



Analysis status WOM

- Physics and Engineering of the WOM
- Problems
- Bioluminescence: Lomb Scargle analysis in frequency space
- Bioluminescence: statistical analysis of rise and fall times
- Base rate and 40-K
- A glimpse at analyses of „special data sets“
- Summary

P-One Meeting
Dec 14-15, 2020

Lutz Köpke
University of Mainz

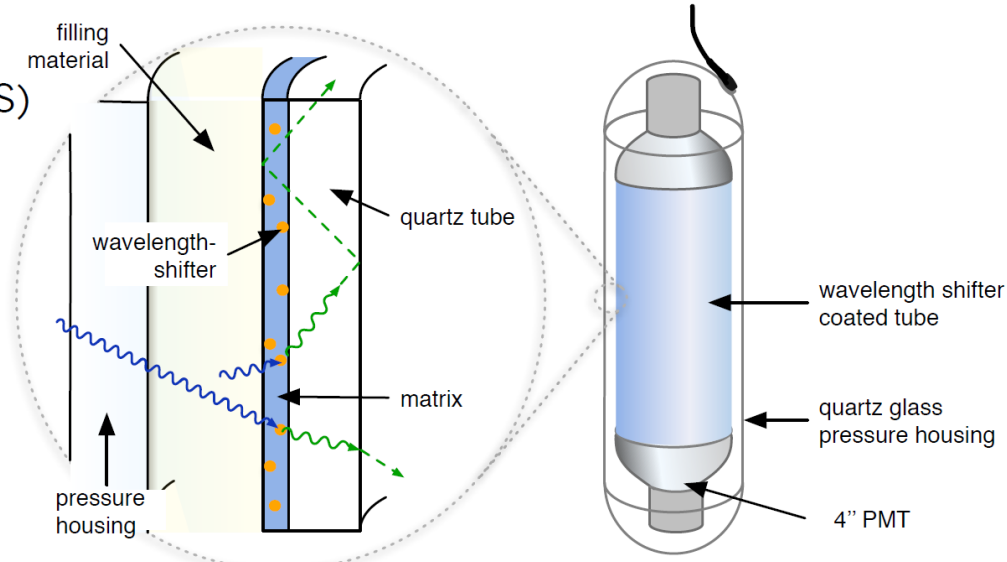
THE PHYSICS IDEA BEHIND THE WOM ...

Basic concept

- Wavelength shifters (WLS)
 - ▶ concentrate light

Features

- large collection area
- better UV sensitivity
- low noise rate (few Hz)
- cost effective



Disadvantages

- Larger absorption length
- Somewhat worse timing

n=1.33	n≈1.5	n=1.33	n≈1.5	n=1.0
ice	quartz	opt. fill	quartz	air

3.5"

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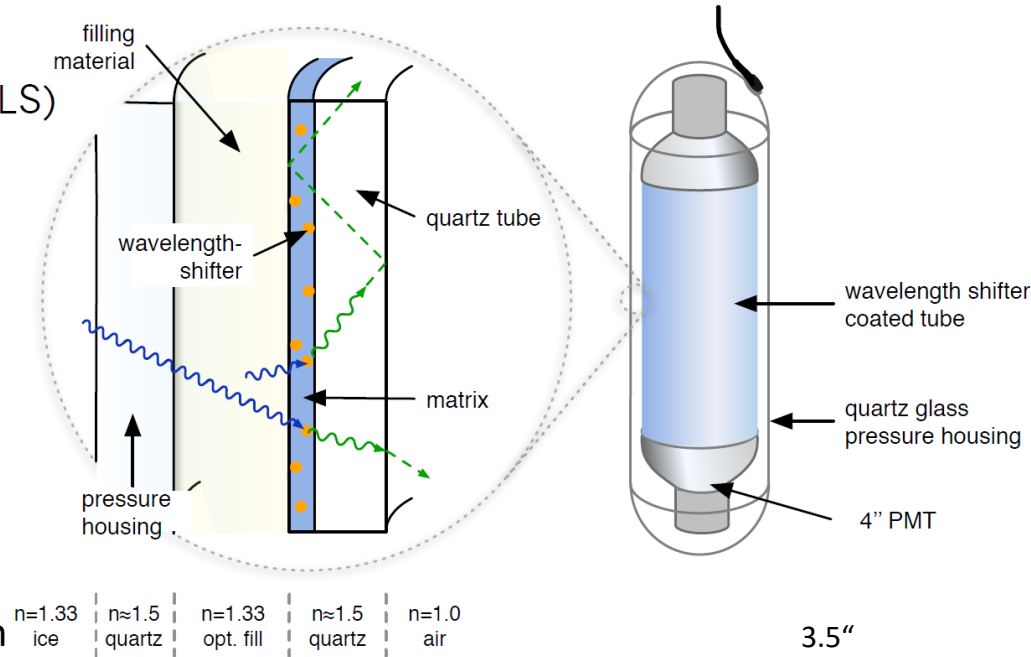
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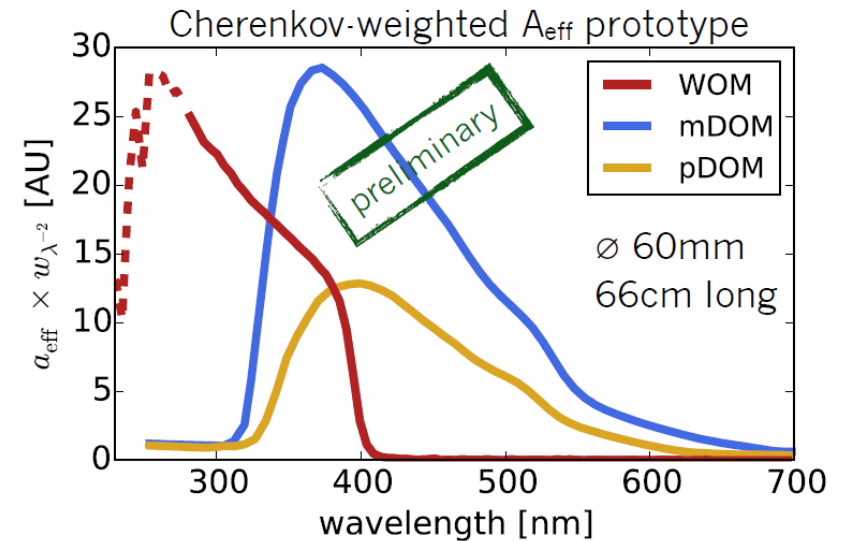
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Effective area comparison between

- straw-b prototype,
- 10" IceCube and
- 24-PMT I3-upgrade optical module



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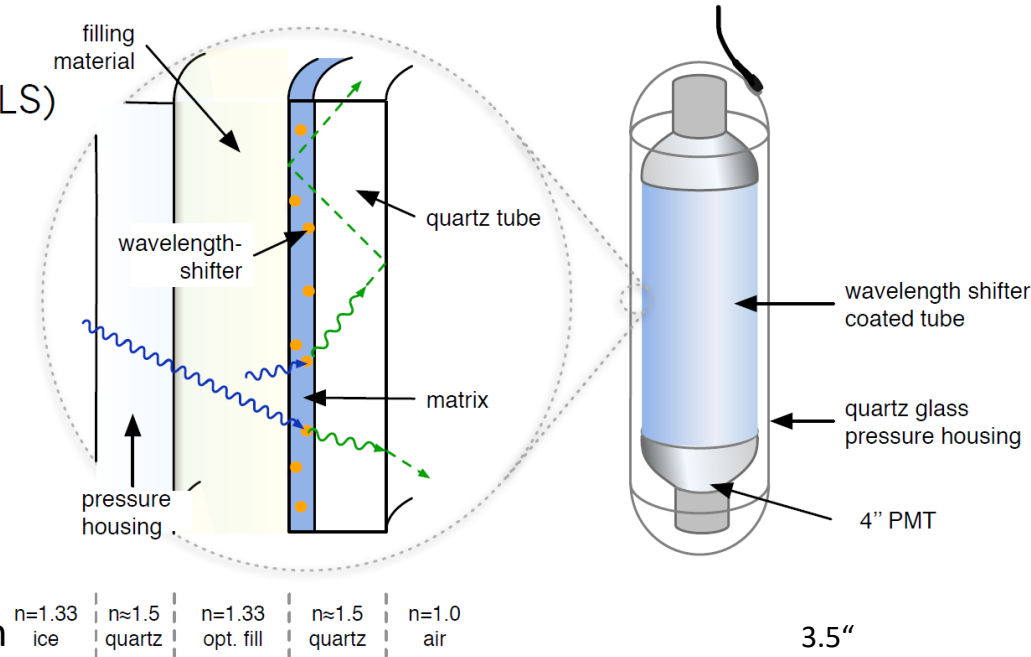
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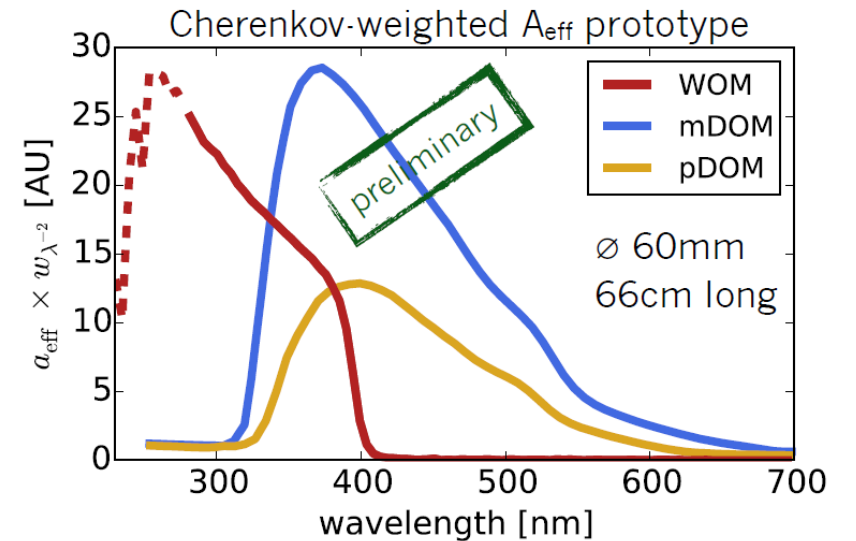
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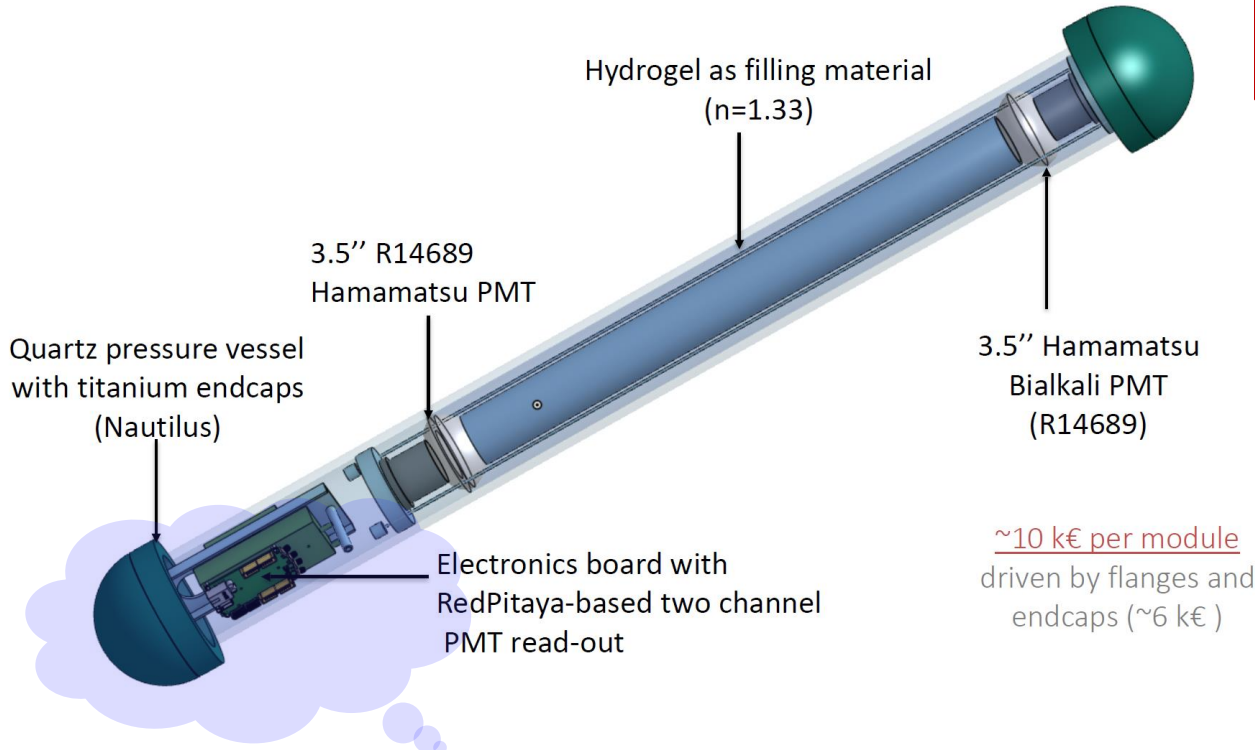


Why is test in STRAW-b important?

- First deployment of WOM: understand properties in situ (pressure/rates/effective volume ...)
- Study bioluminescence in UV regime (pretty much uncharted ...)

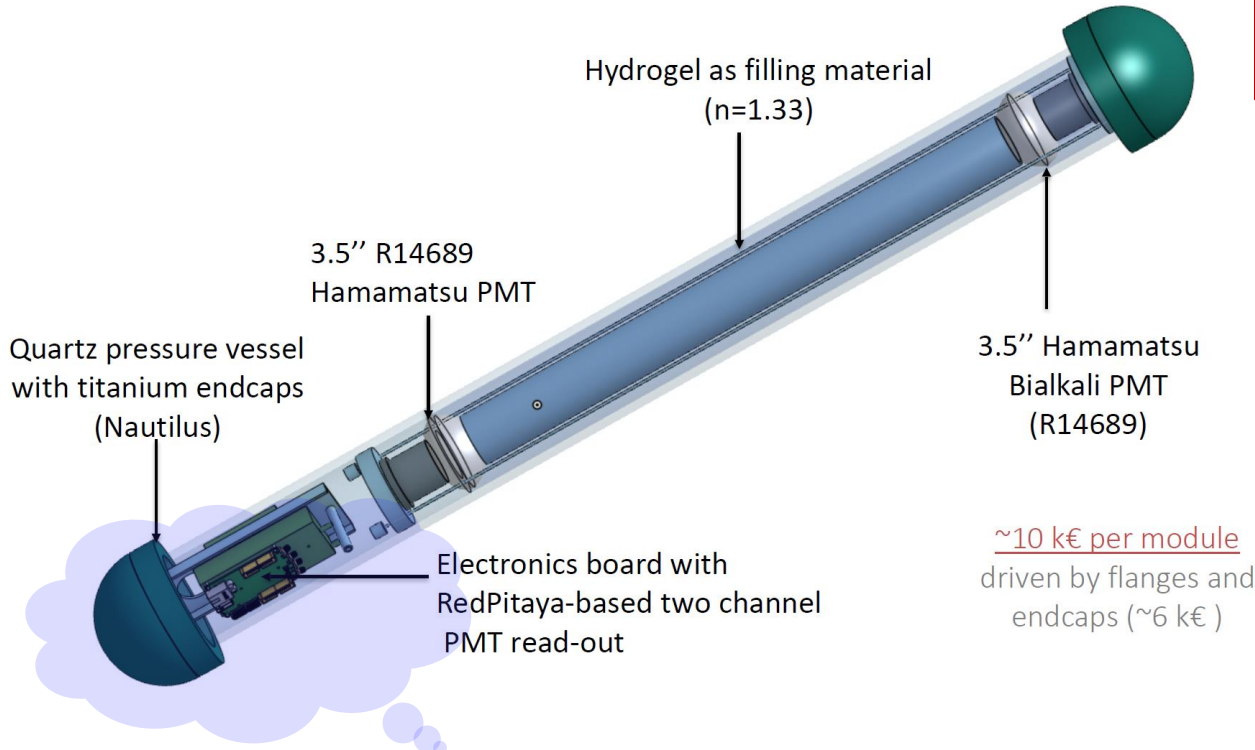
THE ENGINEERING PART

~1.5 m overall length
~0.7 m wave length shifter
~0.1 m outer diameter



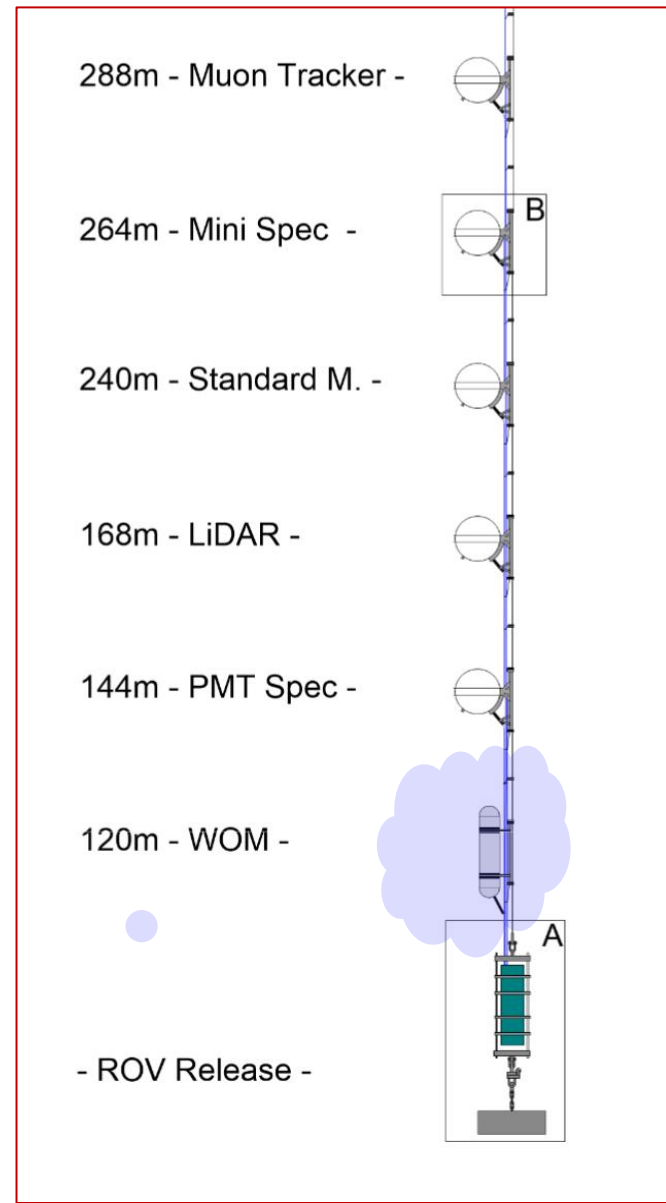
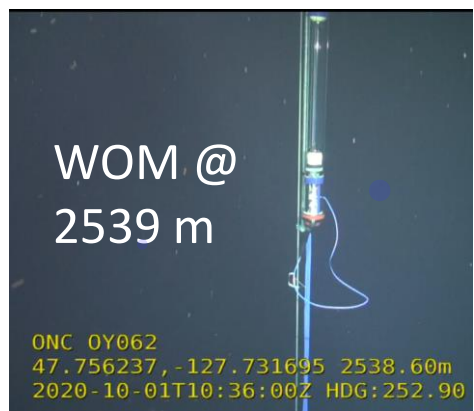
~10 k€ per module
driven by flanges and
endcaps (~6 k€)

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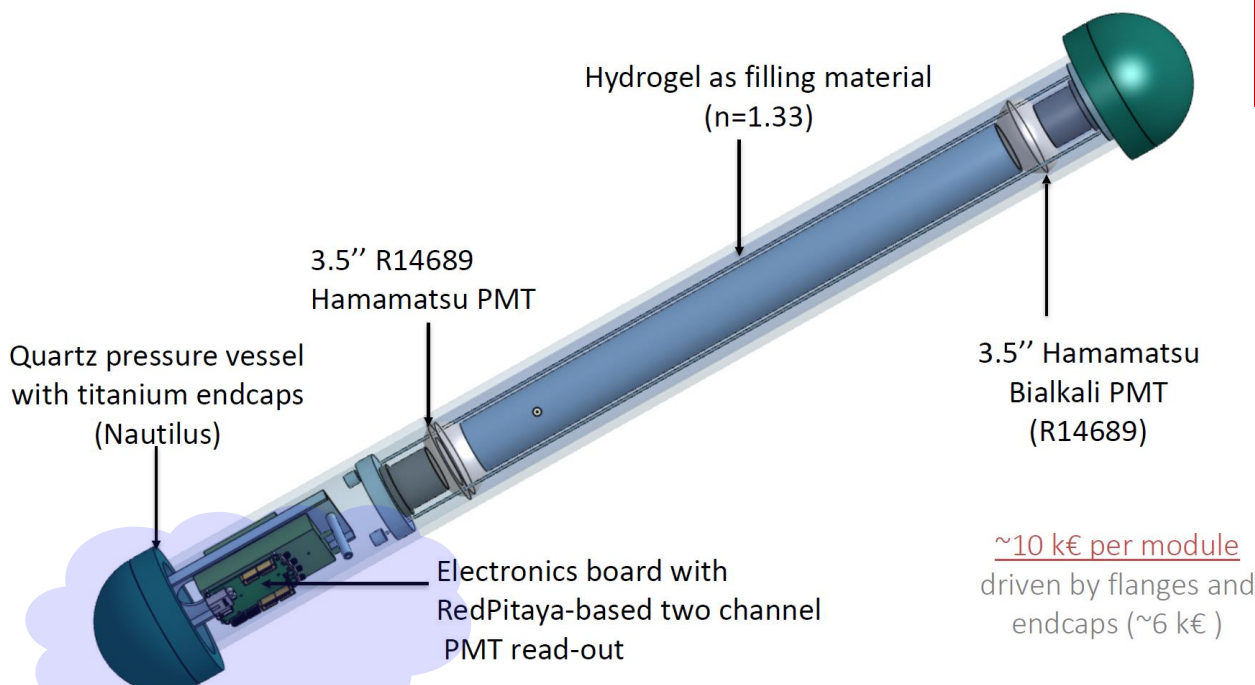


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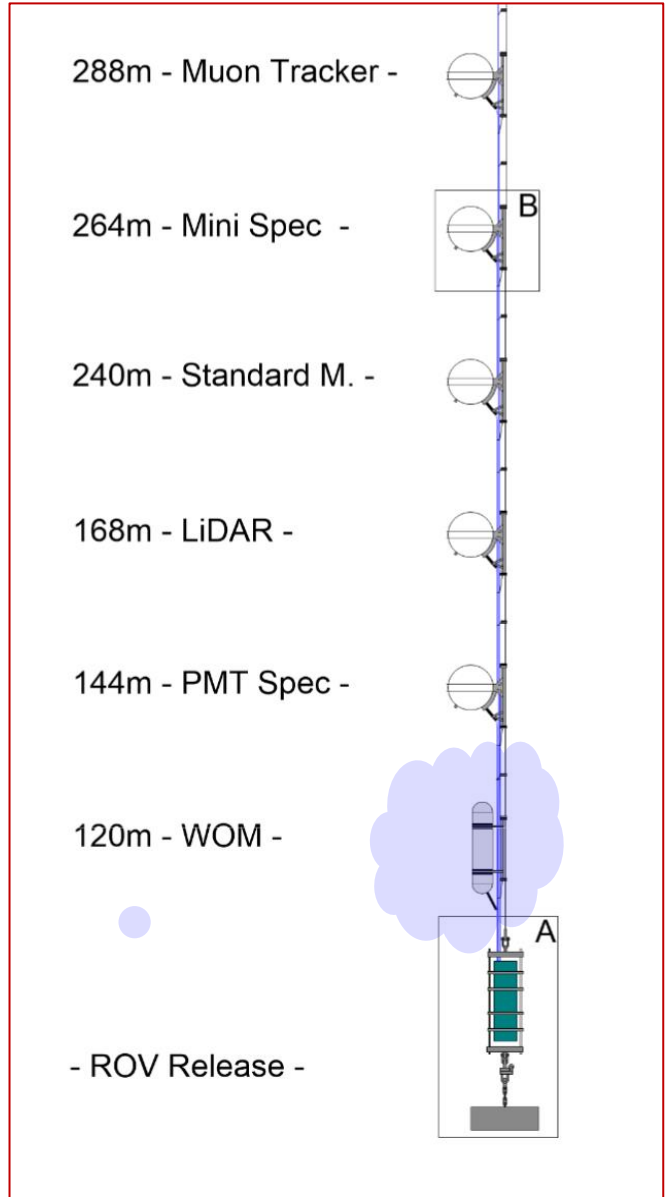
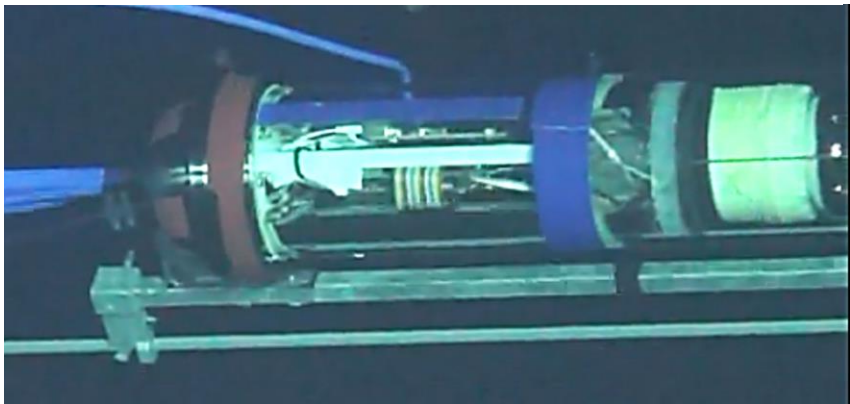
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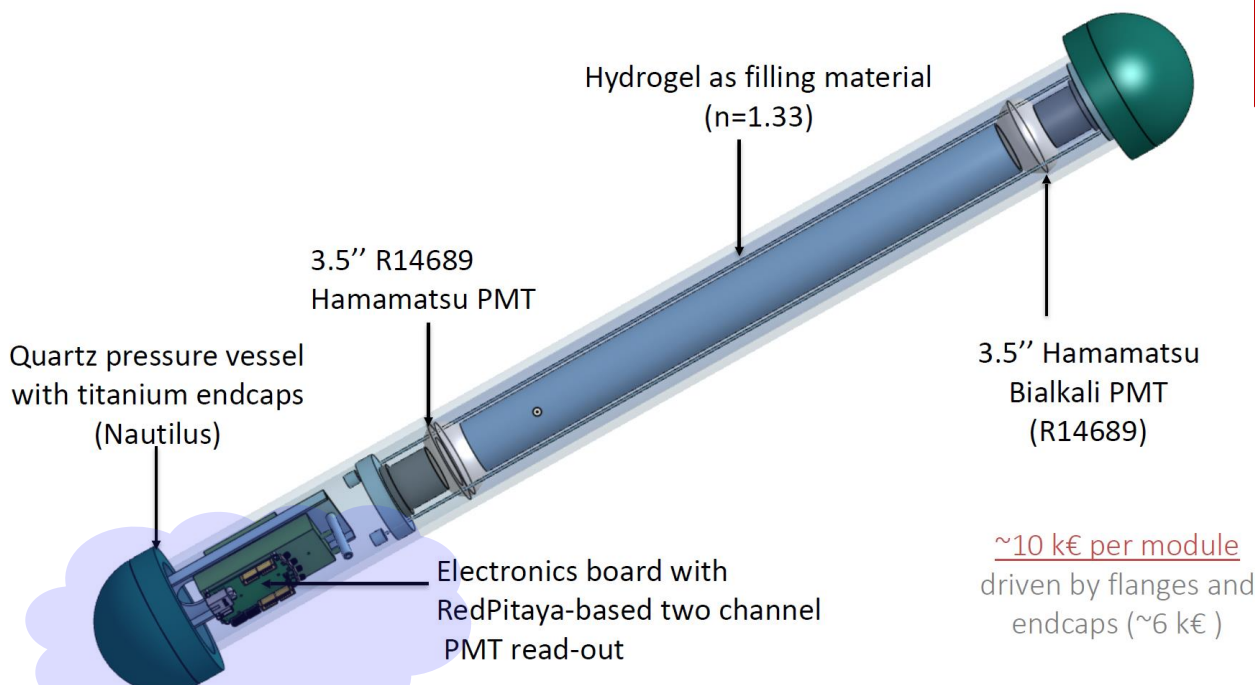
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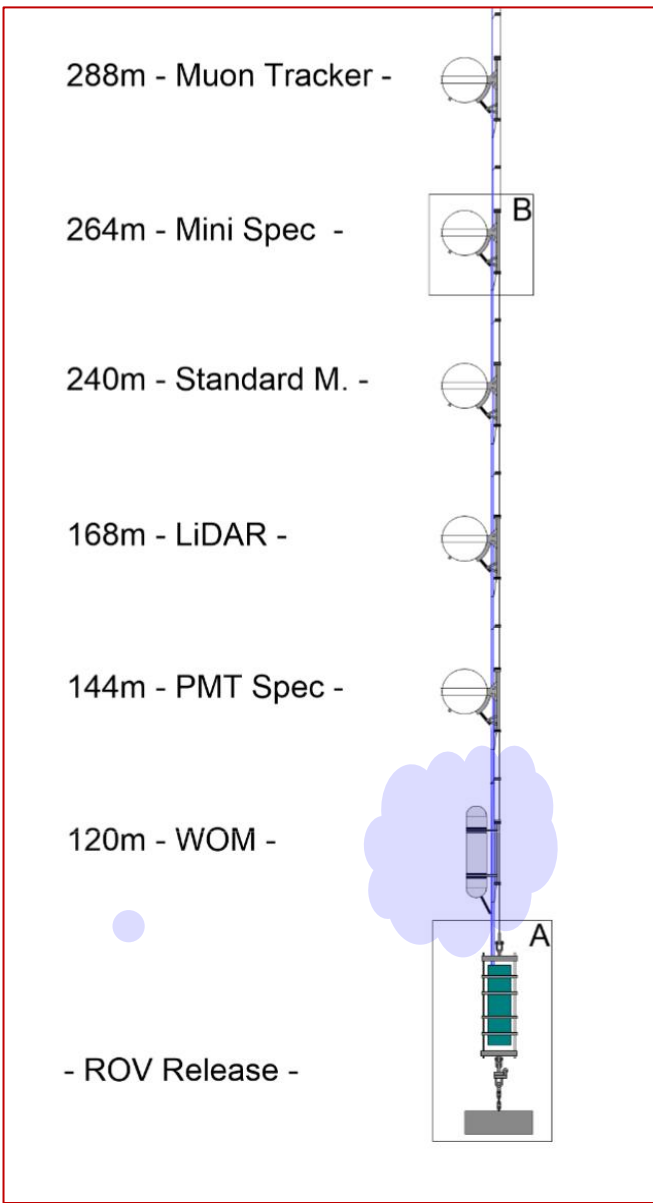
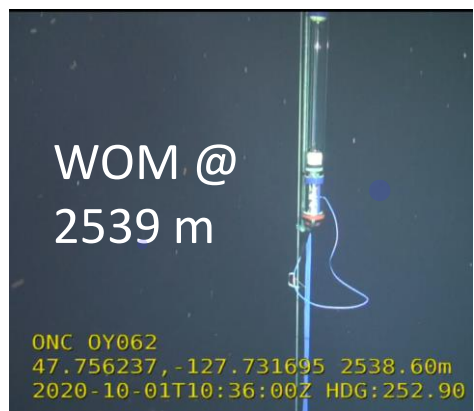
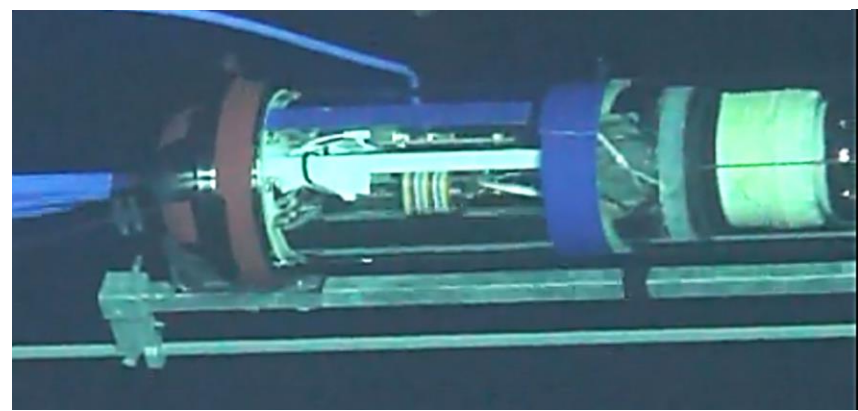


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Zynq/FPGA readout
 pretty configurable:
 →rate/window/threshold
 + coincidence mode

~10 k€ per module
 driven by flanges and endcaps (~6 k€)

enhanced WOMs
 to be deployed in
 IceCube upgrade

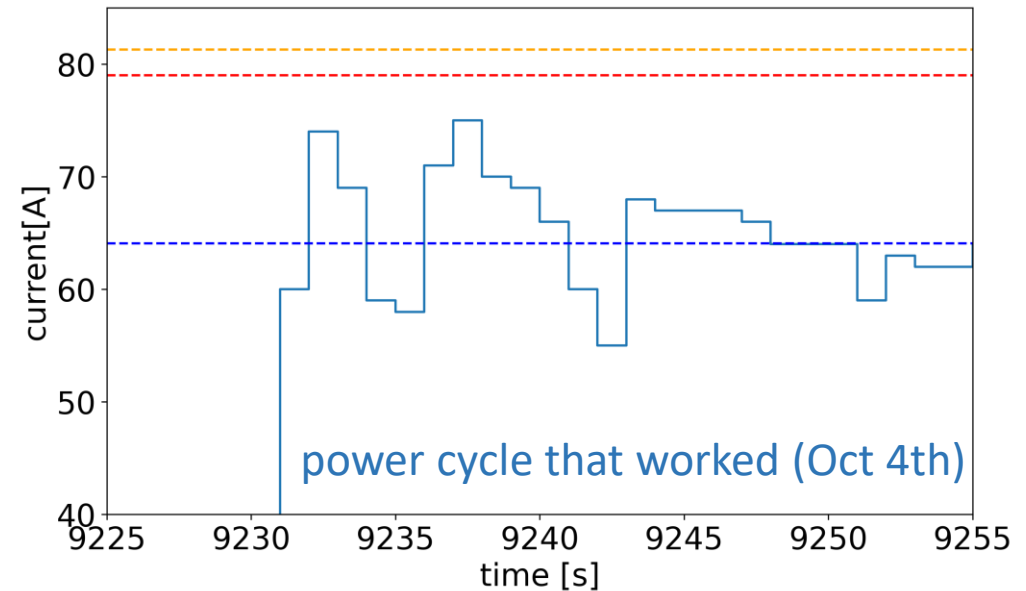
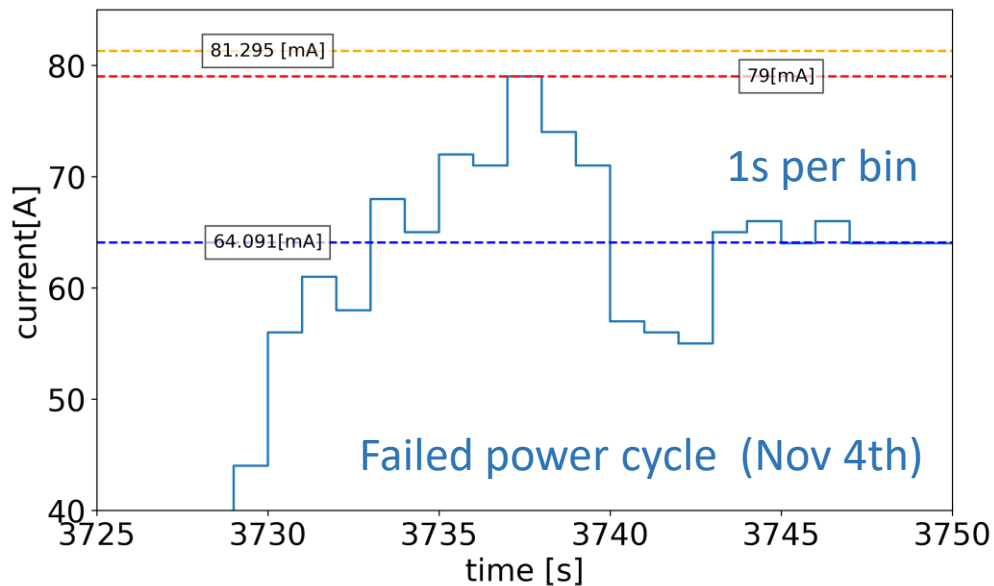


SEVERE PROBLEMS

August: connection to PMT 2 broken (observed in sat water tank test); decided not to repair

Oct 30: Power tripped building main (Port Alberni)

- WOM is no longer responsive to pinging; several attempts to power cycle failed
- WOM still seems to boot (detailed investigations of currents and boot sequence)
- no success with rapid pinging (0.2 s) during power cycle

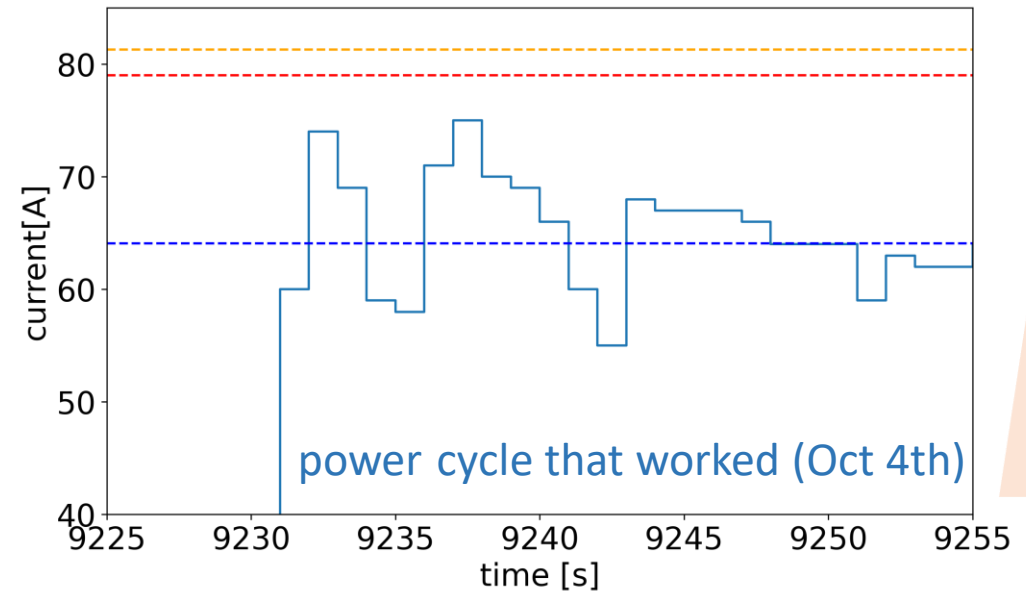
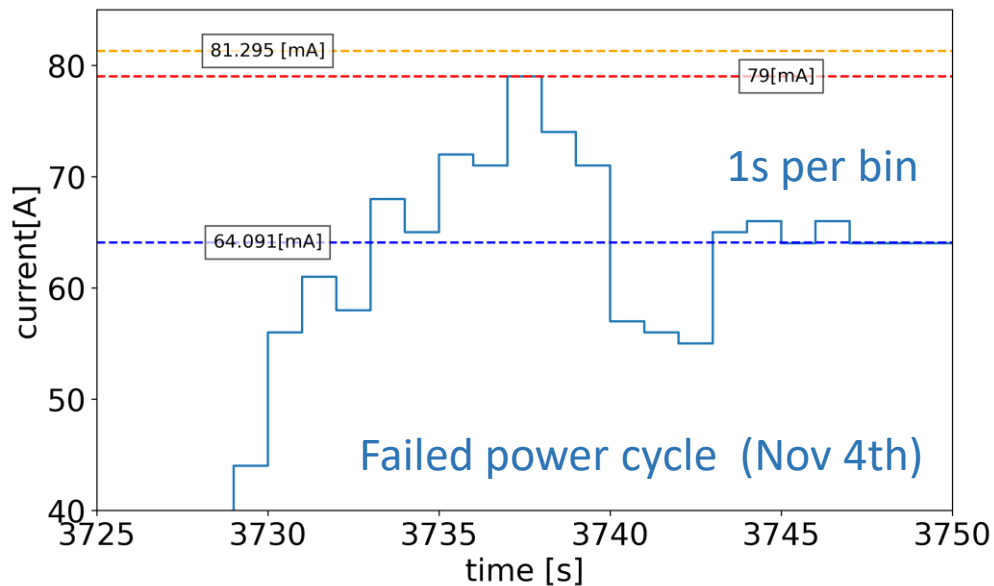


SEVERE PROBLEMS

August: connection to upper PMT 2 broken (seen in salt-water tank test); decided *not* to repair

Oct 30: Power tripped building main (Port Alberni)

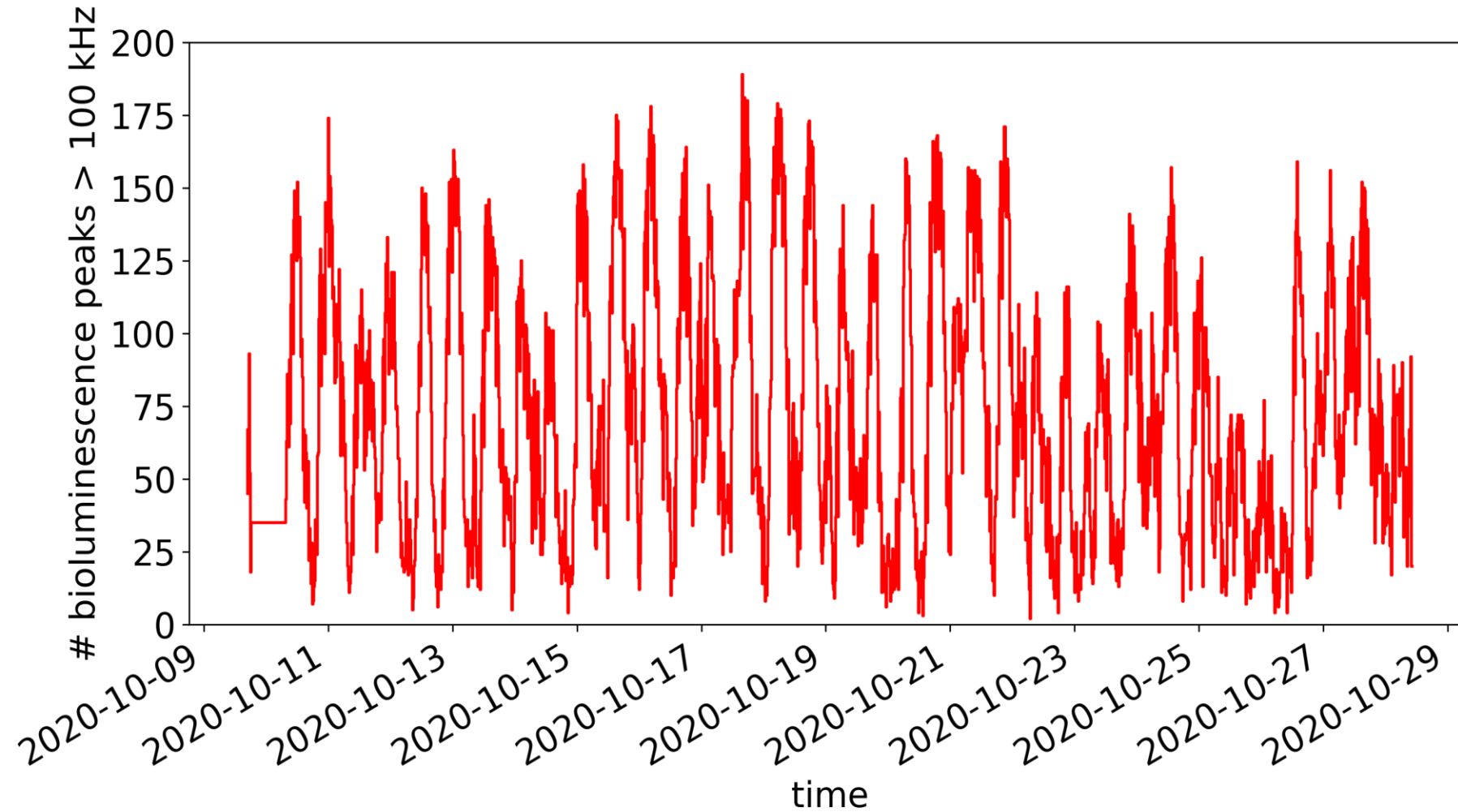
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Left with 18 days of data taken

DATA ANALYSIS

MEASURED RATE OF BIOLUMINESCENCE (18 DAYS)

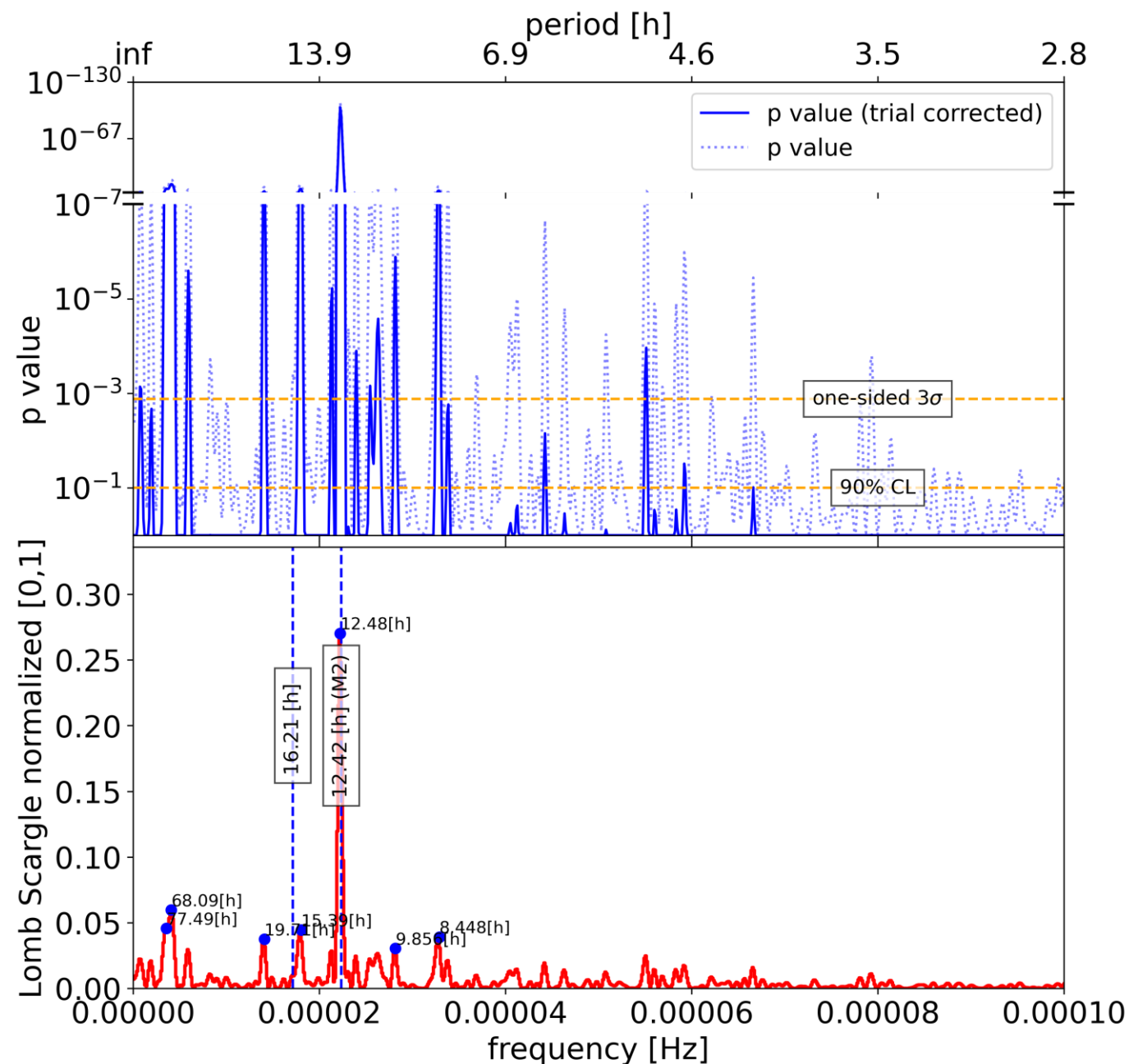


- **Current induced oscillations + bioluminescence variations**
- Latitude ODP 1027C: 47.7567°
- Inertial cycle (Coriolis):
→ $12\text{h}/\sin(\text{latitude})=16.21\text{ h}$
- semidiurnal lunar tide M2:
→ 12.4206012 h
- + many more frequencies!
(beats, harmonics, sun ...)

FREQUENCY SPACE

- Lomb-Scargle transformation
- Estimate p-value [astropy]
- Trial correction still analytical
→ use Toy MC in future

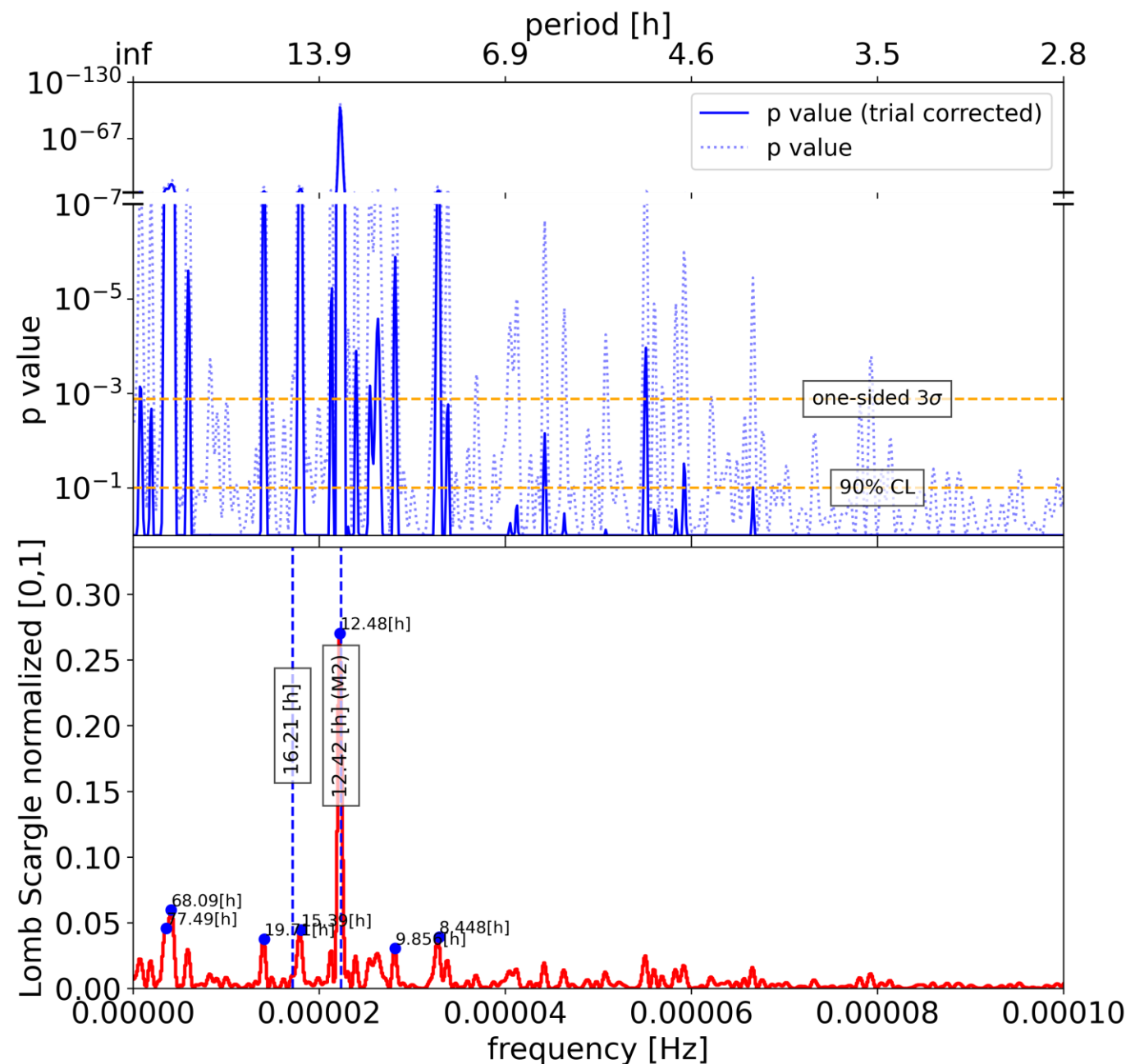
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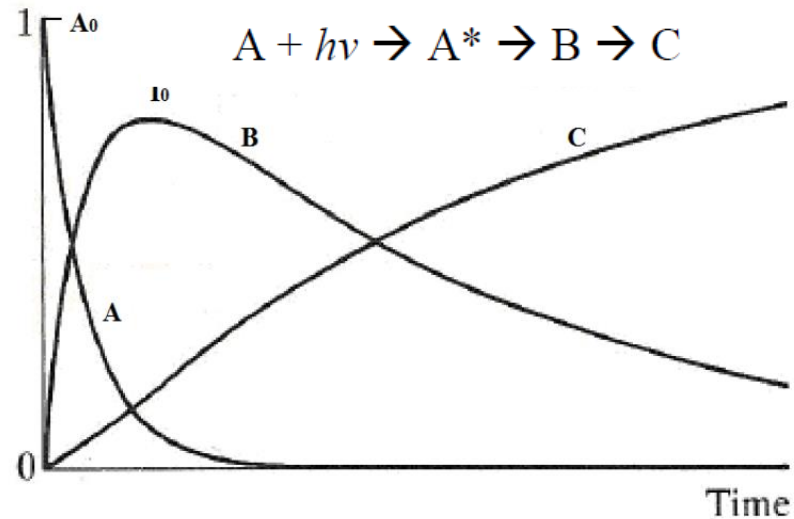
source	T [h] Pred.	T[h] meas.	Beat $f_A - f_B$ T[h]
lunar	12.42	12.42	64.4
inertial	16.21	15.39	



UNDERSTAND „CHEMISTRY“ OF BIOLUMINESCENCE

- Observation: some „light curves“ look rather clean, others are complex due to multiple emissions
- Fit „simple curves“ to minimal kinetic „ABC model“:

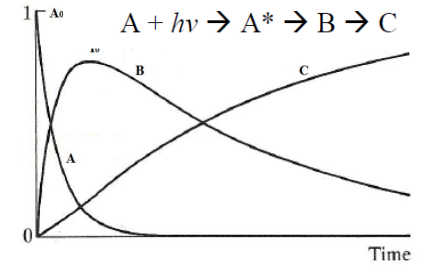
$$B = k_A \cdot A_0 \{ \exp(-k_A \cdot t) - \exp(-k_B \cdot t) \} / \{ k_B - k_A \}$$



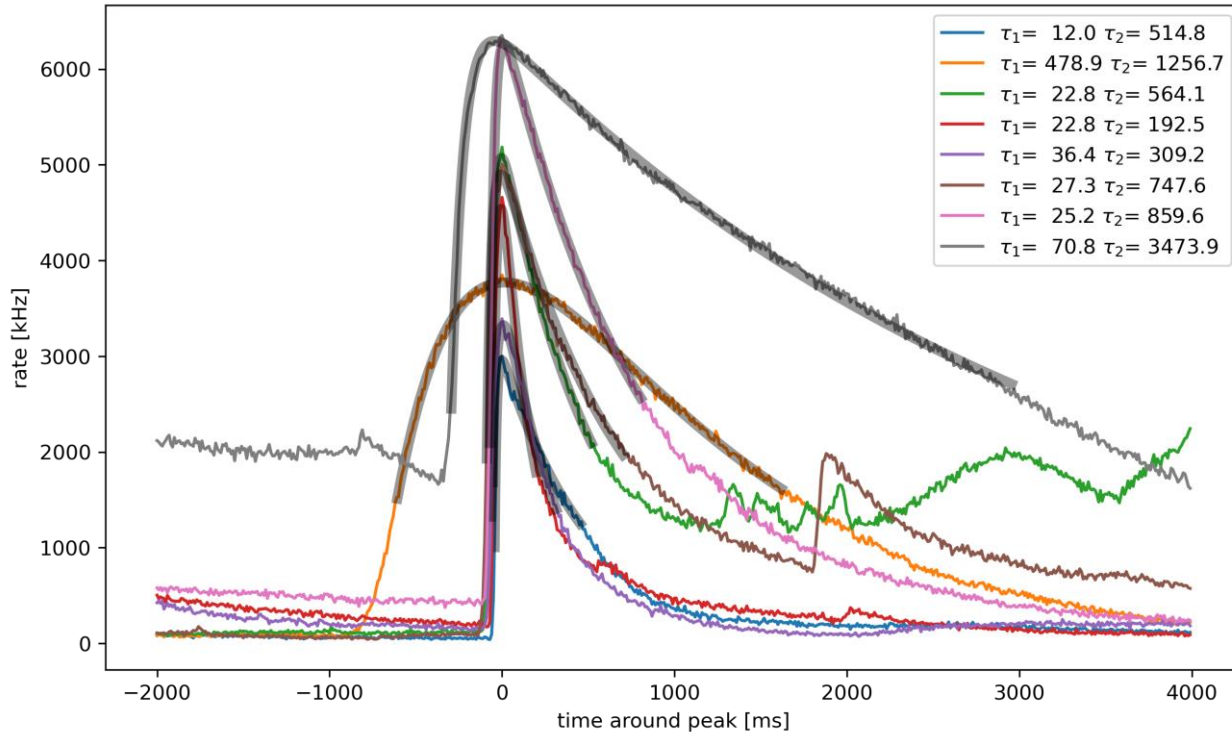
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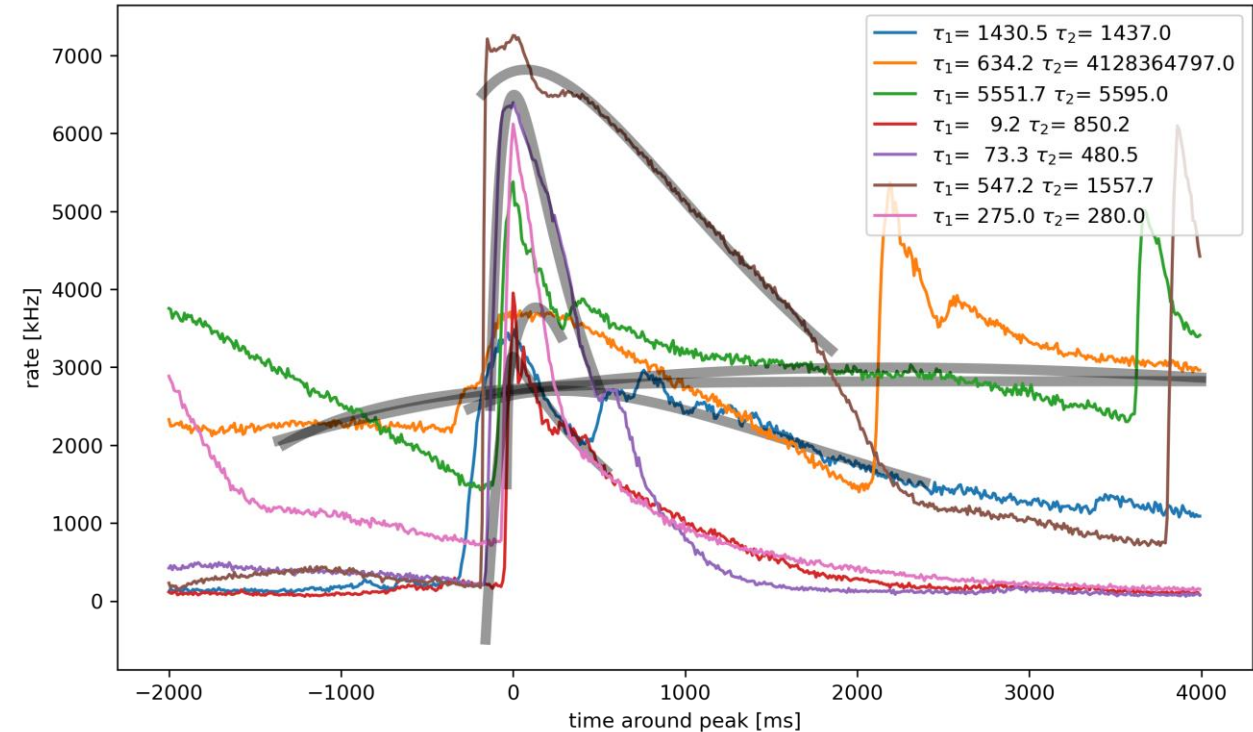
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„simple“ wave forms (good fit χ^2)

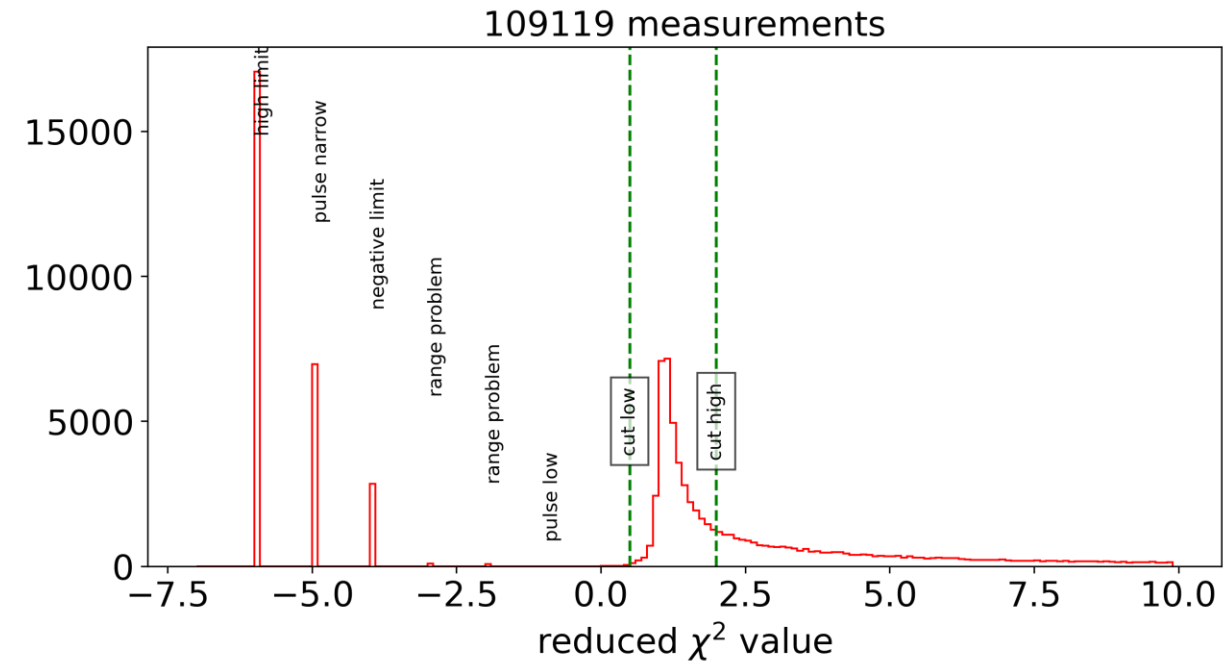


„complex“ wave forms (bad fit χ^2)



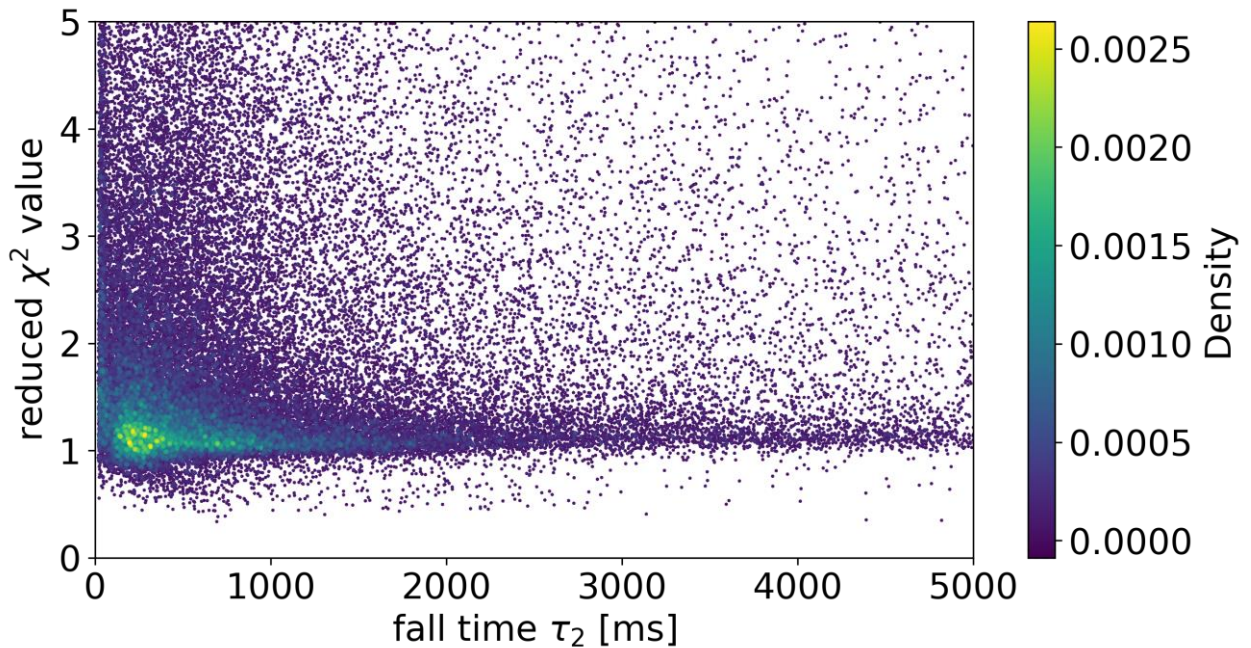
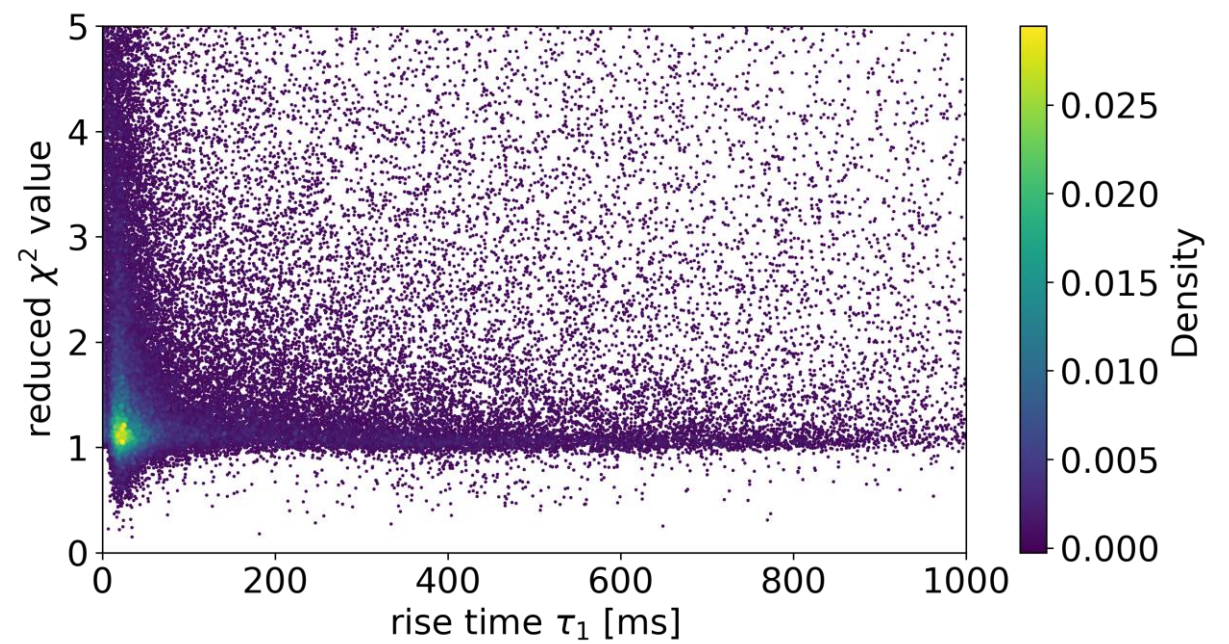
FIT RESULTS

- $\sim 30\%$ of fits have reasonable χ^2
- Some „edge classes“ rejected
- fits typically better for rise time τ_1 than for fall time τ_2
- some tail in χ^2 observed

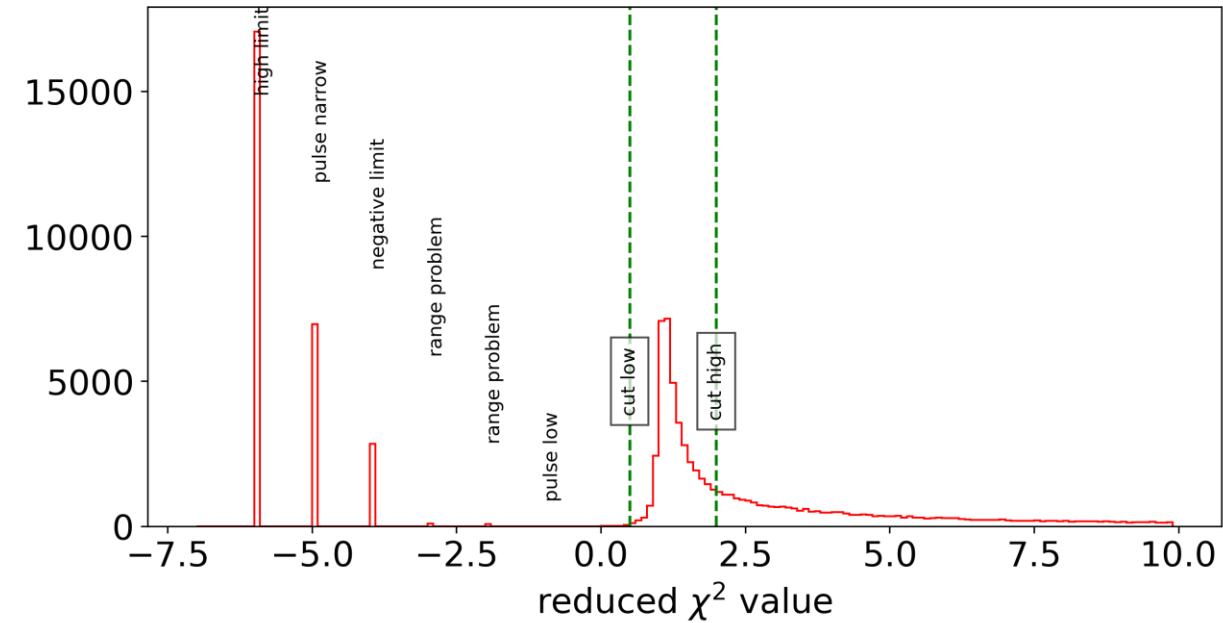


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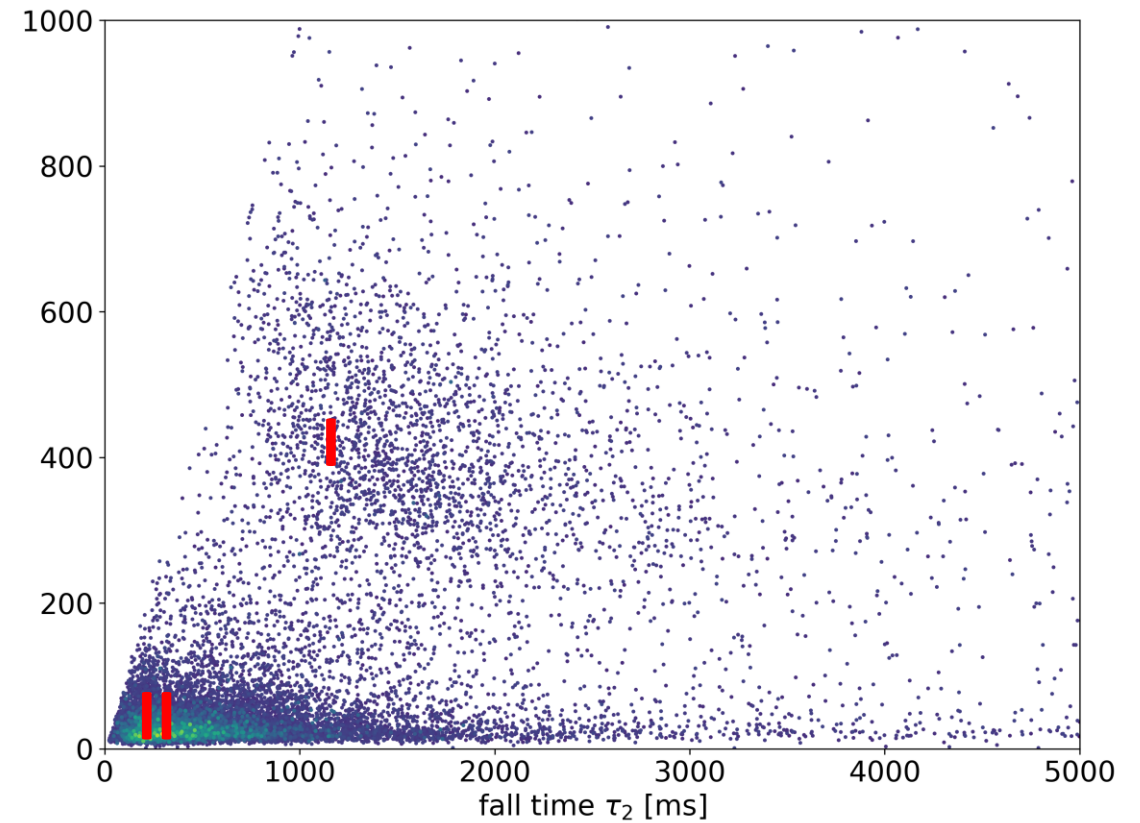
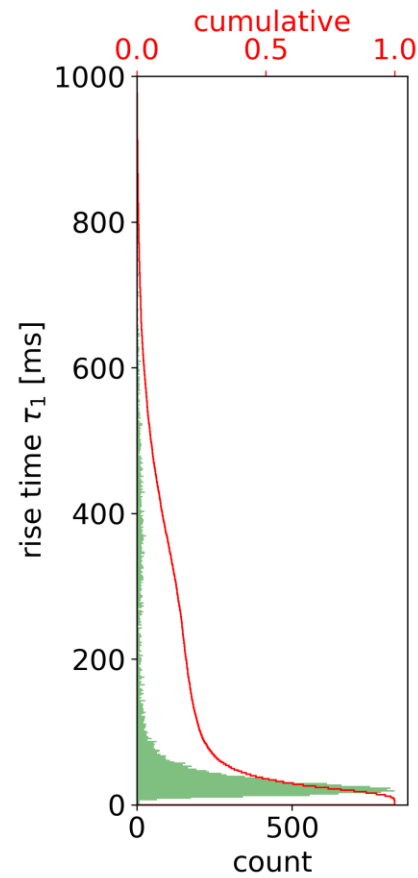
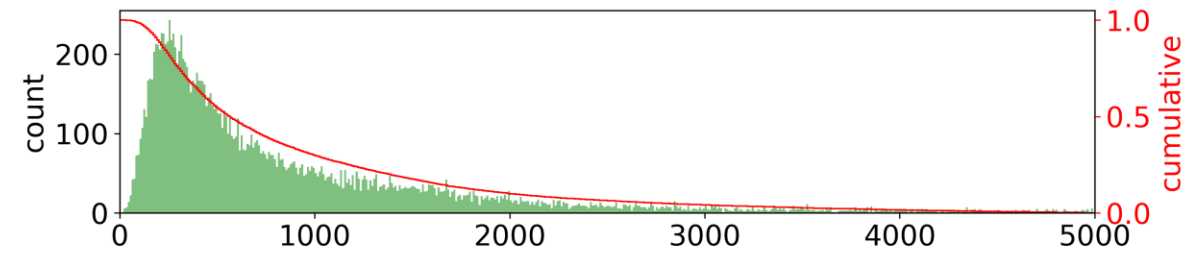
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109119 measurements

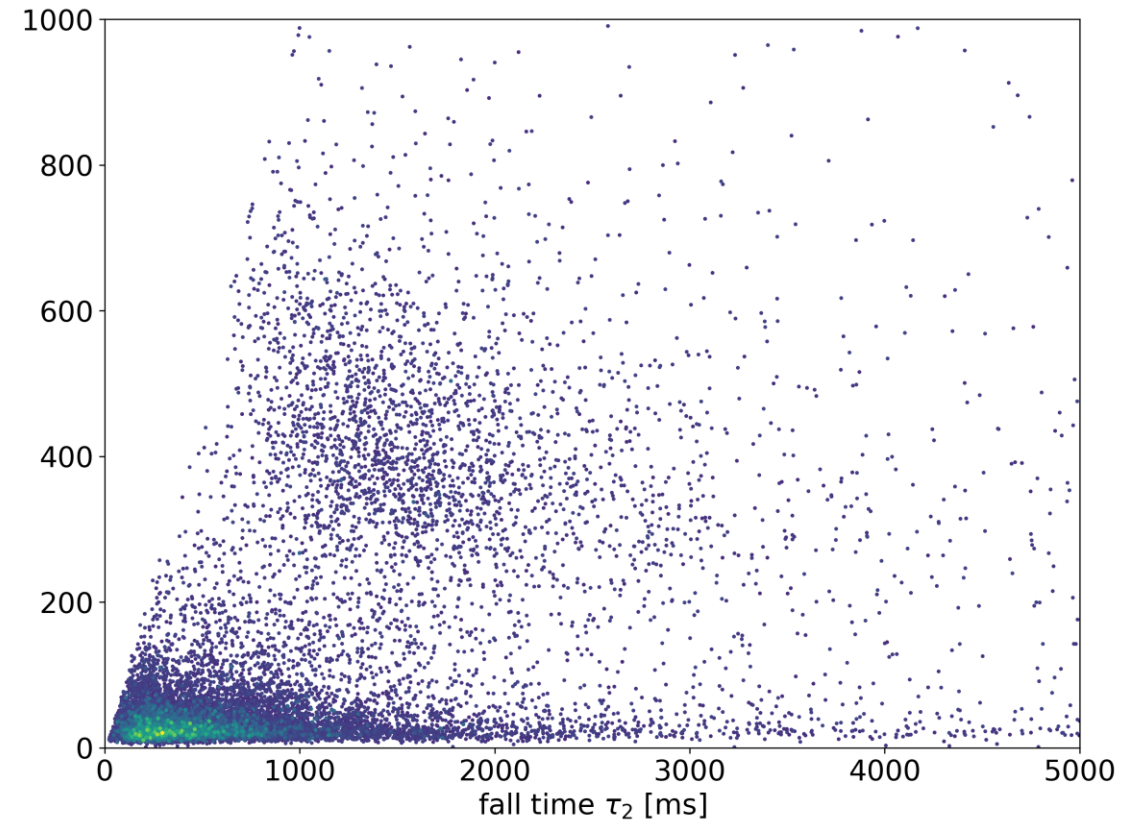
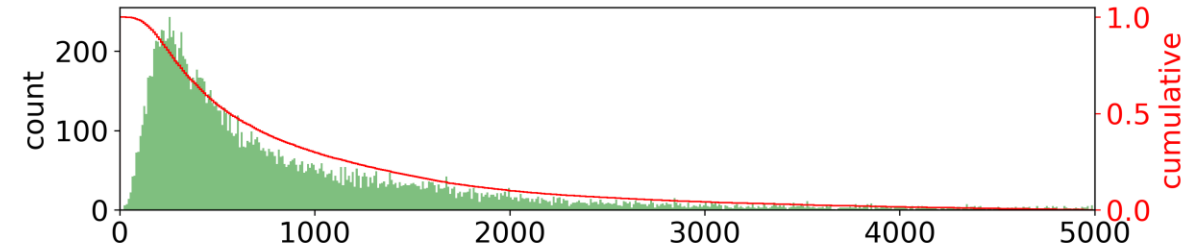
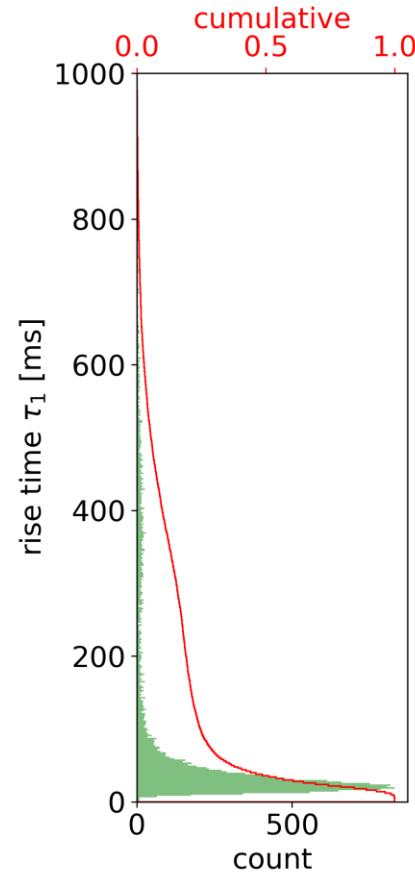


CORRELATION $\tau_1 - \tau_2$



CORRELATION τ_1 - τ_2

- From $\sim 10^5$ attempted fits, $\sim 15\%$
- selected for analysis (~ 16000 fits)
- **2 populations I, II:**
 - I: $\langle \tau_1 \rangle \sim 36$ ms, $\langle \tau_2 \rangle \sim 660$ ms
 - II: $\langle \tau_1 \rangle \sim 560$ ms, $\langle \tau_2 \rangle \sim 1950$ ms ($\sim 18\%$)
- Only seen in fits with good χ^2
- Method can still be improved

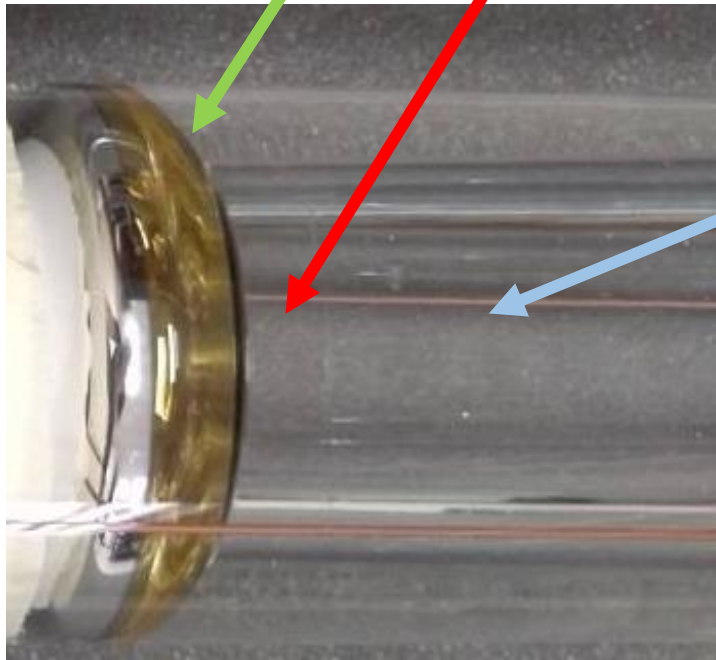


DO WE REALLY SEE LUMINESCENCE IN UV?

Direct light on PMT (~55 %)

Direct light on PMT for $\lambda > 400$ nm through WOM tube (~< 20%?)

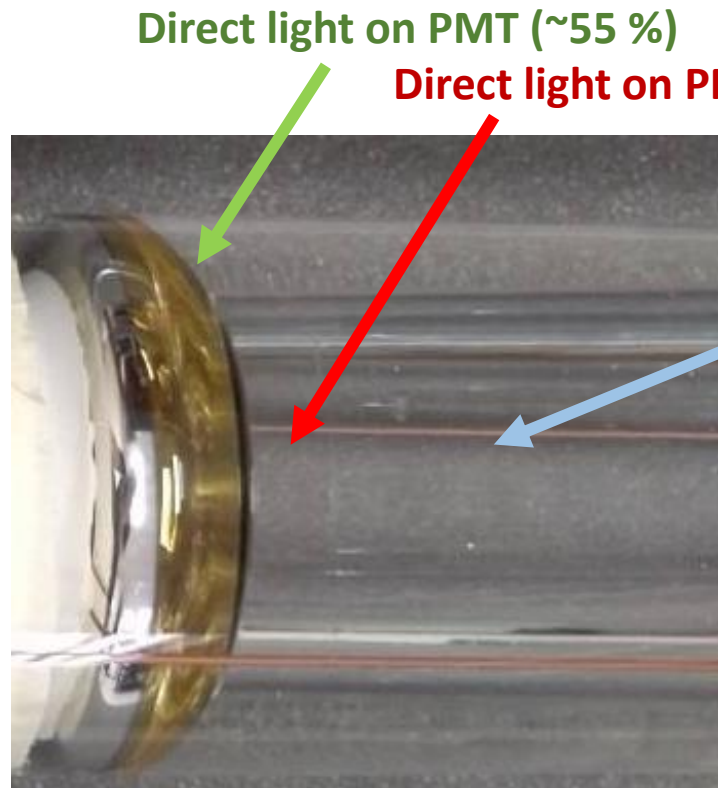
Light < 400 nm caught in WOM tube (per PMT ~400 %)



Active diameter 3.5" R14689 PMT :	81 mm
Length WOM tube:	630 mm
Diameter WOM tube:	55 mm

Expect roughly **4 x PMT acceptance through WOM tube per PMT**

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WOM quartz tube glued to PMT with NOA61 glue – micro cracks?

K-40 FOR ABSOLUTE CALIBRATION?

Implications of optical properties of ocean, lake, and ice for ultrahigh-energy neutrino detection

P. Buford Price

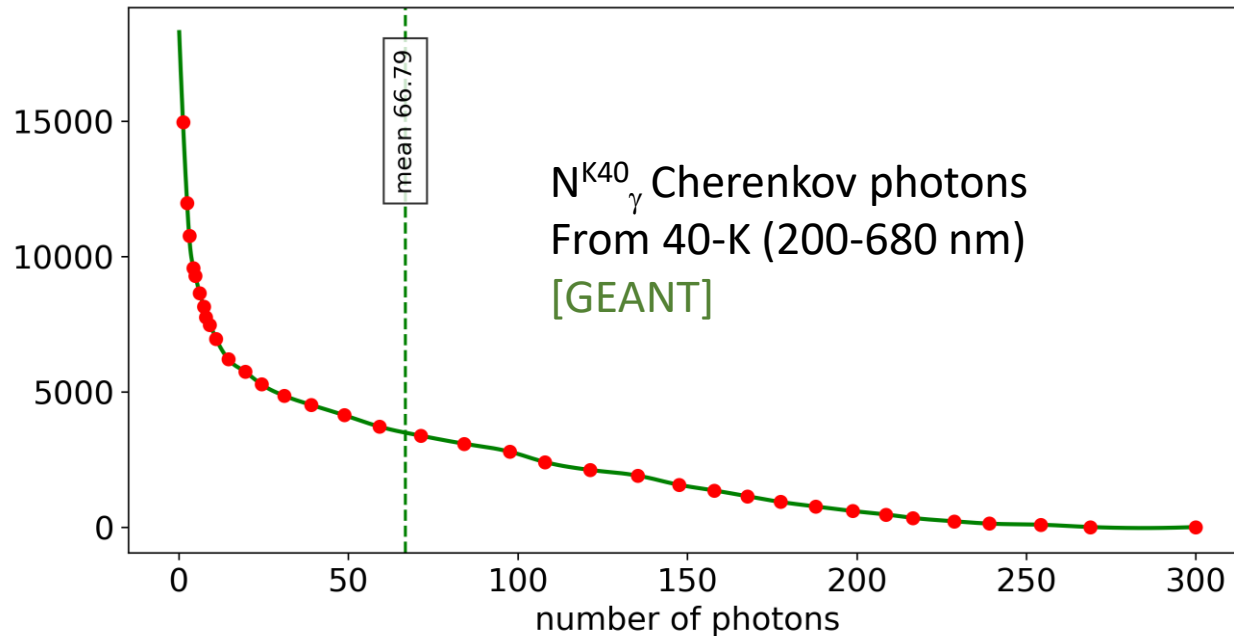
Constant density emitters ($\rho=11000/\text{m}^3$):

$$R_\gamma = N_{\gamma}^{K40} \rho A_{\text{eff}} / \pi a(\lambda) \times (1 - \exp(-a(\lambda) \cdot d_{\text{max}})) \quad [\text{absorption only}]$$

diffuse case more complicated

back on the envelop estimate

See also talk by Mathew this morning ...



K-40 FOR ABSOLUTE CALIBRATION?

Implications of optical properties of ocean, lake, and ice for ultrahigh-energy neutrino detection

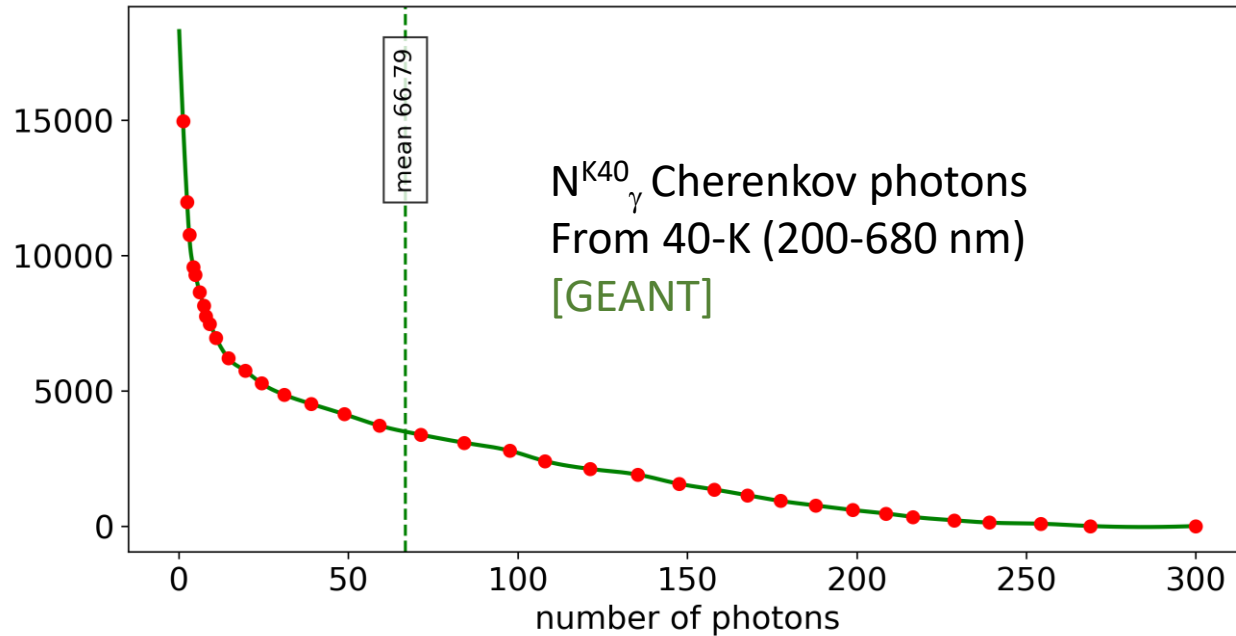
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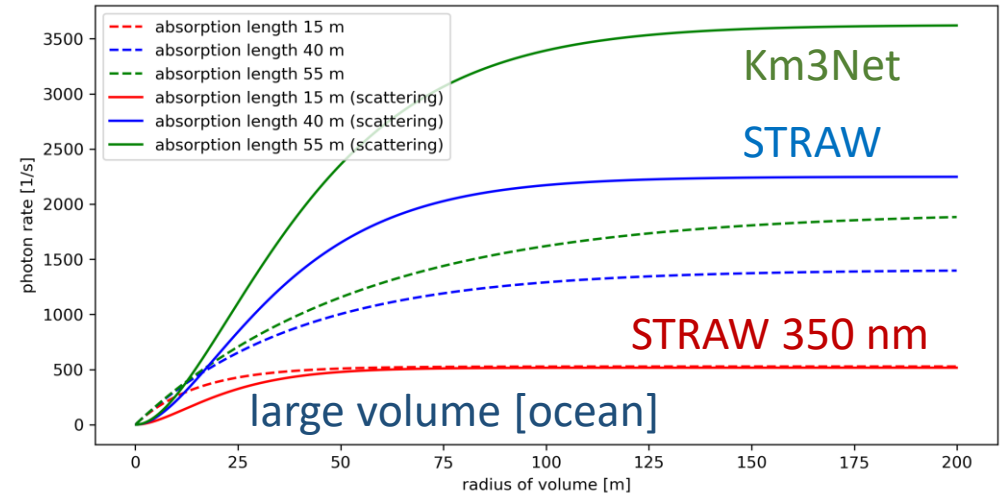
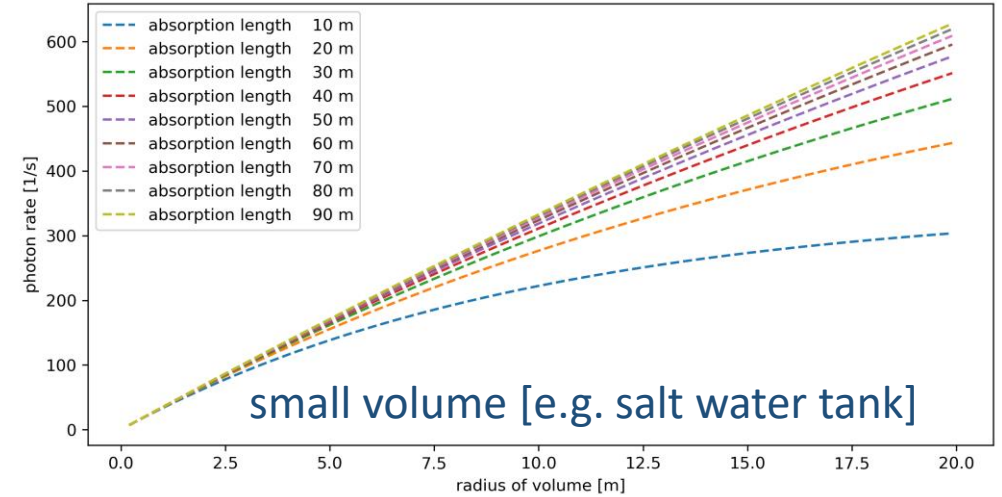
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diffuse case more complicated

Back on the envelop kind of estimate

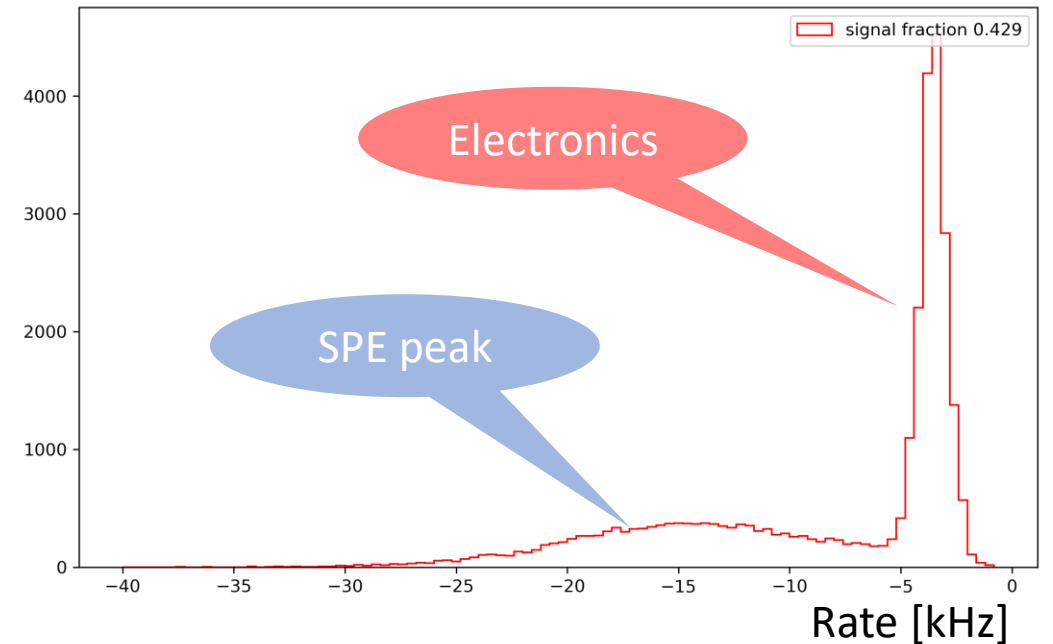
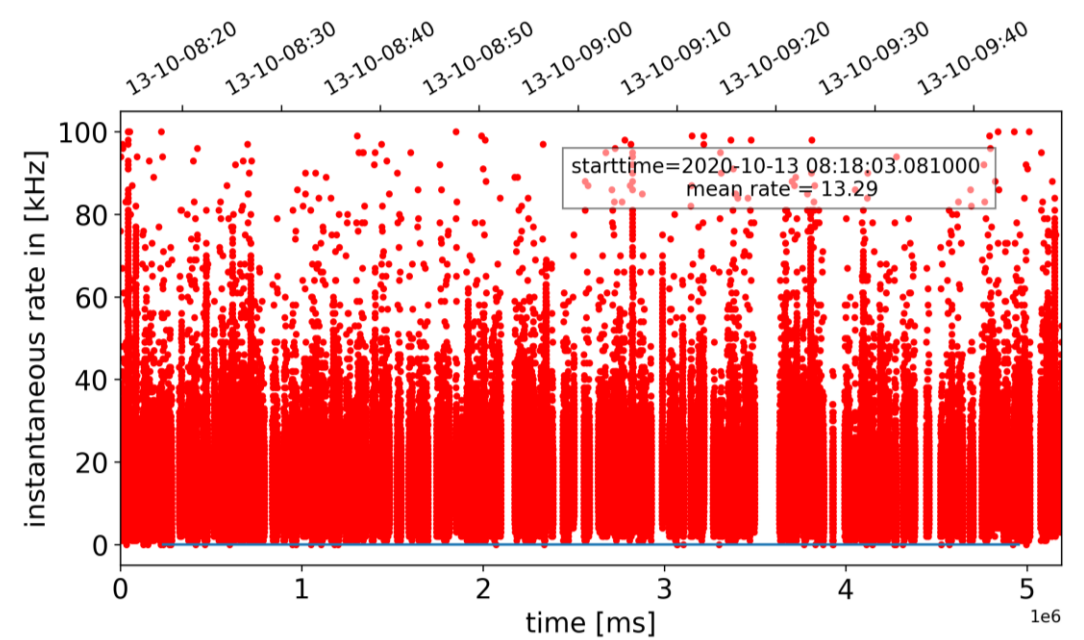
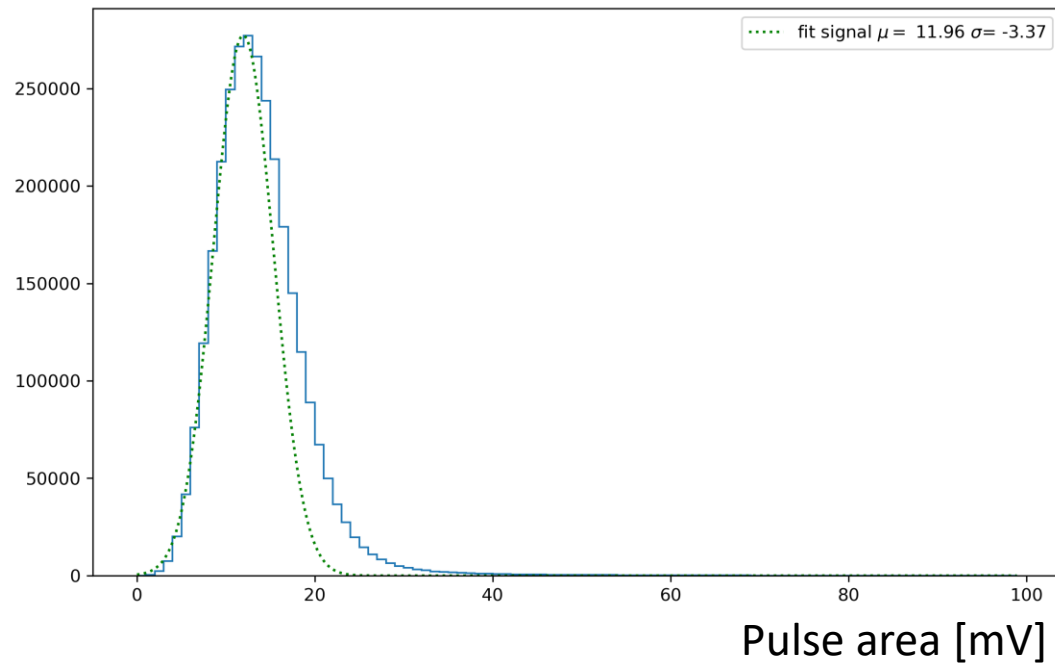


Expected rates for 3" PMT 11 kBq/kg

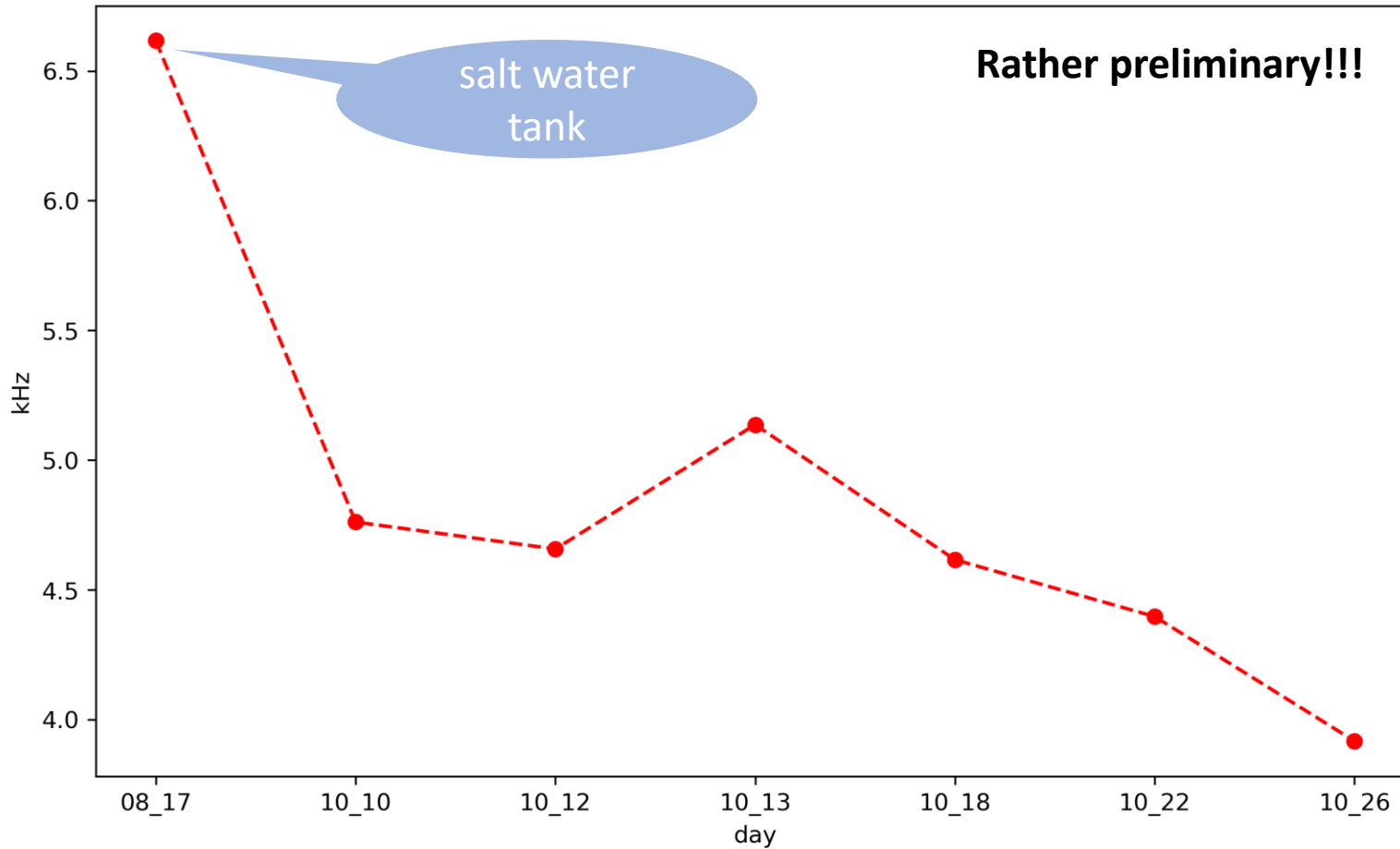


DETERMINE BASELINE

- Only investigate low bioluminescence runs
- De-select bioluminescence events
- Fit lower side of base rate with Gaussian
- Determine SPE fraction



40-K RATE + PMT-NOISE VERSUS TIME



where and what	K-40 [kHz]
Km3Net / 3" PMT [Geant]	4.12
STRAW / 3" PMT [-1.25 kHz noise]	3.8#?
WOM [-0.4 kHz noise]	4.3

Many open questions:

- Why is WOM rate so low (direct light)?
- Noise due to glass correct?
- Was tank light tight?

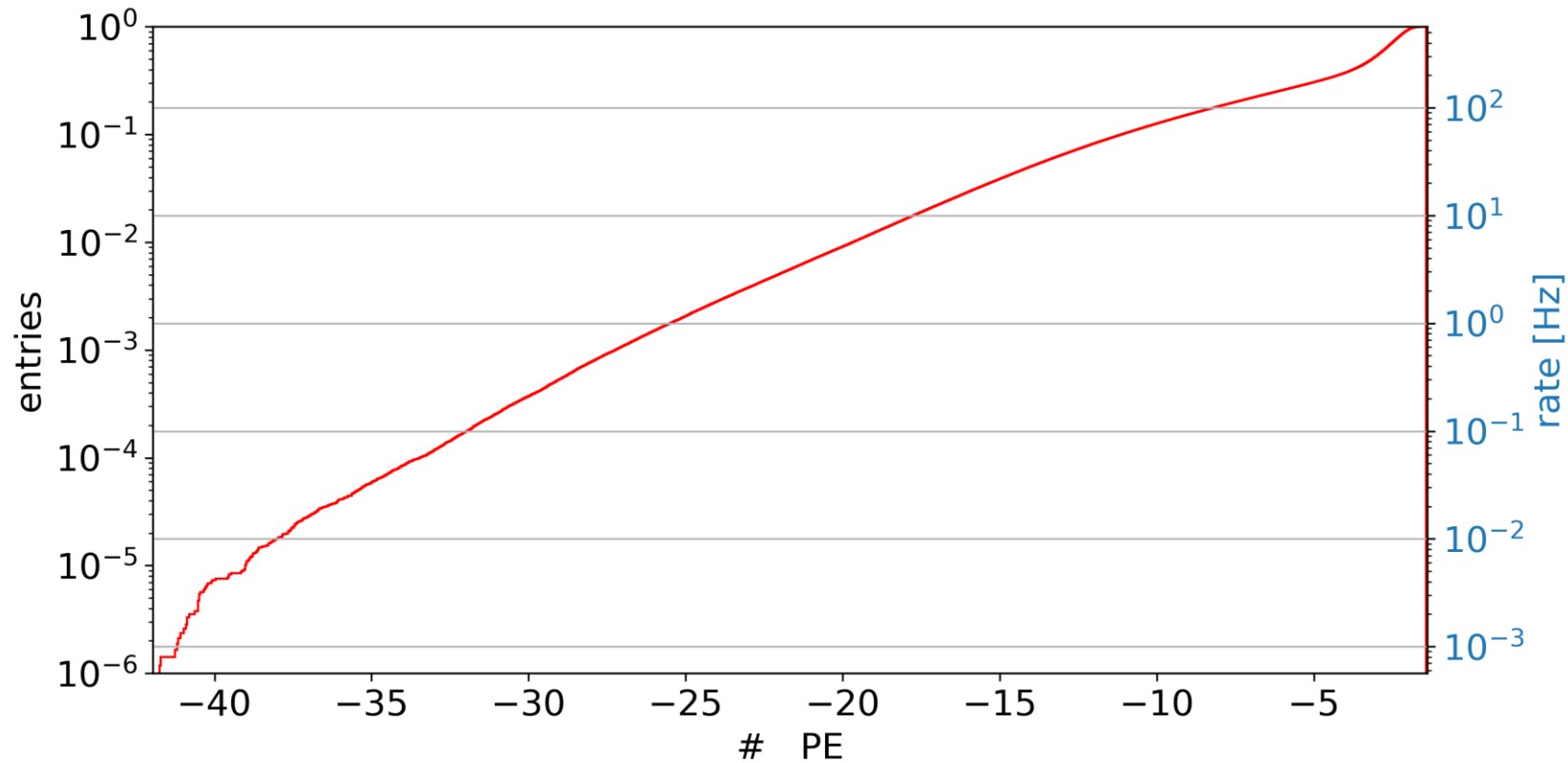
Maybe easier to study closeby 40-K decays via pulse height spectrum (see next slide)

see talk by Immacolata this morning

SPECIAL DATA SETS

PULSE HEIGHT DISTRIBUTION (CUT @ 2 PE)

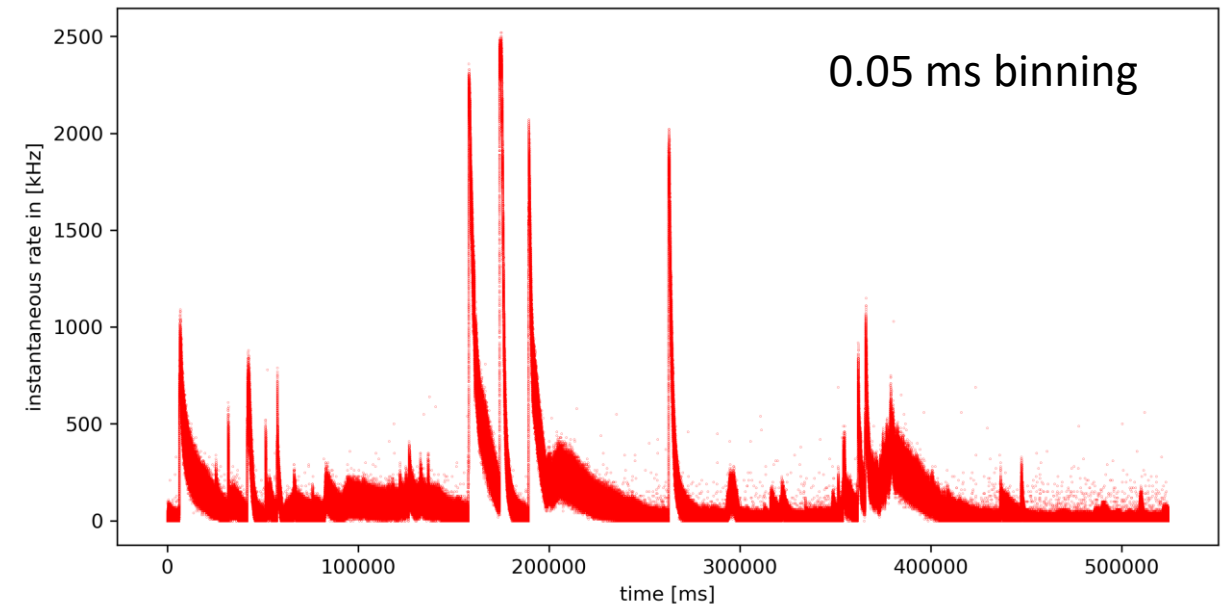
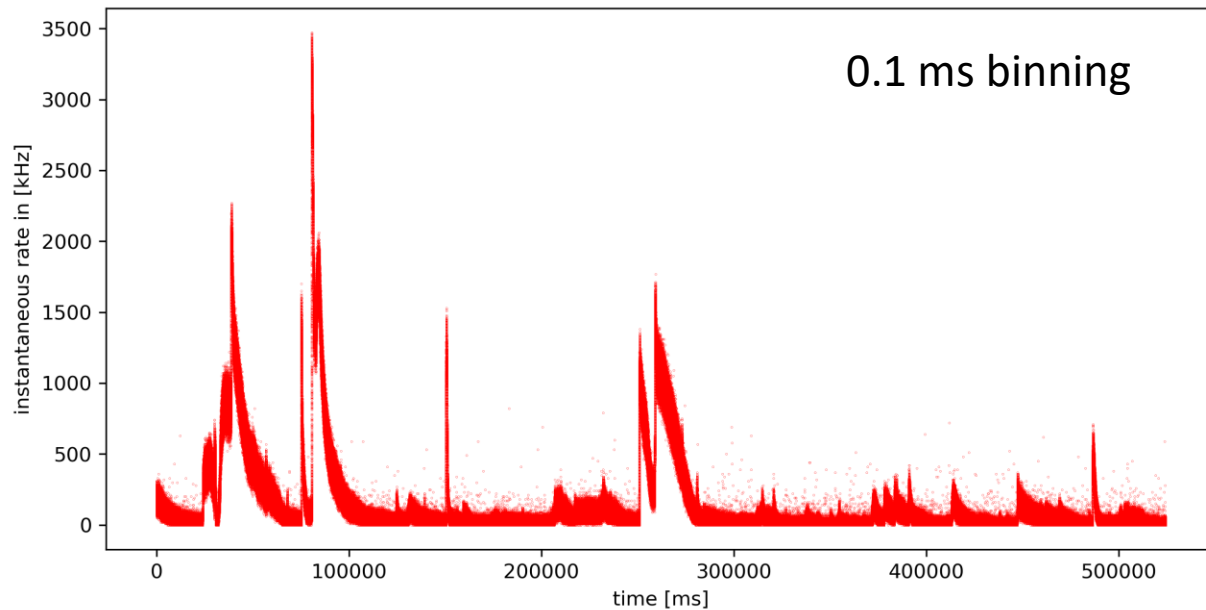
- 5400 s of data taken with cut at 2 PE
- Compare this to 40-K Monte Carlo? (Bioluminescence mainly 1 PE)



FINER RATE BINNINGS

Available runs: 4 x 5400 s (1 ms resolution), 2x400s (0.1 ms), 1x250s (0.05 ms), 3x90s (0.001 ms)

Examples for rate as function of time:

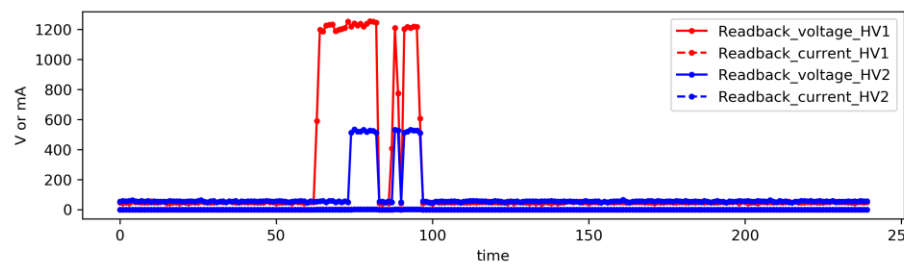
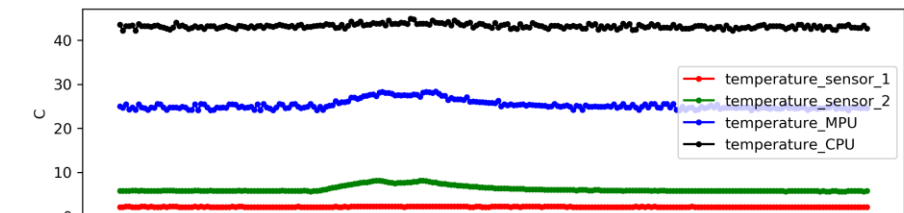
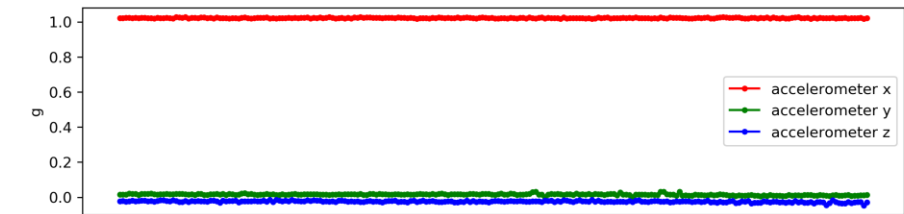
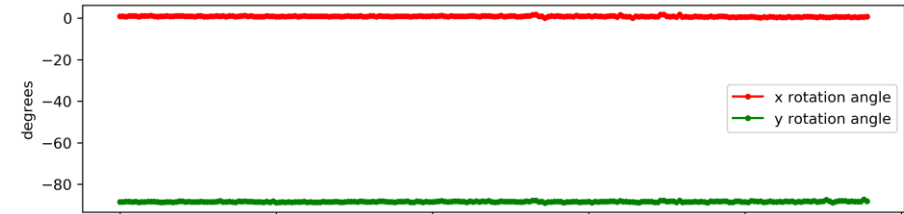
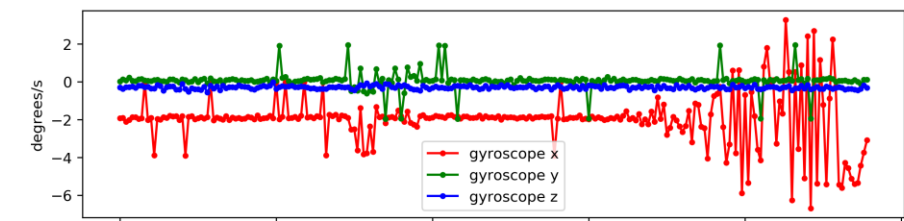


Spurious high rates in single bins are deadtime artifact of DAQ....

ENVIRONMENTAL SENSORS

- 24 days of data in ocean
- 19 sensors read out every minute
- Gyro, magnetometer, acceleration, temperature (4), HV
- Data not in Oceans 2.0 Plotting Utility (but stored in Mainz)

- Not studied in detail yet
- No comparison to other ONC information and WOM rate yet
- Work is under way



SUMMARY

- WOM deployment successful, but connectivity unfortunately lost
- Continuous data taking with 1 ms rate resolution for 18 days
- Many special runs and environmental sensor readings
- First results on bioluminescence based on 18 days
 - frequency space (inertial frequency low?)
 - rise and fall times (how to interpret these results?)
- First ideas on effective area determination in-situ (40-K)
 - K-40 rates too low ?

many open questions !!!

APPENDIX

K-40 FOR ABSOLUTE CALIBRATION?

Implications of optical properties of ocean, lake, and ice for ultrahigh-energy neutrino detection

P. Buford Price

Single emitter:

Absorption dominated

$$N_{\gamma}^{\text{PMT}} = L(E_e) \frac{1}{\pi} \int T(\lambda)\eta(\lambda)2\pi\alpha_e(1 - 1/\beta^2m^2) \times \frac{A_{\text{PMT}}}{4\pi d^2\lambda^2} \exp[-a(\lambda)d]d\lambda,$$

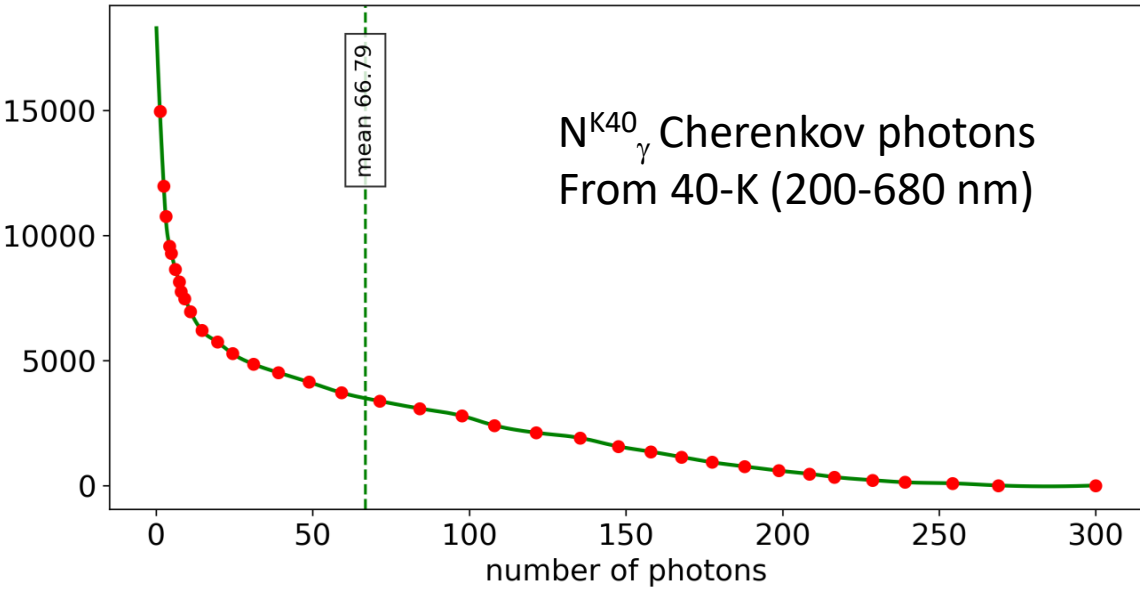
diffusion regime:

$$N_{\gamma}^{\text{PMT}} = L(E_e) \frac{1}{\pi} \int T(\lambda)\eta(\lambda)6\pi\alpha_e(1 - \beta^{-2}m^{-2}) \times \frac{A_{\text{PMT}}\Sigma[b_i(1 - \tau_i)]}{16\pi d\lambda^2} \exp[-\alpha(\lambda)d]d\lambda.$$

Constant density emitters ($\rho=11000/\text{m}^3$):

$$R_{\gamma} = N^{\text{K40}}_{\gamma} \rho A_{\text{eff}} / \pi / a(\lambda) \times (1 - \exp(-a(\lambda) * d_{\text{max}}))$$

diffuse case accordingly



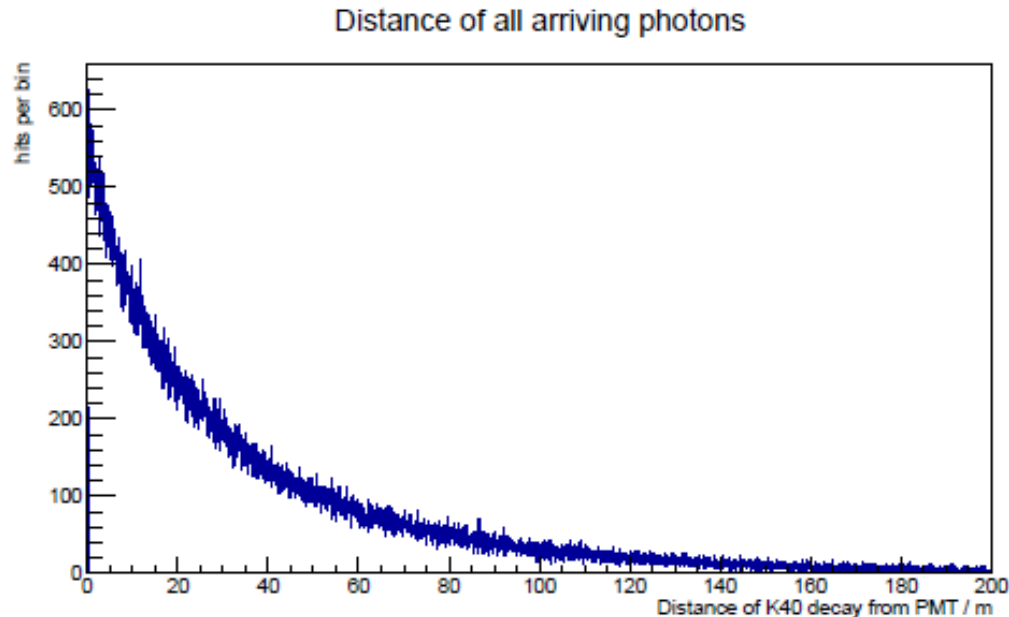
KM3NET

Björn Herold, Simulation and measurement of optical background in the deep sea using a multi-PMT optical module, dissertation 2017:

Assume 13 kBq per m³, 11250 photons per KM3Net PMT hitting cathode surface (before QE and DAQ!)

Expect **128.2 kHz** for full KM3Net Module with 31 PMTs of 76 mm cathode diameter (measurements higher)

4.13 kHz per 3" PMT



Detect photons from far away:
→ Rate depends on absorption!

Look at coincidences or pulse height spectrum instead (sensitive to close 40-K decays!)

Can't do coincidences in straw-b, but maybe pulse height?

Salinity @ Cascadia Basin: 34.83 g /kg
in Mediterranean probably around 38 g/kg; 40-K rate lower in Canada

EXPECTED FREQUENCIES IN PERIODIGRAM

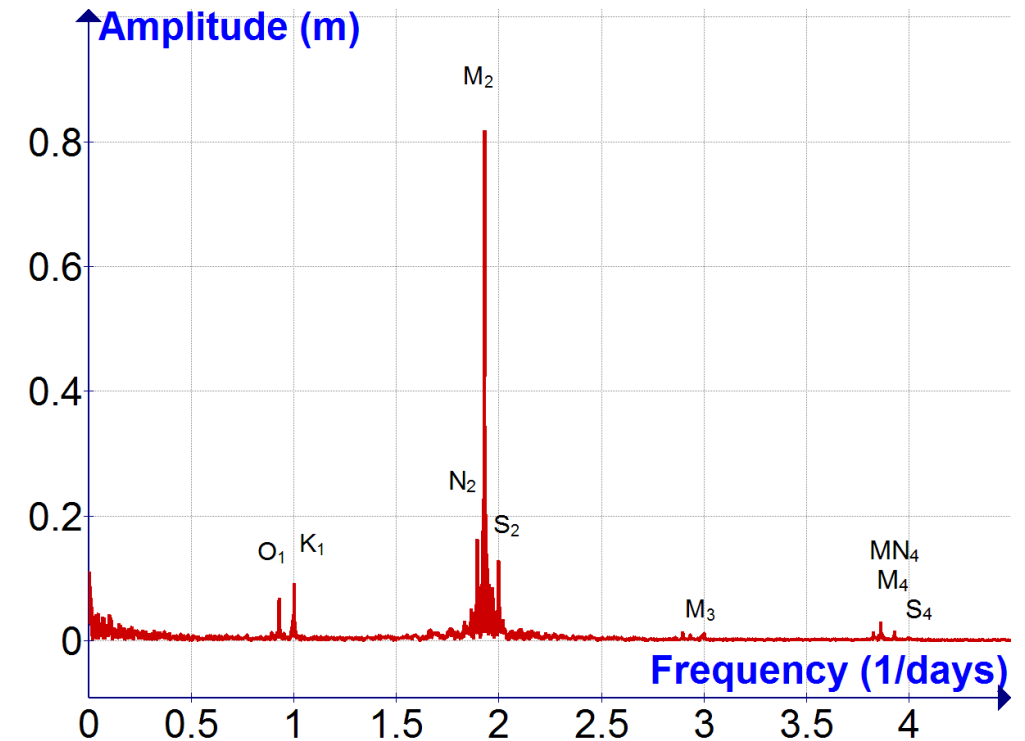
Latitude ODP 1027C: 47.7567°

Dominant frequencies:

semidiurnal lunar tide M2 12.4206012 h
 Inertial cycle 12h/sin(latitude)=16.21 h

https://en.wikipedia.org/wiki/Theory_of_tides

		Amplitudes Kodiak SF		
Principal lunar semidiurnal	M_2	12.4206012	97.3	58.0
Principal solar semidiurnal	S_2	12	32.5	13.7
Larger lunar elliptic semidiurnal	N_2	12.65834751	20.1	12.3
Lunar diurnal	K_1	23.93447213	39.8	36.8
Lunar diurnal	O_1	25.81933871	25.9	23.0



SALT WATER TANK



Water temperature between 4-6 degrees; however, sensor temperature rising to 30 degrees, when running

