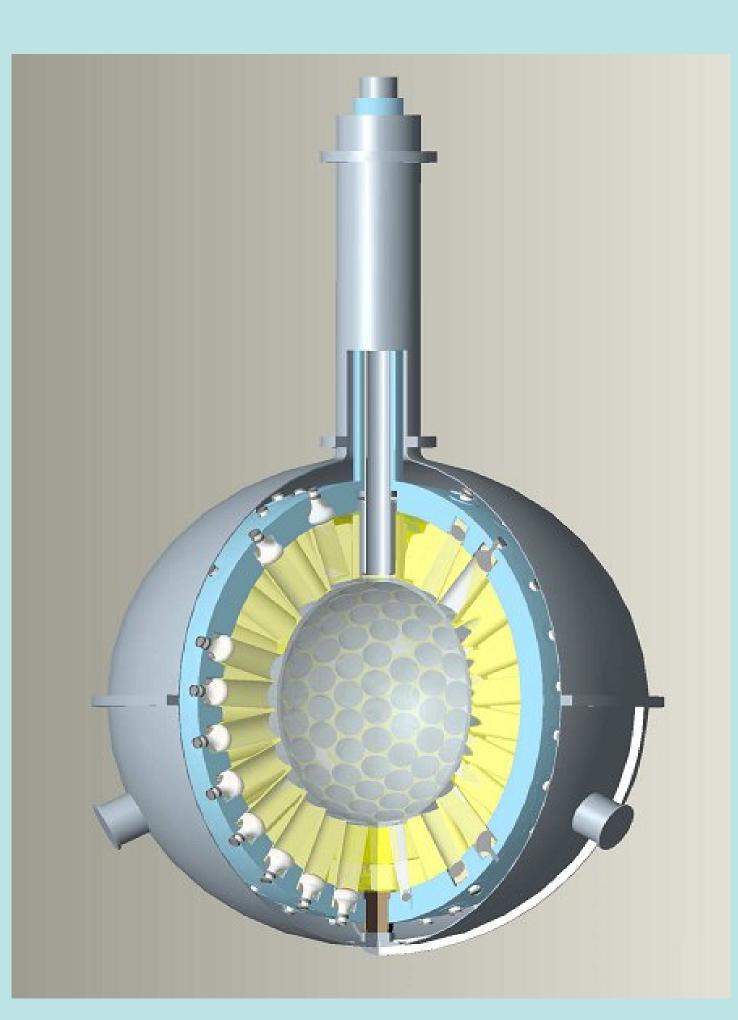


# **Detection Of Heavy Dark Matter Particles In DEAP-3600 Michela Lai on behalf of DEAP-3600 Collaboration**

# **DEAP-3600** is a dark matter detector filled with Atmospheric liquid argon, designed for the WIMP search.

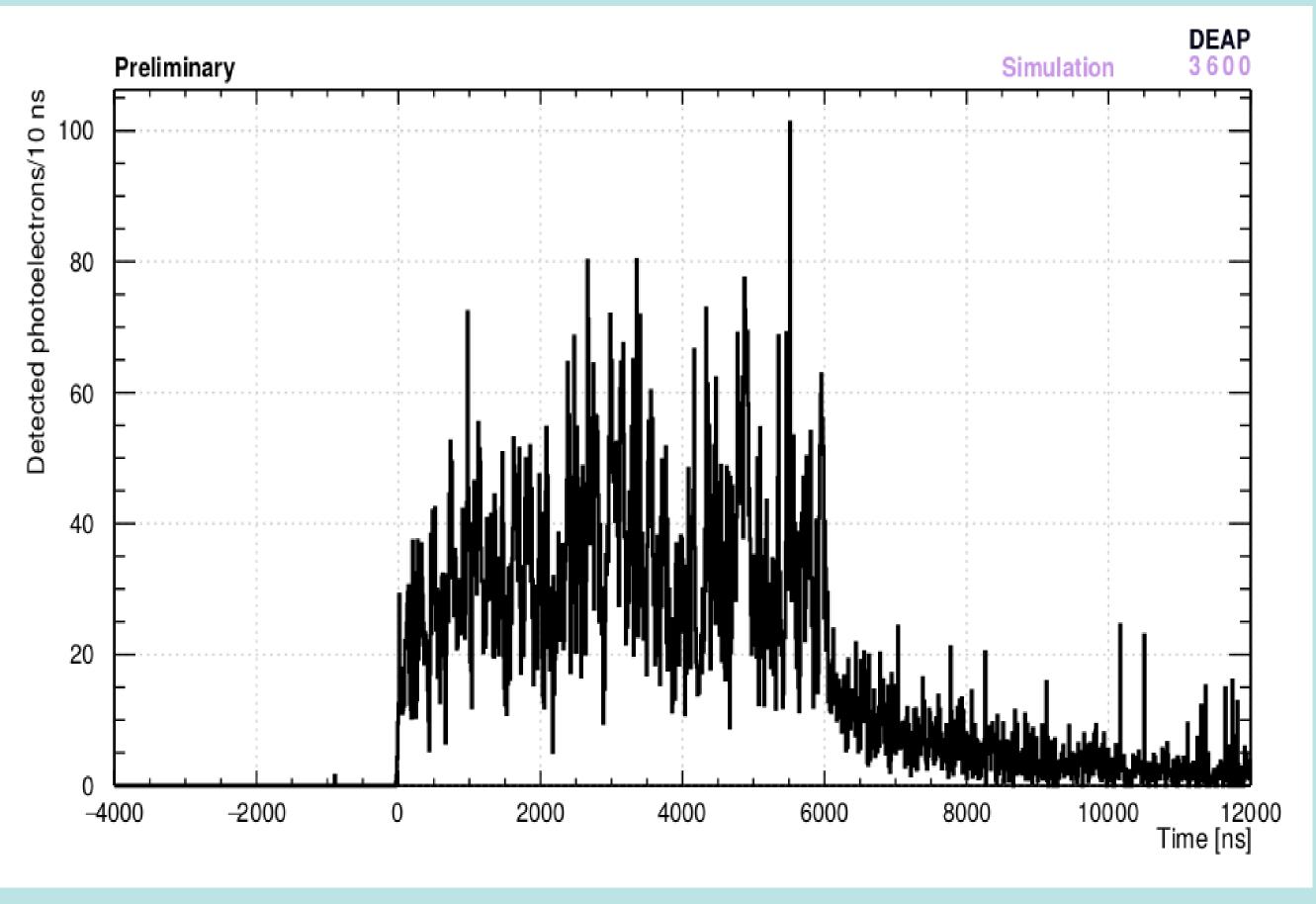


"Design and construction of the DEAP-3600 d ark matter detector", DEAP Collaboration, Ast roparticle Physics, 108, March 2019

Single-phase design **3.3 ton of target mass**, all contained in the acrylic inner vessel. **It is the** largest running dark matter detector based on noble liquids.  $\succ$  connection to the surface through a long neck, which allows for the cooling of the vessel and eventual operations light yield of 7 photoelectrons/keV Photodetection performed by 255 photomultipliers tubes (PMTs), coupled to the vessel via acrylic guides All contained in a stainless-steel sphere All submerged in a cylindrical tank filled with ultrapure water, observed by 48 PMTs

<sup>(6</sup> m.w.e).





The number of subevents, so **the outstanding peaks along** a waveform, identifies multiple recoils in the same acquisition window, as for the present dark matter candidate.

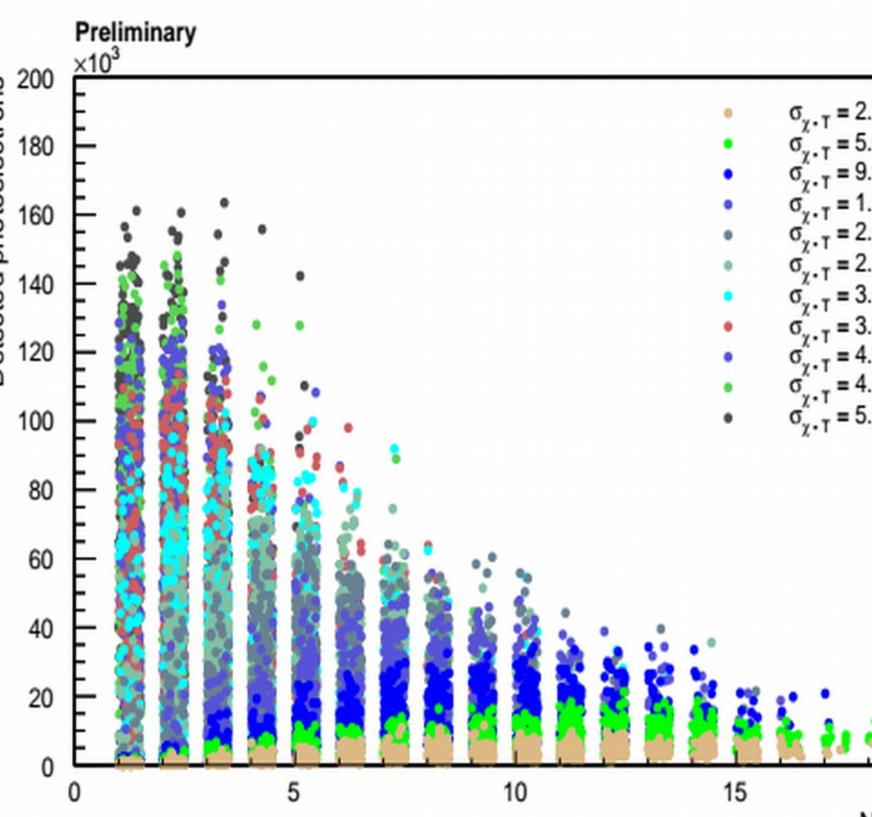
### Set at SNOLAB, under **2 km of rocks**

## **Identikit of Multi-Scattering Heavy particles:**

➢ Mass >> 100 TeV, up to the Planck Mass  $\rightarrow$  Dark matter-target **cross-section** > 5 x 10<sup>-25</sup> cm<sup>2</sup>  $\succ$  produced in the early universe via several mechanisms, like out-of-equilibrium production or pre-heating > the high mass allows to **reach underground** detectors, such as DEAP-3600 ➢ the high cross-section determines multiple **scatterings** in the detector.



The response of the detector to this signal is simulated in RAT, assuming the present geometry of the detector.



At such high cross-sections, the particle performs hundreds, even thousands of scatterings. The distribution of the subevents vs the number of detected photoelectrons for several cross-sections was compared to that from the expected background.

requiring more than one subevent.

200 200 180 160 160 120 120 100 100 100 100 100 100 100 10	-	 	
180	<b>F</b>	 	
160	F	 	
10	F	 	
120	F	 	
100	E	 	
80	F	 	
60	F	 	
40	F	 	
20	E	 	
0	E	 000	

Three Regions (ROIs) Interests ar determined, according the expected to background level.

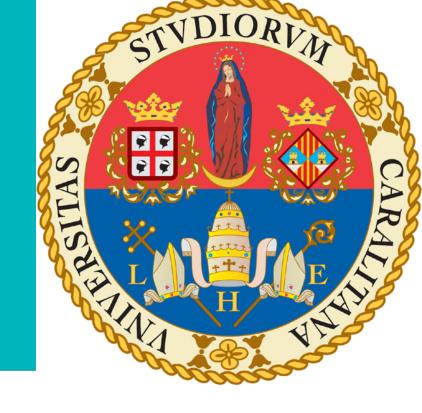
A custom selection cut in the number of subevents was hence optimized up to 10 MeV, divided in ROI #1 and ROI #2 in order to reject our multi-scattering dark matter particles from pile-ups.

No pile-ups are expected above 10 MeV, so 70 kqPE. Here the dominant background are **muons** entering the inner vessel. **This is ROI #3**, where the background suppression is mainly done by the rejection of the events in coincidence with the water tank trigger.

After all the selection cuts, the expected background level in each ROI will be far below 1. Thanks to the high exposure DEAP-3600 will be the first direct detection experiments scanning for heavy multi-scattering dark matter particles up to the Planck Mass.

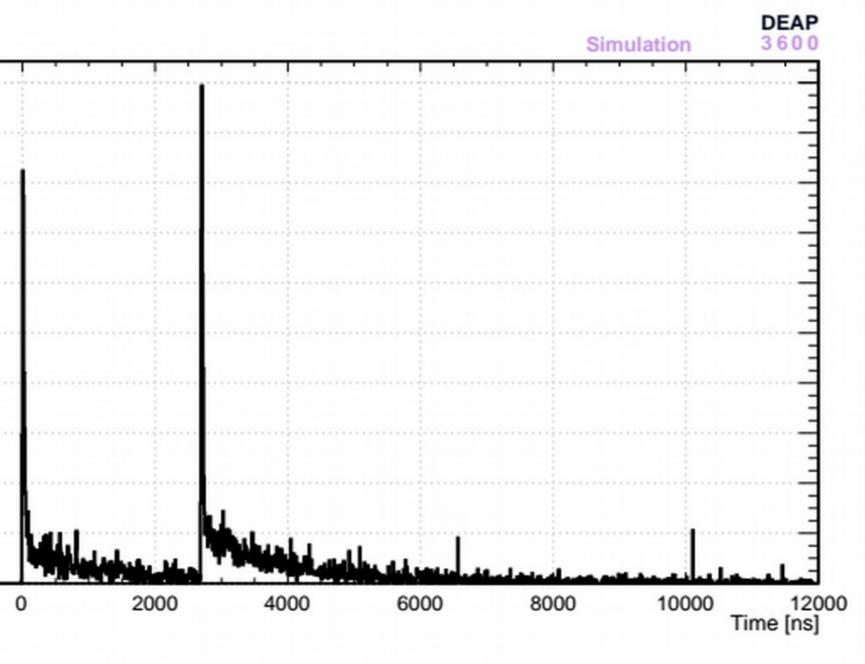
**The unblinding** procedure will start pretty soon, so ...

BEAB = 2.5 x 10<sup>-23</sup> cm<sup>2</sup> = 5.0 x 10<sup>-23</sup> cm<sup>2</sup> = 9.9 x 10<sup>-23</sup> cm<sup>2</sup> = 1.5 x 10<sup>-22</sup> cm<sup>2</sup> = 2.0 x 10<sup>-22</sup> cm<sup>2</sup> = 2.5 x 10<sup>-22</sup> cm<sup>2</sup> - = 3.0 x 10<sup>-22</sup> cm T = 3.5 x 10<sup>-22</sup> cm<sup>2</sup>  $r = 4.0 \times 10^{-22} \text{ cm}^2$  $r = 4.5 \times 10^{-22} \text{ cm}^2$  $r = 5.0 \times 10^{-22} \text{ cm}^2$ Number of detected subevents



Any **background** event performing one single scatter in the detector will be **rejected** by the present analysis by

Still, more background recoils can happen in the same acquisition window. These are referred to as "**pile-up**" events, the dominant background for the present analysis.



of	ROI	Detected			
<b>e</b>		photoelectrons			
g	1	4000-20000			
d	2	20000-70000			
	3	70000- 4 x 10 <sup>8</sup>			