

Fun with axions and axion strings

Anson Hook

University of Maryland

Axions

- Solves the strong CP problem
 - Explains why the neutron eDM is small
- Natural dark matter candidate
 - Unique experimental signatures
- Measurement of the fundamental value of electric charge

Anomalies

$$\frac{\alpha_{\text{em}}}{4\pi} \left(\mathcal{A} \frac{a}{f} + \theta \right) F \tilde{F}$$

\mathcal{A} is the 't Hooft anomaly of the axion

Axion is periodic but coupling is not

$$\frac{a}{f} = \frac{a}{f} + 2\pi \quad \longrightarrow \quad \theta = \theta + 2\pi \mathcal{A}$$

Periodicity of theta

What does this periodicity teach us?

$$\theta = \theta + 2\pi\mathcal{A}$$

Witten effect

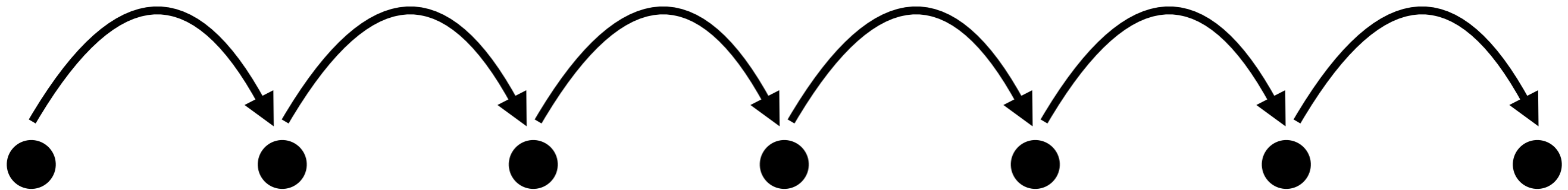
$$q_e = \left(\frac{eg}{2\pi}\right) \frac{\theta e}{2\pi} = \frac{e}{e_{\min}} \frac{\theta e}{2\pi}$$

Not periodic!

Need many monopoles that are exchanged

Periodicity of theta

$$\theta \rightarrow \theta + 2\pi\mathcal{A}$$



$$q_{-2}(\theta) \quad q_{-1}(\theta) \quad q_0(\theta) \quad q_1(\theta) \quad q_2(\theta)$$

$$q_n(\theta + 2\pi\mathcal{A}) = q_{n+1}(\theta)$$

Periodicity of theta

$$q_n(\theta + 2\pi\mathcal{A}) = q_{n+1}(\theta)$$

Periodicity gives difference in charge between various monopoles

$$\Delta q_e = \frac{e}{e_{\min}} \mathcal{A} e$$

Fundamental electric charge of monopoles

$$\Delta q_e = e_{\min} = \sqrt{\mathcal{A}} e$$

Periodicity of theta

Without any assumptions

$$\mathcal{A} = \mathbb{Z} e_{\min}^2 = \sum_f N_f N_a Q_f^2$$

Fractional anomaly necessarily implies fractional electric charge is the minimal charge

How to use the Axionic coupling

$$\frac{A\alpha_{\text{em}}}{4\pi} \frac{a}{f} F \tilde{F}$$

Anomaly gives us information about the quantization of electric charge

Can we measure it?

How to use the Axionic coupling

$$\frac{A\alpha_{\text{em}}}{4\pi} \frac{a}{f} F \tilde{F}$$

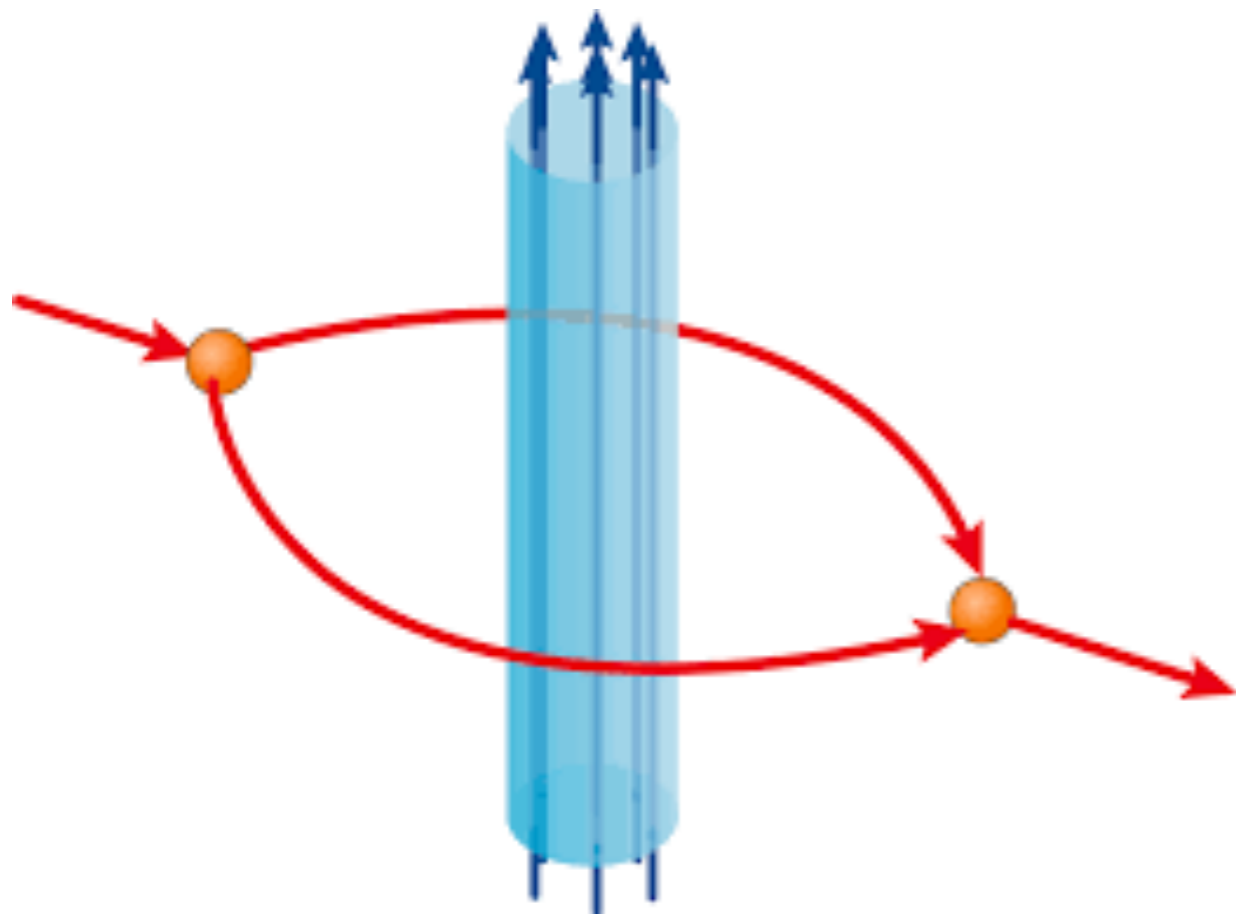
Particle approach **HARD**

Anomaly - decay constant ambiguity and
ambiguity coming from mixing effects

Decay constant RG runs : Wavefunction renormalization

Need something independent of f

How to use the Axionic coupling

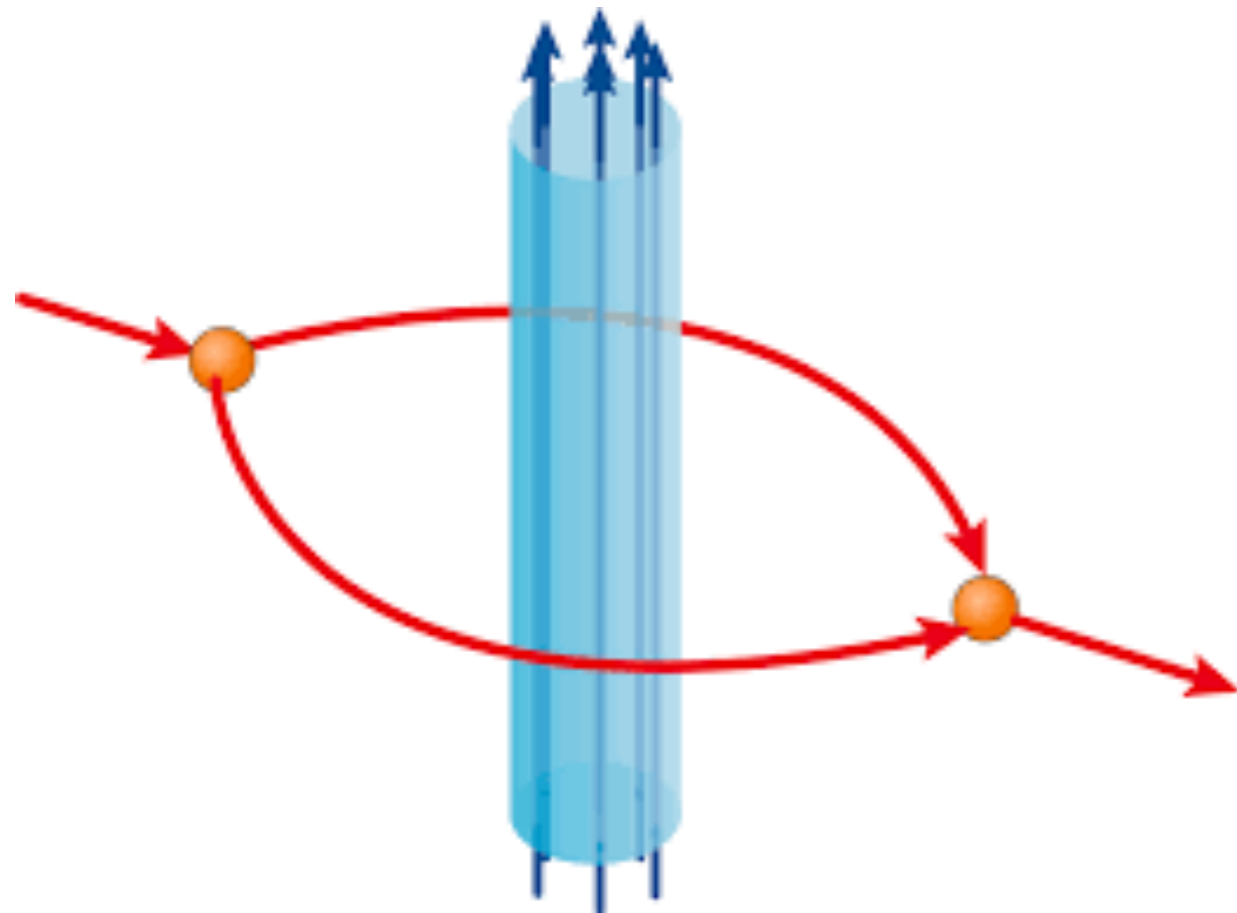


Strings!

$$\frac{a}{f} \rightarrow \frac{a}{f} + 2\pi (\mathbb{Z})$$

How to use the Axionic coupling

Many string properties are topological

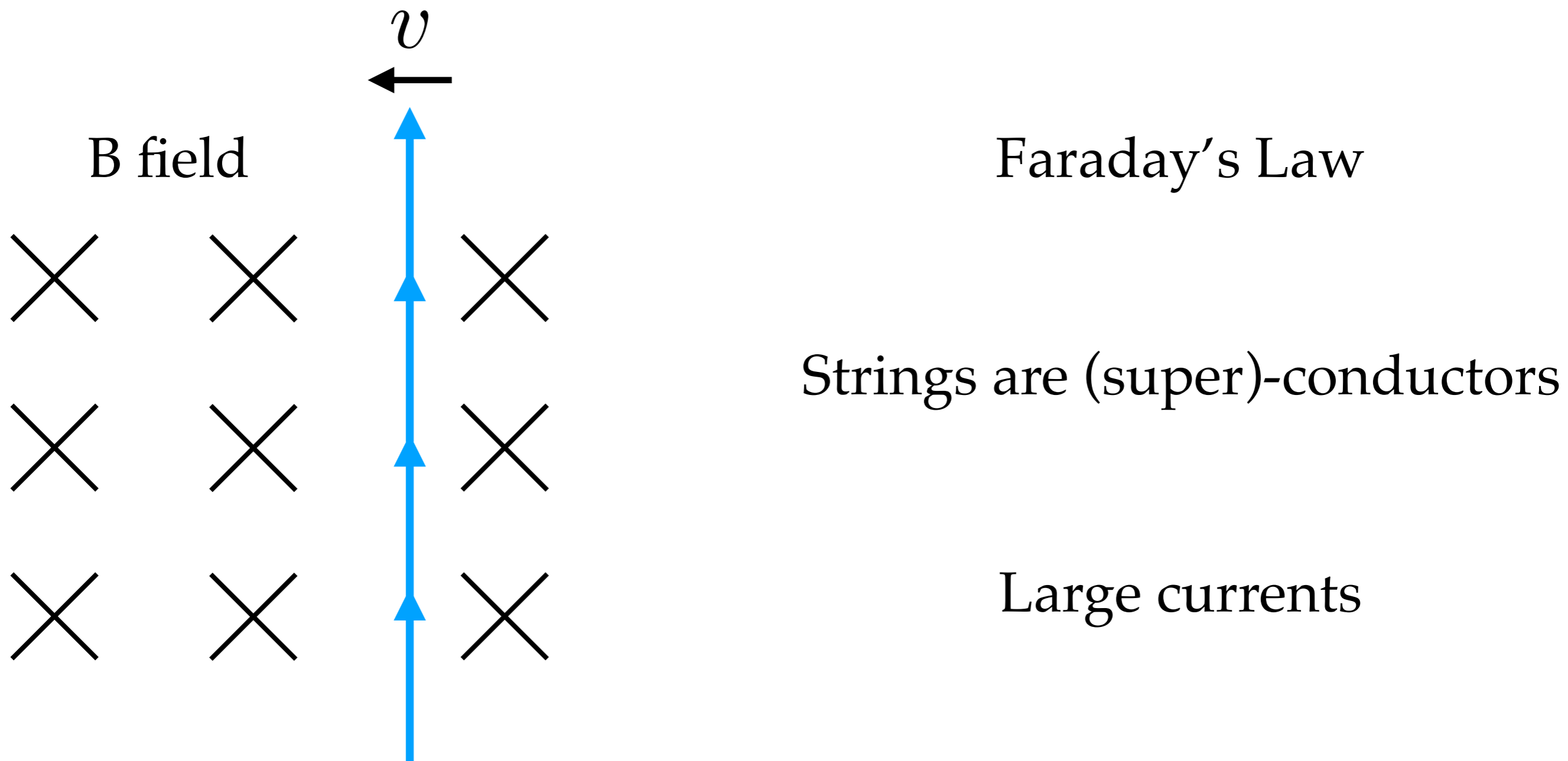


Aharonov-Bohm Effect

Photons acquire phase as it moves around an axion string

How to use the Axionic coupling

Many string properties are topological



Outline

- How strings affect photons
 - Arahamov-Bohm effects
 - CMB ways of looking for strings
- How photons affect strings
 - Superconducting strings
 - Colliding currents

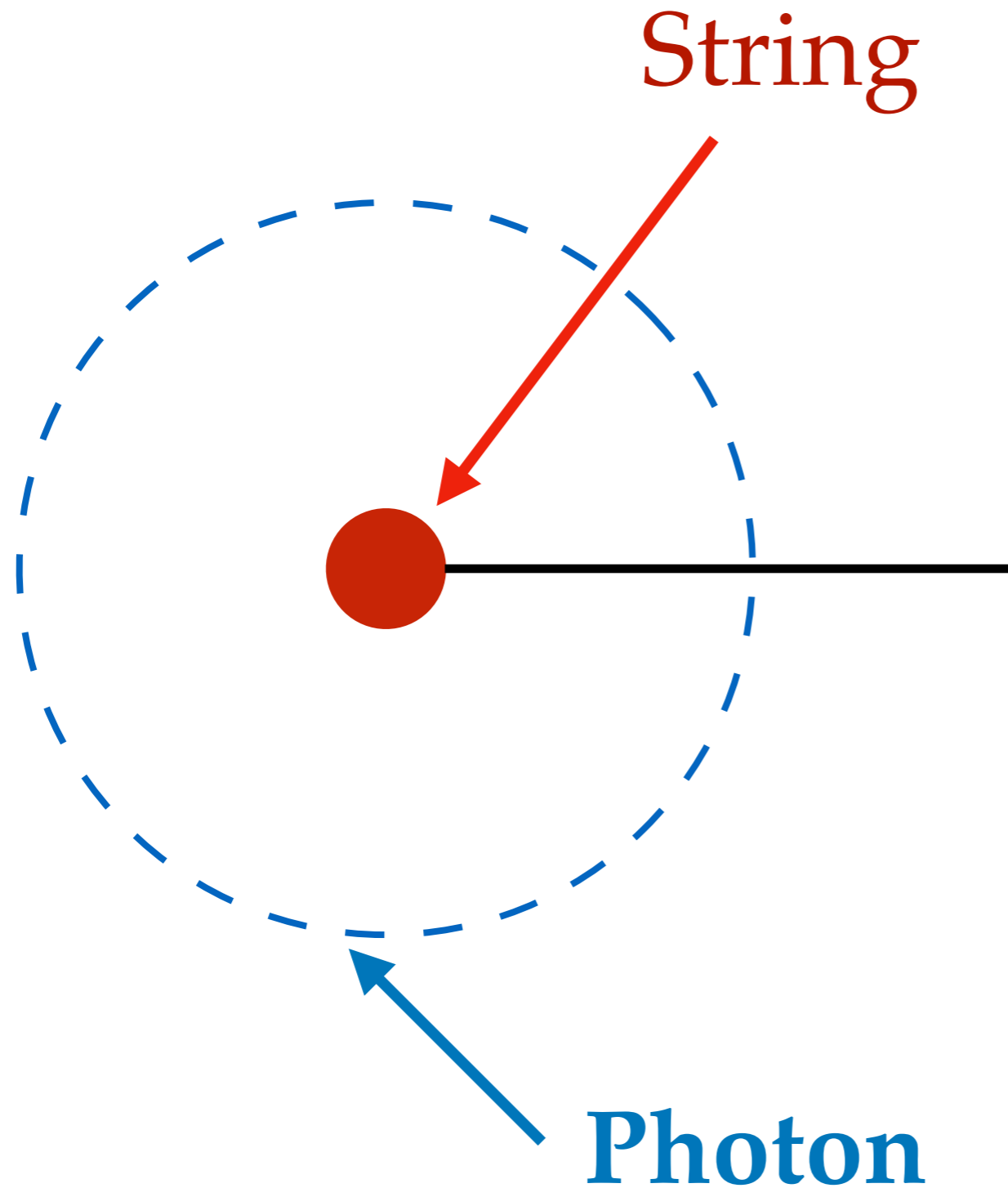
Polarization Rotation

$$\Phi = \int v dt = \mathcal{A} \alpha_{\text{em}} \frac{\Delta a}{2\pi f}$$

Phase rotation of circularly polarized light is a polarization rotation of linearly polarized light

Topological - depends only on initial and final value of the axion

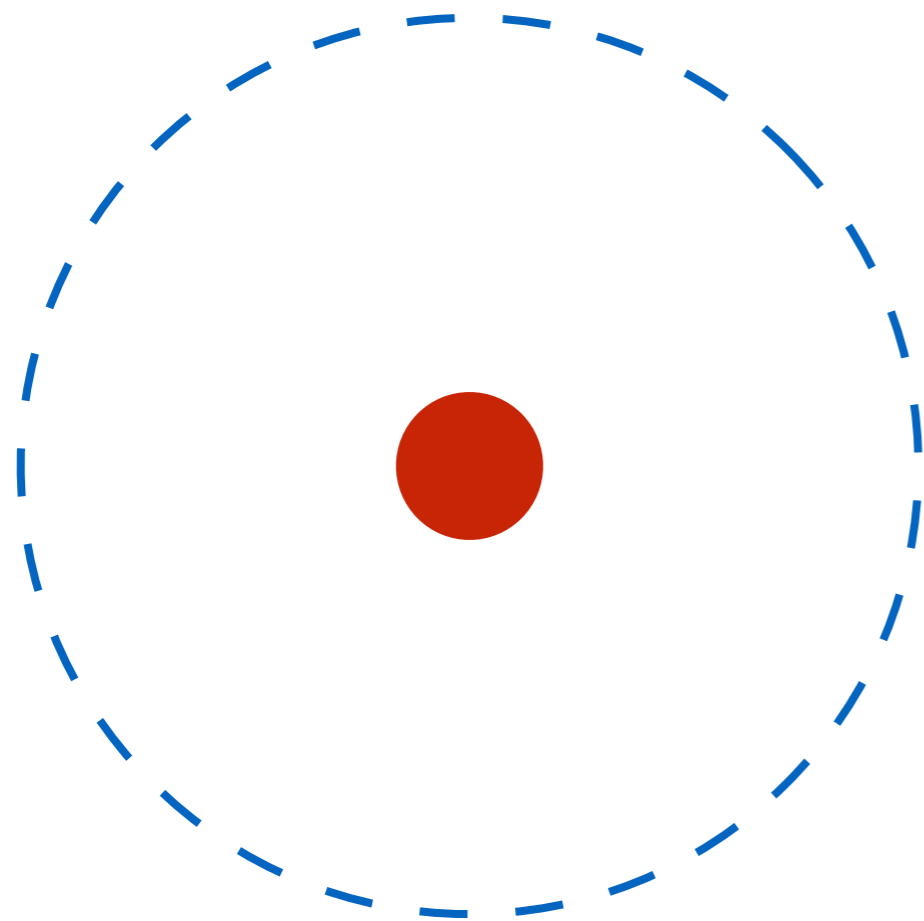
Strings and photons - a love story



$$\theta = \frac{a}{f}$$

Strings and photons - a love story

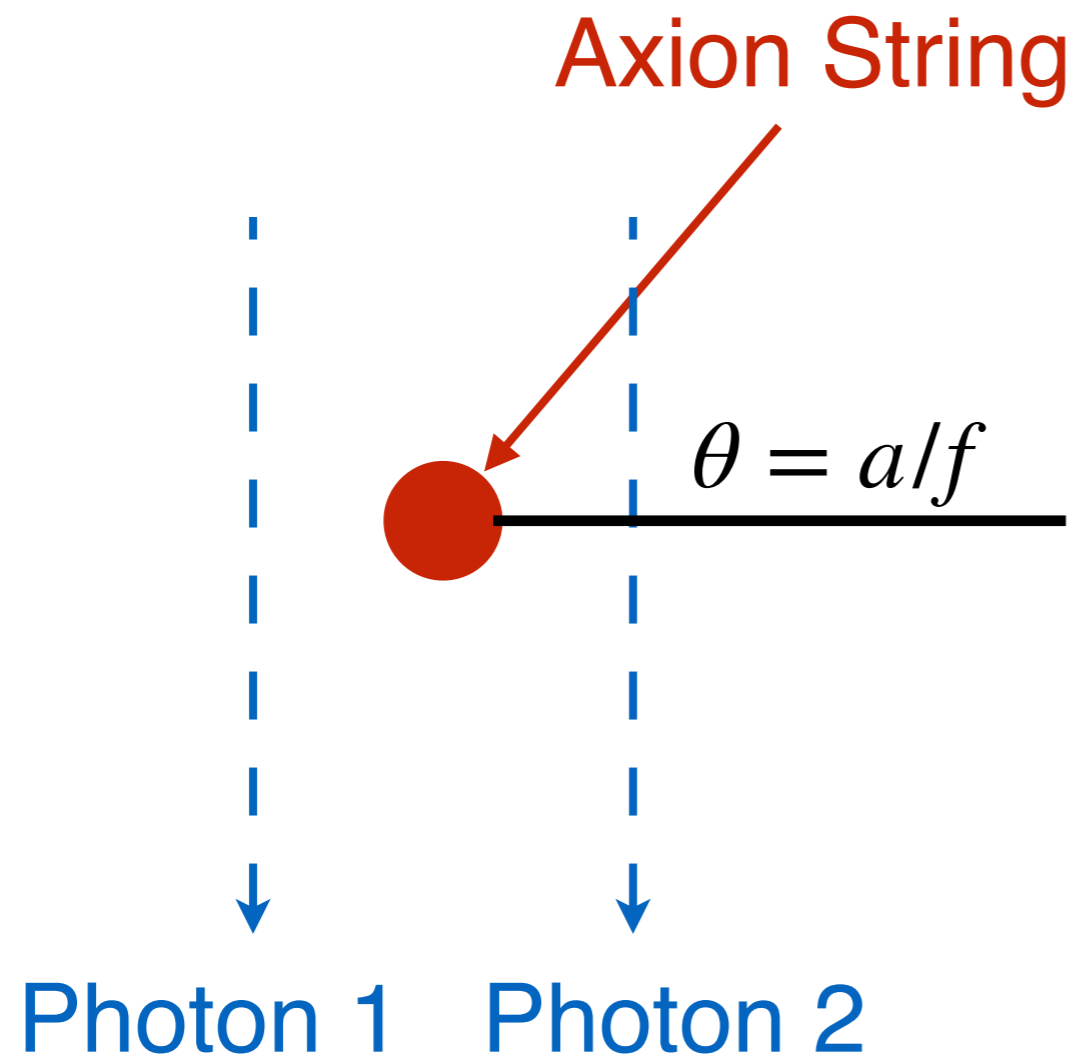
Topological effect



Large polarization rotation
from traveling around a string!

$$\Phi = \frac{\mathcal{A}\alpha_{\text{em}}}{2\pi} \frac{\Delta a}{f} = \mathcal{A}\alpha_{\text{em}}$$

Strings and photons - a love story



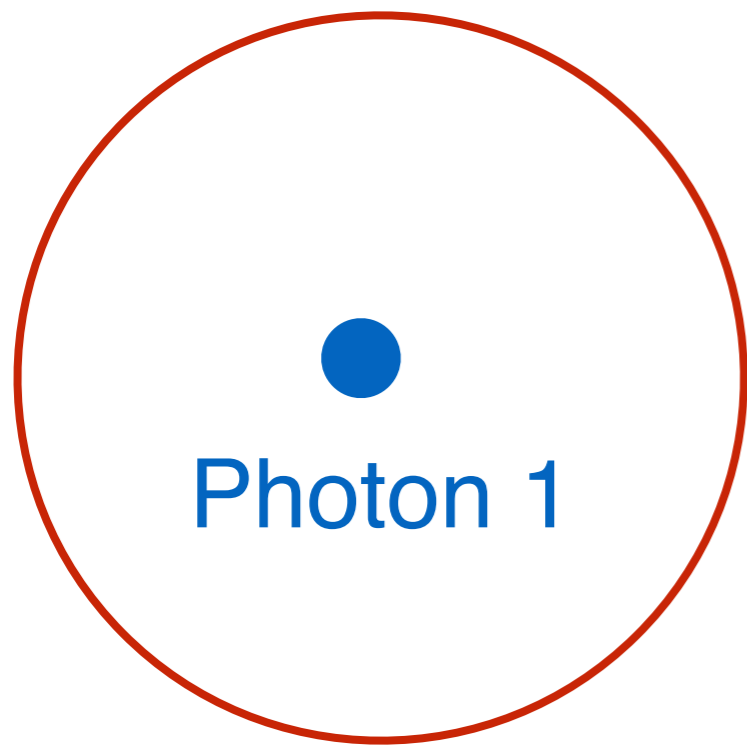
$$\Phi_1 = \frac{\mathcal{A}\alpha_{\text{em}}}{2\pi} \frac{\Delta a}{f} = \frac{\mathcal{A}\alpha_{\text{em}}}{2}$$

$$\Phi_2 = -\frac{\mathcal{A}\alpha_{\text{em}}}{2}$$

$$\Delta\Phi = \mathcal{A}\alpha_{\text{em}}$$

Strings and photons - a love story

String



$$\Phi_2 = 0$$

Photon 1

Photon 2

Photons that go through a string loop acquire a rotation

Photons that do not go through a string loop do not acquire a rotation

$$\Phi_1 = \pm \mathcal{A} \alpha_{\text{em}}$$

Outline

- How strings affect photons
 - ~~Arahanov-Bohm effects~~
 - CMB ways of looking for strings
- How photons affect strings
 - Superconducting strings
 - Colliding currents

String evolution

$$\text{String density } \rho_{\text{string}} = \xi(t) H^2 f^2 \log(f/H)$$

- Scaling solution : $\xi = \text{constant}$
- Scaling violation : $1 \lesssim \xi \lesssim 1000$

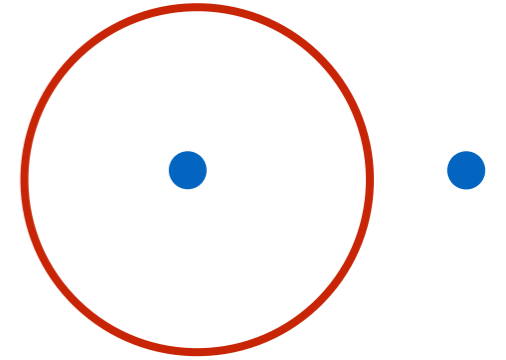
CMB

CMB acts as a backlight which is rotated by strings

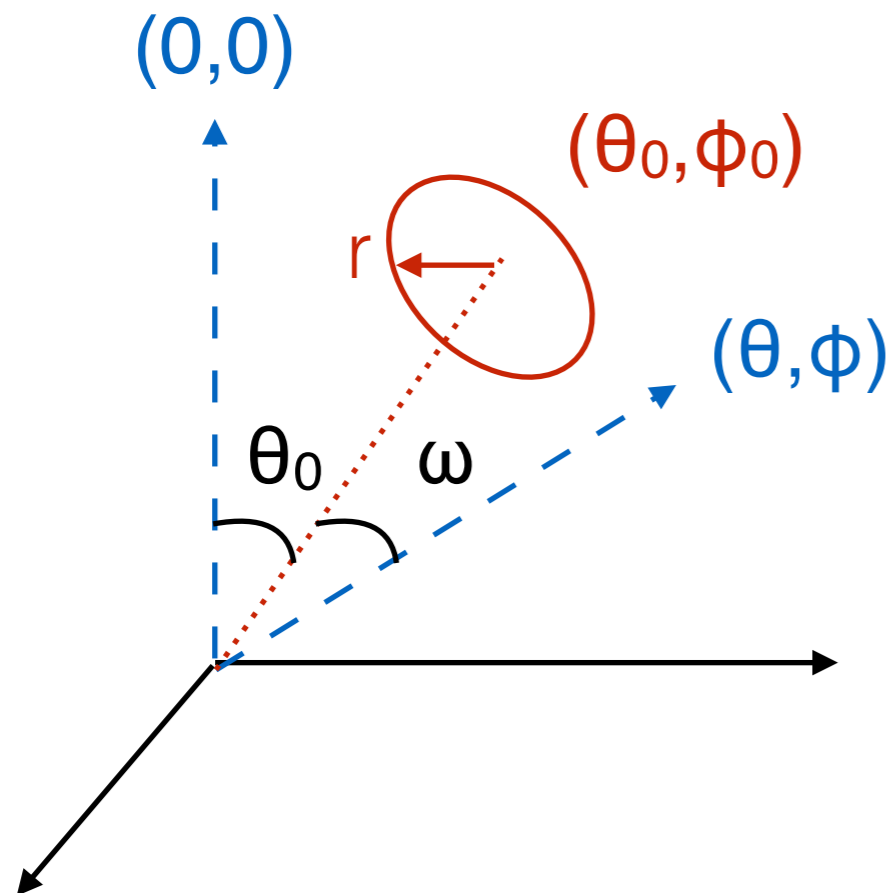
Power spectrum analysis

Look for discontinuous jumps
in the rotation angle

Toy analytics

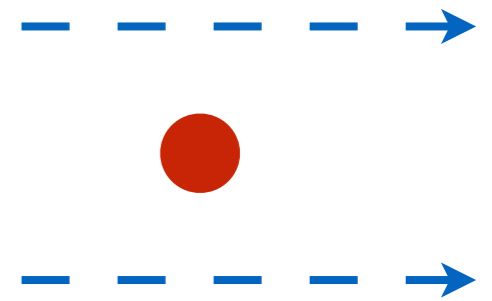


$$\langle \Phi(\hat{\gamma}) \Phi(\hat{\gamma}') \rangle = (\mathcal{A} \alpha_{\text{em}})^2 \int d\eta \int d^2 \hat{s} \int d^2 \hat{k} (\eta_0 - \eta)^2 f(\eta) \\ \times \Theta \left(\frac{\eta}{2} - d(\hat{s}, \hat{\gamma}, \hat{k}, \eta) \right) \Theta \left(\frac{\eta}{2} - d(\hat{s}, \hat{\gamma}', \hat{k}, \eta) \right)$$

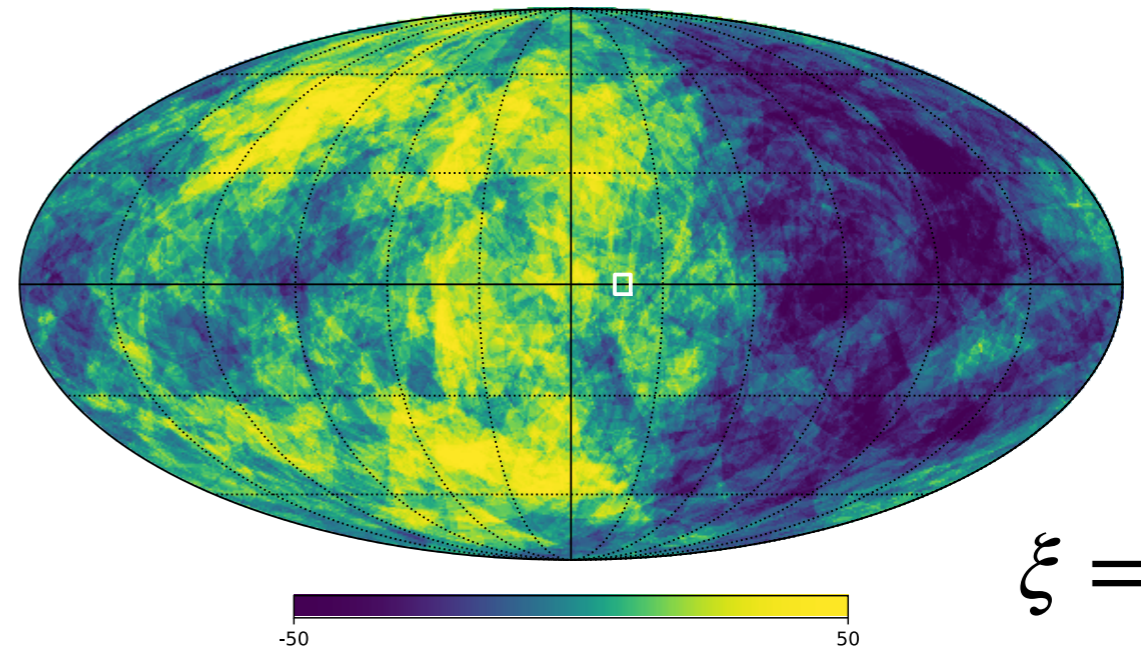


- Take strings to be Hubble sized loops
- Pass through loop acquire a phase $\Phi = \pm \mathcal{A} \alpha_{\text{em}}$
- Integrate over all orientations and directions

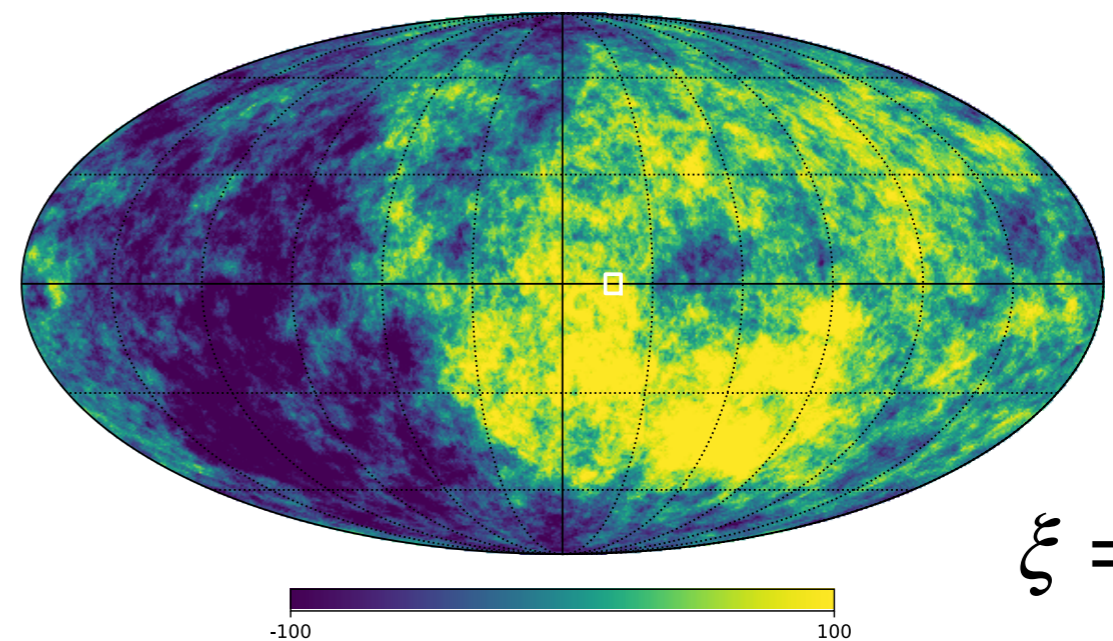
Toy simulation



- Randomly throw down infinitely long straight strings
- Remove strings as time goes on to maintain correct number density
- Trace CMB photon paths

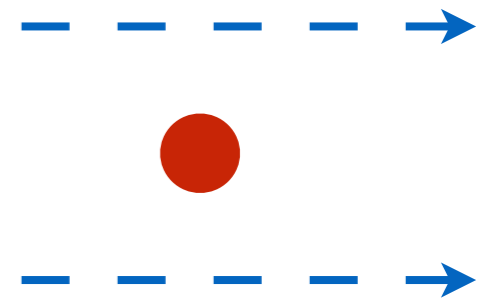


$\xi = 10$

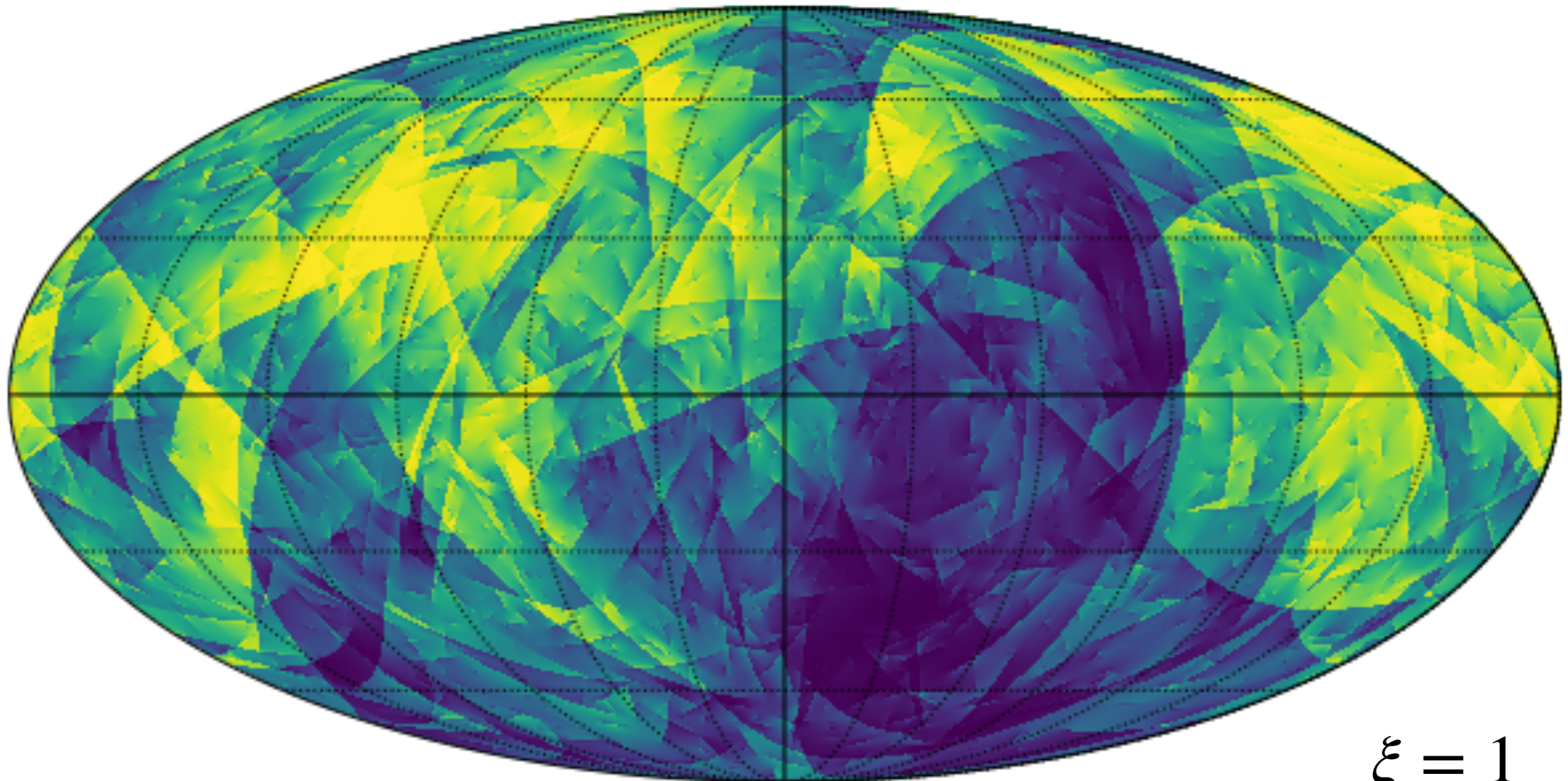


$\xi = 100$

Toy simulation

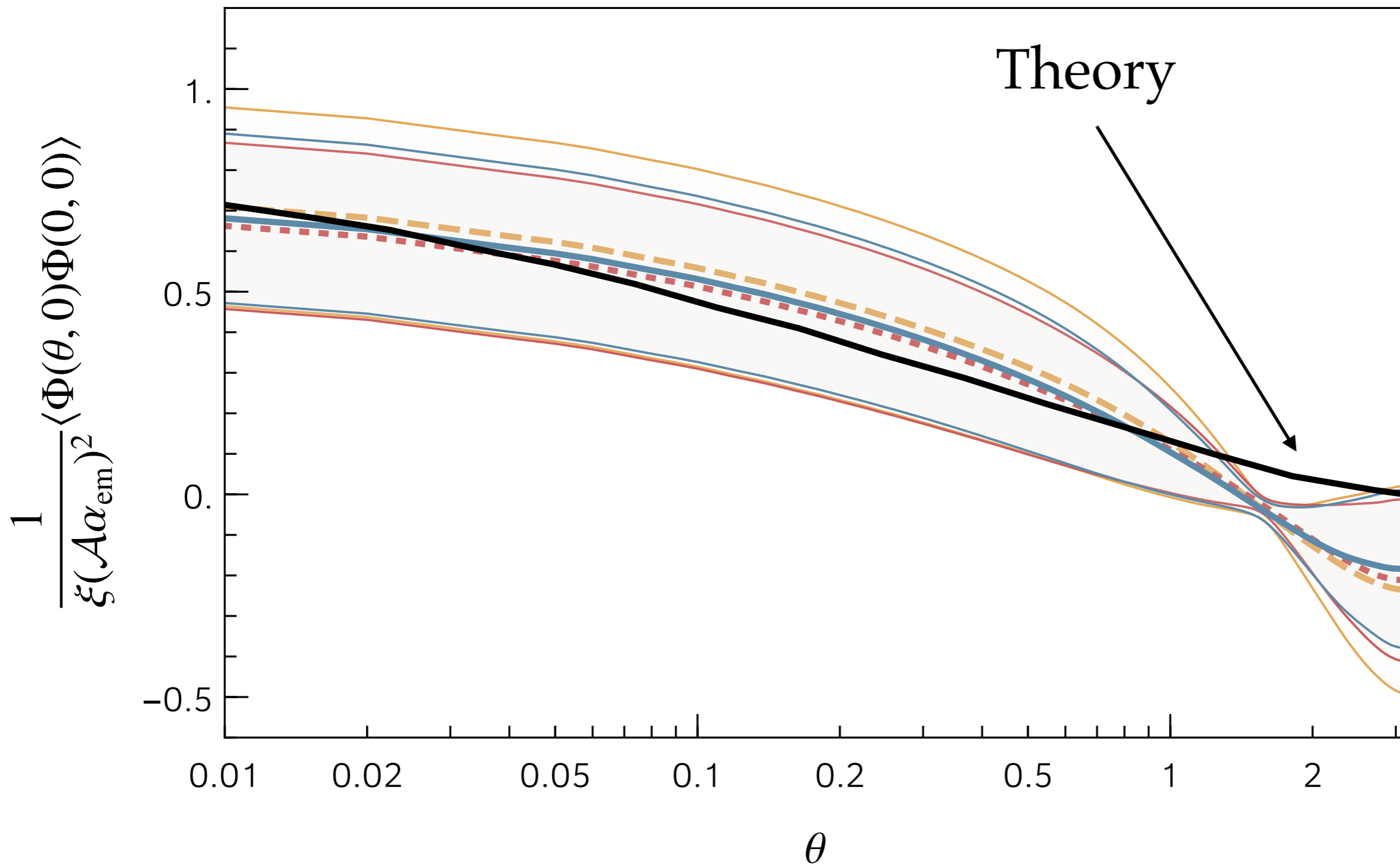
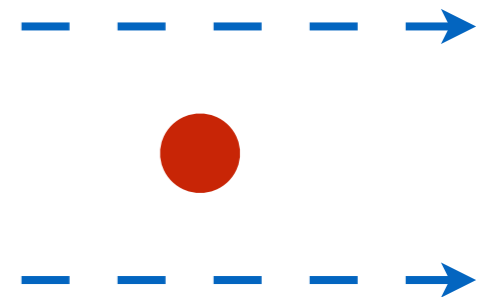


Can see both long and short strings

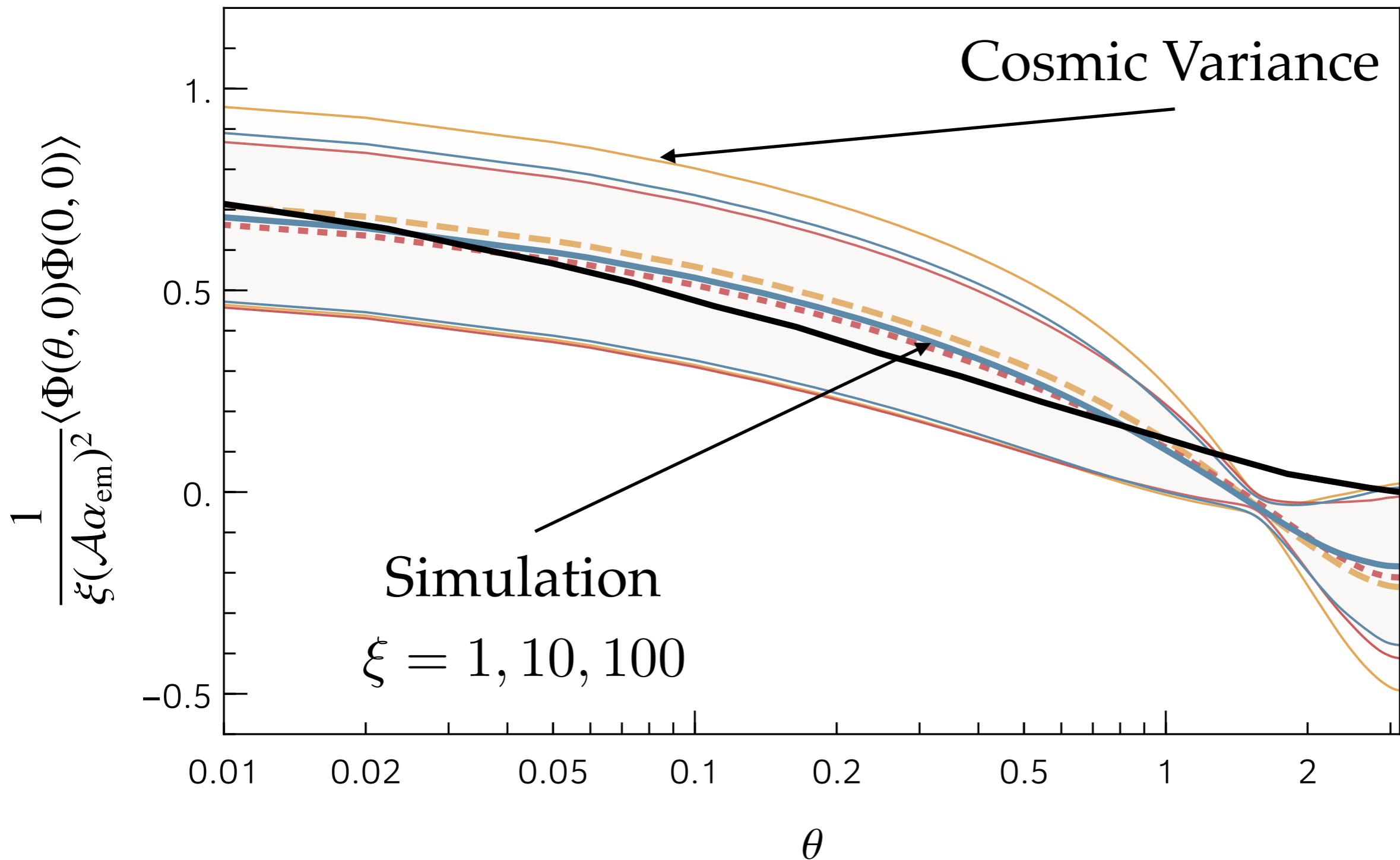
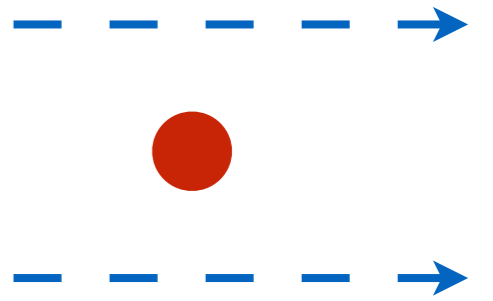


$\xi = 1$

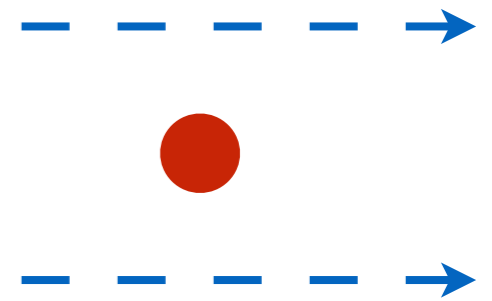
Simulation Comparison



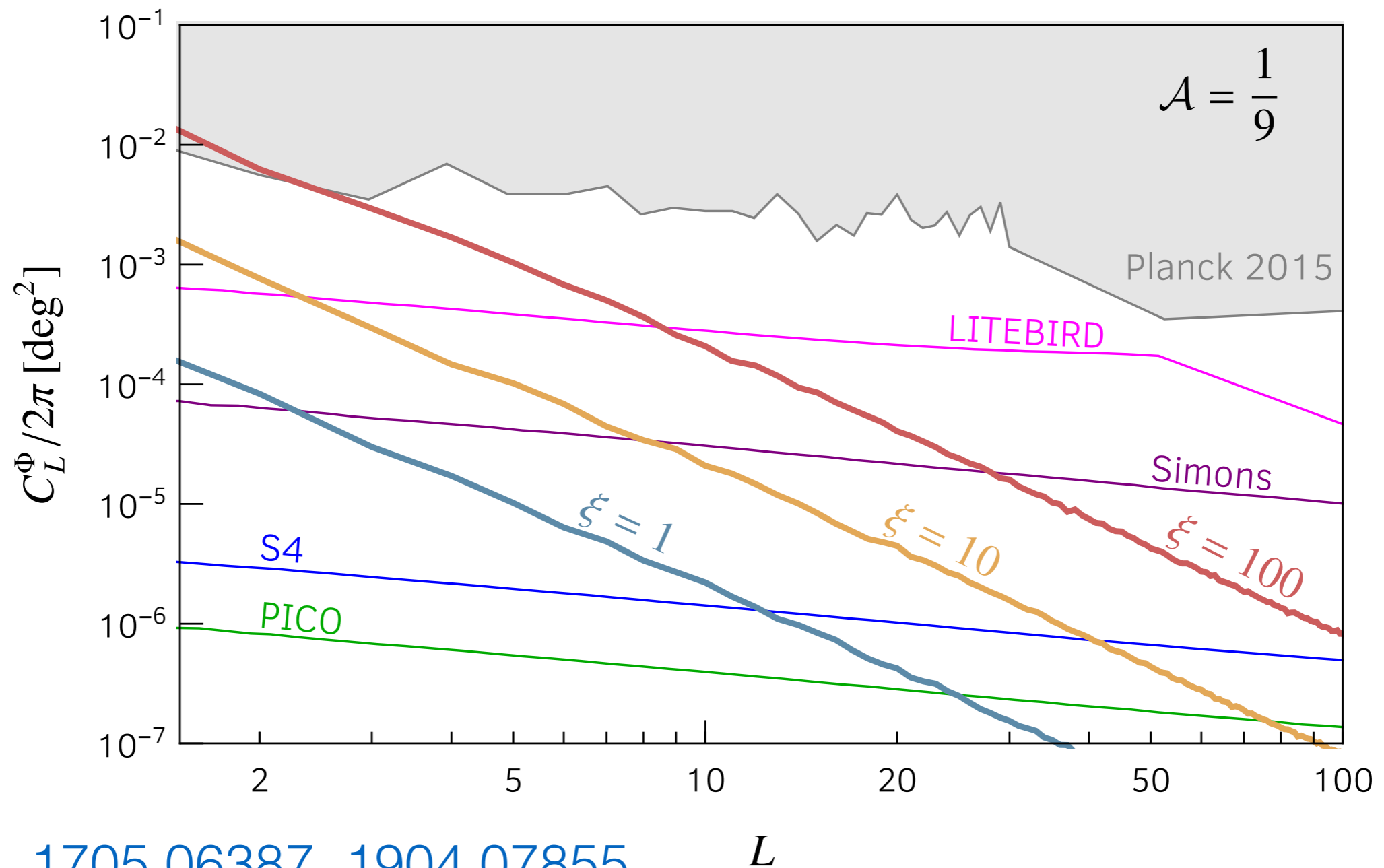
Simulation Comparison



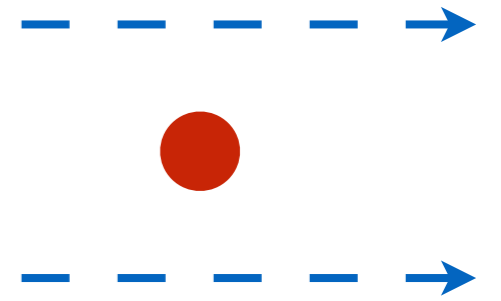
CMB Constraints(ish)



Constraints scale as $\xi \mathcal{A}^2$

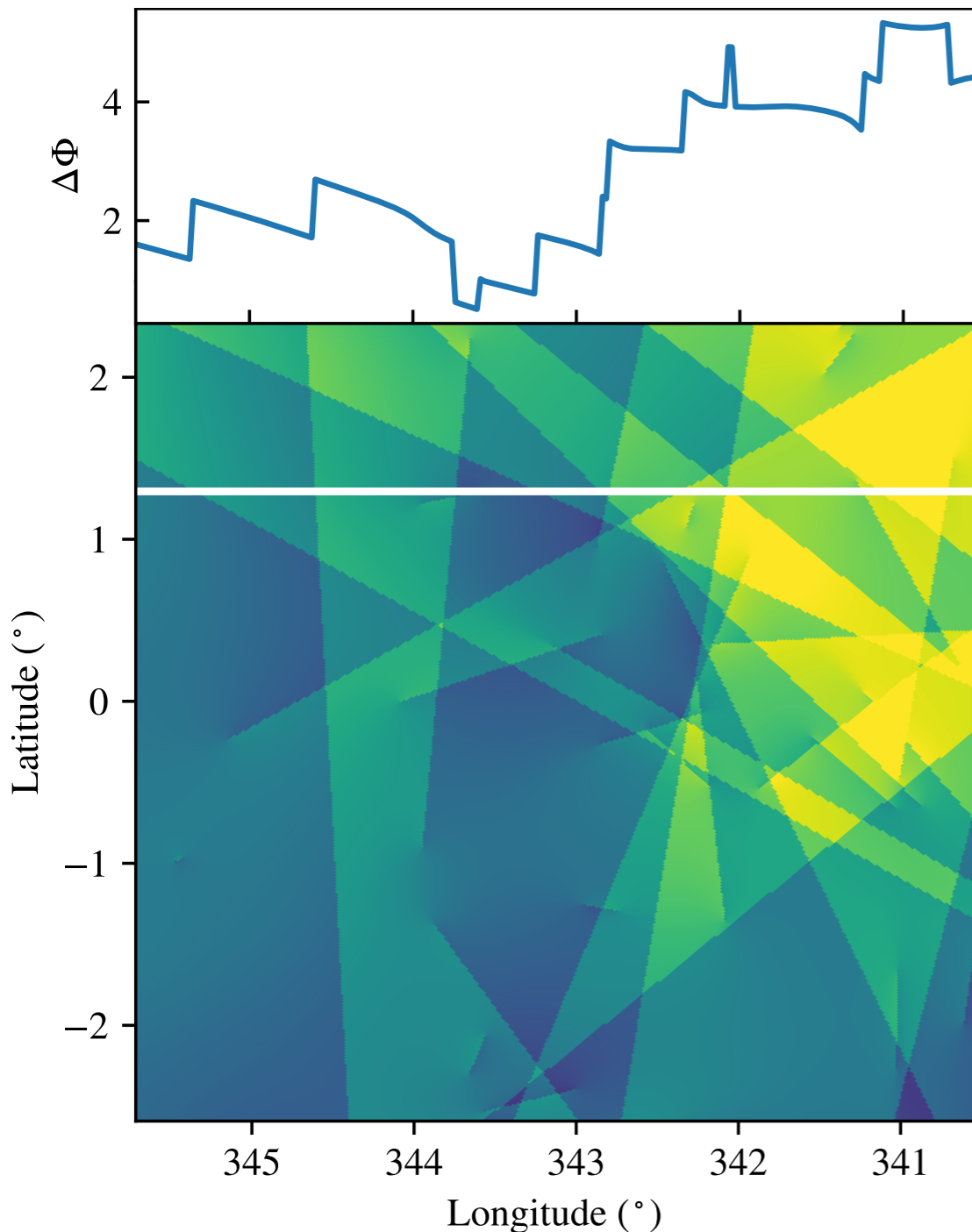


Edge Detection

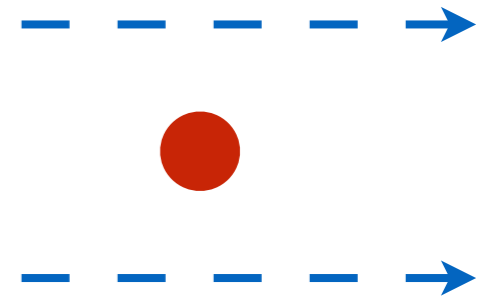


Coollest feature are the edges

- Position space search
- Important elements
 - Angular resolution $\xi \gg 1$
 - Accuracy $\mathcal{A} \ll 1$
- BICEP / KEK / SPT / Polarbear
 - Angular resolution \sim arcmin
 - Accuracy \sim percent



Edge Detection

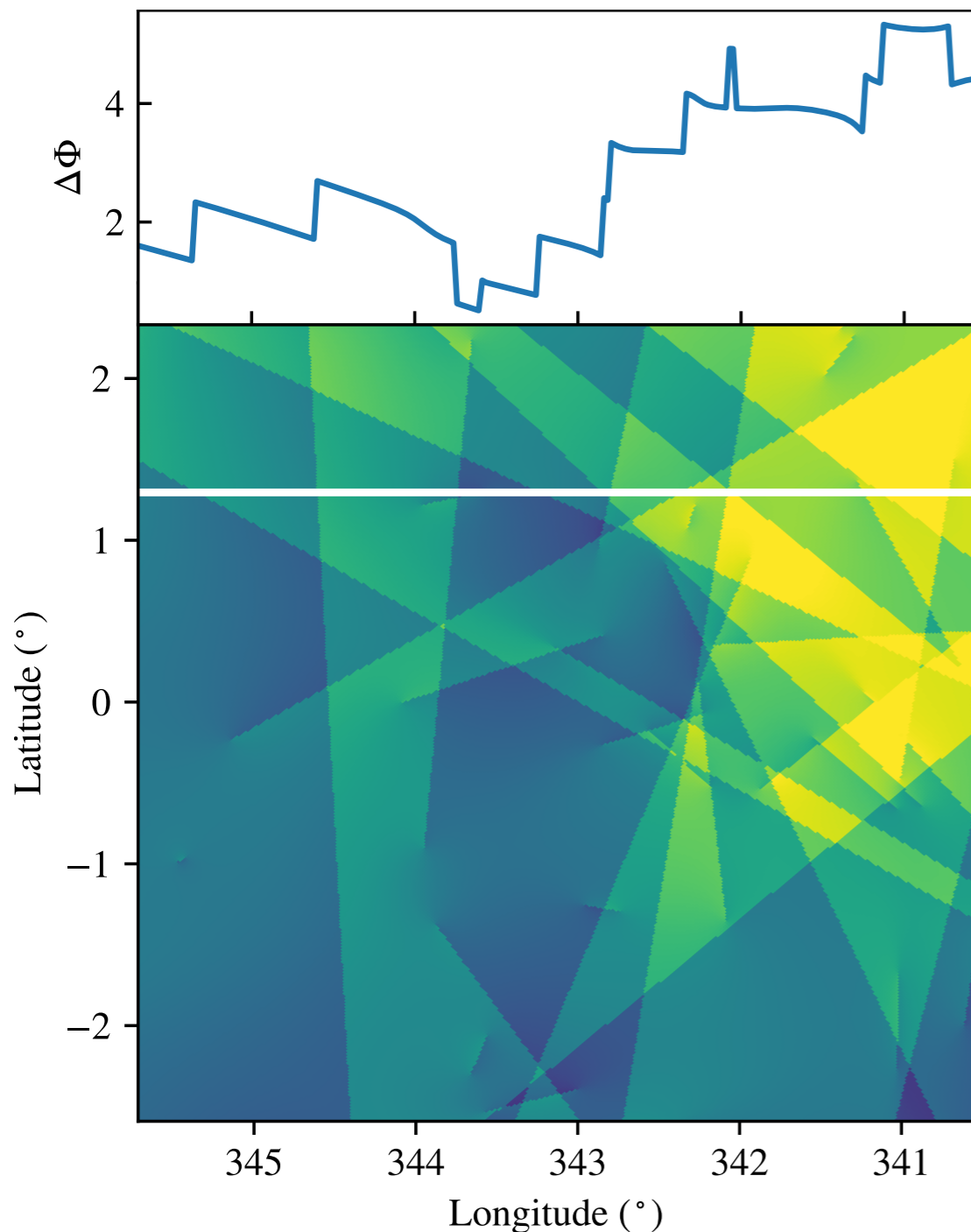


Coollest feature are the edges

Likely edges are most useful
when there are only a few strings

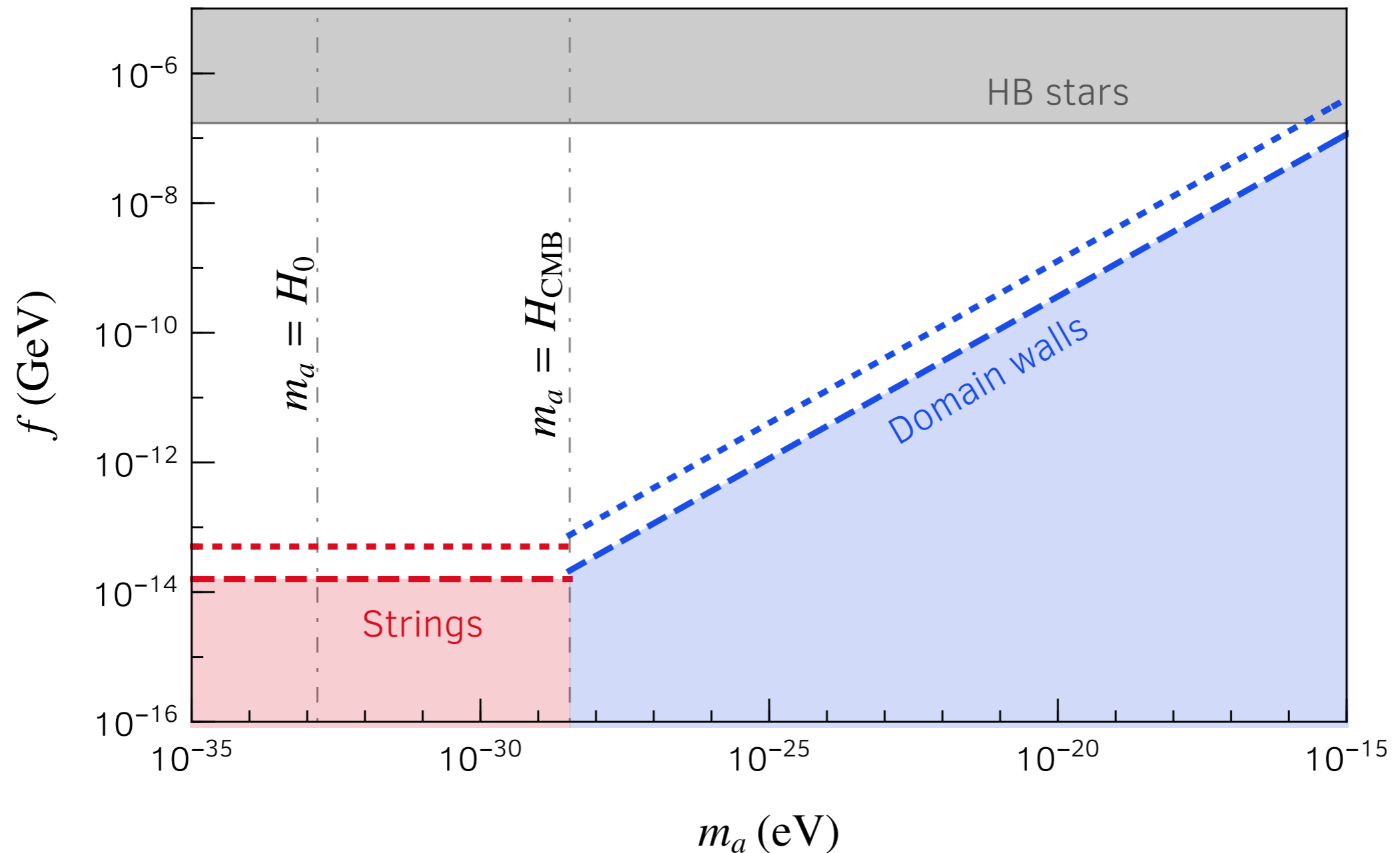
Few strings => production
before / during inflation

Rare chance at testing
pre-inflationary physics

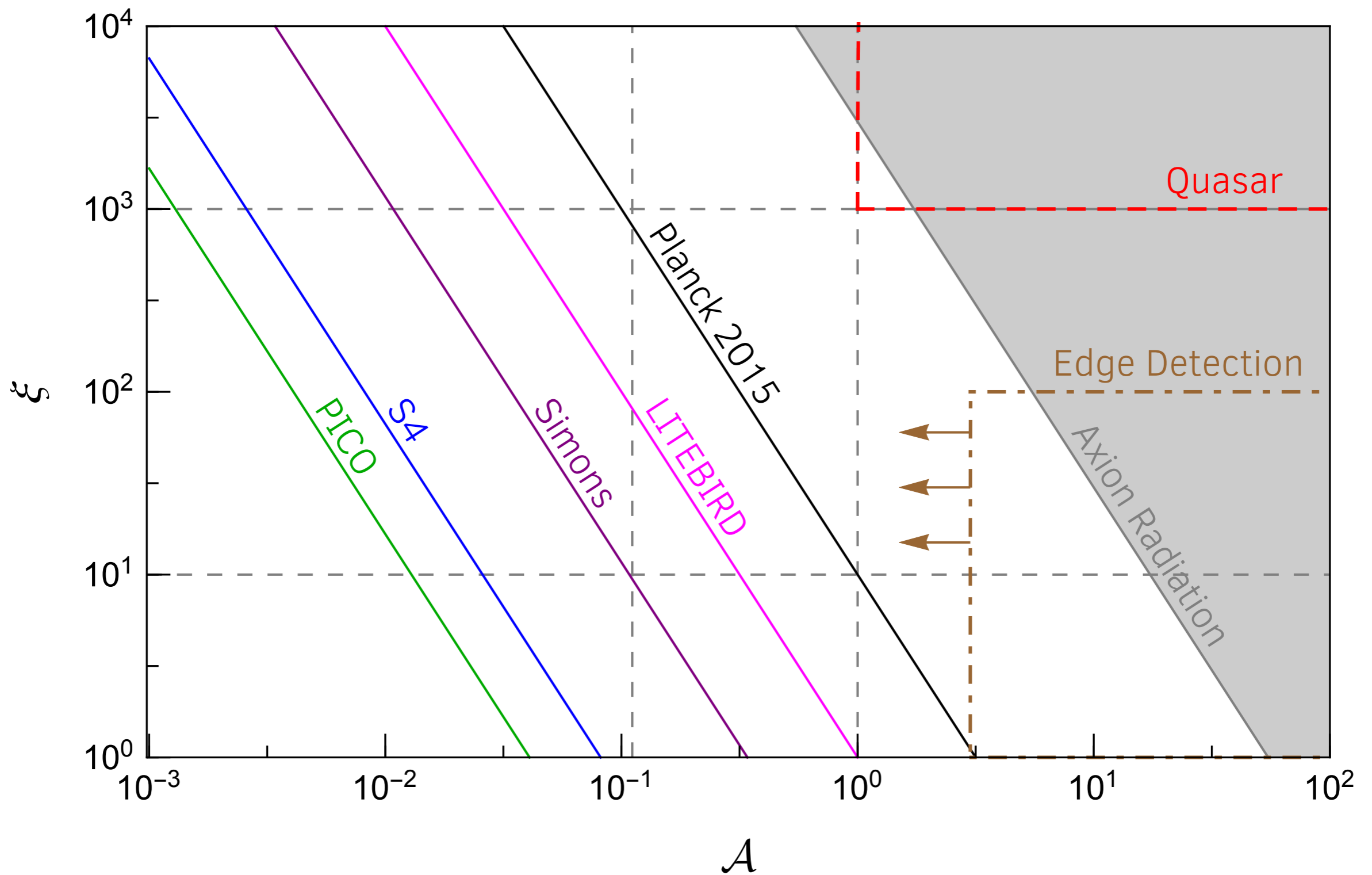


Usual Constraints

Our results are completely independent of decay constant



Constraint(ish)



Outline

- ~~How strings affect photons~~
 - ~~Arahanov-Bohm effects~~
 - ~~CMB ways of looking for strings~~
- How photons affect strings
 - Superconducting strings
 - Colliding currents
- Millikan (What do we learn from anomalies)

Superconducting Strings

Anomaly inflow - anomaly on the string to
restore gauge invariance

$$\frac{A\alpha_{\text{em}}}{4\pi} \frac{a}{f} F \tilde{F}$$

In the presence of strings, this
is not gauge invariant

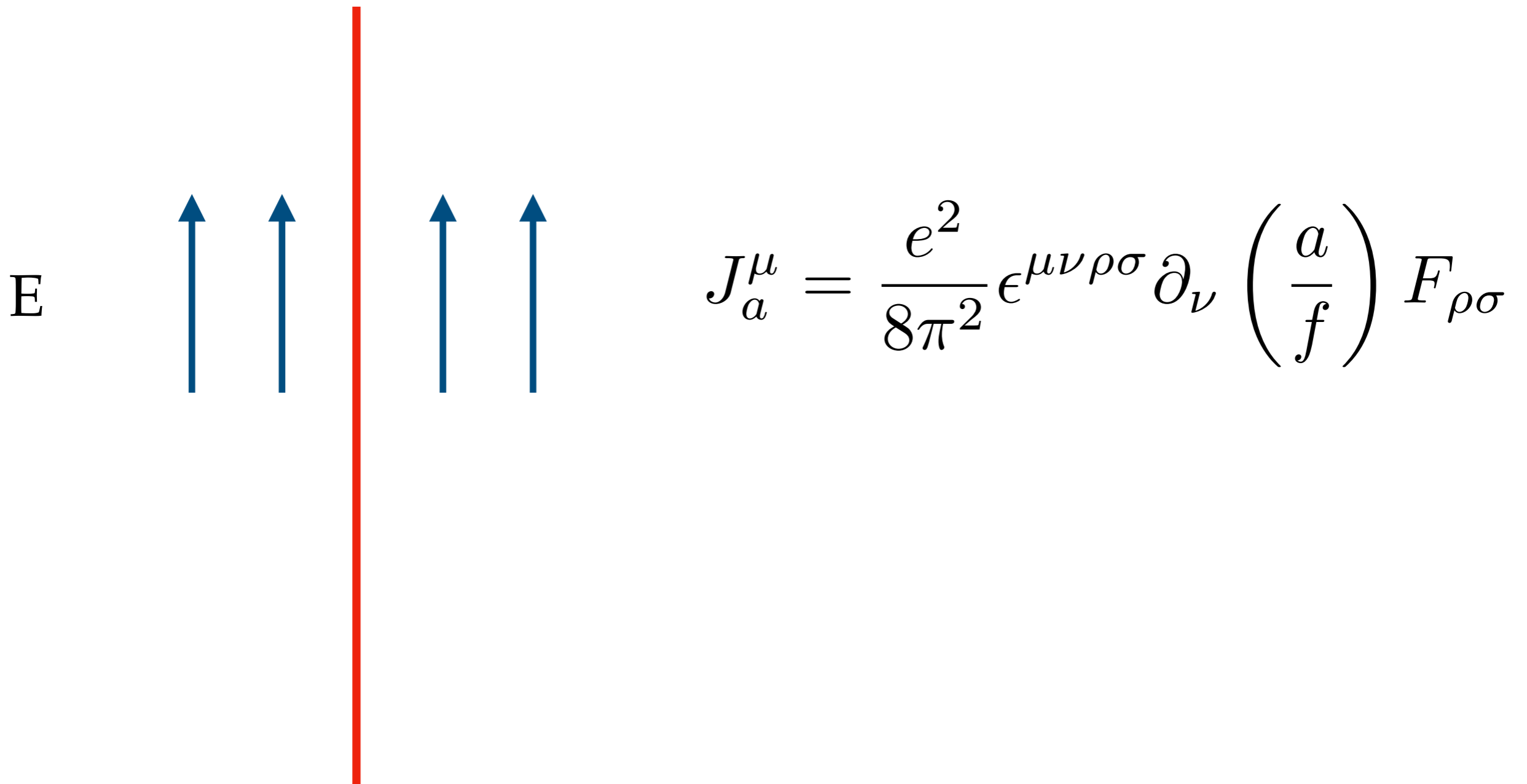
Conservation of charge

$$J_a^\mu = \frac{e^2}{8\pi^2} \epsilon^{\mu\nu\rho\sigma} \partial_\nu \left(\frac{a}{f} \right) F_{\rho\sigma}$$

In a background axion field, electric and magnetic fields generate additional charge and current

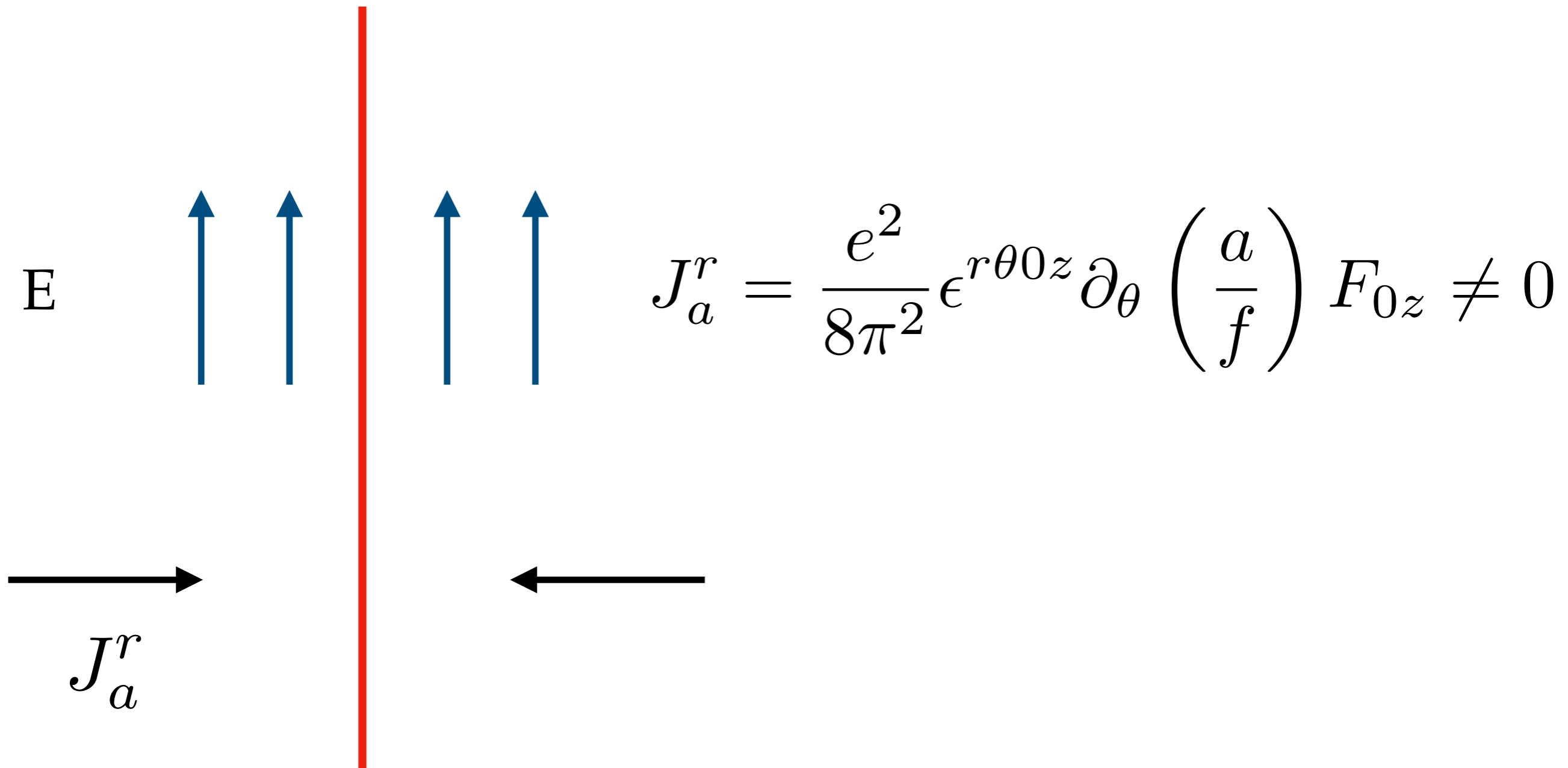
Conservation of charge

String in a background electric field



Conservation of charge

Charge is flowing onto the string!



Conservation of charge

Arbitrarily small electric field must cause charge to appear on the string

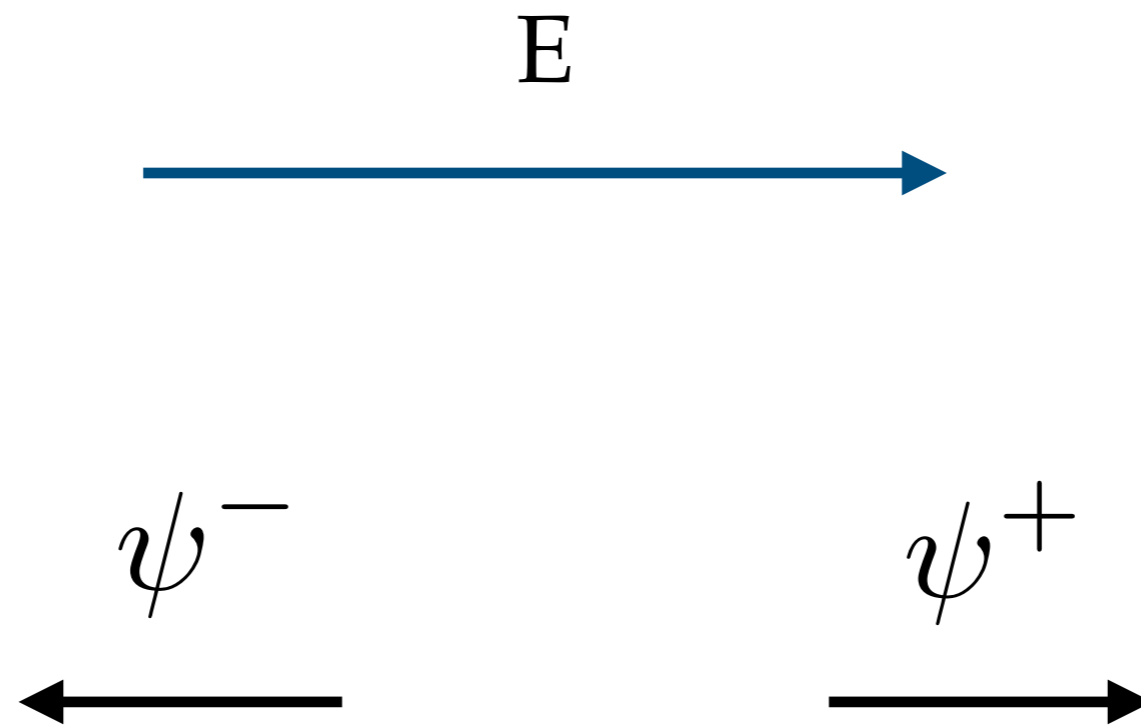
Schwinger pair production : Electric field produces particles

$$\frac{dQ}{dt} \propto e^{-\pi \frac{m^2}{eE}}$$

No exponential in production : must be massless charged particles!

Pair production

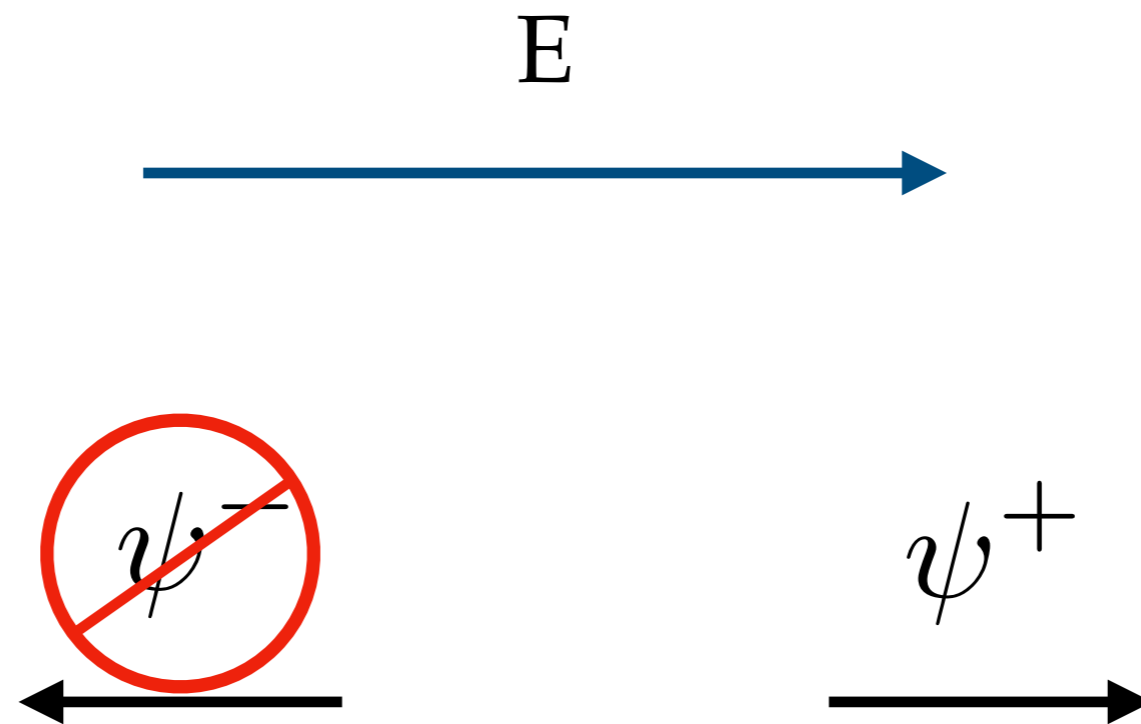
Schwinger pair production



Conserves charge

~~Pair~~ production

Schwinger ~~pair~~ production



Violates charge

Conservation of charge

$$\partial_{\mu} J_{\text{string}}^{\mu} = -\frac{\mathcal{A}e^2}{4\pi} \epsilon_{ab} F^{ab}$$

Conservation of total charge

Violation of string charge

Mismatch between left movers than right movers

Minimal situation : (anti-)particles move in
only one directions

Superconducting strings

$$\partial_{\mu} J_{\text{string}}^{\mu} = -\frac{Ae^2}{4\pi} \epsilon_{ab} F^{ab}$$

Consider a uniform string with no electric field

$$\frac{d\lambda}{dt} \propto \frac{dI}{dt} = 0$$

No resistance = superconductor!

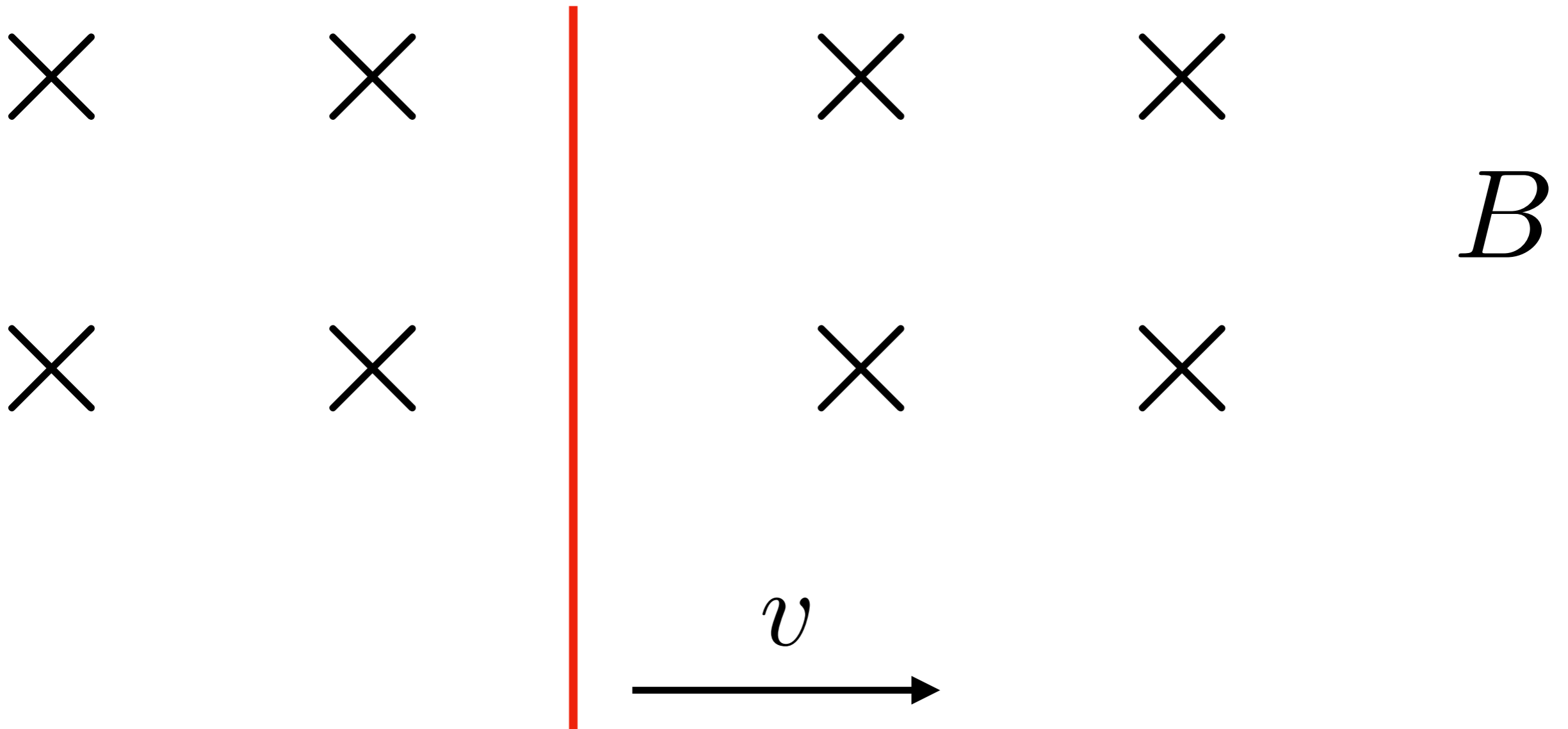
Penetration length comparable with string size

Outline

- ~~How strings affect photons~~
 - ~~Arahanov-Bohm effects~~
 - ~~CMB ways of looking for strings~~
- How photons affect strings
 - ~~Superconducting strings~~
 - Colliding currents

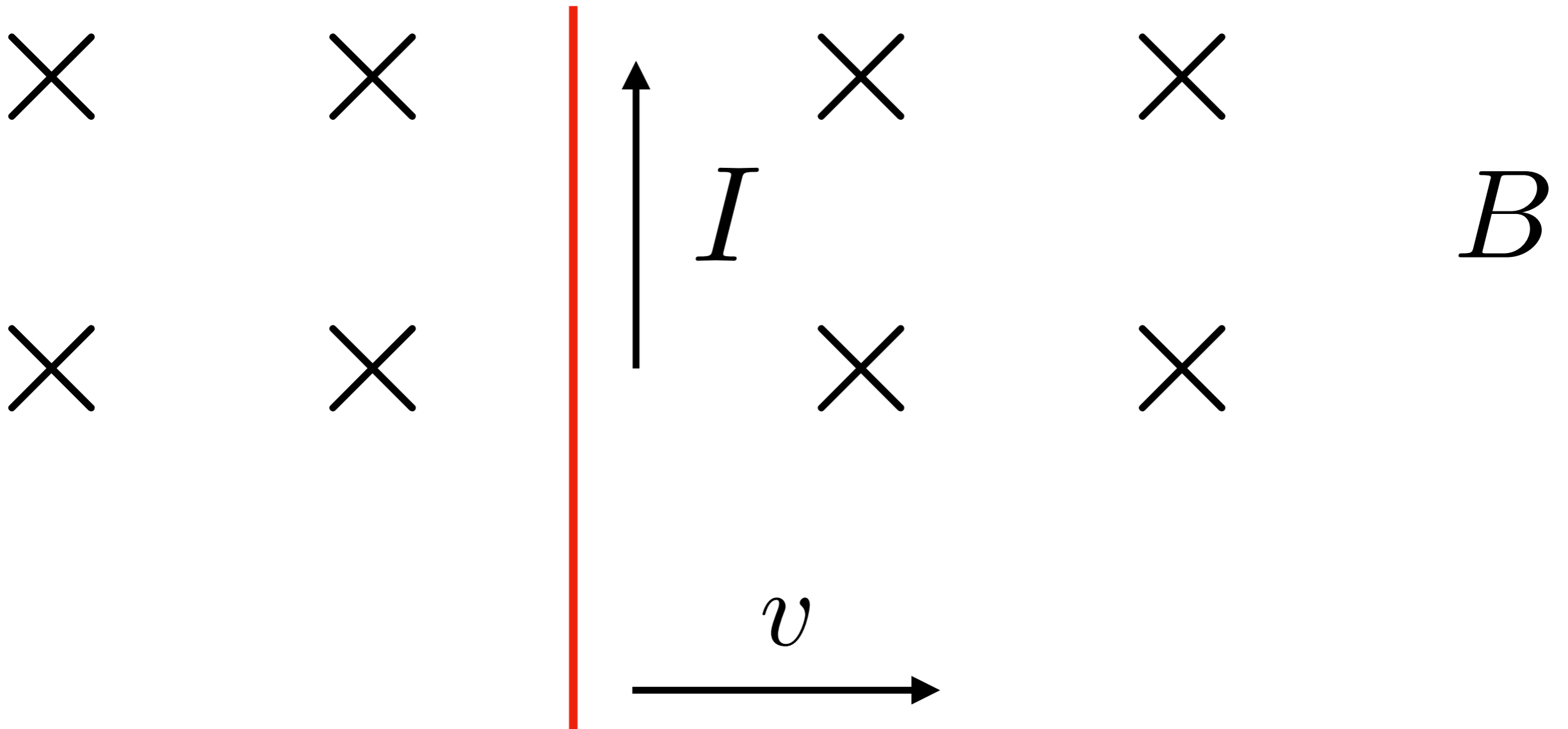
Screening of Flux

String moving through a B field of a galaxy



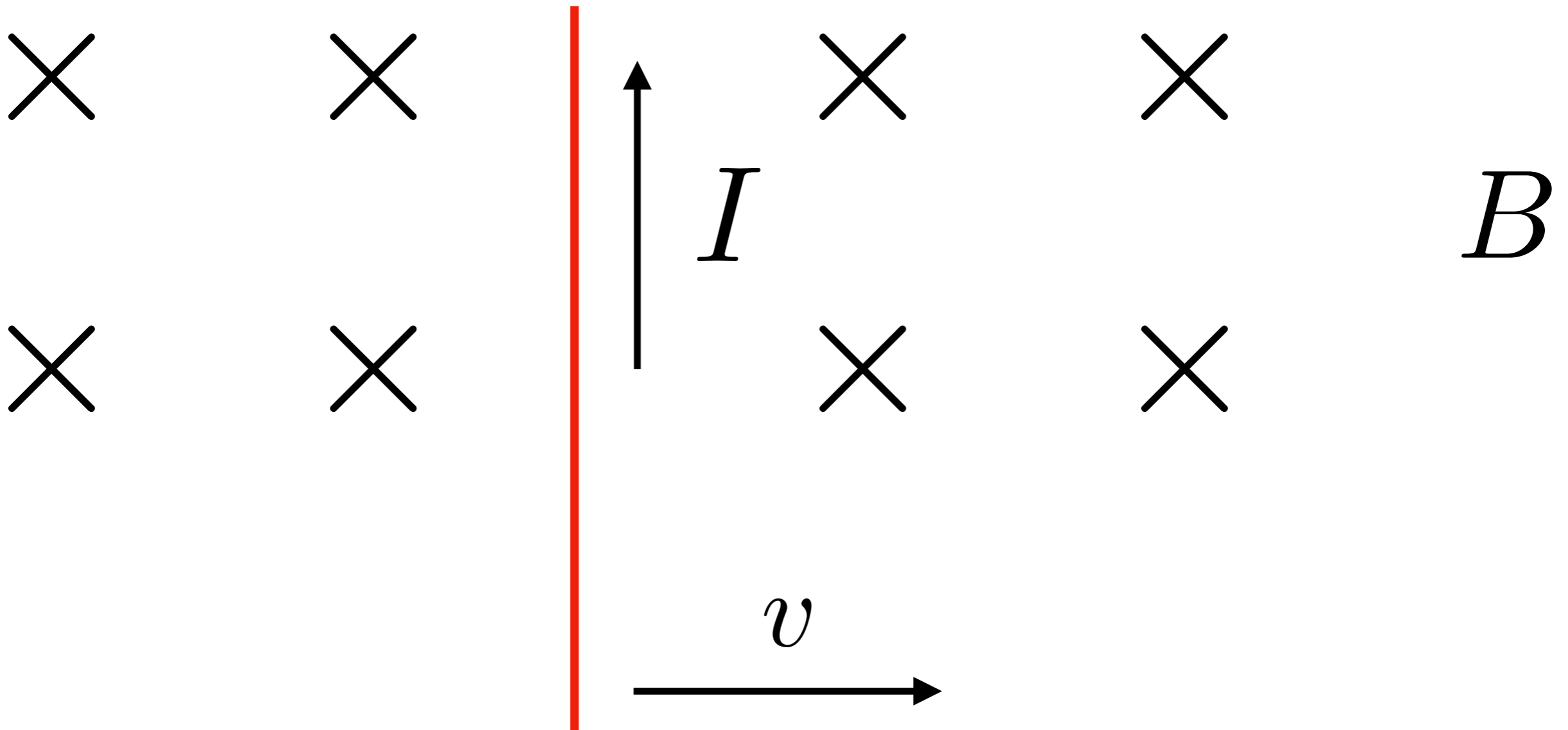
Screening of Flux

Lenz law current



Screening of Flux

Superconductor screens all flux



Screening of Flux

$$\begin{aligned}\Delta Q &= \int dt \dot{Q} = \int d^2x \frac{d\lambda_Q}{dt} = \int d^2x \partial_a j_{\text{string}}^a = -\frac{e^2}{4\pi} \int d^2x \epsilon^{ab} F_{ab} \\ &= -\frac{e^2}{2\pi} \int d^2x E = \frac{e^2}{2\pi} \int dt \frac{d\Phi}{dt} = \frac{e^2 \Delta\Phi}{2\pi}\end{aligned}$$

Take d/dz of above and assuming only left movers

$$I = \frac{\mathcal{A}e^2}{2\pi} \frac{d\Phi}{dz}$$

Screening of Flux

$$I = \frac{\mathcal{A}e^2}{2\pi} \frac{d\Phi}{dz} \qquad \frac{d\Phi_I}{dz} = \frac{I}{2\pi} \log(\Lambda R)$$

Flux gives current

Current gives flux

$$\Phi = \Phi_0 - \Phi_I$$

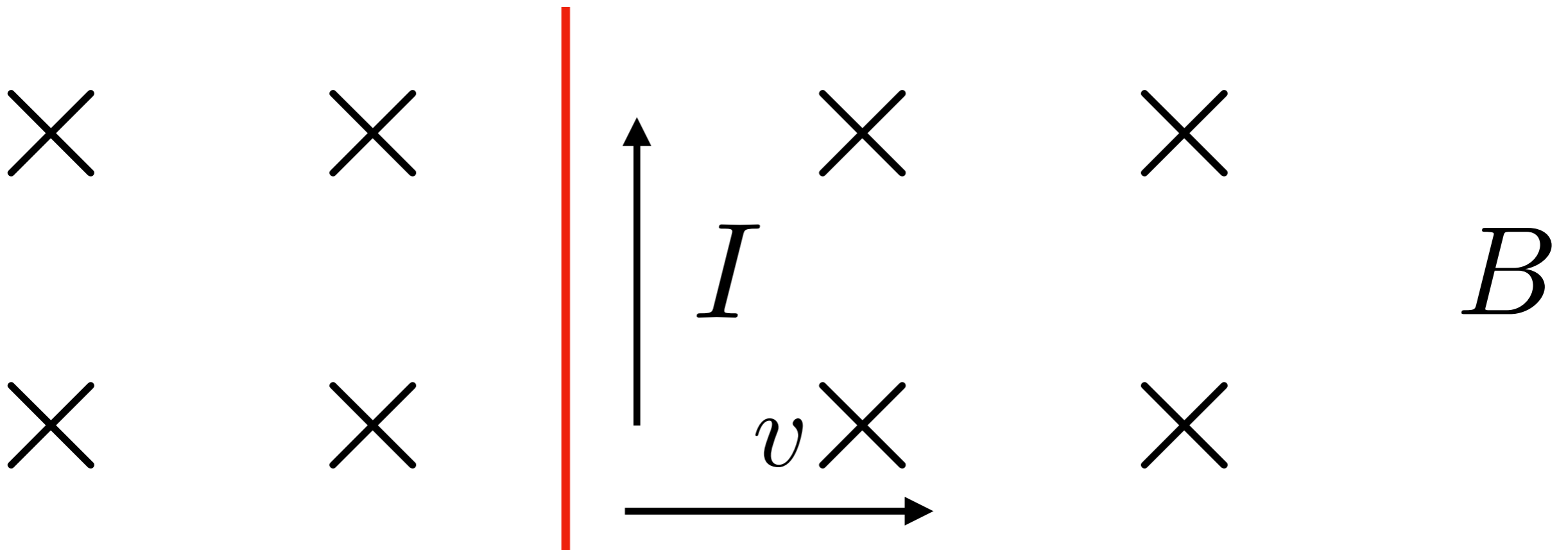
$$\Phi = \frac{\Phi_0}{1 + \frac{\mathcal{A}e^2}{4\pi} \log(\Lambda R)}$$

Flux is not perfectly screened

Currents coming from Galaxies

$$I = \frac{e^2}{2\pi} \frac{1}{1 + \frac{e^2}{4\pi} \log(\Lambda R)} B d v_s \approx 10^{8-9} \text{ GeV}$$

Axion string moving through a magnetic field



Collisions



$$P \sim 10^{40} \text{ erg/s} \sim 10^7 \mathcal{L}_{\text{Sun}}$$

Collisions are very bright!

How often do they occur?

Collisions



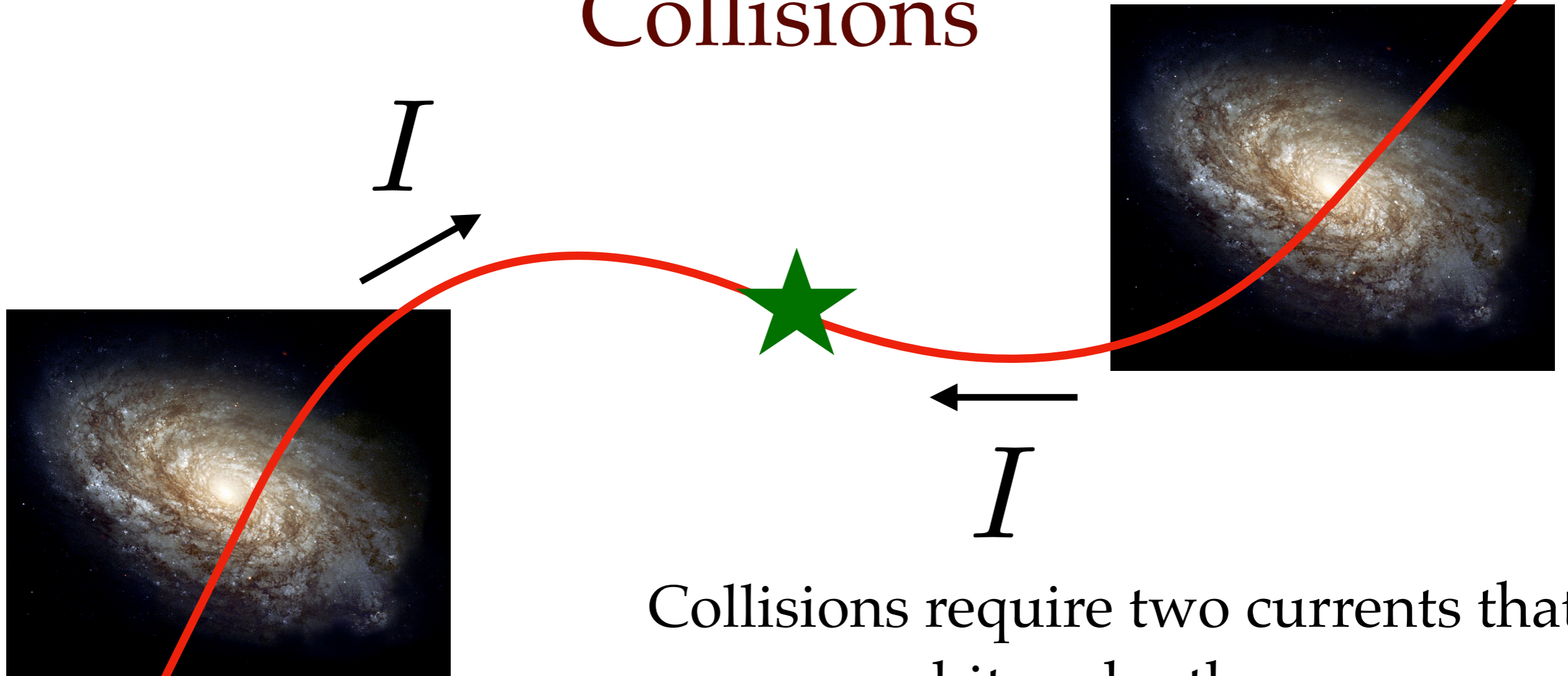
I

Currents are not generated very often

$$\begin{aligned} N_K &\simeq \xi H^3 L_{\text{string}} N_{\text{galaxy}} A_{\text{galaxy}} \\ &\approx 100 \left(\frac{\xi}{10} \right) \left(\frac{N_{\text{galaxy}}}{10^{12}} \right) \left(\frac{A_{\text{galaxy}}}{(10 \text{ kpc})^2} \right) \end{aligned}$$

O(100) String -
Galaxy crossings

Collisions

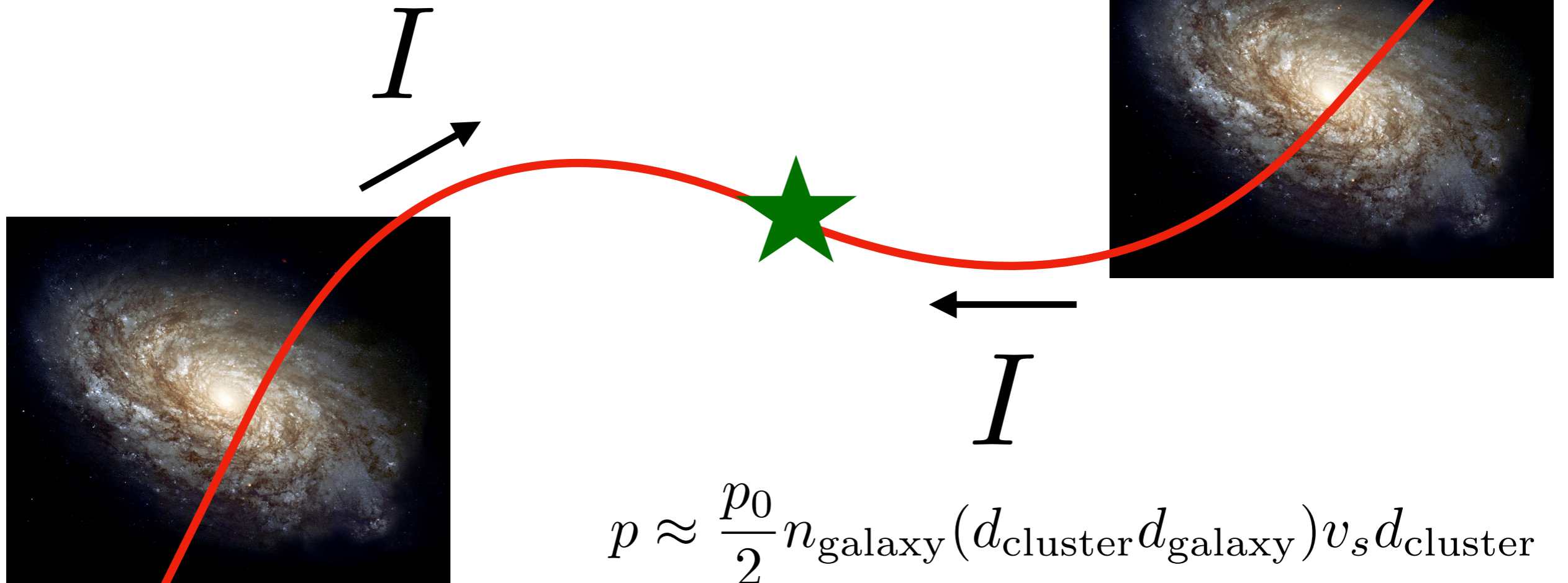


Collisions require two currents that hit each other

Happens if string passes through multiple galaxies in a galaxy cluster

Galaxy cluster $\sim 10^2 - 10^3$ galaxies in 10 Mpc distances

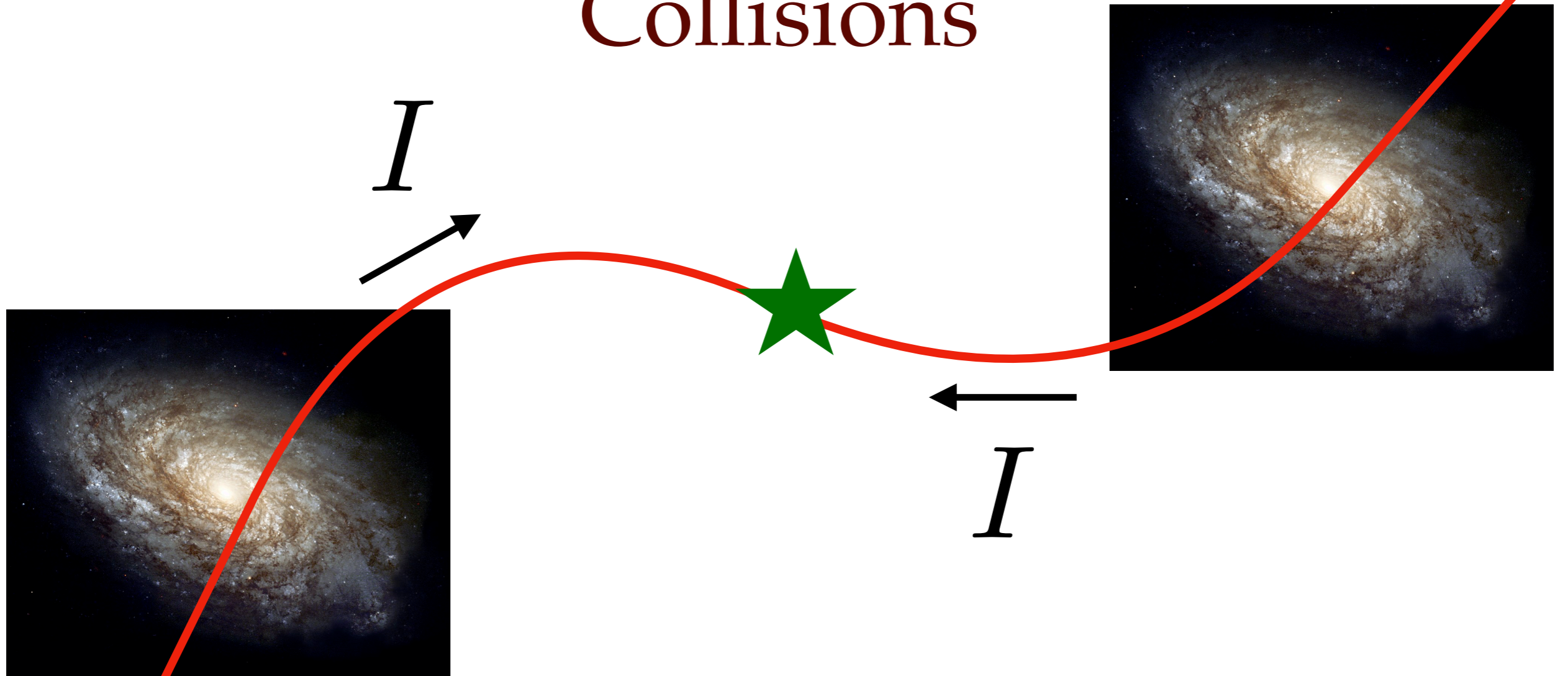
Collisions



$$p \approx \frac{p_0}{2} n_{\text{galaxy}} (d_{\text{cluster}} d_{\text{galaxy}}) v_s d_{\text{cluster}}$$
$$\approx 0.1 \left(\frac{v_s}{0.1} \right) \left(\frac{p_0 N_{\text{gc}}}{10^3} \right) \left(\frac{d_{\text{galaxy}}}{10 \text{ kpc}} \right) \left(\frac{10 \text{ Mpc}}{d_{\text{cluster}}} \right)$$

10% of the $O(100) = O(10)$ collisions at any given time

Collisions

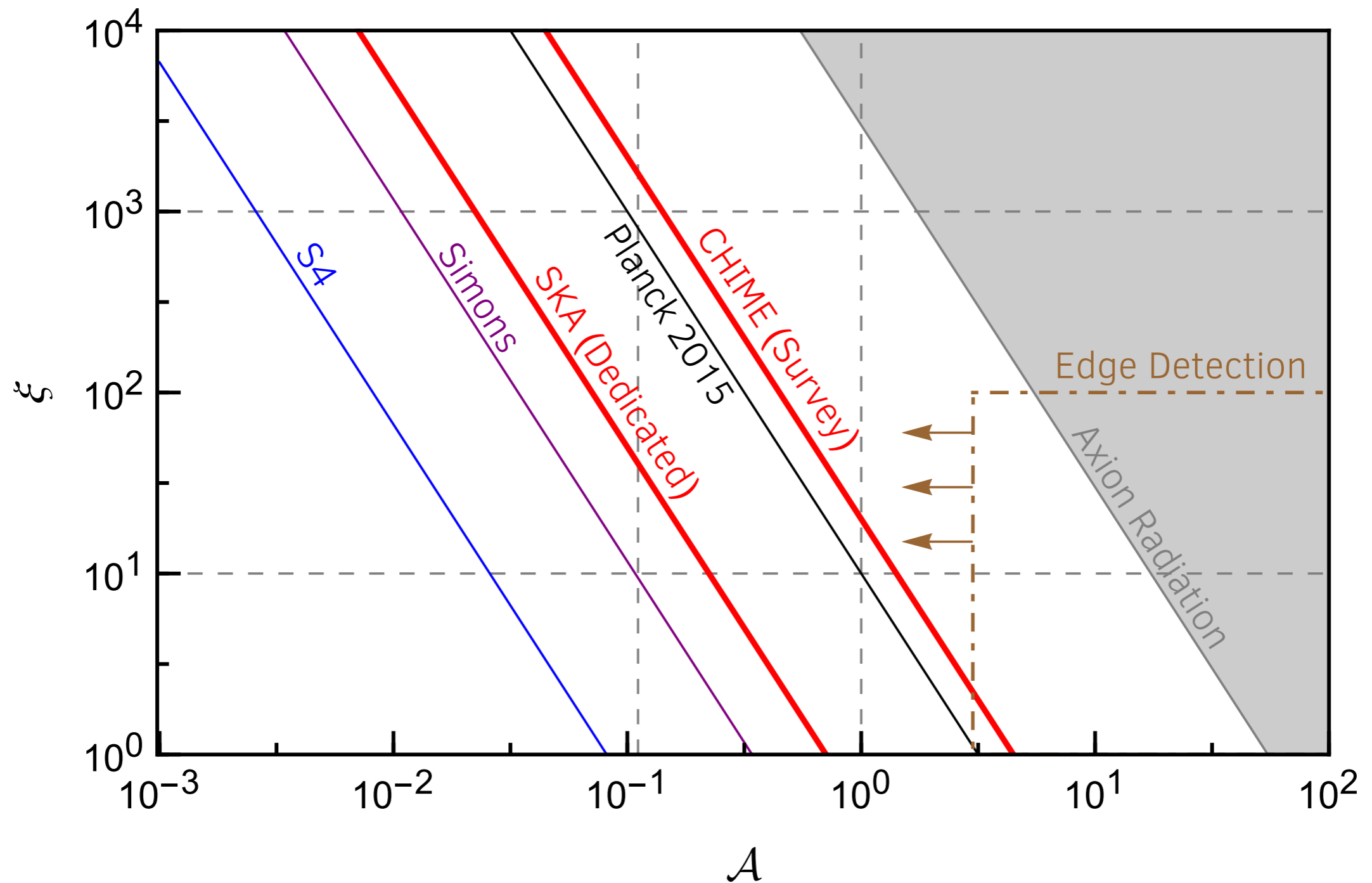


$$P \sim 10^{40} \text{ erg/s} \sim 10^7 \mathcal{L}_{\text{Sun}}$$

But order cosmological distances away

Sensitivity

If these collisions emit 10^{-3} of their energy into radio



Conclusion

Strings

$$\frac{A\alpha_{\text{em}}}{4\pi} \frac{a}{f} F \tilde{F}$$

Arahanov-Bohm features in the CMB

Power spectrum at the
edge of sensitivity

Edges in the CMB

Strings are superconductors

Faraday's Law : currents

Very bright signal coming from collisions