v}_2^{i} - \frac{1}{2} r^2 S_1^{ij} \mathbf{a}_2^{j} \mathbf{v}_1^{i} \right] \\ + \frac{G^2 m_2}{r^4} \mathbf{r}^{i} \left[-\left(m_1 + 2m_2 \right) S_1^{i0} + \left(m_1 - \frac{m_2}{2} \right) S_1^{ij} \mathbf{v}_1^{j} + \frac{5m_2}{2} S_1^{ij} \mathbf{v}_2^{j} \right] + 1 \leftrightarrow 2 \right]$$

$$\mathbf{a}_{\text{reduced}}^{(2,\text{SPN})} = \frac{G}{m\nu r^4} \left\{ \mathbf{n}^i \left[\mathbf{S} \cdot \mathbf{L} \left(-\frac{Gm}{r} (42 + 29\nu) + 3(-1 + 10\nu)\mathbf{v}^2 - 30\nu \dot{r}^2 \right) - \frac{\delta m}{m} \mathbf{\Sigma} \cdot \mathbf{L} \left(\frac{Gm}{r} \left(22 + \frac{33}{2}\nu \right) + 3(1 - 5\nu)\mathbf{v}^2 + 15\nu \dot{r}^2 \right) \right] + 3\dot{r} \mathbf{v}^i \left[3\mathbf{S} \cdot \mathbf{L}(-1 + \nu) + \frac{\delta m}{m} \mathbf{\Sigma} \cdot \mathbf{L}(-1 + 2\nu) \right] - 2\frac{Gm}{r} \mathbf{L}^i \left[\mathbf{S} \cdot \mathbf{n}(1 + 2\nu) + \frac{\delta m}{m} \mathbf{\Sigma} \cdot \mathbf{n}(1 + \nu) \right] \right\} + \frac{G}{r^3} \left[(\mathbf{S} \times \mathbf{n})^i \dot{r} \left[\frac{Gm}{r} (26 + 25\nu) + \frac{3}{2} (1 - 15\nu)\mathbf{v}^2 + \frac{45}{2}\nu \dot{r}^2 \right] + \frac{\delta m}{m} (\mathbf{\Sigma} \times \mathbf{n})^i \dot{r} \left[\frac{Gm}{r} (10 + \frac{27}{2}\nu) + \left(\frac{3}{2} - 12\nu \right) \mathbf{v}^2 + 15\nu \dot{r}^2 \right] + (\mathbf{S} \times \mathbf{v})^i \left[-\frac{\delta m}{r} (\mathbf{\Sigma} \times \mathbf{v})^i \left[-\frac{\delta m}{r} (10 + \frac{15}{2}\nu) + \left(\frac{3}{2} - 8\nu \right) \mathbf{v}^2 + 9\nu \dot{r}^2 \right] \right\}$$

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NLO Spin Effects

• Energy:
$$E = \sum_{A=1}^{2} \sum_{n} \mathbf{p}_{\mathbf{x}_{A}^{(n)}} \cdot \mathbf{x}_{A}^{(n+1)} + V + E_{\text{reduced}}$$
$$\mathbf{p}_{\mathbf{x}_{A}^{(n)}} = -\sum_{A=1}^{2} \sum_{k=n+1}^{2} \left(-\frac{d}{dt}\right)^{k-n-1} \frac{\partial V}{\partial \mathbf{x}_{A}^{(k)}}$$

• Orbital angular momentum: $\mathbf{L} = \epsilon^{ijk} \sum_{A=1}^{2} \sum_{n} \mathbf{x}_{A}^{j(n)} \mathbf{p}_{\mathbf{x}_{A}^{(n)}}^{k} + \mathbf{L}_{\text{reduced}}$

Nontrivial check:
$$\frac{d}{dt}(\mathbf{L} + \mathbf{S}) = 0$$

 \sim

• Center of mass correction from 1-point function:

$$\mathbf{r}_{\rm cm}^i = \frac{1}{m} \int d^3 x \, \mathbf{x}^i \, T^{00}(\mathbf{x}, t)$$

• Results with corresponding results in literature were in agreement!

• Energy Loss:

$$\frac{dE}{dt} = -\frac{G}{5} \left(I_{ij}^{(3)} I_{ij}^{(3)} + \frac{16}{9} J_{ij}^{(3)} J_{ij}^{(3)} + \frac{5}{189} I_{ijk}^{(4)} I_{ijk}^{(4)} + \frac{5}{84} J_{ijk}^{(4)} J_{ijk}^{(4)} + \cdots \right)$$

• Angular momentum loss:

$$\frac{d\boldsymbol{J}^{i}}{dt} = -G\epsilon^{iab}\left(\frac{2}{5}I_{aj}^{(2)}I_{bj}^{(3)} + \frac{32}{45}J_{aj}^{(2)}J_{bj}^{(3)} + \frac{1}{63}I_{ajk}^{(3)}I_{bjk}^{(4)} + \frac{1}{28}J_{ajk}^{(3)}J_{bjk}^{(4)} + \cdots\right)$$

• Momentum Loss:

$$\frac{d\mathbf{P}^{i}}{dt} = -G\left(\frac{2}{63}I_{ijk}^{(4)}I_{jk}^{(3)} + \frac{16}{45}\epsilon_{ijk}I_{jl}^{(3)}J_{kl}^{(3)} + \frac{1}{126}\epsilon_{ijk}I_{jlm}^{(4)}J_{klm}^{(4)} + \frac{4}{63}J_{ijk}^{(4)}J_{jk}^{(3)} + \cdots\right)$$

• Center of mass flux:

$$\frac{d\mathbf{G}^{i}}{dt} = \mathbf{P}^{i} - G\left(\frac{1}{21}\left(I_{ijk}^{(3)}I_{jk}^{(3)} - I_{ijk}^{(4)}I_{jk}^{(2)}\right) + \frac{2}{21}\left(J_{ijk}^{(3)}J_{jk}^{(3)} - J_{ijk}^{(4)}J_{jk}^{(2)}\right) + \cdots\right)$$

• Results consistent with literature for NLO spin-orbit and spin-spin!

• Reduction to circular, spin-(anti)aligned orbits:

$$S_{A} \equiv \pm S_{A} \mathscr{C}, \quad \dot{r} = 0, \quad r\omega^{2} = -\langle \mathbf{n} \cdot \mathbf{a} \rangle$$
$$\frac{\dot{\omega}}{\omega} = \frac{3}{2x} \left(\frac{dE[x]}{dx}\right)^{-1} \frac{dE[x]}{dt}, \quad \text{with } x \equiv (Gm\omega)^{2/3}$$

• Accumulated phase:

$$\phi = \int dt \, \omega = \int d\omega \, \frac{\omega}{\dot{\omega}}$$

$$\begin{split} \phi &= \phi_0 - \frac{32}{\nu} \left\{ x^{-5/2} + x^{-3/2} \left[\frac{3715}{1008} + \frac{55}{12} \nu \right] + \frac{x^{-1}}{Gm^2} \left[\frac{125}{8} \frac{\delta m}{m} \Sigma_{\ell}^c + \frac{235}{6} S_{\ell}^c \right] \right. \\ &+ x^{-1/2} \left[\frac{15293365}{1016064} + \frac{27145}{1008} \nu + \frac{3085}{144} \nu^2 \right] - \frac{\log x}{Gm^2} \left[\left(\frac{41745}{448} - \frac{15}{8} \nu \right) \frac{\delta m}{m} \Sigma_{\ell}^c + \left(\frac{554345}{2016} + \frac{55}{8} \nu \right) S_{\ell}^c \right] \right\} \\ \bullet \quad \text{Recoil Velocity:} \quad \mathbf{V}_{\text{kick}}^i = \frac{1}{m} \int_{-\infty}^t dt \left(\frac{d\mathbf{P}^i}{dt} \right) \end{split}$$

$$\left(V_{\text{kick}}^{i}\right)_{\text{SO}} = \frac{8\nu^{2}x^{9/2}}{15Gm^{2}}\mathbf{n}^{i}\left\{-2\left(\Sigma_{\text{c}}\ell\right) + \frac{1}{21}x\left(-940\delta\left(S_{\text{c}}\ell\right) + 3(-67+412\nu)(\Sigma_{\text{c}}\ell)\right)\right\}$$

• Results consistent with literature up to NLO spin-orbit and spin-spin!

- Computed NLO spin effects, including phase and recoil velocity
- Computed NLO spin pieces needed for templates
- Completed pieces needed at NNLO
- Showed correspondence between literature results and EFT
 - Conservative sector
 - Radiative sector (new!)
- NNLO in the works!

Thank you!