

# Dark Matter at future $e^+e^-$ colliders

## chances of detection vs DM spin

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based on

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**Dark-matter-spin effects at future  $e^+e^-$  colliders**

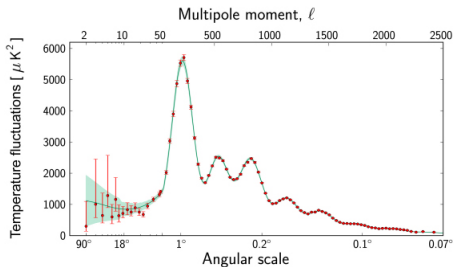
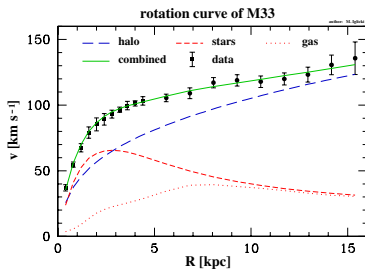
DOI:[10.1007/JHEP08\(2020\)052](https://doi.org/10.1007/JHEP08(2020)052), arXiv:[2003.06719](https://arxiv.org/abs/2003.06719)

**International Workshop on Future Linear Colliders**  
15-18 March 2021

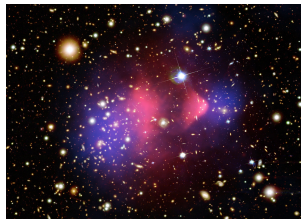
# Introduction: dark matter (DM)

## Evidence for dark matter

- rotation curves
- CMB fluctuations  $\Rightarrow$  ~~MOND~~
- gravitational lensing
- colliding clusters
- ...



source: <http://sci.esa.int/planck>



source: <https://apod.nasa.gov/apod>

DM believed to be beyond the Standard Model

## Collider search for dark matter

- Experimental approach: missing-energy analysis
- Signal more clear at  $e^+e^-$  than at hadron colliders
- Near-future plans:
  - ILC (international)
  - CEPC (China)
  - FCC-ee (Europe)
  - CLIC (Europe)

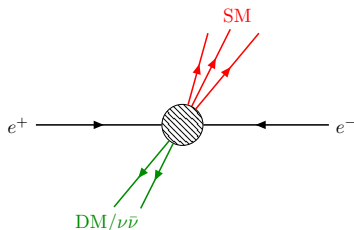
$$\sqrt{s} \approx 250 \text{ GeV}$$

ILC 1903.01629

CEPC 1811.10545

FCC CERN-ACC-2018-0057

CLIC 1812.06018



- 3 simple (but not simplified) DM models of different spins: 0, 1, 1/2
- Consider a mechanism of DM production at  $e^+e^-$  colliders
- Take current experimental constraints into account
- Check...
  - ...what range of parameters is still allowed?
  - ...how many DM-production events can we expect at future colliders\*?
  - ...whether it is possible to disentangle models of different spins?

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\*during calculations, collider parameters of ILC are used

The **main properties** of the models:

- Simple, but **renormalizable** and **QFT-consistent**
- Common **parameter space**
- **DM** connected to **SM** by the **Higgs portal**:  $\kappa|H|^2|S|^2$
- **Mixing** in the scalar sector

**real-part fluctuations** of a new singlet  $S$  and the Higgs doublet  $H$

$$\times \begin{bmatrix} \cos \alpha & -\sin \alpha \\ \sin \alpha & \cos \alpha \end{bmatrix}$$

two Higgs-like **mass eigenstates**:  $h_1$  ( $m_1 = 125$  GeV) and  $h_2$  ( $m_2 = ?$ )

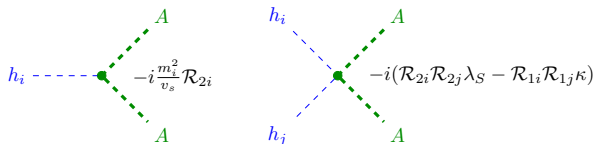
The **models**:

- **pseudo-Goldstone** (scalar) DM model ( $s = 0$ )  $\rightarrow$  dark scalar  $A$
- **vector** DM model ( $s = 1$ )  $\rightarrow$  dark gauge vector  $X_\mu$
- **fermion** DM model ( $s = 1/2$ )  $\rightarrow$  dark Majorana fermion  $\psi$

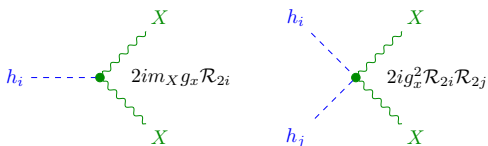
# Theoretical models

## DM interactions

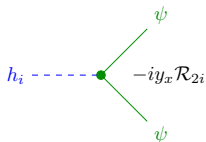
- pGDM



- VDM



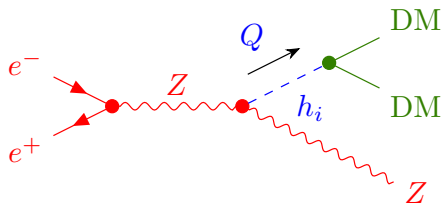
- FDM



Input parameters:  $\underbrace{v, m_1, v_s, m_2}_{\text{assumed to be SM-like}}, \sin \alpha, m_{\text{DM}} \longrightarrow \kappa \equiv \frac{(m_1^2 - m_2^2) \sin 2\alpha}{2vv_s}$

$\Downarrow$   
 DM **purely gravitational**  
 if  $m_1 = m_2$

# Considered process



$$Q^2 = s - 2E_Z\sqrt{s} + m_Z^2$$

$$\equiv m_{\text{rec}}^2$$

$$\text{DM} = A, X_\mu, \psi$$

**note:** in principle  $\Gamma_{1,2}$   
model-dependent

the differential cross-section:

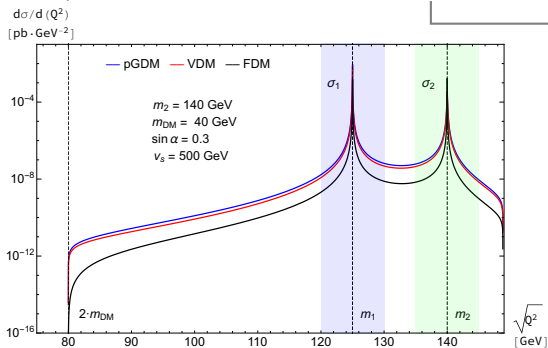
$$\frac{d\sigma}{dQ^2} \sim \sigma_{\text{SM}}(Q^2) \cdot \sin^2 \alpha \cos^2 \alpha \cdot \left| \frac{1}{Q^2 - m_1^2 + im_1\Gamma_1} - \frac{1}{Q^2 - m_2^2 + im_2\Gamma_2} \right|^2$$

$$\times \sqrt{1 - 4\frac{m_{\text{DM}}^2}{Q^2}} \cdot \begin{cases} 1 & (\text{pGDM}) \\ 1 - 4\frac{m_{\text{DM}}^2}{Q^2} + 12\left(\frac{m_{\text{DM}}^2}{Q^2}\right)^2 & (\text{VDM}) \\ 2\frac{m_{\text{DM}}^2}{Q^2} \left(1 - 4\frac{m_{\text{DM}}^2}{Q^2}\right) & (\text{FDM}) \end{cases}$$

# Considered process: shape of the differential cross section

$$\frac{d\sigma}{dQ^2} \sim \sigma_{SM}(Q^2) \cdot \sin^2 \alpha \cos^2 \alpha \cdot \left| \frac{1}{Q^2 - m_1^2 + im_1\Gamma_1} - \frac{1}{Q^2 - m_2^2 + im_2\Gamma_2} \right|^2$$

$$\times \sqrt{1 - 4 \frac{m_{DM}^2}{Q^2}} \cdot \begin{cases} 1 & \text{(pGDM)} \\ 1 - 4 \frac{m_{DM}^2}{Q^2} + 12 \left( \frac{m_{DM}^2}{Q^2} \right)^2 & \text{(VDM)} \\ 2 \frac{m_{DM}^2}{Q^2} \left( 1 - 4 \frac{m_{DM}^2}{Q^2} \right) & \text{(FDM)} \end{cases}$$



- **Shape fitting**  
⇒ spin **hard to check**
- **Total  $\sigma$  as a function of  $\sqrt{s}$**   
( $\sqrt{Q^2} = 2m_{DM}$  threshold)  
⇒ mass of DM possible to determine

after integration:  $\sigma \approx \sigma_1 \cdot \mathbb{1}_{2m_{DM} < m_1 < \sqrt{s} - m_Z} + \sigma_2 \cdot \mathbb{1}_{2m_{DM} < m_2 < \sqrt{s} - m_Z}$

$$\sigma_1 \equiv \sigma_{SM}(m_1) \cdot \cos^2 \alpha \cdot \text{BR}(h_1 \rightarrow \text{DM})$$

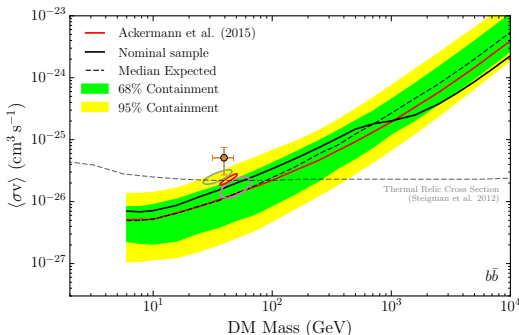
$$\sigma_2 \equiv \sigma_{SM}(m_2) \cdot \sin^2 \alpha \cdot \text{BR}(h_2 \rightarrow \text{DM})$$

# Current limits and constraints

1. **Perturbativity** ( $h_i$ -DM):  $g_x, y_x < 4\pi \Rightarrow v_s > m_{\text{DM}}/4\pi$

2.  $h^2 \Omega_0^{\text{DM}} = 0.12 \pm 0.0012 \Rightarrow$  **constraint on  $\langle\sigma v\rangle_{\text{ann}}$**   $\Rightarrow$  **constraint on  $\kappa^2$**  Planck 1807.06209

3. **ID experiments**  $\Rightarrow$  **limit on  $\langle\sigma v\rangle_{\text{ann}}$**   $\Rightarrow$  **limit on  $m_{\text{DM}}$**  Fermi-LAT 1611.03184



4. **DD experiments and  $\nu$  telescopes**  $\Rightarrow$  **limit on  $\sigma^{\text{DD}}$**   $\Rightarrow$  **limit on  $\kappa^2$**  XENON1T 1805.12562  
IceCube 1601.00653

5. **LHC**  $\Rightarrow |\sin \alpha| \lesssim 0.3$  Robens, Stefaniak 1601.07880

6. **LHC**  $\Rightarrow \text{BR}(h_1 \rightarrow \text{DM}) < 19\%$  CMS 1809.05937

- Constraints:

- $|\sin \alpha| \lesssim 0.3$
  - relic density ( $\Omega_0^{\text{DM}}$ ) constraint on  $\kappa^2$
  - $v_s < \frac{m_{\text{DM}}}{4\pi}$
  - direct detection (DD) limit on  $\kappa^2$
  - $\text{BR}(h_1 \rightarrow \text{DM}) < 19\%$
- } taken into account **numerically**

- Free parameters:  $m_2, m_{\text{DM}}, \sin \alpha, v_s \rightarrow \kappa^2$

- $\kappa^2 = \kappa^2(\Omega_0^{\text{DM}}) \Rightarrow$  3 parameters left

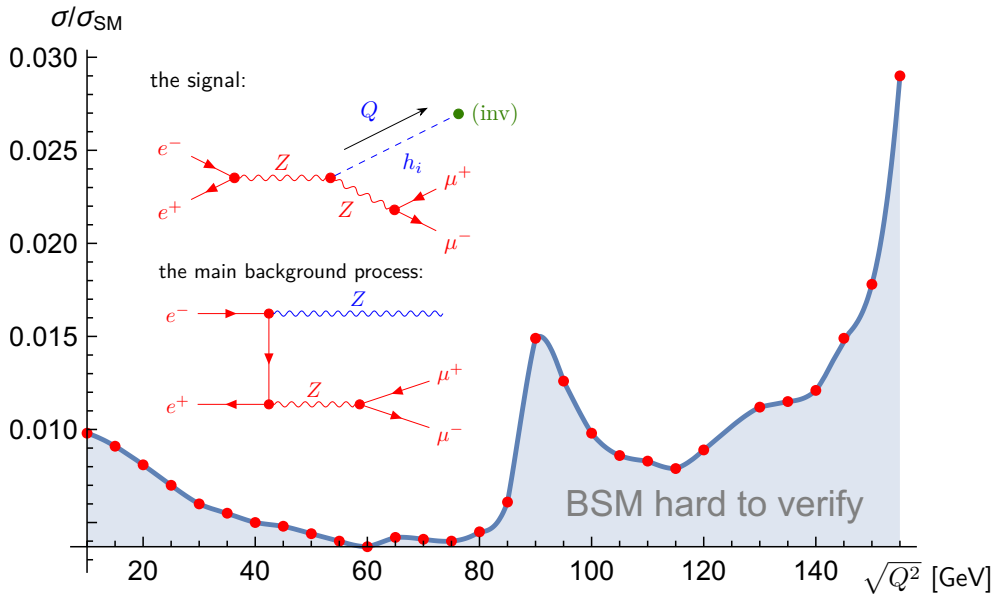
- Cross section maximized with respect to  $\sin \alpha$

$\Rightarrow \sin \alpha \rightarrow \sin \alpha_{\text{max}} = 0.3$ , free parameters:  $m_2, m_{\text{DM}}$

- Low  $\sqrt{s}$  gives high cross section

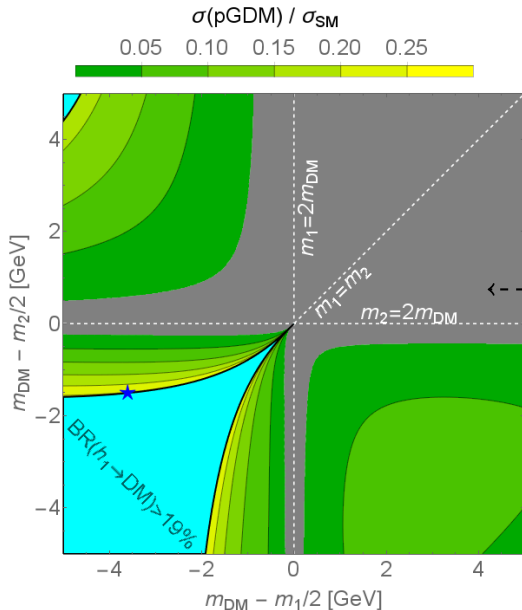
(ILC:  $\sqrt{s} = 250$  GeV, 20-year luminosity:  $\int \mathcal{L} dt = 2000 \text{ fb}^{-1}$ )

ILC: 1903.01629



(see also [Krzysztof Mękała's talk: CLIC sensitivity to invisible scalar decays](#))

# Results: maximal cross section for pGDM



parameters of ★

$$m_2 = 120.8 \text{ GeV}$$

$$m_{\text{DM}} = 58.9 \text{ GeV}$$

$$\sin \alpha = 0.30$$

$$v_s = 646 \text{ GeV}$$

$$\Gamma_1 = 7.4 \cdot 10^{-3} \text{ GeV}$$

$$\Gamma_2 = 9.8 \cdot 10^{-3} \text{ GeV}$$

$$\text{BR}(h_1 \rightarrow \text{DM}) = 19\%$$

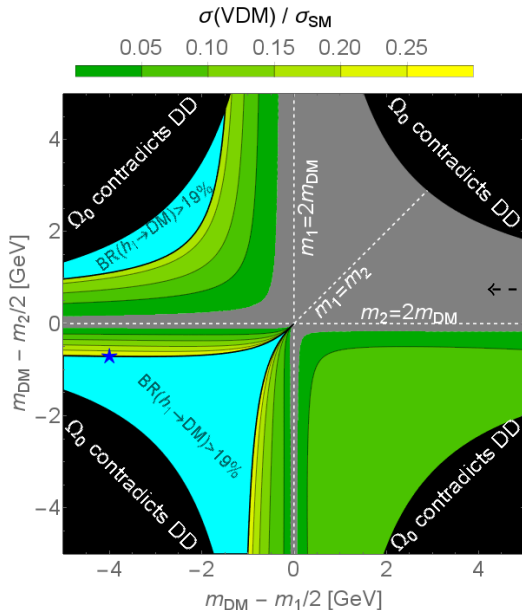
$$\text{BR}(h_2 \rightarrow \text{DM}) = 95\%$$

$$\sigma = 62 \text{ fb} \rightarrow 1.24 \cdot 10^5 \text{ events}$$

below the 95% CL  
limit for the ILC

1903.01629

# Results: maximal cross section for VDM



parameters of ★

$$m_2 = 118.4 \text{ GeV}$$

$$m_{\text{DM}} = 58.5 \text{ GeV}$$

$$\sin \alpha = 0.30$$

$$v_s = 561 \text{ GeV}$$

$$\Gamma_1 = 7.4 \cdot 10^{-3} \text{ GeV}$$

$$\Gamma_2 = 6.4 \cdot 10^{-3} \text{ GeV}$$

$$\text{BR}(h_1 \rightarrow \text{DM}) = 18\%$$

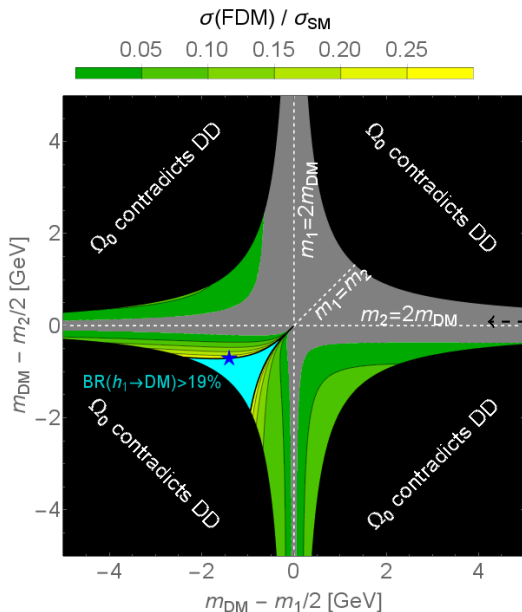
$$\text{BR}(h_2 \rightarrow \text{DM}) = 92\%$$

$$\sigma = 61 \text{ fb} \rightarrow 1.22 \cdot 10^5 \text{ events}$$

below the 95% CL  
limit for the ILC

1903.01629

# Results: maximal cross section for FDM



parameters of ★

$$m_2 = 123.6 \text{ GeV}$$

$$m_{\text{DM}} = 61.1 \text{ GeV}$$

$$\sin \alpha = 0.30$$

$$v_s = 76 \text{ GeV}$$

$$\Gamma_1 = 7.4 \cdot 10^{-3} \text{ GeV}$$

$$\Gamma_2 = 5.9 \cdot 10^{-3} \text{ GeV}$$

$$BR(h_1 \rightarrow \text{DM}) = 18\%$$

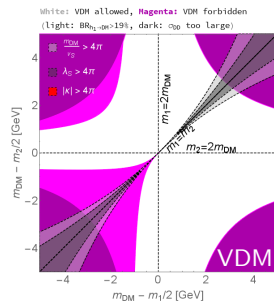
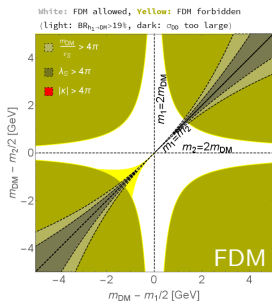
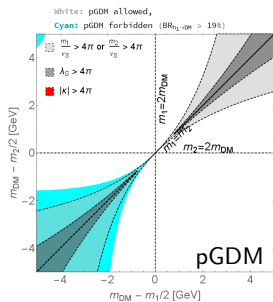
$$BR(h_2 \rightarrow \text{DM}) = 91\%$$

$$\sigma = 59 \text{ fb} \rightarrow 1.18 \cdot 10^5 \text{ events}$$

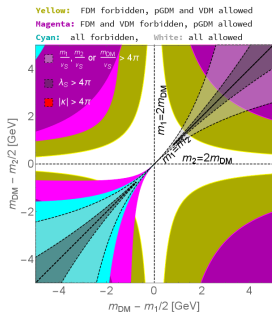
below the 95% CL  
limit for the ILC

1903.01629

# Results: allowed and forbidden regions

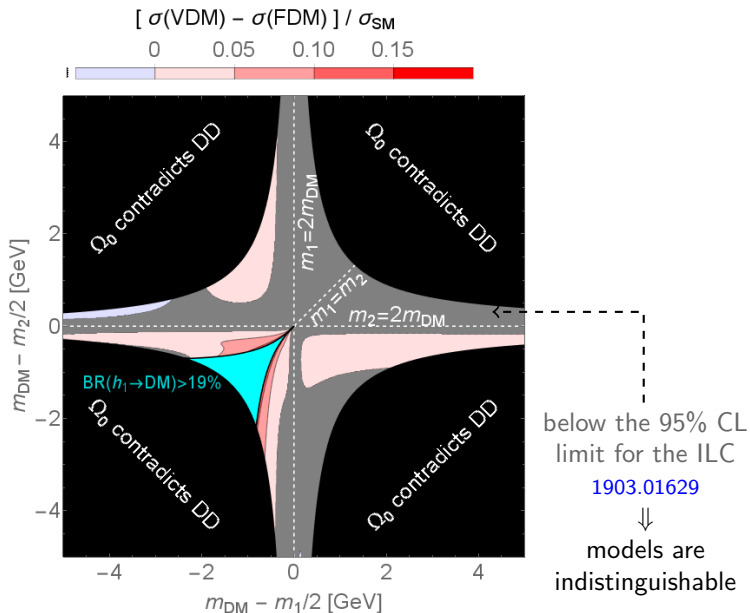


all plots combined:

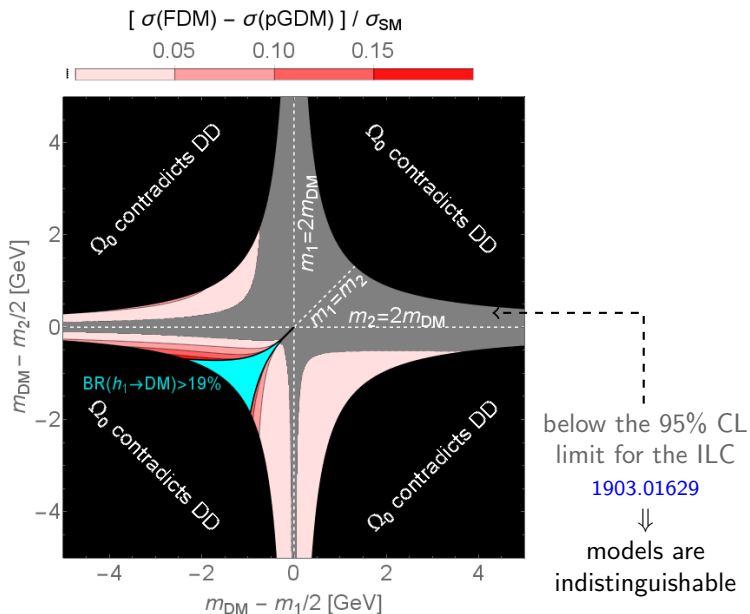


- ⇒ there exist regions where
- only pGDM and VDM are allowed
  - only pGDM is allowed
  - nothing is allowed

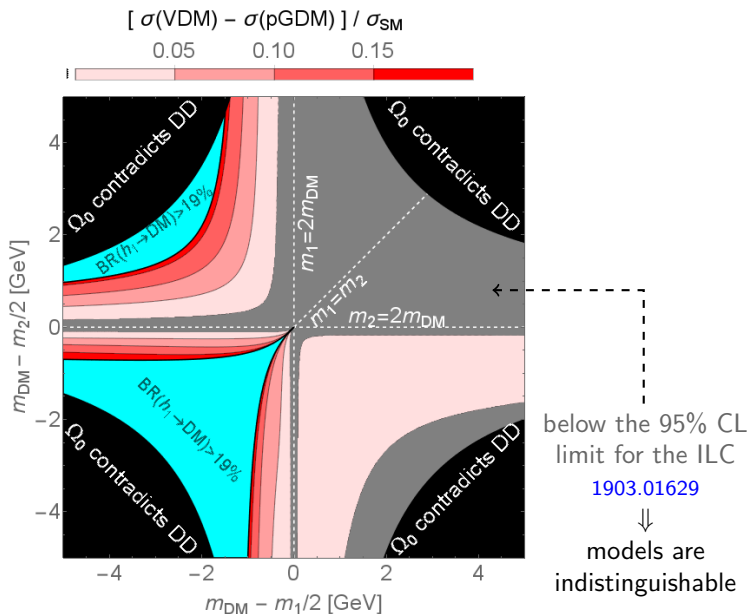
# Results: fermion vs vector



# Results: pseudo-Goldstone vs fermion



# Results: pseudo-Goldstone vs vector



We introduce 3 simple models:

- pseudo-Goldstone dark matter model ( $s = 0$ )
- vector dark matter model ( $s = 1$ )
- fermion dark matter model ( $s = 1/2$ )

which are QFT-consistent and share common parameter space

Comparing models of various spins we conclude that...

- ...maximal cross section is similar in all 3 cases ( $\sim 1.2 \cdot 10^5$  events / 20 y)
- ...signal-to-background ratio can be  $\sim 10\%$
- ...allowed range of parameters is largest for pGDM
- ...there are regions where cases of different spins could be disentangled

But...

- ...if parameters far from optimal, DM of any spin hard to be detected

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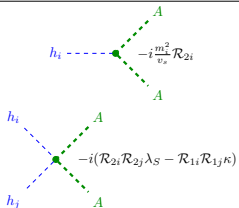
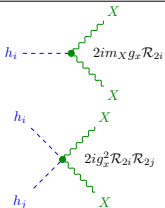
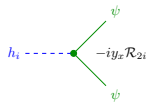
Thank you!

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a lot of

# BACKUP SLIDES

# Theoretical models – comparison

	PSEUDO-GOLDSTONE DM MODEL	VECTOR DM MODEL	FERMION DM MODEL
symmetry group	$\mathcal{G} = \mathcal{G}_{SM} \times \mathbb{Z}_2 \times U(1)_X$ <small>↖ global, softly broken</small>	$\mathcal{G} = \mathcal{G}_{SM} \times U(1)_X$ <small>↖ local</small>	$\mathcal{G} = \mathcal{G}_{SM} \times \mathbb{Z}_4$
new states ( $\mathcal{G}_{SM}$ -even)	complex scalar $S$ ( $q = (1, 1)$ )	gauge vector $X_\mu$ complex scalar $S$ ( $q = 1$ )	LH fermion $\chi$ ( $q = 1$ ) real scalar $S$ ( $q = 2$ )
Lagrangian	$\mathcal{L} = \mathcal{L}_{SM} + \mathcal{L}_{DM} - V(H, S)$		
scalar potential	$V(H, S) = -\mu_H^2  H ^2 + \lambda_H  H ^4$ $-\mu_S^2  S ^2 + \lambda_S  S ^4$ $+\kappa  H ^2  S ^2 + \mu^2 (S^2 + S^{*2})$	$V(H, S) = -\mu_H^2  H ^2 + \lambda_H  H ^4$ $-\mu_S^2  S ^2 + \lambda_S  S ^4$ $+\kappa  H ^2  S ^2$	$V(H, S) = -\mu_H^2  H ^2 + \lambda_H  H ^4$ $-\frac{\mu_S^2}{2} S^2 + \frac{\lambda_S}{4} S^4$ $+\frac{\kappa}{2}  H ^2 S^2$
SSB	$H \rightarrow (\pi^+, v + h + i\pi^0)^T / \sqrt{2}$ $S \rightarrow (v_s + \phi + iA) / \sqrt{2}$	$H \rightarrow (\pi^+, v + h + i\pi^0)^T / \sqrt{2}$ $S \rightarrow (v_s + \phi + i\sigma) / \sqrt{2}$	$H \rightarrow (\pi^+, v + h + i\pi^0)^T / \sqrt{2}$ $S \rightarrow v_s + \phi$
Higgs sector mixing	$\begin{pmatrix} h \\ \phi \end{pmatrix} = \mathcal{R} \begin{pmatrix} h_1 \\ h_2 \end{pmatrix} = \begin{bmatrix} \cos \alpha & -\sin \alpha \\ \sin \alpha & \cos \alpha \end{bmatrix} \begin{pmatrix} h_1 \\ h_2 \end{pmatrix} \quad \text{tg } 2\alpha = \frac{\kappa v v_s}{\lambda_H v^2 - \lambda_S v_s^2}$		
dark state	$A \equiv \sqrt{2} \text{Im } S$ ( $m_A^2 = -4\mu^2$ )	$X_\mu$ ( $m_X^2 = g_X^2 v_s^2$ )	$\psi \equiv \chi + \chi^c$ ( $m_\psi^2 = y_x^2 v_s^2$ )
dark matter interactions			

input parameters:  $\underbrace{v, m_1, v_s, m_2}_{\text{assumed to be SM-like}}, \sin \alpha, m_{DM} \rightarrow \kappa \equiv \frac{(m_1^2 - m_2^2) \sin 2\alpha}{2v v_s} \Rightarrow \text{DM purely gravitational if } m_1 = m_2$

- Gauge group:  $\mathcal{G} = \underbrace{SU(3)_c \times SU(2)_L \times U(1)_Y}_{\text{Standard Model gauge group}} \times \mathbb{Z}_2 \times \overbrace{U(1)_X}^{\text{global symmetry}}$
- $U(1)_X$  softly broken to  $\mathbb{Z}'_2$
- Complex scalar  $S$  introduced,  $\mathbb{Z}_2 : S \rightarrow S^*$ ,  $\mathbb{Z}'_2 : S \rightarrow -S$
- Lagrangian

$$\mathcal{L} = \mathcal{L}_{\text{SM}} + (\partial^\mu S)^* (\partial_\mu S) - V(H, S)$$

$$V(H, S) = -\mu_H^2 |H|^2 + \lambda_H |H|^4 - \mu_S^2 |S|^2 + \lambda_S |S|^4 + \underbrace{\kappa |H|^2 |S|^2}_{\text{Higgs portal coupling}} + \overbrace{\mu^2 (S^2 + S^{*2})}^{U(1)_X\text{-breaking term}}$$

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- SSB  $H \rightarrow (\pi^+, v + h + i\pi^0)^T / \sqrt{2}$   $S \rightarrow (v_s + \phi + iA) / \sqrt{2}$
- We introduce  $h_1, h_2$

$$\begin{pmatrix} h \\ \phi \end{pmatrix} = \underbrace{\begin{bmatrix} \cos \alpha & -\sin \alpha \\ \sin \alpha & \cos \alpha \end{bmatrix}}_{\mathcal{R}} \begin{pmatrix} h_1 \\ h_2 \end{pmatrix}$$

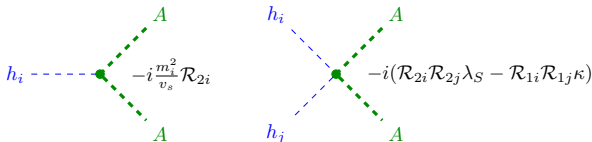
$$\text{tg } 2\alpha = \kappa v v_s / (\lambda_H v^2 - \lambda_S v_s^2)$$

$$m_1^2 = \lambda_H v^2 (1 + \sec 2\alpha) + \lambda_S v_s^2 (1 - \sec 2\alpha)$$

$$m_2^2 = \lambda_H v^2 (1 - \sec 2\alpha) + \lambda_S v_s^2 (1 + \sec 2\alpha)$$

$$\begin{aligned} V(H, S) \longrightarrow & \frac{1}{2} m_1^2 h_1^2 + \frac{1}{2} m_2^2 h_2^2 - 2\mu^2 A^2 + \frac{m_1^2 \sin^2 \alpha + m_2^2 \cos^2 \alpha}{8v_s^2} A^4 \\ & - \sum_{i=1,2} \frac{1}{2} \frac{m_i^2 \mathcal{R}_{2i}}{v_s} A^2 h_i + \sum_{i,j=1,2} \frac{1}{4} (\mathcal{R}_{2i} \mathcal{R}_{2j} \lambda_S - \mathcal{R}_{1i} \mathcal{R}_{1j} \kappa) A^2 h_i h_j \\ & + \text{const} + \mathcal{O}([h_1, h_2]^3) \end{aligned}$$

- $A \equiv \sqrt{2} \text{Im}(S)$  is a massive dark particle,  $m_A^2 = -4\mu^2$



- Gauge group:  $\mathcal{G} = \underbrace{SU(3)_c \times SU(2)_L \times U(1)_Y}_{\text{Standard Model gauge group}} \times \overbrace{U(1)_X}^{\text{gauge symmetry}}$
- $U(1)_X$  gauge vector  $X_\mu$  and complex scalar  $S$  introduced
- Discrete  $\mathbb{Z}_2$  symmetry:  $X_\mu \rightarrow -X_\mu, S \rightarrow S^*$   
 $\Rightarrow$  no kinetic mixing with SM fields
- Lagrangian

$$\mathcal{L} = \mathcal{L}_{\text{SM}} + \mathcal{L}_{\text{DM}} - V(H, S)$$

$$\mathcal{L}_{\text{DM}} = -\frac{1}{4} \mathcal{F}_{\mu\nu} \mathcal{F}^{\mu\nu} + (\mathcal{D}^\mu S)^* (\mathcal{D}_\mu S)$$

$$V(H, S) = -\mu_H^2 |H|^2 + \lambda_H |H|^4 - \mu_S^2 |S|^2 + \lambda_S |S|^4 + \underbrace{\kappa |H|^2 |S|^2}_{\text{Higgs portal coupling}}$$

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- We introduce  $h_1, h_2$

$$\begin{pmatrix} h \\ \phi \end{pmatrix} = \underbrace{\begin{bmatrix} \cos \alpha & -\sin \alpha \\ \sin \alpha & \cos \alpha \end{bmatrix}}_{\mathcal{R}} \begin{pmatrix} h_1 \\ h_2 \end{pmatrix}$$

$$\text{tg } 2\alpha = \kappa v v_s / (\lambda_H v^2 - \lambda_S v_s^2)$$

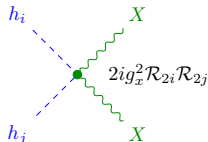
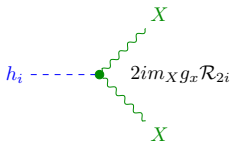
$$m_1^2 = \lambda_H v^2 (1 + \sec 2\alpha) + \lambda_S v_s^2 (1 - \sec 2\alpha)$$

$$m_2^2 = \lambda_H v^2 (1 - \sec 2\alpha) + \lambda_S v_s^2 (1 + \sec 2\alpha)$$

$$\mathcal{L}_{\text{DM}} \longrightarrow -\frac{1}{4}\mathcal{F}_{\mu\nu}\mathcal{F}^{\mu\nu} + \frac{1}{2}\partial^\mu \phi \partial_\mu \phi + \frac{g_x^2 v_s^2}{2} X^\mu X_\mu + \frac{g_x^2}{2} X^\mu X_\mu \phi^2 + g_x^2 v_s X^\mu X_\mu \phi$$

$$V(H, S) \longrightarrow \frac{1}{2}m_1^2 h_1^2 + \frac{1}{2}m_2^2 h_2^2 + \text{const} + \mathcal{O}([\text{field}]^3)$$

- $X_\mu$  is a massive dark particle,  $m_X^2 = g_x^2 v_s^2$



- Gauge group:  $\mathcal{G} = \underbrace{SU(3)_c \times SU(2)_L \times U(1)_Y}_{\text{Standard Model gauge group}} \times \mathbb{Z}_4$

- Left-handed fermion  $\chi$  and real scalar  $S$  introduced,  
 $\mathbb{Z}_4 : \chi \rightarrow i\chi, S \rightarrow -S$

- Lagrangian

$$\mathcal{L} = \mathcal{L}_{\text{SM}} + \mathcal{L}_{\text{DM}} - V(H, S)$$

$$\mathcal{L}_{\text{DM}} = i\bar{\chi}\not{\partial}\chi + \frac{1}{2}\partial^\mu S\partial_\mu S - \frac{y_\chi}{2}(\bar{\chi}^c\chi + \bar{\chi}\chi^c)S$$

$$V(H, S) = -\mu_H^2|H|^2 + \lambda_H|H|^4 - \frac{\mu_S^2}{2}S^2 + \frac{\lambda_S}{4}S^4 + \underbrace{\frac{\kappa}{2}|H|^2S^2}_{\text{Higgs portal coupling}}$$

$$\mathcal{L}_{\text{DM}} = i\bar{\chi}\not{\partial}\chi + \frac{1}{2}\partial^\mu S \partial_\mu S - \frac{y_x}{2}(\bar{\chi}^c\chi + \bar{\chi}\chi^c)S$$

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- SSB  $H \rightarrow (\pi^+, v + h + i\pi^0)^T/\sqrt{2}$   $S \rightarrow v_s + \phi$
- We introduce  $h_1, h_2$

$$\begin{pmatrix} h \\ \phi \end{pmatrix} = \underbrace{\begin{bmatrix} \cos \alpha & -\sin \alpha \\ \sin \alpha & \cos \alpha \end{bmatrix}}_{\mathcal{R}} \begin{pmatrix} h_1 \\ h_2 \end{pmatrix}$$

$$\text{tg } 2\alpha = \kappa v v_s / (\lambda_H v^2 - \lambda_S v_s^2)$$

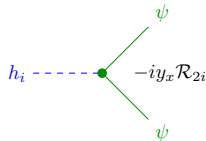
$$m_1^2 = \lambda_H v^2 (1 + \sec 2\alpha) + \lambda_S v_s^2 (1 - \sec 2\alpha)$$

$$m_2^2 = \lambda_H v^2 (1 - \sec 2\alpha) + \lambda_S v_s^2 (1 + \sec 2\alpha)$$

$$\mathcal{L}_{\text{DM}} \longrightarrow \frac{i}{2}\bar{\psi}\not{\partial}\psi + \frac{1}{2}\partial^\mu \phi \partial_\mu \phi - \frac{y_x v_s}{2}\bar{\psi}\psi - \frac{y_x}{2}\bar{\psi}\psi\phi$$

$$V(H, S) \longrightarrow \frac{1}{2}m_1^2 h_1^2 + \frac{1}{2}m_2^2 h_2^2 + \text{const} + \mathcal{O}([\text{field}]^3)$$

- Majorana fermion  $\psi \equiv \chi + \chi^c$  is a massive dark particle,  $m_\psi^2 = y_x^2 v_s^2$



$$V_{\text{pGDM}}(H, S) = -\mu_H^2 |H|^2 + \lambda_H |H|^4 - \mu_S^2 |S|^2 + \lambda_S |S|^4 + \kappa |H|^2 |S|^2 + \mu^2 (S^2 + S^{*2})$$

$$V_{\text{VDM}}(H, S) = -\mu_H^2 |H|^2 + \lambda_H |H|^4 - \mu_S^2 |S|^2 + \lambda_S |S|^4 + \kappa |H|^2 |S|^2$$

$$V_{\text{FDM}}(H, S) = -\mu_H^2 |H|^2 + \lambda_H |H|^4 - \frac{\mu_S^2}{2} S^2 + \frac{\lambda_S}{4} S^4 + \frac{\kappa}{2} |H|^2 S^2$$

- asymptotic positivity of the scalar potential

- $\lambda_S > 0$ ,  $\lambda_H > 0$
- $\kappa > -2\sqrt{\lambda_H \lambda_S}$  ← always satisfied:

$$\kappa = \frac{(m_1^2 - m_2^2) \sin 2\alpha}{2vv_s} > -\sqrt{\left[ \frac{(m_1^2 - m_2^2) \sin 2\alpha}{2vv_s} \right]^2 + \frac{m_1^2 m_2^2}{v^2 v_s^2}} = -2\sqrt{\lambda_H \lambda_S}$$

- minimum at  $v, v_s > 0$  ( $\mu^2$  only for pGDM):

- $\kappa^2 < 4\lambda_H \lambda_S$
- $\mu^2 < 0$
- $2\lambda_S \mu_H^2 - \kappa(\mu_S^2 - 2\mu^2) > 0$
- $2\lambda_H (\mu_S^2 - 2\mu^2) - \kappa \mu_H^2 > 0$

# Input parameters

$$V_{\text{pGDM}}(H, S) = -\mu_H^2 |H|^2 + \lambda_H |H|^4 - \mu_S^2 |S|^2 + \lambda_S |S|^4 + \kappa |H|^2 |S|^2 + \mu^2 (S^2 + S^{*2})$$

$$V_{\text{VDM}}(H, S) = -\mu_H^2 |H|^2 + \lambda_H |H|^4 - \mu_S^2 |S|^2 + \lambda_S |S|^4 + \kappa |H|^2 |S|^2$$

$$V_{\text{FDM}}(H, S) = -\mu_H^2 |H|^2 + \lambda_H |H|^4 - \frac{\mu_S^2}{2} S^2 + \frac{\lambda_S}{4} S^4 + \frac{\kappa}{2} |H|^2 S^2$$

- Input parameters:  $v, m_1, v_s, m_2, \sin \alpha, m_{\text{DM}}$

assumed to be SM-like

- Other parameters of the models in terms of the input parameters:

$$\mu^2 = -\frac{1}{4} m_{\text{DM}}^2 \quad (\text{pGDM}) \quad g_x = \frac{m_{\text{DM}}}{v_s} \quad (\text{VDM}) \quad y_x = \frac{m_{\text{DM}}}{v_s} \quad (\text{FDM})$$

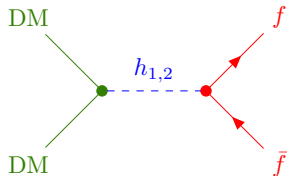
$$\lambda_H = \frac{m_1^2 \cos^2 \alpha + m_2^2 \sin^2 \alpha}{2v^2} \quad \lambda_S = \frac{m_1^2 \sin^2 \alpha + m_2^2 \cos^2 \alpha}{2v_s^2}$$

$$\mu_H^2 = \frac{1}{2} m_1^2 \cos^2 \alpha + \frac{1}{2} m_2^2 \sin^2 \alpha + \frac{1}{4} \frac{v_s}{v} (m_1^2 - m_2^2) \sin 2\alpha$$

$$\mu_S^2 = \frac{1}{2} m_1^2 \sin^2 \alpha + \frac{1}{2} m_2^2 \cos^2 \alpha + \frac{1}{4} \frac{v}{v_s} (m_1^2 - m_2^2) \sin 2\alpha$$

$$\kappa = \frac{(m_1^2 - m_2^2) \sin 2\alpha}{2vv_s}$$

- $m_2 = m_1 \Rightarrow \kappa = 0 \Rightarrow$  no Higgs portal  $\Rightarrow$  DM completely decoupled



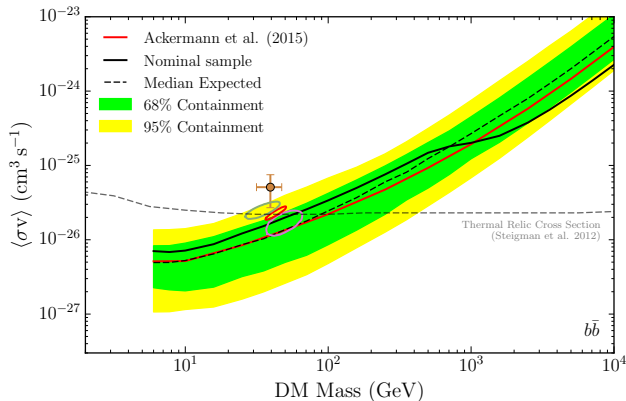
$$\langle\sigma v\rangle_{f\bar{f}}^{\text{ID}} = \frac{m_{\text{DM}} m_f^2 \kappa^2 \cdot (m_{\text{DM}}^2 - m_f^2)^{3/2}}{\pi D_1 D_2} \times \begin{cases} 12 & \text{pGDM} \\ 1 & \text{VDM} \\ \frac{9}{4} \left(\frac{m_{\text{DM}}}{T}\right)^{-1} & \text{FDM} \end{cases} + [\text{higher orders in } (m_{\text{DM}}/T)^{-1}]$$

$D_i = (4m_{\text{DM}}^2 - m_i)^2 + m_i^2 \Gamma_i^2$

$m_f^2$  factor  $\Rightarrow b\bar{b}$  contribution dominates.

- If  $\langle\sigma v\rangle^{\text{ID}} = \sigma_0 (m_{\text{DM}}/T)^{-n}$  then  $h^2 \Omega_0^{\text{DM}} \sim (n+1) (m_{\text{DM}}/T_f)^{n+1} / \sigma_0$
- Correct  $h^2 \Omega_0^{\text{DM}} \longleftrightarrow \langle\sigma v\rangle|_{\text{now}} = (T_0/T_f)^n \cdot \underbrace{(n+1) \cdot 1.9 \cdot 10^{-9} \text{ GeV}^{-2}}_{\langle\sigma v\rangle|_{\text{freeze out}}}$
- $T_f \sim m_{\text{DM}}/25$ , hence

$$\kappa^2 = 3.5 \cdot 10^{-10} \text{ GeV}^{-4} \frac{D_1 D_2}{m_{\text{DM}} (m_{\text{DM}}^2 - m_b^2)^{3/2}} \cdot \begin{cases} 1/12 & \text{pGDM} \\ 1 & \text{VDM} \\ \frac{8m_\psi}{9T_f} \approx 22 & \text{FDM} \end{cases}$$



value of  $\langle\sigma v\rangle|_{\text{freeze out}}$   
corresponding to  
correct  $\Omega_0^{\text{DM}}$

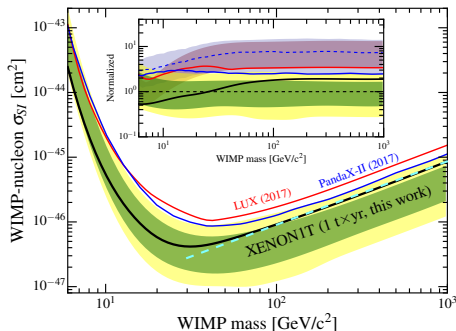
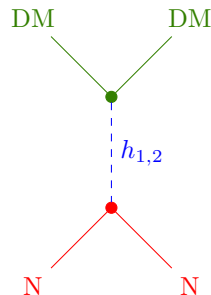
$$\langle\sigma v\rangle|_{\text{now}} = (T_0/T_f)^n \cdot \langle\sigma v\rangle|_{\text{freeze out}}$$

- pGDM, VDM:  $n = 0 \Rightarrow m_{\text{DM}} \gtrsim 30 \text{ GeV}$
- FDM:  $n = 1 \Rightarrow \langle\sigma v\rangle|_{\text{now}}$  orders of magnitude lower than the limit

$\sigma_{SD} = 0$  (no axial couplings in our models)

$$\sigma_{SI} \sim \kappa^2 \cdot \frac{m_{DM}^2}{m_2^4} \begin{cases} \left[ \frac{\mathcal{A}}{64\pi^2 v v_s^2} \right]^2 & (\text{pGDM}) \\ 1 & (\text{VDM}), (\text{FDM}) \end{cases}$$

$\mathcal{A}$  – combination of 1-loop integrals  
(the tree level **vanishes** for pGDM)



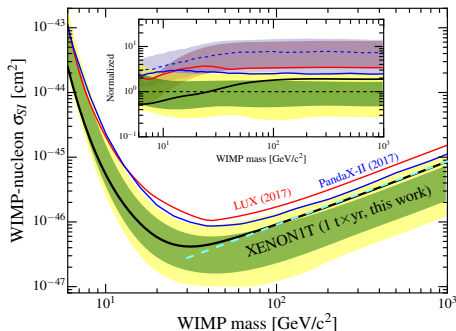
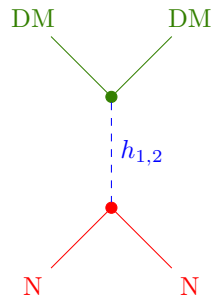
$$\sigma_{SI} \lesssim \frac{1 \text{ cm}^2}{1 \text{ GeV}} \cdot 10^{-48.05}$$

$$\kappa^2 < \frac{m_2^4}{m_{DM}} \cdot 2.5 \cdot 10^{-11} \text{ GeV}^{-3} \times \begin{cases} \left[ \frac{\mathcal{A}}{64\pi^2 v v_s^2} \right]^{-2} & (\text{pGDM}) \\ 1 & (\text{VDM}), (\text{FDM}) \end{cases}$$

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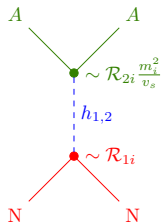
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huge!

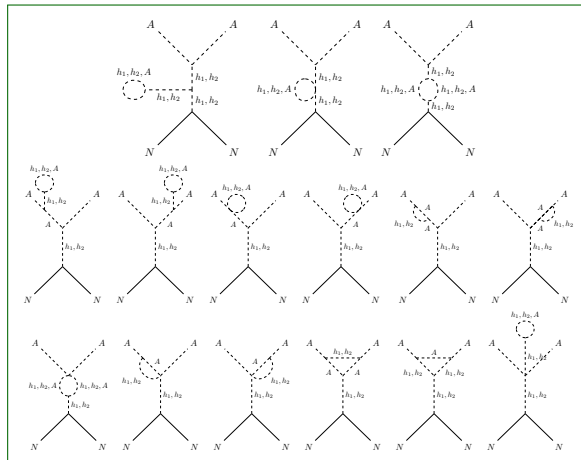


the tree level:

$$\sigma \sim \frac{(\cos \alpha \sin \alpha)^2}{v_s^2} \left| \frac{m_1^2}{Q^2 - m_1^2} - \frac{m_2^2}{Q^2 - m_2^2} \right|^2 \xrightarrow{Q^2 \rightarrow 0} 0$$

the 1-loop level:

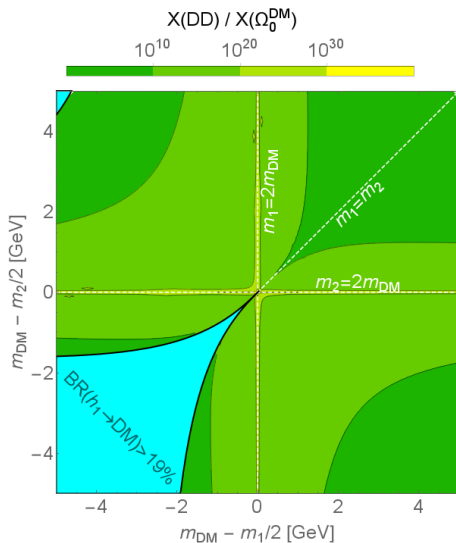
$$\sigma \sim \mathcal{A}^2$$



# DD vs $\Omega_0^{\text{DM}}$ limit for pGDM

$$\kappa^2(\text{DD}) \equiv \frac{m_2^4}{m_{\text{DM}}} \cdot 2.5 \cdot 10^{-11} \text{ GeV}^{-3} \begin{cases} \left[ \frac{\mathcal{A}}{64\pi^2 v v_s^2} \right]^{-2} & (\text{pGDM}) \\ 1 & (\text{VDM}), (\text{FDM}) \end{cases}$$

↑ huge!



- DM annihilation rate in the Sun:  $\Gamma_A$
- Capture rate:  $C_C$
- Change of DM's amount  $N$  in the Sun

$$\dot{N} = C_C - C_A N^2$$

where  $C_A \equiv 2 \cdot \Gamma_A / N^2 = \text{const}$

- Solution

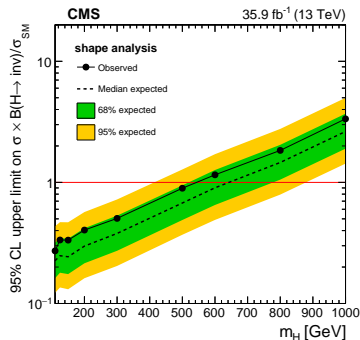
$$N(t) = \sqrt{C_C / C_A} \cdot \tanh(t \sqrt{C_C C_A})$$

- For large  $t$

$$N(t) \rightarrow \sqrt{C_C / C_A}$$

$$\Rightarrow \boxed{\Gamma_A = \frac{1}{2} C_C}$$

- $C_C$  expressible in terms of DM-nucleon cross section
- $\Gamma_A$  calculable from muon flux measurements



$$\sigma \approx \sigma_1 \cdot \mathbb{1}_{2 m_{DM} < m_1 < \sqrt{s} - m_Z} + \sigma_2 \cdot \mathbb{1}_{2 m_{DM} < m_2 < \sqrt{s} - m_Z}$$

$$\sigma_1 \equiv \sigma_{SM}(m_1) \cdot \cos^2 \alpha \cdot \text{BR}(h_1 \rightarrow \text{DM})$$

$$\sigma_2 \equiv \sigma_{SM}(m_2) \cdot \sin^2 \alpha \cdot \text{BR}(h_2 \rightarrow \text{DM})$$

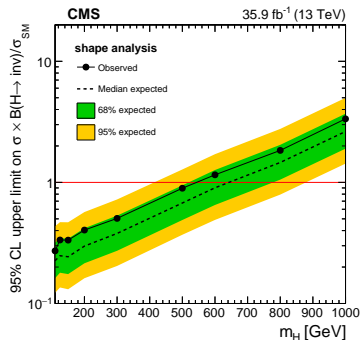
Conditions:

$$(1) \quad \sigma_1 < 0.19 \sigma_{SM}(m_1)$$

$$(2) \quad \log \left[ \frac{\sigma_2}{\sigma_{SM}(m_2)} \right] < 0.0011 \cdot \frac{m_2}{1 \text{ GeV}} - 0.63 \quad (\sqrt{s} = 13 \text{ TeV})$$

For  $\sin \alpha < 0.3$  condition (2) always satisfied:

$$\left. \frac{\sigma_2}{\sigma_{SM}(m_2)} \right|_{\sqrt{s}=13 \text{ TeV}} = \sin^2 \alpha \cdot \text{BR}(h_2 \rightarrow \text{DM}) < 10^{-1} < 10^{0.0011 \cdot \frac{m_2}{1 \text{ GeV}} - 0.63}$$



$$\sigma \approx \sigma_1 \cdot \mathbb{1}_{2 m_{DM} < m_1 < \sqrt{s} - m_Z} + \sigma_2 \cdot \mathbb{1}_{2 m_{DM} < m_2 < \sqrt{s} - m_Z}$$

$$\sigma_1 \equiv \sigma_{SM}(m_1) \cdot \cos^2 \alpha \cdot \text{BR}(h_1 \rightarrow \text{DM})$$

$$\sigma_2 \equiv \sigma_{SM}(m_2) \cdot \sin^2 \alpha \cdot \text{BR}(h_2 \rightarrow \text{DM})$$

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# Considered process: degenerated case

$$\frac{d\sigma}{dQ^2} = \frac{\sigma_{\text{SM}}(Q^2)v^2}{32\pi^2} \frac{Q^4 \cdot \mathcal{X}}{[(Q^2 - m_1^2)^2 + (m_1\Gamma_1)^2][(Q^2 - m_2^2)^2 + (m_2\Gamma_2)^2]} \times$$

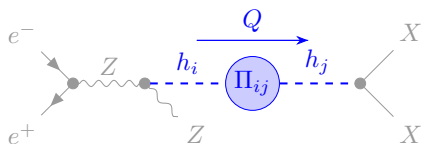
$$\times \sqrt{1 - 4\frac{m_{\text{DM}}^2}{Q^2}} \cdot \begin{cases} 1 & \text{(pGDM)} \\ 1 - 4\frac{m_{\text{DM}}^2}{Q^2} + 12\left(\frac{m_{\text{DM}}^2}{Q^2}\right)^2 & \text{(VDM)} \\ 2\frac{m_{\text{DM}}^2}{Q^2}\left(1 - 4\frac{m_{\text{DM}}^2}{Q^2}\right) & \text{(FDM)} \end{cases}$$

$$\left. \begin{aligned} \Gamma_1 &\xrightarrow{m_2 \rightarrow m_1} \cos^2 \alpha \Gamma_{h \rightarrow \text{SM}} + \sin^2 \alpha \Gamma_{h \rightarrow \text{DM}} \\ \Gamma_2 &\xrightarrow{m_2 \rightarrow m_1} \sin^2 \alpha \Gamma_{h \rightarrow \text{SM}} + \cos^2 \alpha \Gamma_{h \rightarrow \text{DM}} \end{aligned} \right\}$$

$$\mathcal{X} \equiv (\cos \alpha \sin \alpha)^2 \frac{(m_1^2 - m_2^2)^2 + (m_1\Gamma_1 - m_2\Gamma_2)^2}{v^2 v_s^2} \Big|_{m_1=m_2} \neq 0$$

$$\text{but } \kappa = \frac{(m_1^2 - m_2^2) \sin 2\alpha}{2vv_s} \xrightarrow{m_2 \rightarrow m_1} 0 !$$

# Considered process: degenerated case. The propagator



$$\Pi_{ij} \equiv \Pi_{ij}(Q^2) \equiv i \cdot \text{Im} [\text{self energy}]$$

$$\Pi_{ii}(m_i^2) = im_i \Gamma_i$$

the matrix element:

$$\begin{aligned} \mathcal{M} &= \mathcal{M}_{e^+e^- \rightarrow Zh_i}(Q^2) \cdot \Delta_{ij}(Q^2) \cdot \mathcal{M}_{h_j \rightarrow XX}(Q^2) = \\ &= \mathcal{M}_{e^+e^- \rightarrow Zh}(Q^2) \cdot \underbrace{\mathcal{R}_{1i} \cdot \Delta_{ij}(Q^2) \cdot \mathcal{R}_{2j}}_{\hat{\Delta}(Q^2)} \cdot \mathcal{M}_{h \rightarrow XX}(Q^2), \end{aligned}$$

the propagator:

$$\begin{aligned} \hat{\Delta}(Q^2) &= \mathcal{R}_{1i} \mathcal{R}_{2j} \cdot \frac{1}{\det D} \overbrace{\begin{bmatrix} Q^2 - m_2^2 + \Pi_{22} & -\Pi_{12} \\ -\Pi_{21} & s - m_1^2 + \Pi_{11} \end{bmatrix}}^D_{ij} = \\ &= \sin \alpha \cos \alpha \cdot \frac{(m_1^2 - m_2^2) - [(\Pi_{11} - \Pi_{22}) - (\text{tg } \alpha \cdot \Pi_{12} - \text{ctg } \alpha \cdot \Pi_{21})]}{(Q^2 - m_1^2 + \Pi_{11})(Q^2 - m_2^2 + \Pi_{22}) - \Pi_{12}\Pi_{21}} \end{aligned}$$

$$\hat{\Delta}(Q^2) = \sin \alpha \cos \alpha \cdot \frac{(m_1^2 - m_2^2) - [(\Pi_{11} - \Pi_{22}) - (\text{tg } \alpha \cdot \Pi_{12} - \text{ctg } \alpha \cdot \Pi_{21})]}{(Q^2 - m_1^2 + \Pi_{11})(Q^2 - m_2^2 + \Pi_{22}) - \Pi_{12}\Pi_{21}}$$

- If  $|m_1 - m_2| \gg \Gamma_1, \Gamma_2$ , then ← in practice, only this case relevant

$$\begin{aligned} \hat{\Delta}(Q^2) &\approx \sin \alpha \cos \alpha \cdot \frac{m_1^2 - m_2^2}{(Q^2 - m_1^2 + \Pi_{11})(Q^2 - m_2^2 + \Pi_{22})} \\ &\approx \sin \alpha \cos \alpha \cdot \left[ \frac{1}{Q^2 - m_1^2 + im_1\Gamma_1} - \frac{1}{Q^2 - m_2^2 + im_2\Gamma_2} \right] \\ &\equiv \hat{\Delta}^{(\text{BW})}(Q^2) \end{aligned}$$

- If  $m_1 \sim m_2$ , an explicit calculation shows that

$$\begin{aligned} &[(\Pi_{11} - \Pi_{22}) - (\text{tg } \alpha \cdot \Pi_{12} - \text{ctg } \alpha \cdot \Pi_{21})] \Big|_{m_1=m_2} = 0 \\ &\Rightarrow \hat{\Delta}(Q^2) \Big|_{m_1=m_2} = 0 \end{aligned}$$

$$\Pi_{ij} = \Pi_{ij}^{\text{DM}} + \Pi_{ij}^{W^+W^-} + \Pi_{ij}^{ZZ} + \sum_q \Pi_{ij}^{q\bar{q}} + \sum_l \Pi_{ij}^{l^+l^-} + \sum_{k,l} \Pi_{ij}^{h_k h_l}$$

(DM = A, X,  $\psi$ ; q – SM quarks; l – SM leptons)

by straightforward calculations:

$$\left[ (\Pi_{11}^{ab} - \Pi_{22}^{ab}) - (\text{tg } \alpha \cdot \Pi_{12}^{ab} - \text{ctg } \alpha \cdot \Pi_{21}^{ab}) \right] \Big|_{m_1=m_2} = 0$$

for  $ab = \text{DM}, W^+W^-, ZZ, q\bar{q}, l^+l^-, h_k h_l$  (sum over  $k, l = 1, 2$ )

$$\Rightarrow \left[ (\Pi_{11} - \Pi_{22}) - (\text{tg } \alpha \cdot \Pi_{12} - \text{ctg } \alpha \cdot \Pi_{21}) \right] \Big|_{m_1=m_2} = 0.$$

$$\bullet \Pi_{ij}^{\text{DM}}(Q^2) = I(Q^2, m_{\text{DM}}, m_{\text{DM}}) \frac{\mathcal{R}_{2i}\mathcal{R}_{2j}}{32\pi^2 v_s^2} (m_i m_j)^2 \begin{cases} 1 & (\text{pGDM}) \\ 1 - 2m_{\text{DM}}^2 \frac{4Q^2 - m_i^2 - m_j^2}{(m_i m_j)^2} + 12 \left( \frac{m_{\text{DM}}^2}{(m_i m_j)^2} \right)^2 & (\text{VDM}) \\ 2 \frac{m_{\text{DM}}^2 Q^2}{(m_i m_j)^2} \left( 1 - 4 \frac{m_{\text{DM}}^2}{Q^2} \right) & (\text{FDM}) \end{cases}$$

$$\bullet \Pi_{ij}^{W^+W^-}(Q^2) = I(Q^2, m_W, m_W) \frac{\mathcal{R}_{1i}\mathcal{R}_{1j}}{16\pi^2 v^2} (m_i m_j)^2 \times \left[ 1 - 2m_W^2 \frac{4Q^2 - m_i^2 - m_j^2}{(m_i m_j)^2} + 12 \frac{m_W^4}{(m_i m_j)^2} \right]$$

$$\bullet \Pi_{ij}^{ZZ}(Q^2) = I(Q^2, m_Z, m_Z) \frac{\mathcal{R}_{1i}\mathcal{R}_{1j}}{32\pi^2 v^2} (m_i m_j)^2 \times \left[ 1 - 2m_Z^2 \frac{4Q^2 - m_i^2 - m_j^2}{(m_i m_j)^2} + 12 \frac{m_Z^4}{(m_i m_j)^2} \right]$$

$$\bullet \Pi_{ij}^{q\bar{q}}(Q^2) = I(Q^2, m_q, m_q) \cdot \frac{3\mathcal{R}_{1i}\mathcal{R}_{1j}}{8\pi^2 v^2} m_q^2 Q^2 \left( 1 - 4 \frac{m_q^2}{Q^2} \right)$$

$$\bullet \Pi_{ij}^{h_k h_l}(Q^2) = I(Q^2, m_k, m_l) \cdot \frac{V_{ikl} V_{jkl}}{32\pi^2}$$

$$\bullet \Pi_{ij}^{l^+ l^-}(Q^2) = I(Q^2, m_l, m_l) \cdot \frac{\mathcal{R}_{1i}\mathcal{R}_{1j}}{8\pi^2 v^2} m_l^2 Q^2 \left( 1 - 4 \frac{m_l^2}{Q^2} \right)$$

$$* V_{111} \equiv 3m_1^2 \left( \frac{\sin^3 \alpha}{v_s} + \frac{\cos^3 \alpha}{v} \right)$$

$$* V_{112} = V_{121} = V_{211} \equiv (2m_1^2 + m_2^2) \sin \alpha \cos \alpha \left( \frac{\sin \alpha}{v_s} - \frac{\cos \alpha}{v} \right)$$

$$* V_{222} \equiv 3m_2^2 \left( \frac{\cos^3 \alpha}{v_s} - \frac{\sin^3 \alpha}{v} \right)$$

$$* V_{221} = V_{212} = V_{122} \equiv (m_1^2 + 2m_2^2) \sin \alpha \cos \alpha \left( \frac{\cos \alpha}{v_s} + \frac{\sin \alpha}{v} \right)$$

$$\bullet I(Q^2, m_a, m_b) \equiv i \cdot \mathcal{I}m \left[ \frac{1}{i\pi^2} \int \frac{d^4 l}{(l^2 - m_a^2)[(l+Q)^2 - m_b^2]} \right] = i\pi \cdot \frac{\lambda^{1/2}(Q^2, m_a^2, m_b^2)}{Q^2} \cdot \mathbb{1}_{Q^2 > (m_a + m_b)^2}$$

$$\Rightarrow I(Q^2, m, m) = i\pi \cdot \sqrt{1 - \frac{4m^2}{Q^2}} \cdot \mathbb{1}_{Q^2 > 4m^2} \quad (\lambda(a, b, c) \equiv a^2 + b^2 + c^2 - 2(ab + bc + ca))$$

# The $h_1$ 's and $h_2$ 's decay widths

$$\Gamma_{h_i \rightarrow ab} \equiv \frac{\Pi_{ii}^{ab}(m_i^2)}{im_i}$$

⇓

- $$\Gamma_{h_i \rightarrow \text{DM}} = \frac{\mathcal{R}_{2i}^2}{v_s^2} \frac{m_i^3}{32\pi} \sqrt{1 - \frac{4m_{\text{DM}}^2}{m_i^2}} \begin{cases} 1 & \text{(pGDM)} \\ 1 - 4\frac{m_{\text{DM}}^2}{m_i^2} + 12\left(\frac{m_{\text{DM}}^2}{m_i^2}\right)^2 & \text{(VDM)} \\ 2\frac{m_{\text{DM}}^2}{m_i^2} \left(1 - 4\frac{m_{\text{DM}}^2}{m_i^2}\right) & \text{(FDM)} \end{cases}$$
- $$\Gamma_{h_i \rightarrow \text{SM}} = \mathcal{R}_{1i}^2 \cdot \gamma(m_i) \quad (\gamma - \text{decay width of SM Higgs particle of given mass})$$
- $$\Gamma_{h_1 \rightarrow h_2 h_2} = \sin^2 \alpha \cos^2 \alpha (m_1^2 + 2m_2^2)^2 \left( \frac{\cos \alpha}{v_s} + \frac{\sin \alpha}{v} \right)^2 \frac{\sqrt{m_1^2 - 4m_2^2}}{32\pi m_1^2} \approx$$

$$\approx \frac{\sin^2 \alpha \cos^4 \alpha}{v_s^2} (m_1^2 + 2m_2^2)^2 \frac{\sqrt{m_1^2 - 4m_2^2}}{32\pi m_1^2}$$
- $$\Gamma_{h_2 \rightarrow h_1 h_1} = \sin^2 \alpha \cos^2 \alpha (2m_1^2 + m_2^2)^2 \left( \frac{\sin \alpha}{v_s} - \frac{\cos \alpha}{v} \right)^2 \frac{\sqrt{m_2^2 - 4m_1^2}}{32\pi m_2^2} \approx$$

$$\approx \frac{\sin^2 \alpha \cos^4 \alpha}{v^2} (2m_1^2 + m_2^2)^2 \frac{\sqrt{m_2^2 - 4m_1^2}}{32\pi m_2^2}$$