

Argon-1 @ Carleton

An R&D detector for next generation liquid argon experiments

CAP Congress – June 9th 2021 Speaker: David Gallacher (MSc. Student)

+ Mark Boulay, Michael Perry, Ryan Crampton, Jeff Mason, Vance Strickland, Jason Antsey





Talk Trajectory



- Introduction and Motivation
- Argon-1 Detector and Physics Goals
- Preliminary Detector Data



And The Future of Single-Phase LAr DM Detectors





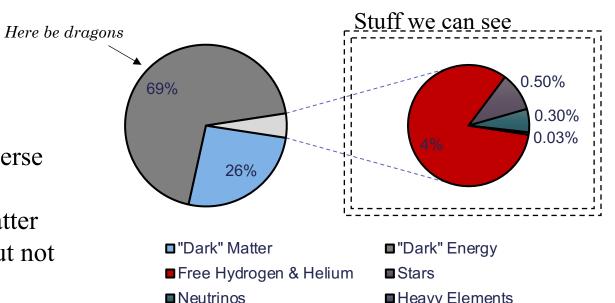
What's the matter?



- Only ~5% of energy in the universe makes up "Normal Matter"
- ~26% in "Dark" or invisible matter
 - Interacts through gravity but not with light
 - Well known from galactic rotation curves and CMBR data (Gravitational signatures)
- Question:

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- What makes up this missing matter?
- Possible candidates:
 - Cold dark baryonic matter?
 - Rogue black holes?
 - New particle(s)?



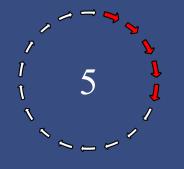
■Heavy Elements



WIMPs: Weaky Interacting Massive Particles. A highly motivated candidate for particle DM



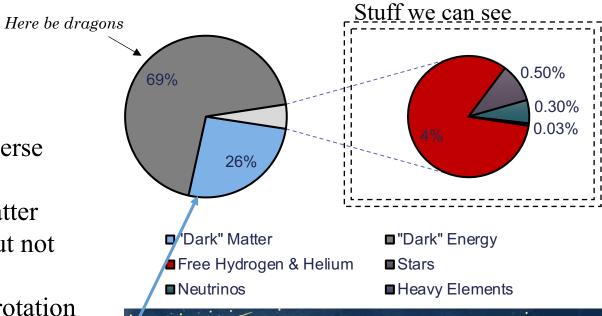
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- What makes up this missing matter?
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 - This is where we search!





WIMPs: Weaky Interacting Massive Particles. A highly motivated candidate for particle DM



DEAP-3600



- DEAP-3600 is a direct dark matter search experiment collecting data at SNOLAB since 2016
- Single Phase liquid argon (LAr) detector:
 - WIMPs interact with Ar nuclei, which creates scintillation light
 - Photons from LAr (128 nm) scintillation are shifted to visible (~ 420 nm) by TPB coating on vessel
- Uses pulse-shape discrimination (PSD) to separate signal-like (WIMP DM) and background-like events (Predominantly Ar39 decays)
 - Singlet decay time $\tau_s \sim 6$ ns
 - Triplet decay time $\mathbf{\tau}_t \sim 1400 \text{ ns}$

~3300 kg LAr

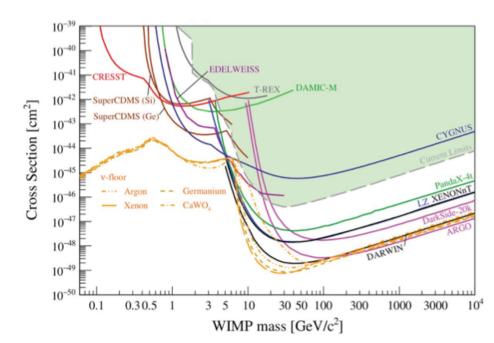


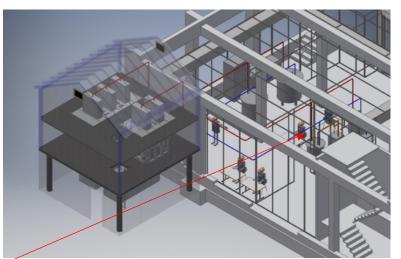
Beyond DEAP



- Researchers from DEAP, Darkside, MiniCLEAN, and ArDM + more, have come together to form the Global Argon Dark Matter Collaboration
- Research and development efforts are directed towards a next generation O(100t) LAr detector
- Proposed experiment 'ARGO' will have sensitivity to the argon neutrino floor
- COLD Lab at Carleton will play a major role in this R&D

Argon-

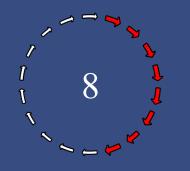




Carleton Noble Liquid Detector Lab



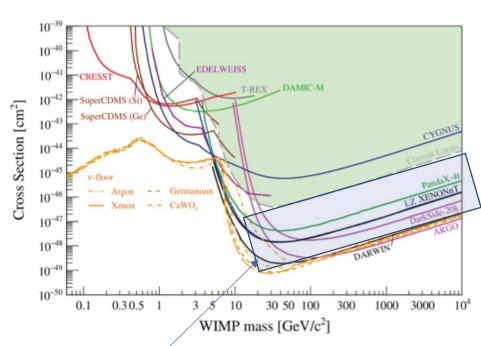
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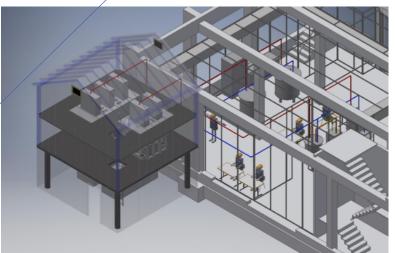


See Mark Boulay's talk (Development of the ARGO dark matter experiment) [Thursday @ 2:30 R-PPD] for more details!

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Carleton Noble Liquid Detector Lab



Argon-1

The Detector and Physics Goals

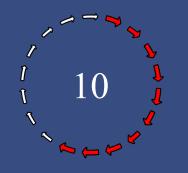


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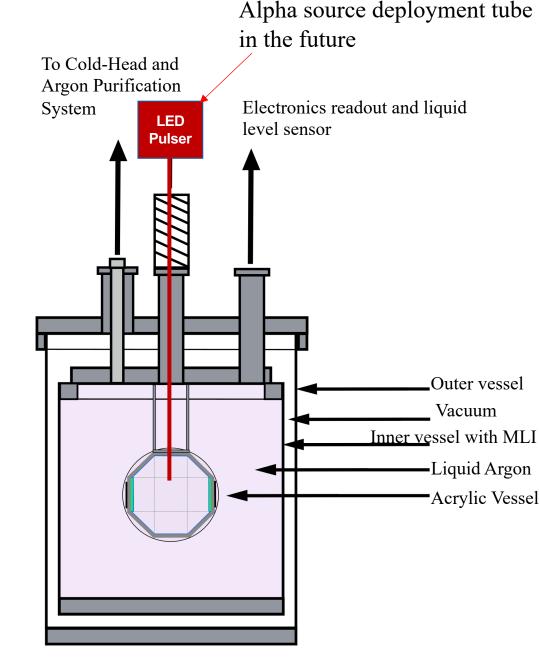


The Detector

Argon-1



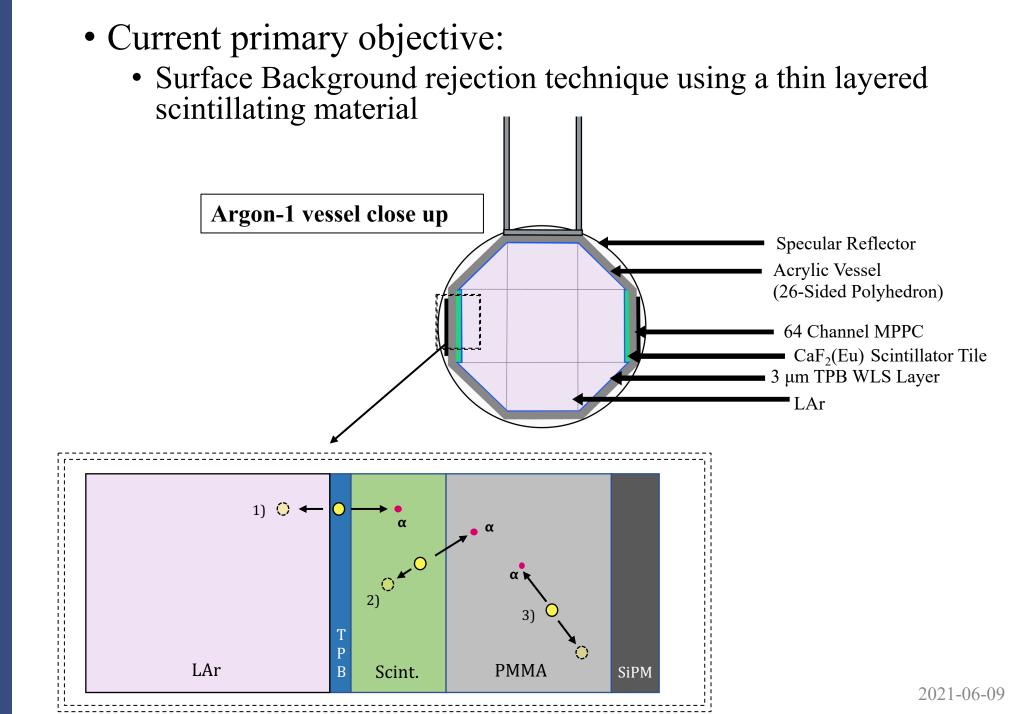
- Argon-1 is a single-phase LAr detector inside of a ~30kg cryostat, LAr scintillation light is shifted to visible by a TPB coating
- The blue TPB light is detected by SiPMs mounted inside the LAr on the sides of the vessel
- Plan to have 17 SiPM arrays surrounding detector for a total of 80 channels
 - Currently operating with 2 SiPMs with 65 channels





The Detector

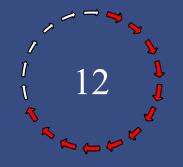
Physics Goals – Surface Background Rejection Technique





The Detector

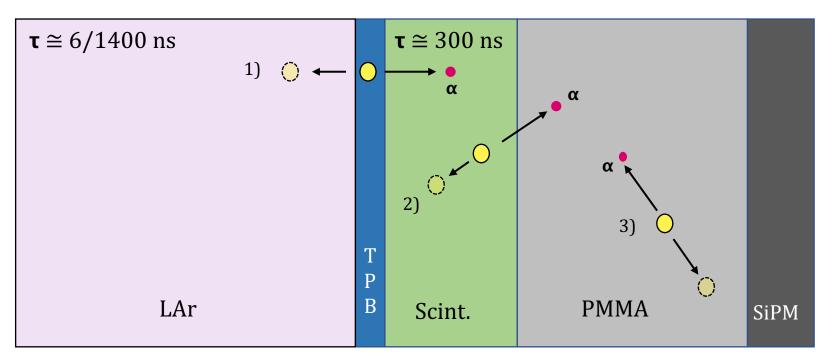
Physics Goals – Surface Background Rejection Technique



• Current primary objective:

See Hicham's Talk at 3:55 PM (R3-6) Characterization of wavelength shifters for background rejection in liquid argon dark matter experiments

• Surface Background rejection technique using a thin layered scintillating material



Using a long-time constant scintillator layer we can use pulse-shape discrimination to reject surface events by shifting them away from our region of interest

- Case 1) Po-210 decay in TPB layer (active)
- Case 2) Po-210 decay in long decay time constant scintillator layer (active)
- Case 3) Po-210 decay in PMMA (non-active)



Monte-Carlo

RAT: Argon-1



- Built a Monte-Carlo simulation using a GEANT4 based software called RAT (Used by DEAP-3600, SNO+ & more)
- MC shows promise for an independent check of surface alpha tagging using light pattern on high channel density SiPM

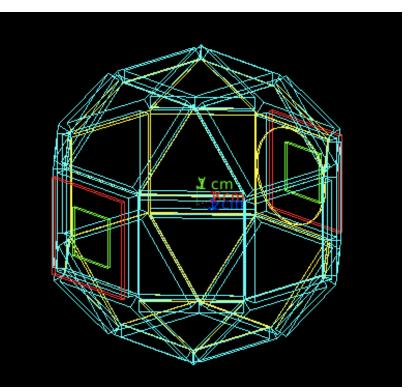
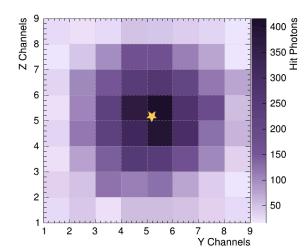


Figure: Simulated Argon-1, Acrylic vessel(Cyan), TPB(Yellow) and SiPMs(Green)



Surface Event



Z Channels

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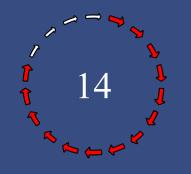
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Y Channels



Monte-Carlo

Detector Response



- Will use MC to help understand detector response for surfacealpha measurement
- First Step: Determine MC detector light yield and calibrate to data

Right: Simulated monoenergetic β 's with 200 keV across the inner vessel

 $Light yield = \frac{Mean of detected PE}{Energy Deposited (keV)}$

Using Gaussian fit: MC LY = **4.7586 PE/keV** Resolution = 20.82 keV (**10.4%**) at **200 keV**

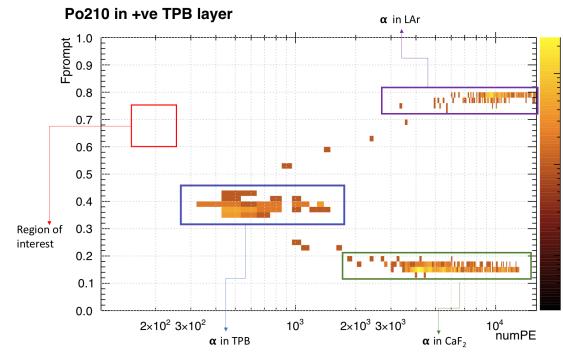
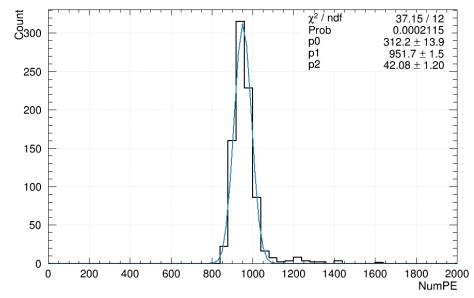


Figure: Example of A1 MC with 2 SiPM arrays and different event topologies, reference of approximate DM ROI included for comparison





Preliminary Data

Fresh off the DAQ

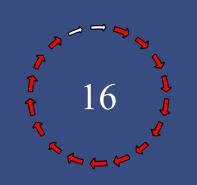


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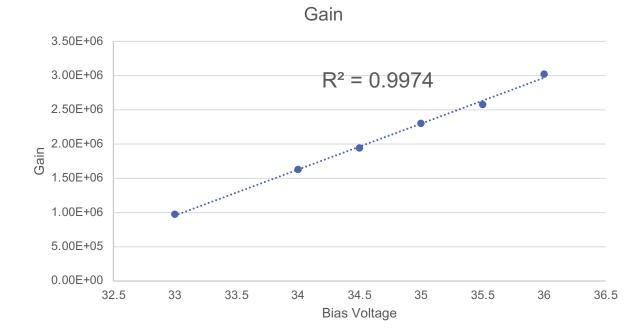


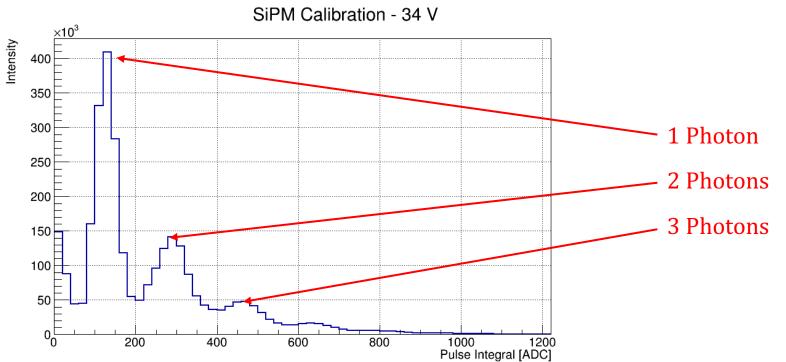
Preliminary Data

SiPM Calibration



Using a pulsed LED light source, we can calibrate the SiPMs



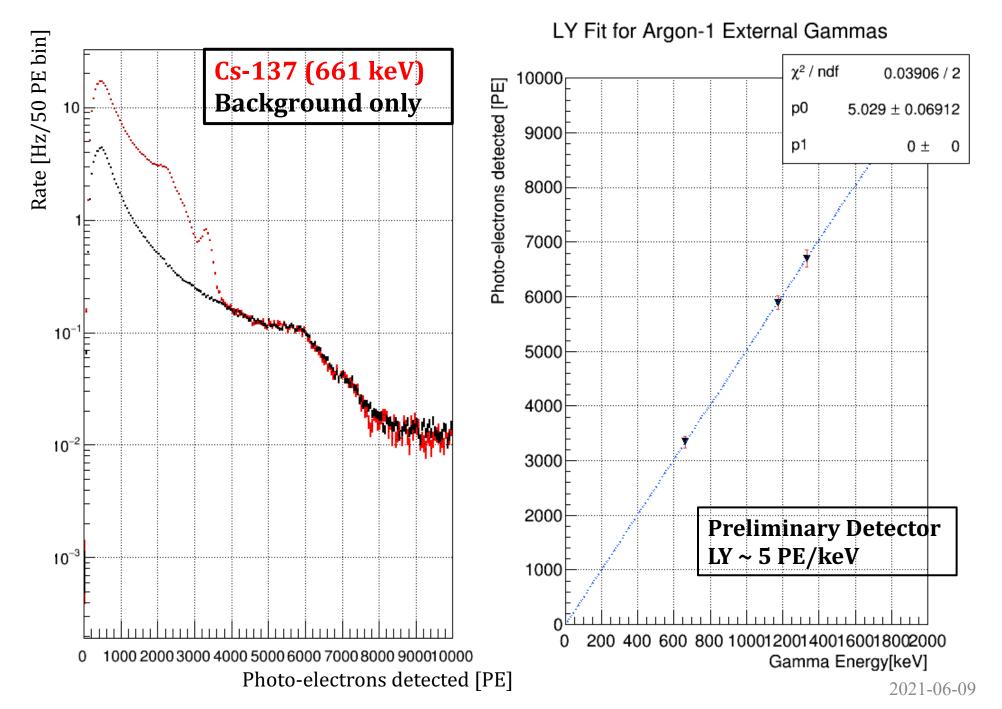




Preliminary Data

External gammas for detector calibration







Conclusion

What's next?



What we've accomplished so far:

- Argon-1 R&D detector successfully commissioned
- DAQ Running with 65 SiPM channels with full waveform readout, and capacity to expand to 130+ in the future
- Monte-Carlo Simulations shows excellent promise for novel background rejection using a layered active surface

The Next Steps:

- Continue detector calibration using external gammas (TI-208, Cs-137 and Na-22 sources) and with internal gammas using Kr-83m
- Ongoing study of SiPM Characterization in LAr conditions
- Reconfigure the detector to perform surface-background discrimination measurement
- Deploy proto-type PDCs from USherbrooke being developed for ARGO, inside Argon-1 for in-situ comparison to analogue SiPMs



Thanks for your attention!

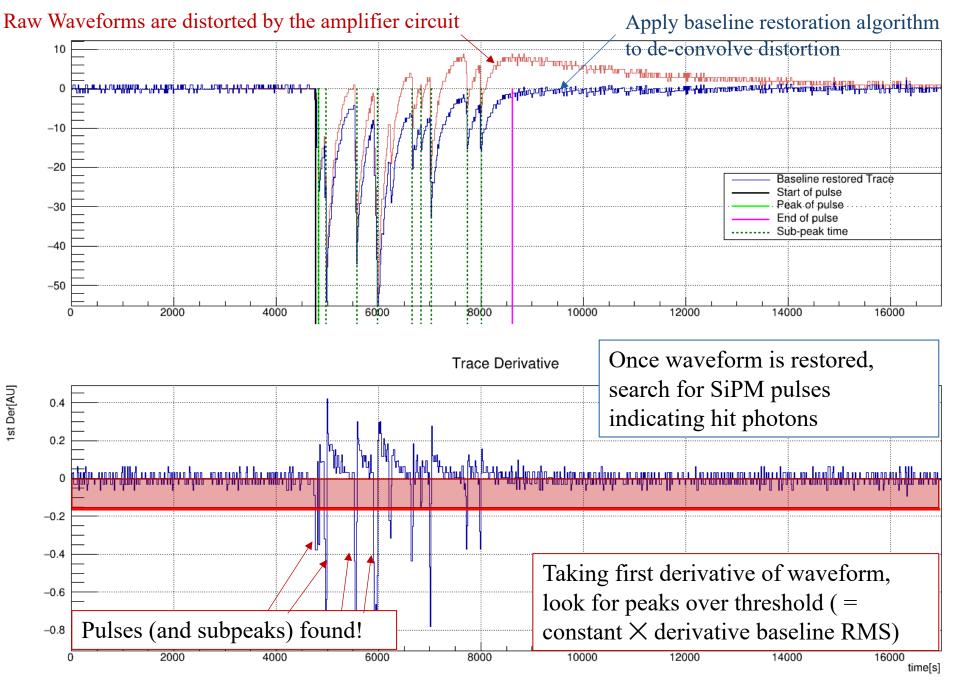






SiPM Pulse-Finding

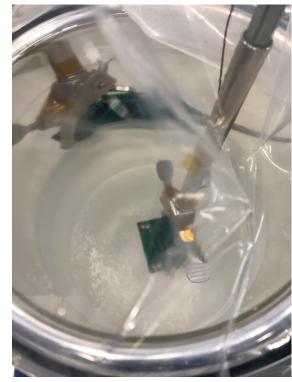
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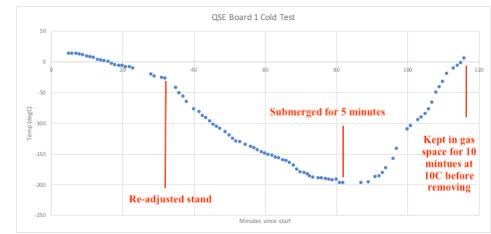


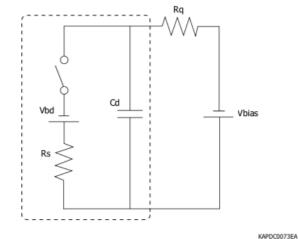
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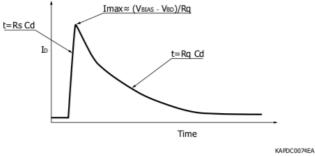
- Ran cryogenic cycle of "QSE"board with adapter attached (To mimic SiPM Soldered)
- No cracking or issues, X-rayed before and after with no noticeable issues
- Continuity good for all channels before and after cooling cycle







[Figure 1-13] Conceptual output pulse of the equivalent circuit



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MIDAS DAQ

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- Front-end software controls V1740 boards for data readout
- FE also controls SiPM Pre-amplifiers and HV supply
- Triggered events are saved to a custom ROOT DS for analysis
- Write out full waveforms first then apply pulsefinding and high-level analysis

Run Status						
Run 192	Start: Mon May 1	0 18:19:51 2021	Running time: 23h23m44s			
Running Stop Pause	Alarms: On	Restart: Off	Data dir: /home/argon1/Data/MidasFiles			

1620766443 16:54:03.437 2021/05/11 [jsroot_server,LOG] Program jsroot_server on host argon1DAQ started

Equipment						
Equipment +	Status	Events	Events[/s]	Data[MB/s]		
FEV1740	Initialized	15.031M	479.4	55.178		
nano03	Running	0	0.0	0.000		
nano04	Running	0	0.0	0.000		
nano07	Running	0	0.0	0.000		
nano08	Running	0	0.0	0.000		
nano09	Running	0	0.0	0.000		

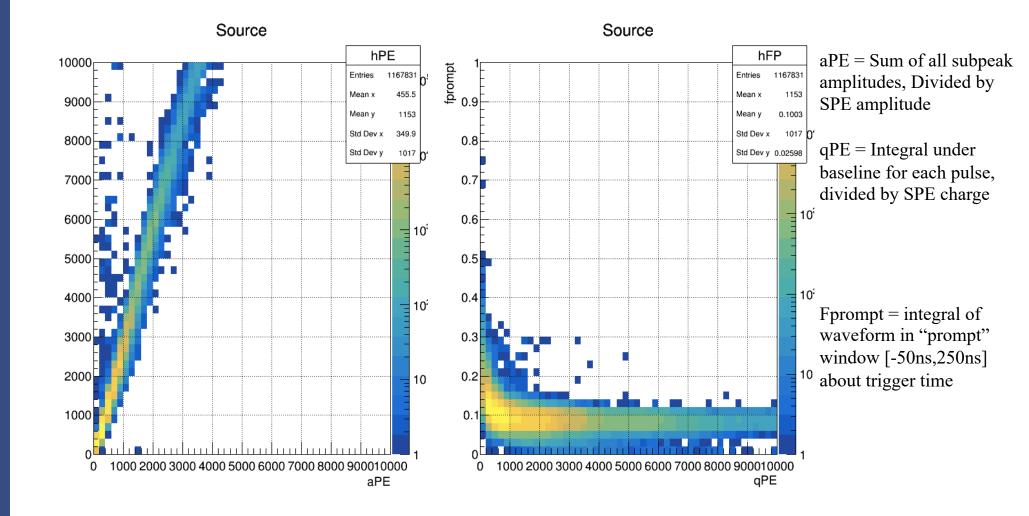
Logging Channels							
Channel	Events	MB written	Compr.	Disk Level			
#0: argon1_run00191.mid.lz4	0	0.000	0.0%	13.9%			
Lazy Label	Progress	File Name	# Files	Total			

	Clients	
mhttpd [argon1DAQ]	Logger [argon1DAQ]	saveload [argon1DAQ]
mserver [argon1DAQ]	nano03 [nanopi03]	nano07 [nanopi07]
nano08 [nanopi08]	nano09 [nanopi09]	nano04 [nanopi04]
feV1740 [argon1DAQ]	jsroot_server [argon1DAQ]	

Figure: Argon-1 DAQ MIDAS FE Interactive webpage



Fprompt vs PE

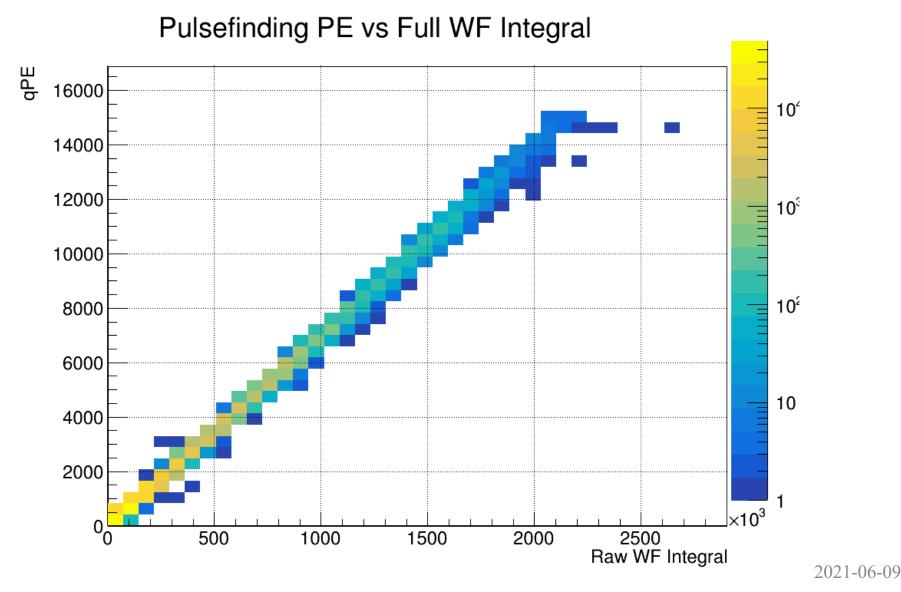




Charge Linearit Y

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• We scale the integrated charge by SPE constants for each channel for each pulse found, but we can look at how linear the charge is to compare





Co60

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• Co60 spectrum is not so clean because of the 2 gammas and Compton edge of 1337 kev gamma combining with peak of 1173 kev gamma

