

Alpha background rejection in DEAP-3600 using pyrene

Shivam Garg for the DEAP-3600 collaboration

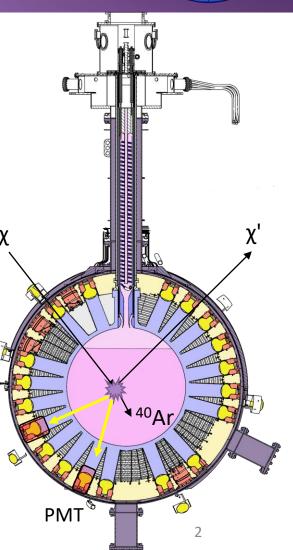
2021 CAP Virtual Congress June 10th, 2021



Carleton

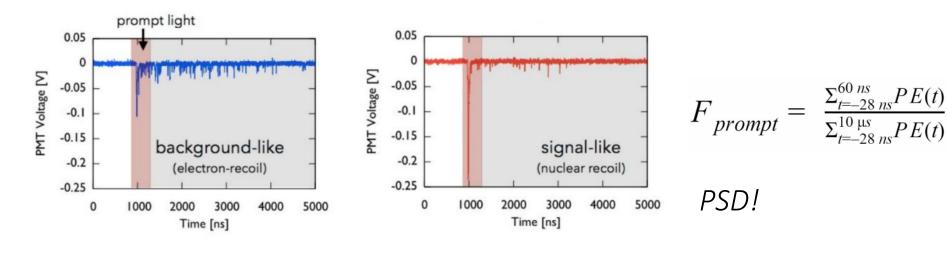


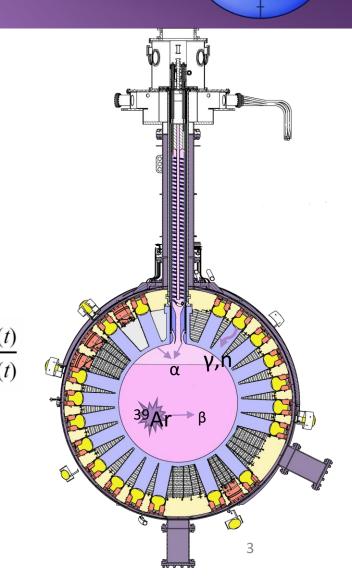
- Liquid argon based dark matter detector
- Located in SNOLAB, Sudbury, Canada
- Dark matter Experiment using Argon Pulseshape discrimination
- >3 tonnes target mass of LAr in acrylic vessel
- Collected data from 2016-2020; upcoming hardware upgrades this year



Backgrounds in DEAP

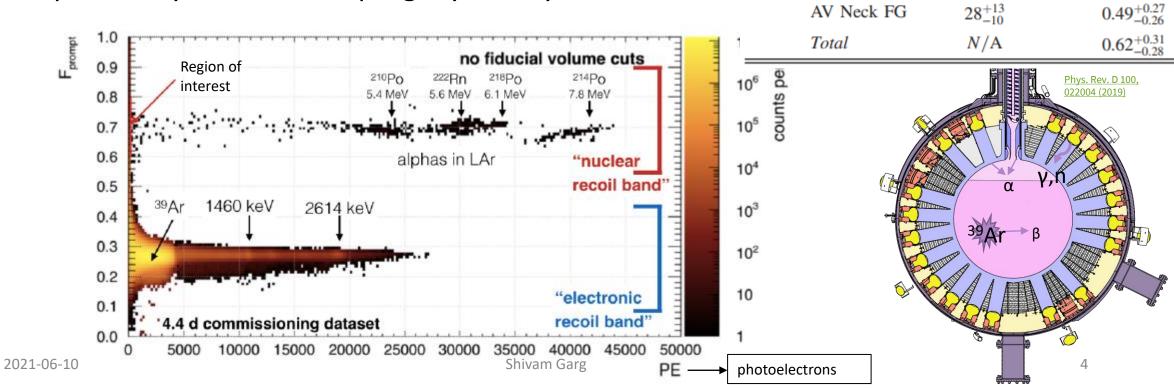
- Beta decays from ³⁹Ar
- Gammas and neutrons from PMT glass
- Alpha decays from radon progeny in acrylic





Backgrounds in DEAP

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N^{ROI}

 0.03 ± 0.01

< 0.14

 $0.10\substack{+0.10\\-0.09}$

< 0.11

< 0.08

 N^{CR}

 2.44×10^{9}

 $<3.3 \times 10^{5}$

 6 ± 4

< 0.2

<3600

Source

Cherenkov

Radiogenic

Cosmogenic

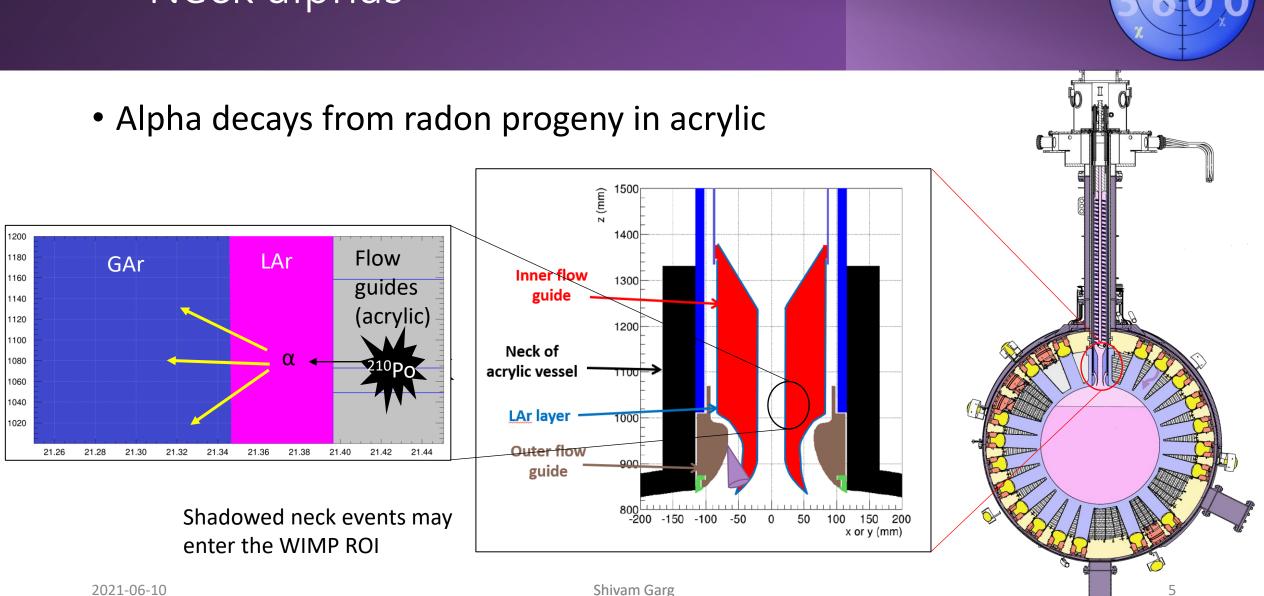
AV surface

ERs

 β/γ 's

n's

α's



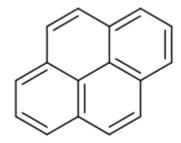
Neck alphas

Shivam Garg

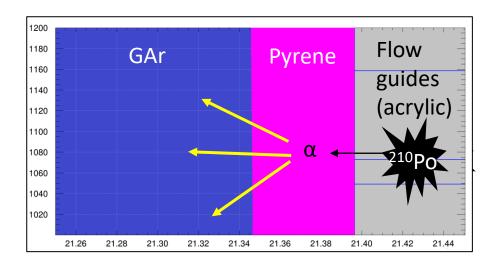


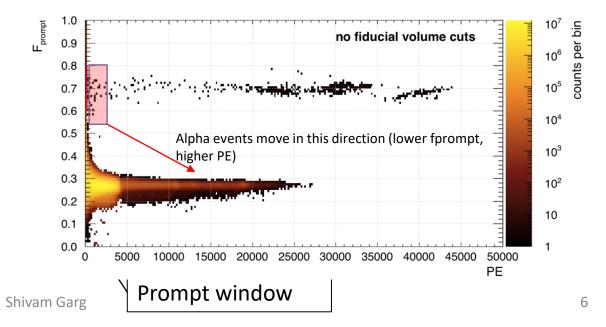


 Pyrene is a slow wavelength shifter slow = long time constant => prompt light will decrease => fprompt decreases



• Pyrene coated flowguides to be installed during hardware upgrades to mitigate neck alpha backgrounds

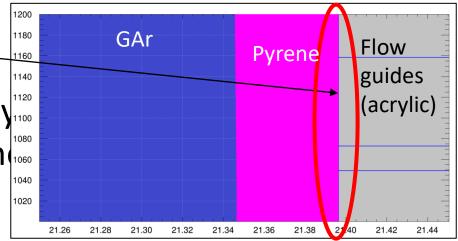






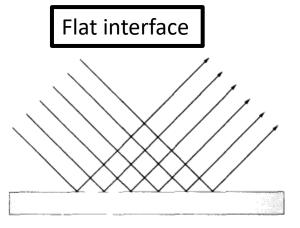


- Qualify the optics of components in the neck
- Relevant parameters
 - Surface roughness
 - Refractive index of pyrene/acrylic
- See Hicham's (R3-6) talk for more about the time constant and light yield
- Wavelength shifting time constant of py¹¹²⁰
 Wavelength shifting light yield of pyren
 - Scintillation light yield of pyrene
 - Scintillation time constant of pyrene
 - Opacity



UNIFIED model in GEANT4





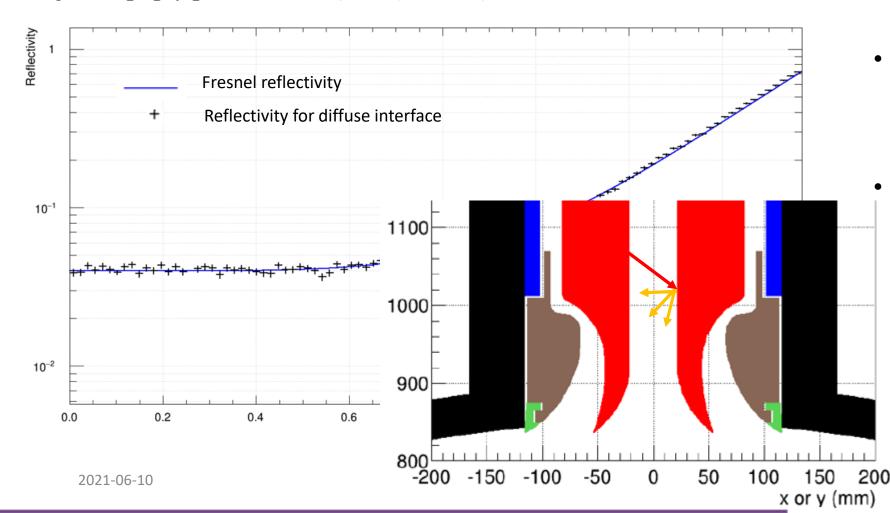
Rough interface Facet normal Global normal

Freshel equations $r_{\perp} \equiv \left(\frac{E_{0r}}{E_{0i}}\right)_{\perp} = \frac{n_i \cos \theta_i - n_t \cos \theta_t}{n_i \cos \theta_i + n_t \cos \theta_t}$

- UNIFIED model's main parameter σ_{α} (more about UNIFIED model here doi: 10.1109/NSSMIC.1996.591410)
- $\alpha \sim \text{Gaus}(\alpha; 0, \sigma_{\alpha}) * \sin(\alpha)$
- High $\sigma_{\alpha} =>$ more diffuse reflections Low $\sigma_{\alpha} =>$ more specular reflections

Why surface roughness matters





- Reflectivity curve changes significantly as a function of sigma_alpha
 - Changing reflectivity affects the path of photons through the neck (which ultimately changes the number of photons reaching the PMTs)

We measure reflectivity (transmittance), fit out sigma_alpha

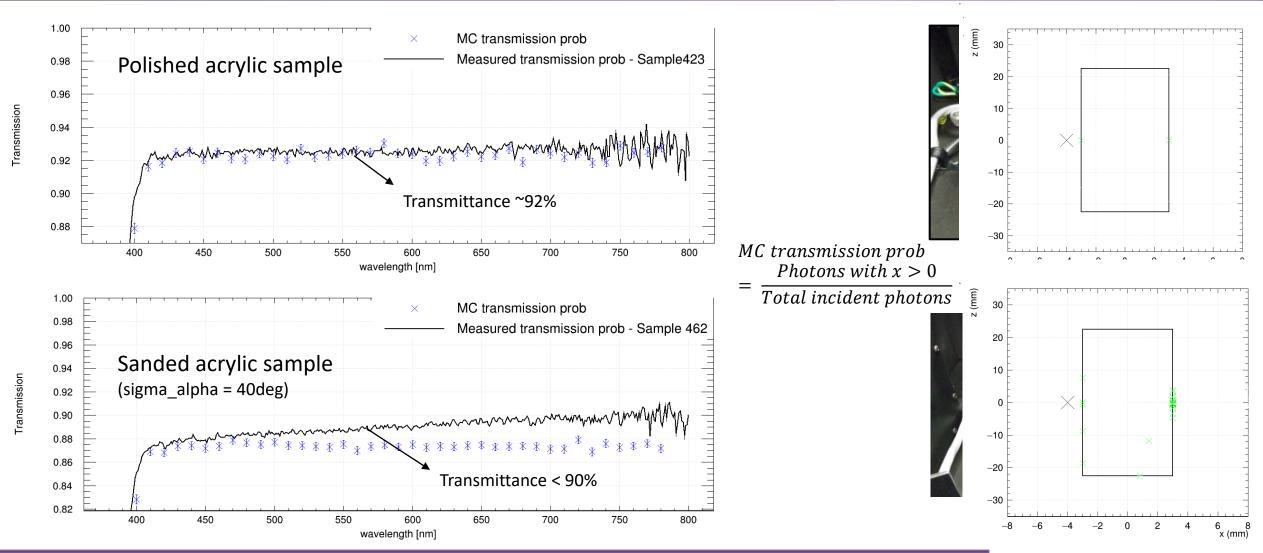
Spectrophotometer



• PerkinElmer Lambda 1050 spectrophotometer at COLD Lab, Carleton (Image credits – Jeff Mason) Integrating sphere module

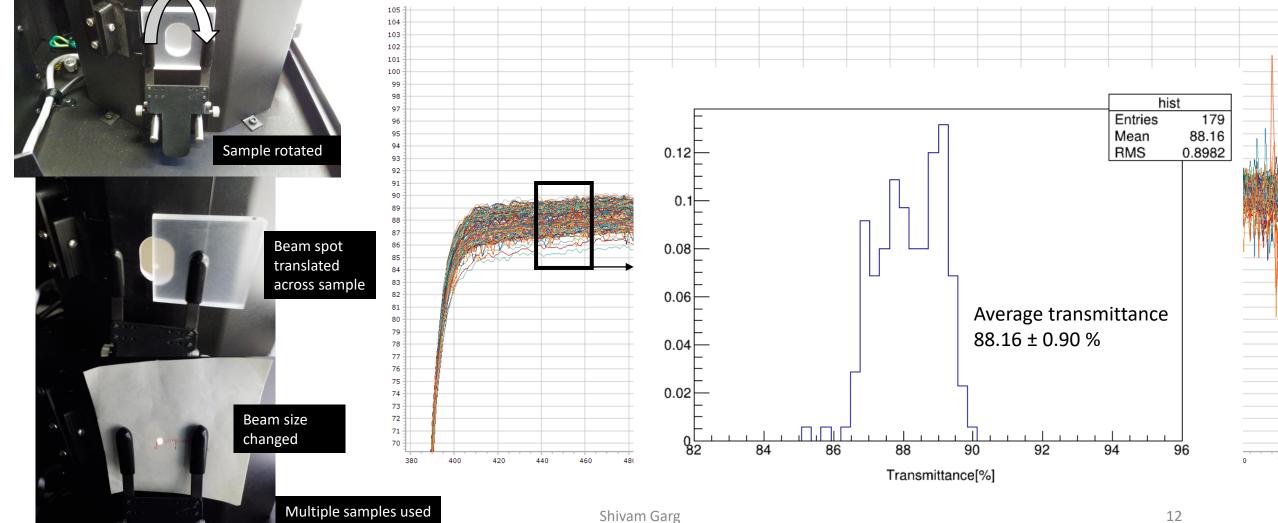
Finding sigma alpha





Finding sigma alpha

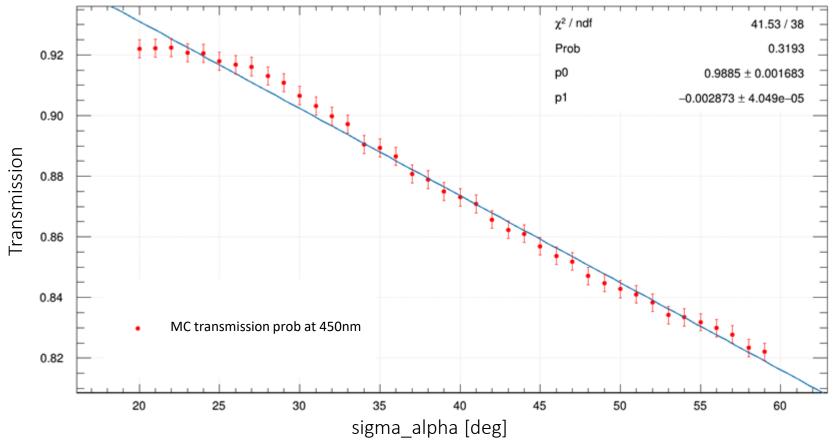




Finding sigma alpha



Transmission for sanded sample versus sigma_alpha (at 450nm)



- Simulated acrylic samples with different sigma_alphas
- sigma_alpha =
 37.2 ± 3.1 degrees





- Slow wavelength shifter on DEAP neck flowguides being implemented as part of hardware upgrades to mitigate the dominant detector background (neck alphas). Allows powerful discrimination of neck backgrounds
- Developed detailed optical model of surfaces and benchmarked at COLD Lab, Carleton
- Preliminary sigma alpha value = 37.2 ± 3.1 degrees describes the acrylic surface well
- More sample data needed to study systematics (ongoing)
- Full optical model with uncertainties will be used to characterize detector backgrounds
- Thanks for listening!