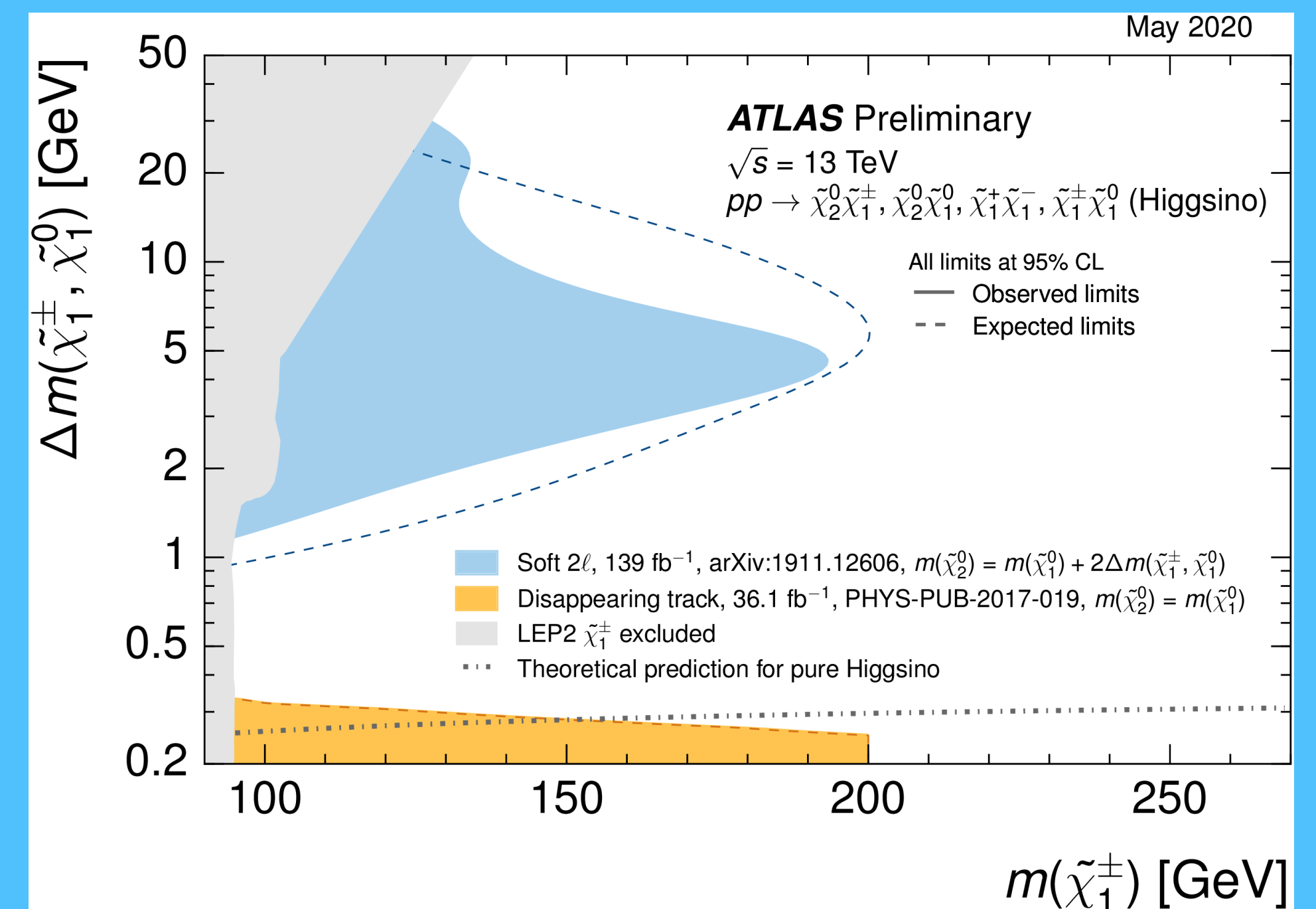
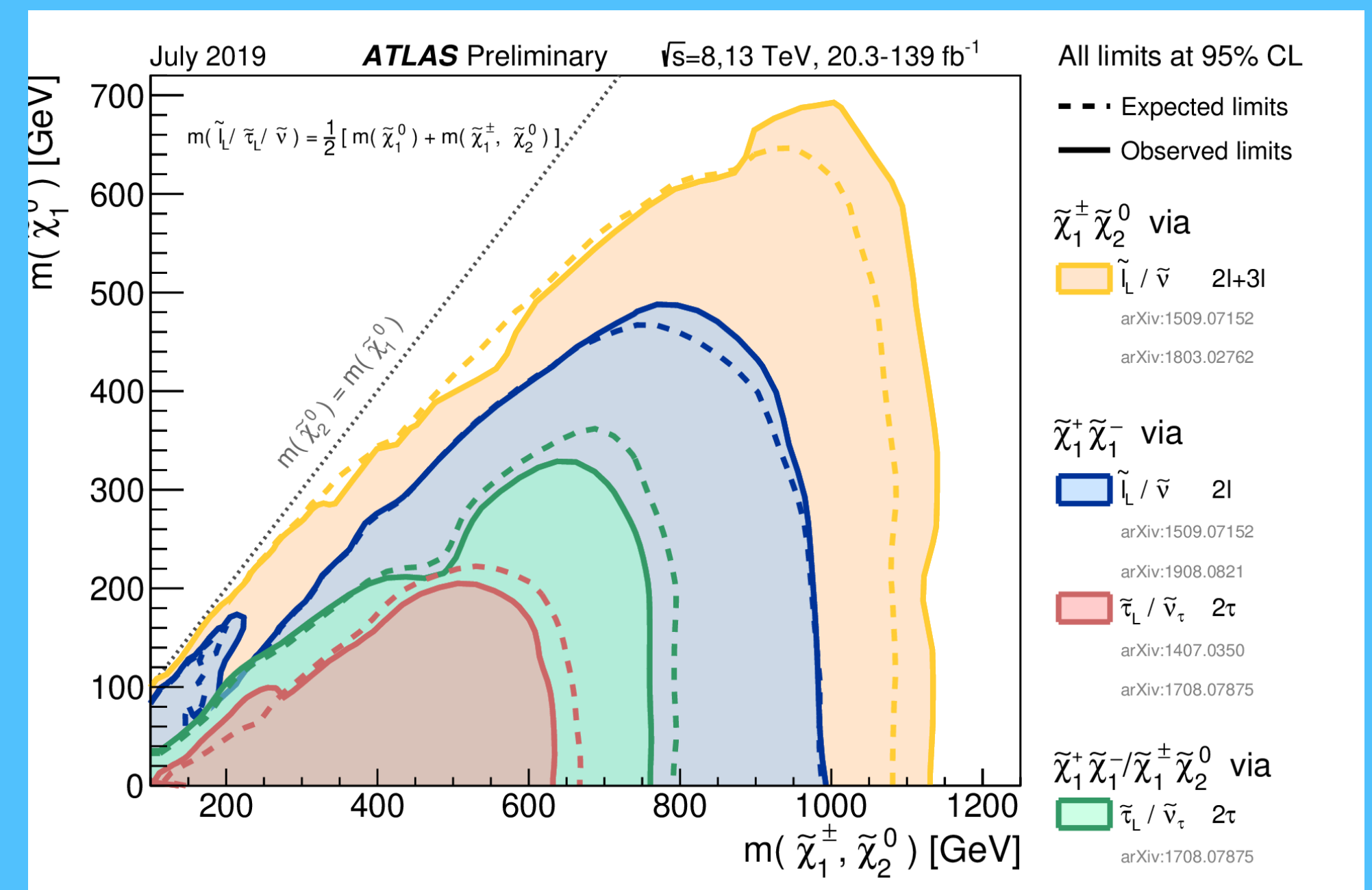


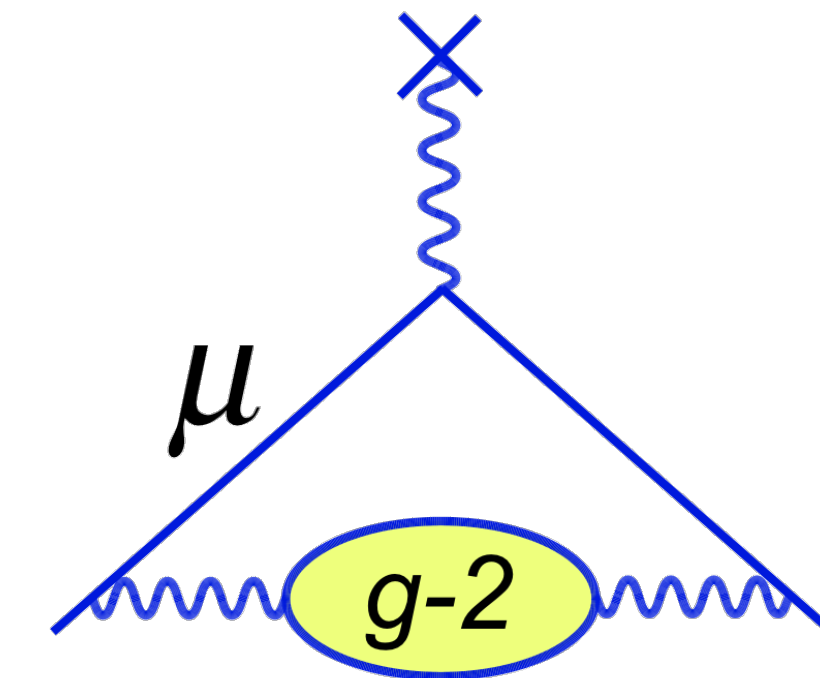
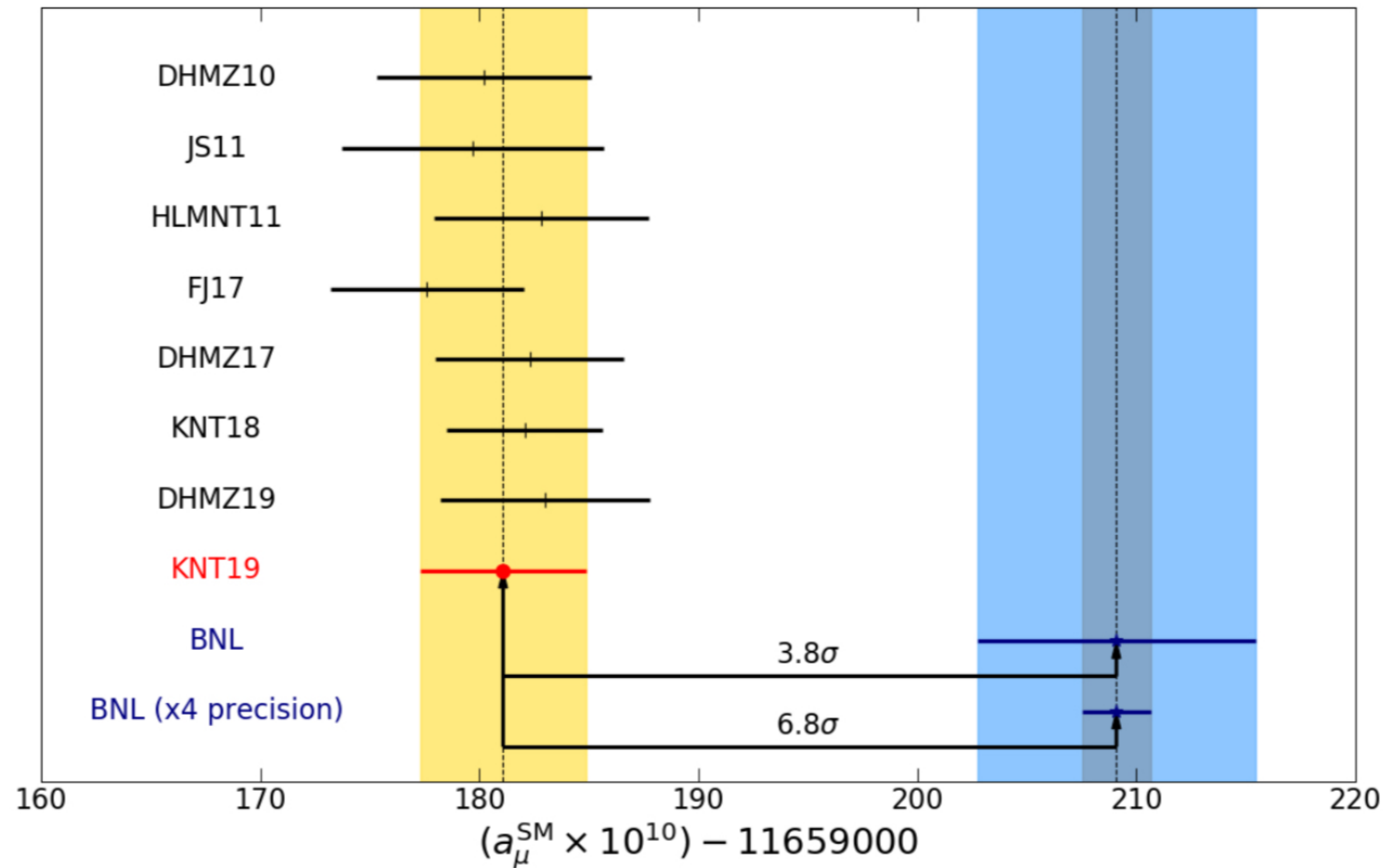
Manimala Chakraborti
AstroCeNT, Warsaw

Improved $(g - 2)_\mu$ measurements and SUSY at existing and future colliders

- EW sector may be hiding key to new physics
- Modest production cross section, mass bounds from the LHC comparably weak
- May show up elsewhere : DM experiments, $(g - 2)_\mu$...
- 3.7σ discrepancy in $(g - 2)_\mu$
- New results from Fermilab 'MUON $(g-2)$ ' coming soon !



Muon ($g-2$)



$$a_\mu^{\text{exp}} - a_\mu^{\text{theo,SM}} = (27.9 \pm 7.6) \times 10^{-10}$$

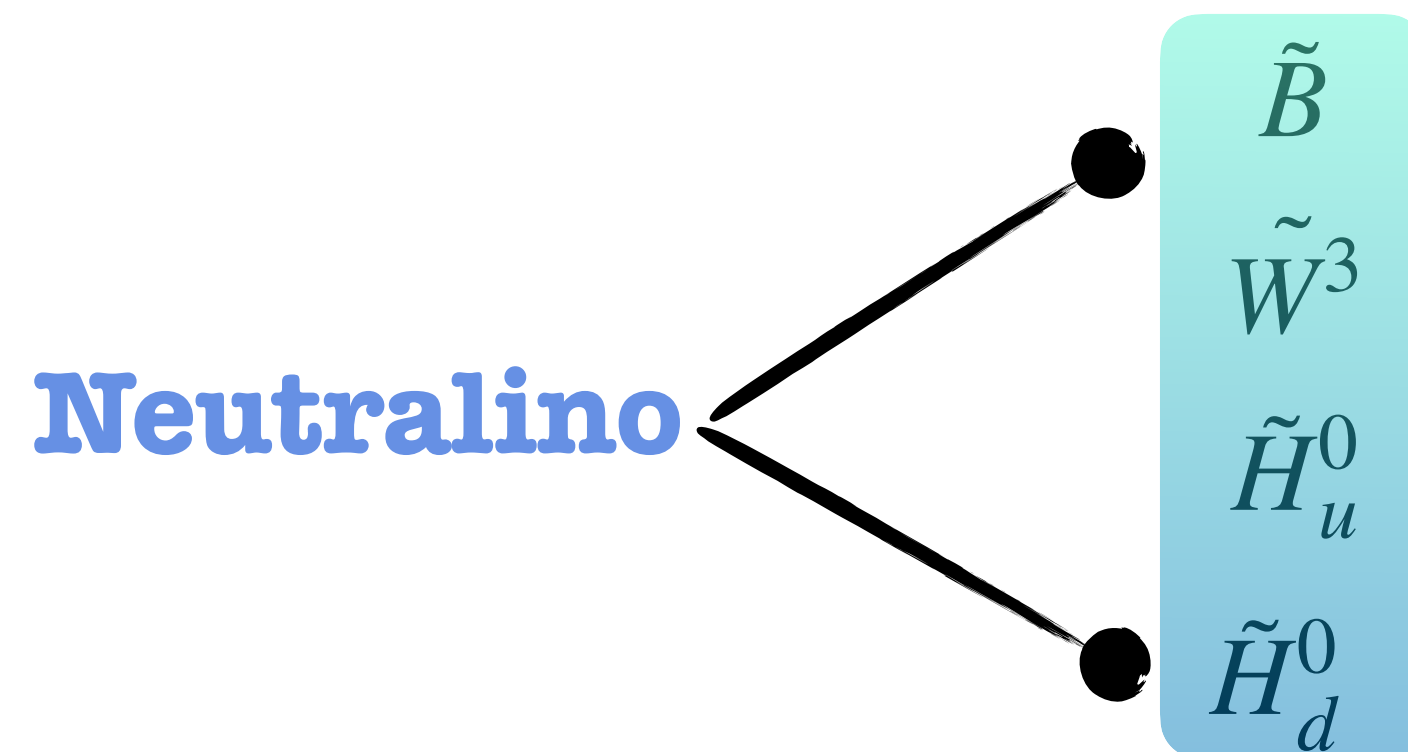
• Aoyama *et al* '20

$\sim 3.7 \sigma$

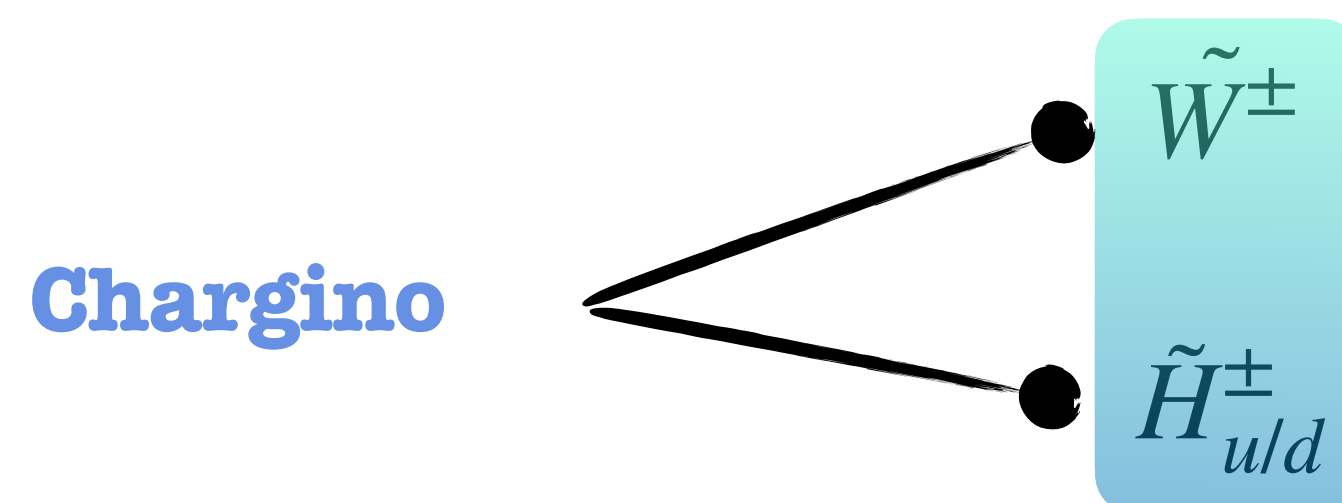
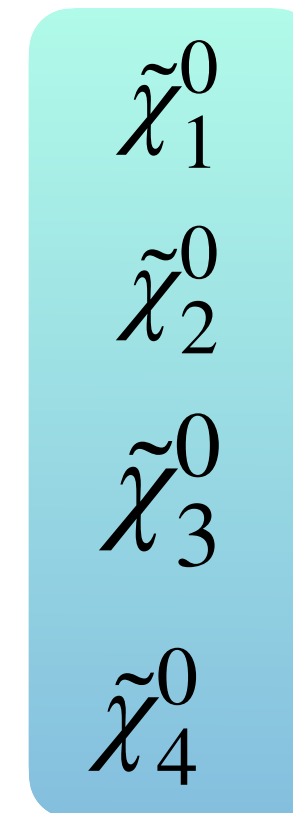
Muon $g-2$ experiment at Fermilab aims at 4 x BNL precision

EW Gauginos

Masses and mixing are determined by U(1) and SU(2) gaugino masses M_1, M_2 and Higgs mass parameter μ .



Mass
Diag. \rightarrow

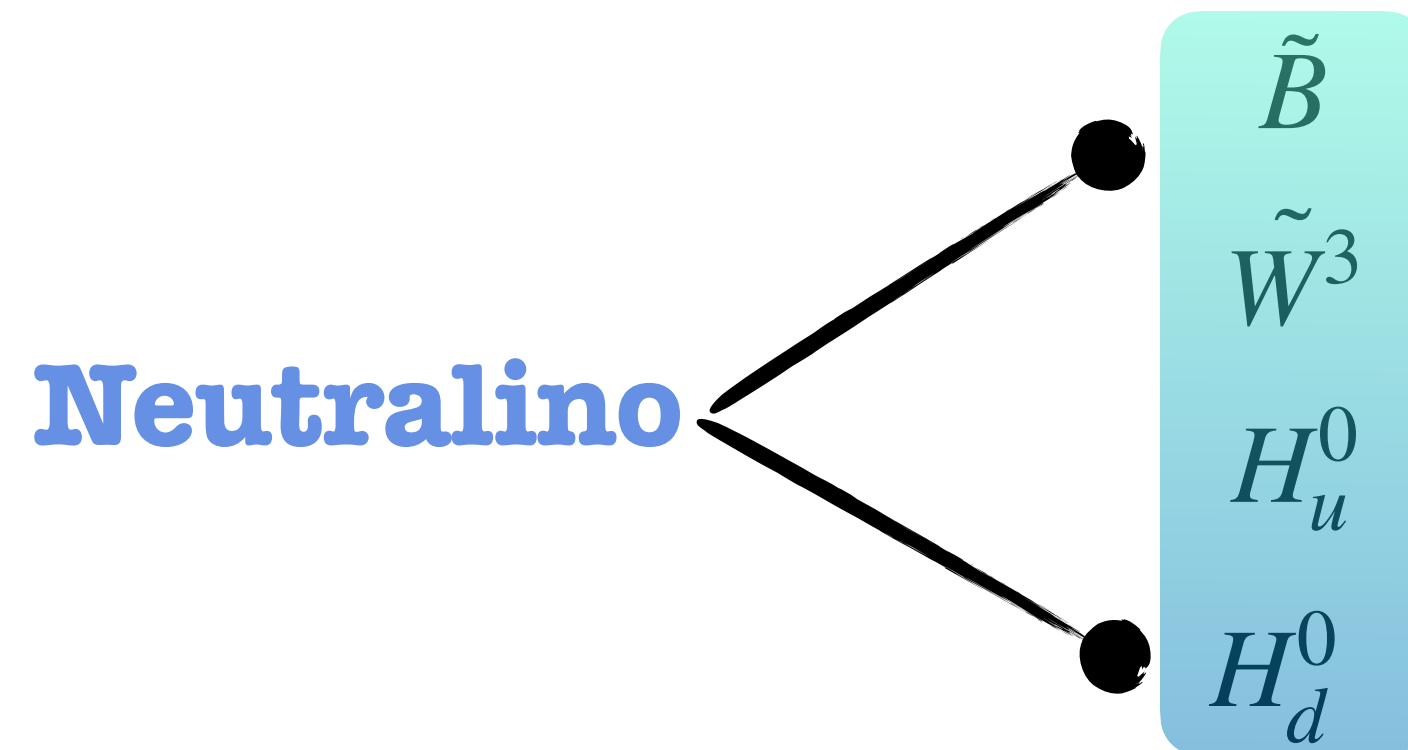


Mass
Diag. \rightarrow



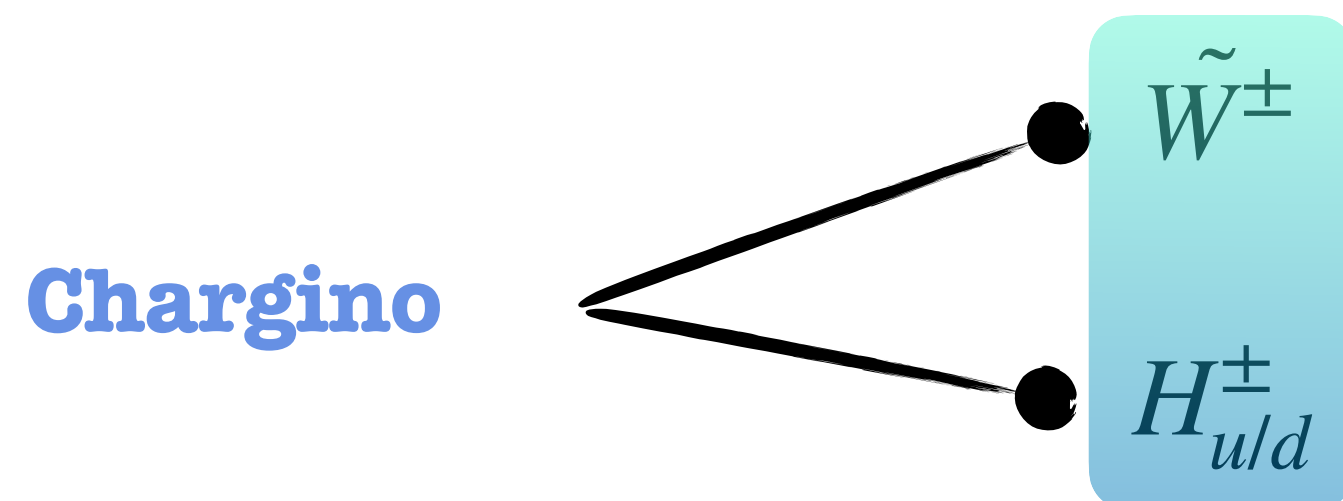
EW Gauginos

Masses and mixing are determined by U(1) and SU(2) gaugino masses M_1, M_2 and Higgs mass parameter μ .



Neutralino Mass Matrix

$$M_N = \begin{pmatrix} M_1 & 0 & -M_Z c_\beta s_W & M_Z s_\beta s_W \\ 0 & M_2 & M_Z c_\beta c_W & -M_Z s_\beta c_W \\ -M_Z c_\beta s_W & M_Z c_\beta c_W & 0 & -\mu \\ M_Z s_\beta s_W & -M_Z s_\beta c_W & -\mu & 0 \end{pmatrix}$$



Chargino Mass Matrix

$$M_C = \begin{pmatrix} M_2 & \sqrt{2}M_W c_\beta \\ \sqrt{2}M_W s_\beta & \mu \end{pmatrix}$$

FOUR PARAMETERS



$M_1, M_2, \mu, \tan \beta$

Sleptons

Slepton Mass Matrix

$$M_{\tilde{L}}^2 = \begin{pmatrix} m_l^2 + m_{LL}^2 & m_l X_l \\ m_l X_l & m_l^2 + m_{RR}^2 \end{pmatrix}$$

$$m_{LL}^2 = m_{\tilde{L}}^2 + (I_l^{3L} - Q_l s_w^2) M_z^2 c_{2\beta}$$

$$m_{RR}^2 = m_{\tilde{R}}^2 + Q_l s_w^2 M_z^2 c_{2\beta}$$

$$X_l = A_l - \mu (\tan \beta)^{2I_l^{3L}}$$

PARAMETERS



$M_1, M_2, \mu, \tan \beta, m_{\tilde{L}}, m_{\tilde{R}}$

First two gens. $m_{\tilde{l}_1} \sim m_{LL}$ $m_{\tilde{l}_2} \sim m_{RR}$

Constraints

Proper recasting is important



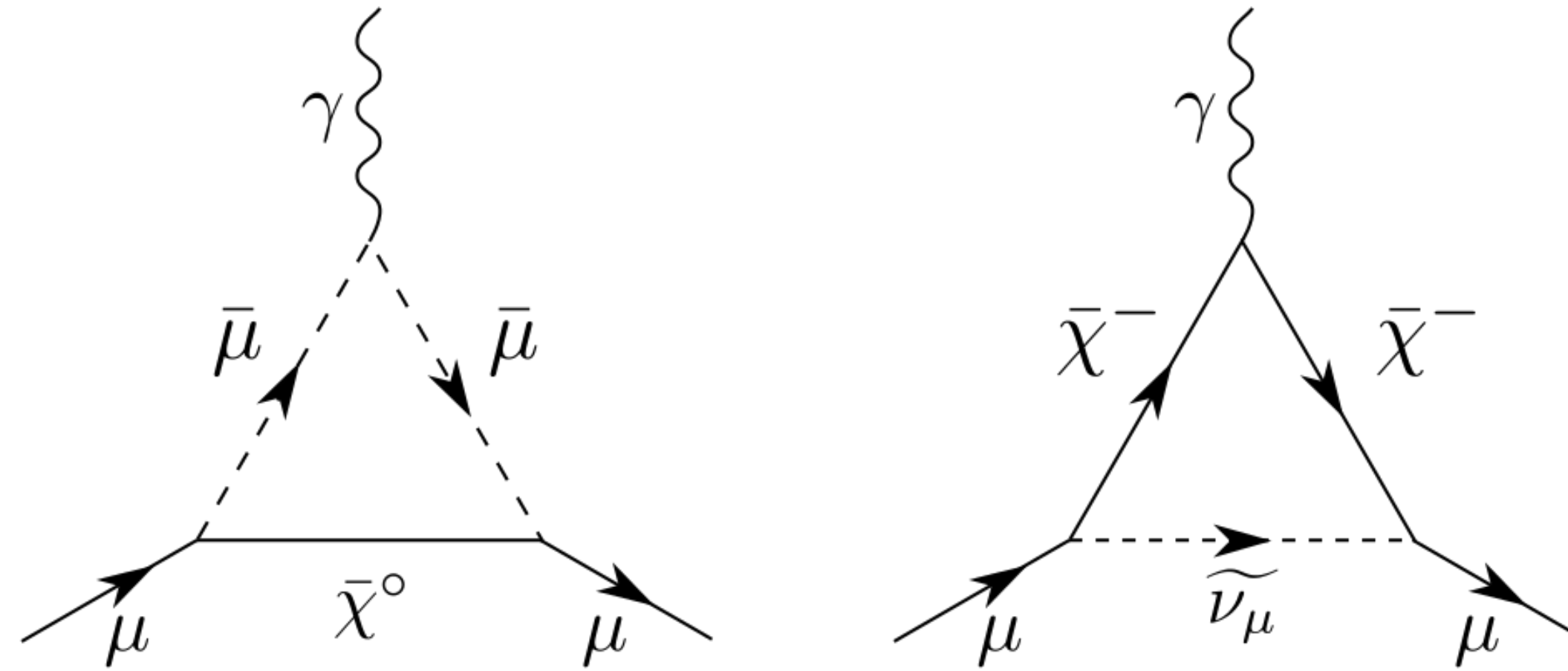
Direct Searches at LHC

- LHC searches restricted to **simplified models**.
- $\tilde{\chi}_1^+$ and $\tilde{\chi}_2^0$ taken to be mass-degenerate and purely wino. $\tilde{\chi}_1^0$ purely bino.
- All three generations of sleptons and sneutrinos assumed mass degenerate.
- Heavier gauginos $\tilde{\chi}_3^0, \tilde{\chi}_4^0, \tilde{\chi}_2^+$ assumed to be decoupled.
- No sensitivity to parameters like $\tan \beta$.

Indirect Constraints

- Muon (g-2).
- WMAP/PLANCK relic density.
- Spin independent direct detection data from XENON/LUX.
- Indirect detection constraints of dark matter.

Muon (g-2) in MSSM



- SUSY contributions from Chargino-Sneutrino and Smuon-Neutralino loop

- SM EW 1 loop : $\frac{\alpha}{\pi} \frac{m_\mu^2}{M_W^2}$. MSSM , 1 loop : $\frac{\alpha}{\pi} \frac{m_\mu^2}{M_{SUSY}^2} \times \tan\beta$

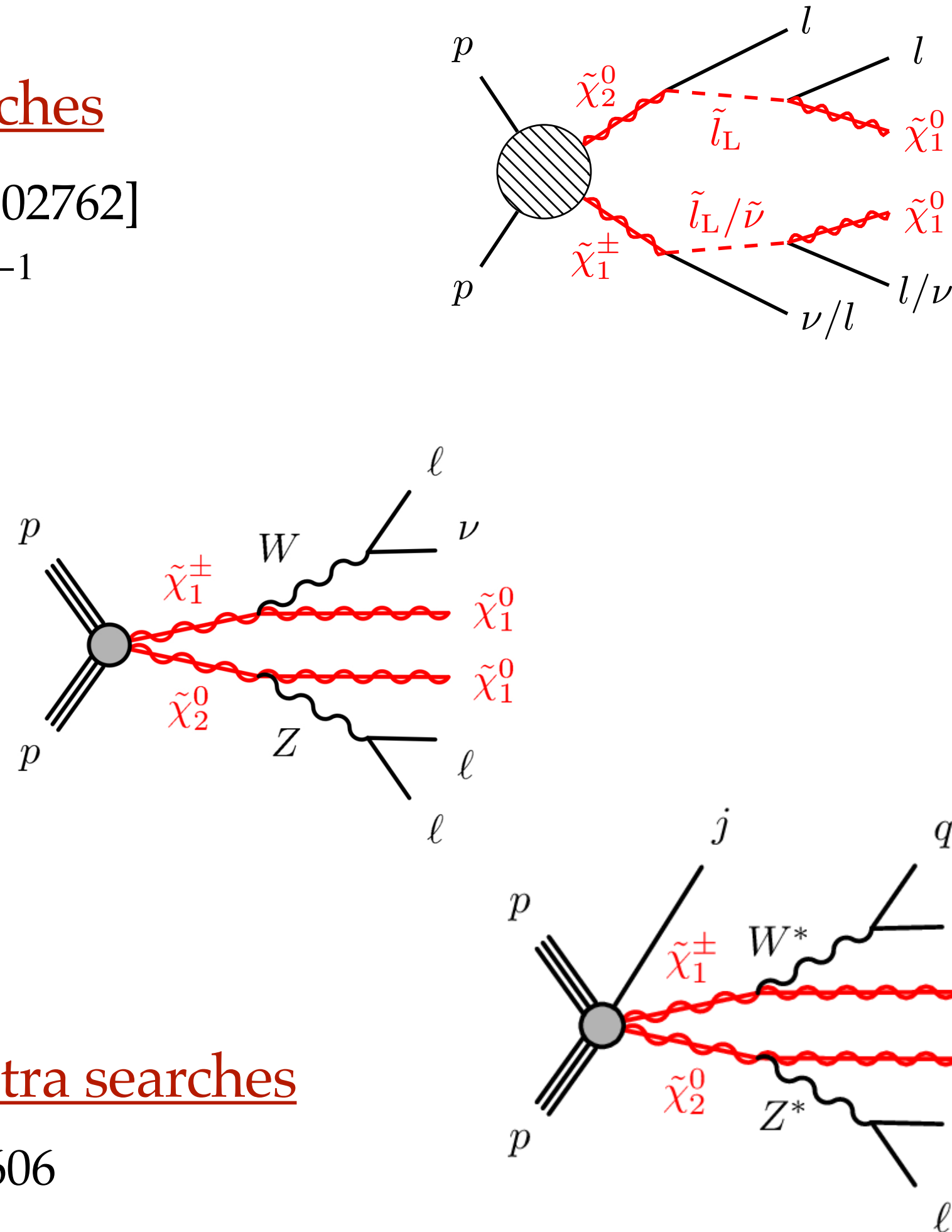
- SUSY can easily explain anomaly !

upper limits on EW super partner masses

Searches at the LHC

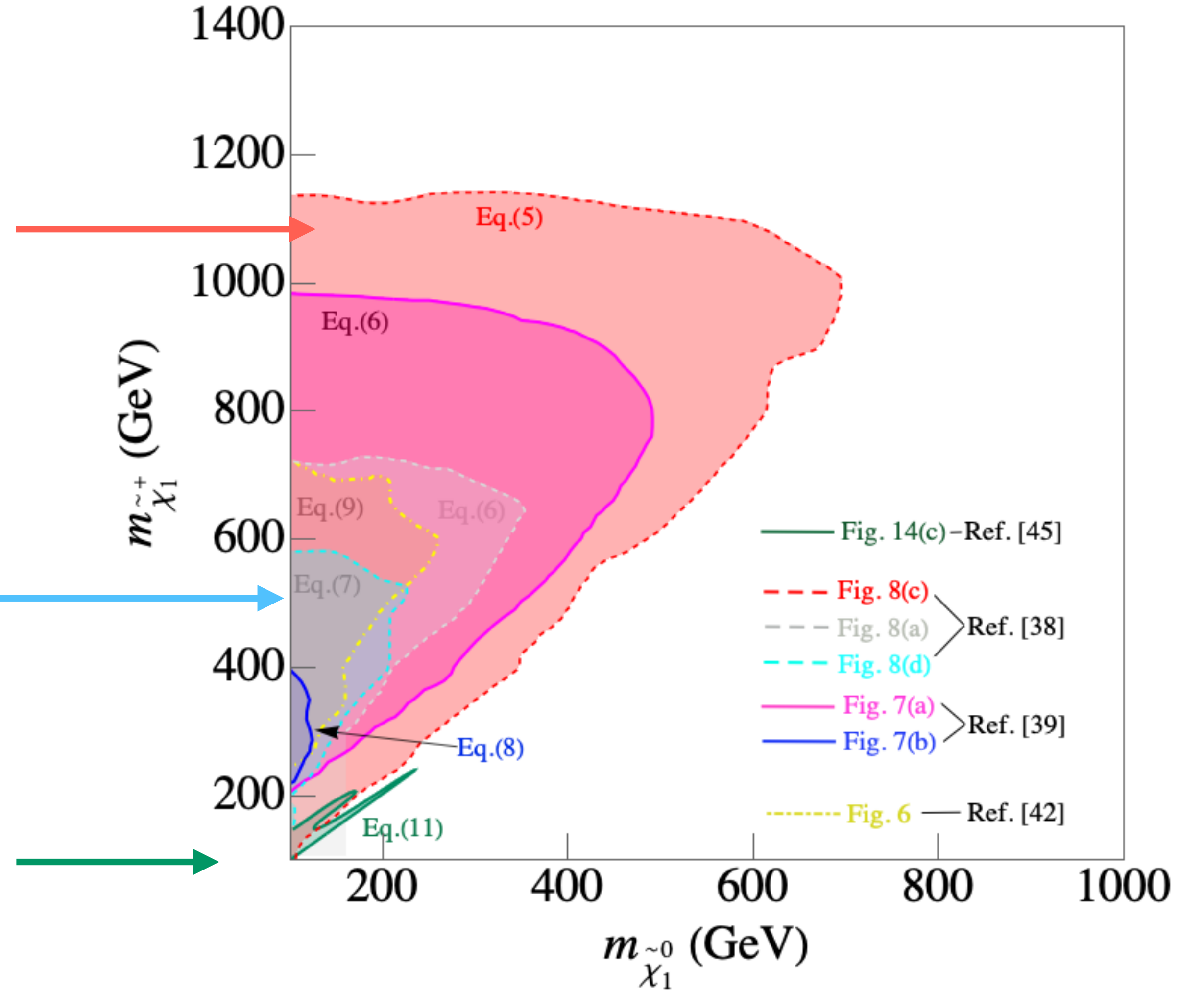
- Trilepton searches

ATLAS [1803.02762]
13 TeV, 36 fb⁻¹



- Compressed spectra searches

ATLAS 1911.12606



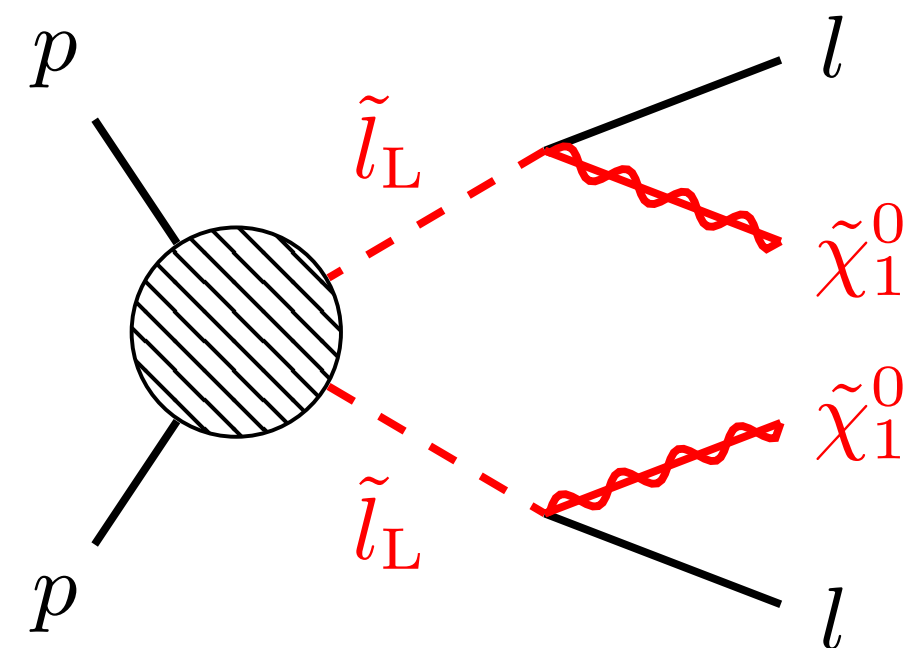
Proper recasting is important → checkMATE

Searches at the LHC

- Slepton pair production

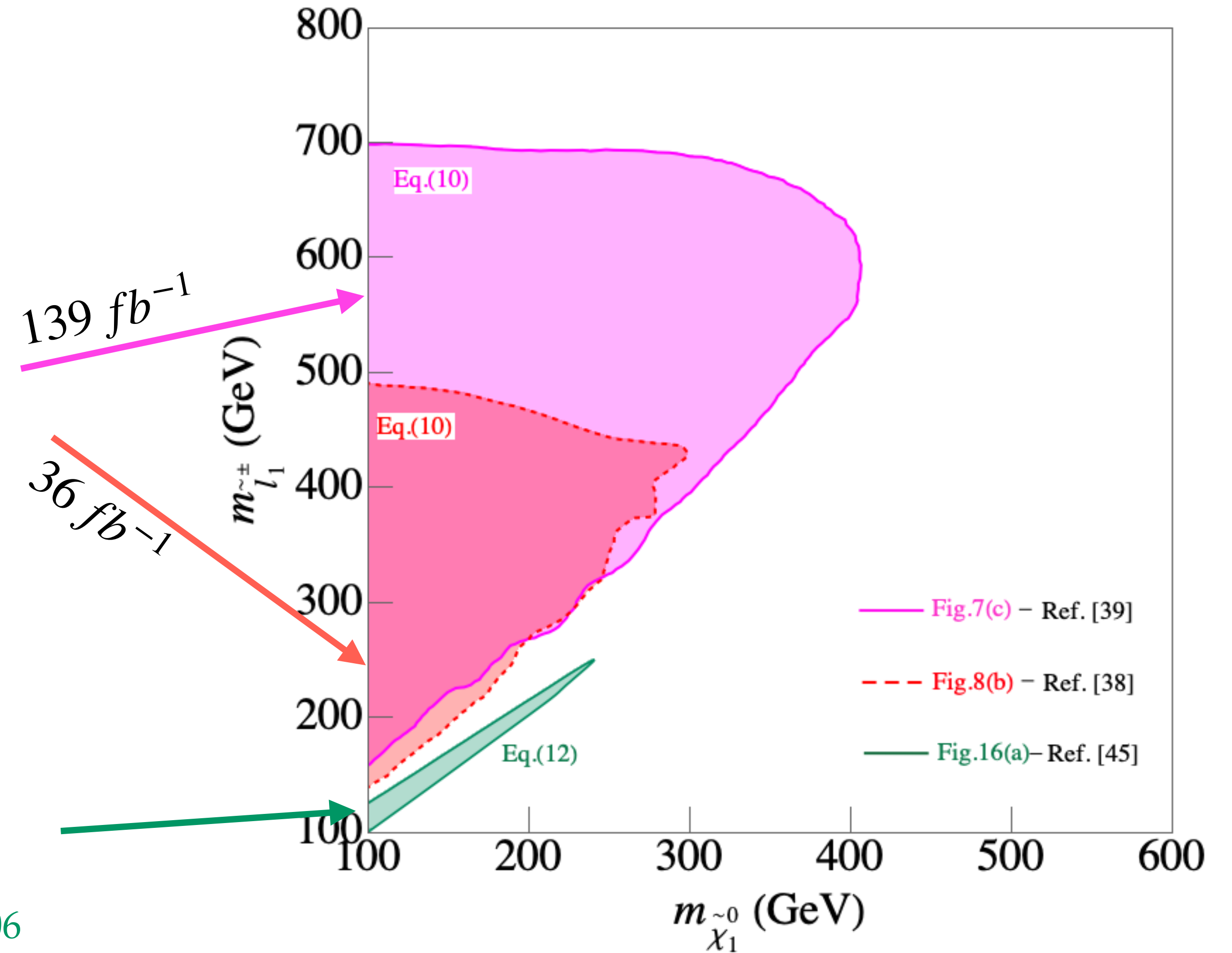
ATLAS [1908.08215]

13 TeV, 139 fb^{-1}



COMPRESSED

ATLAS 1911.12606



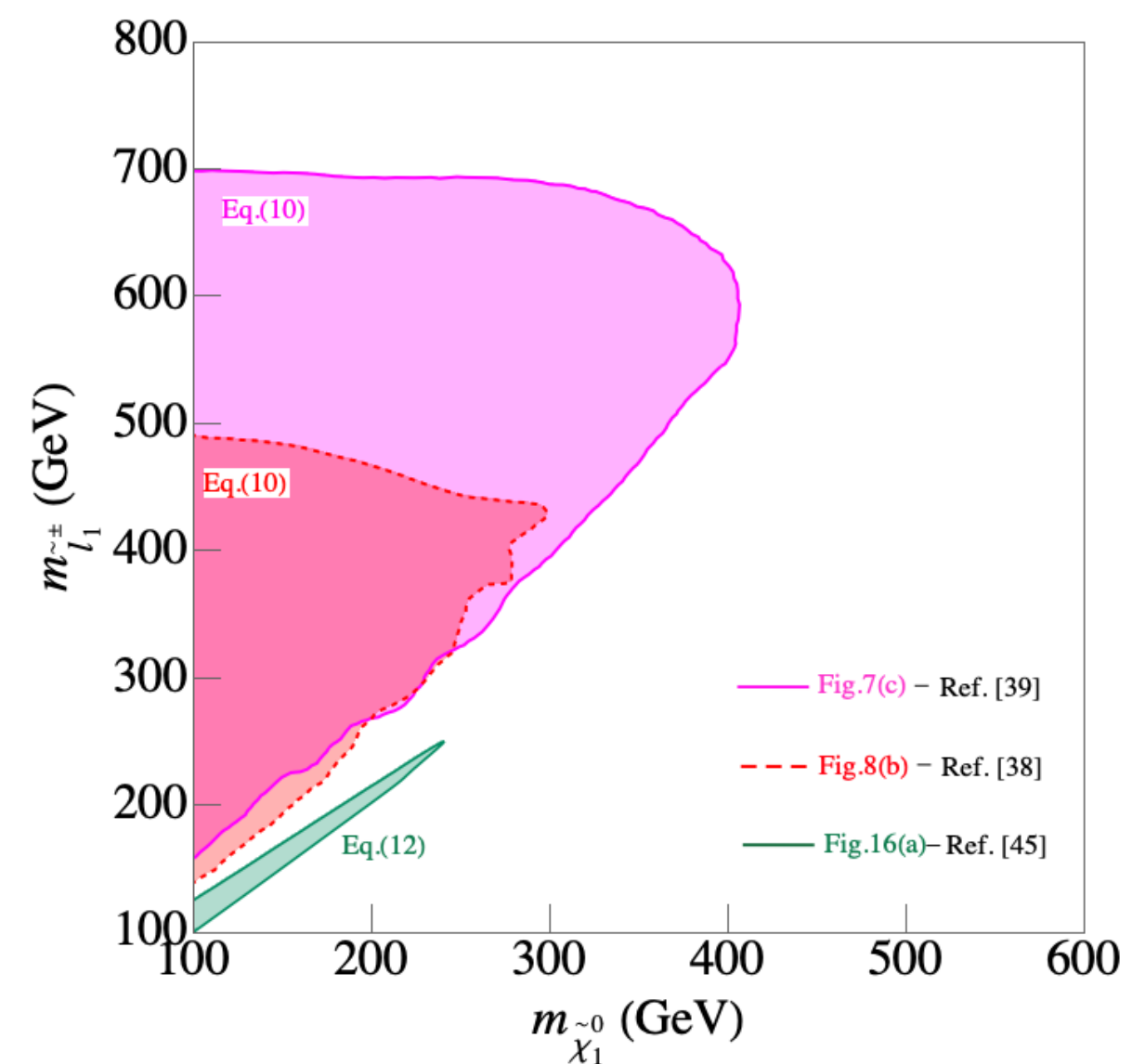
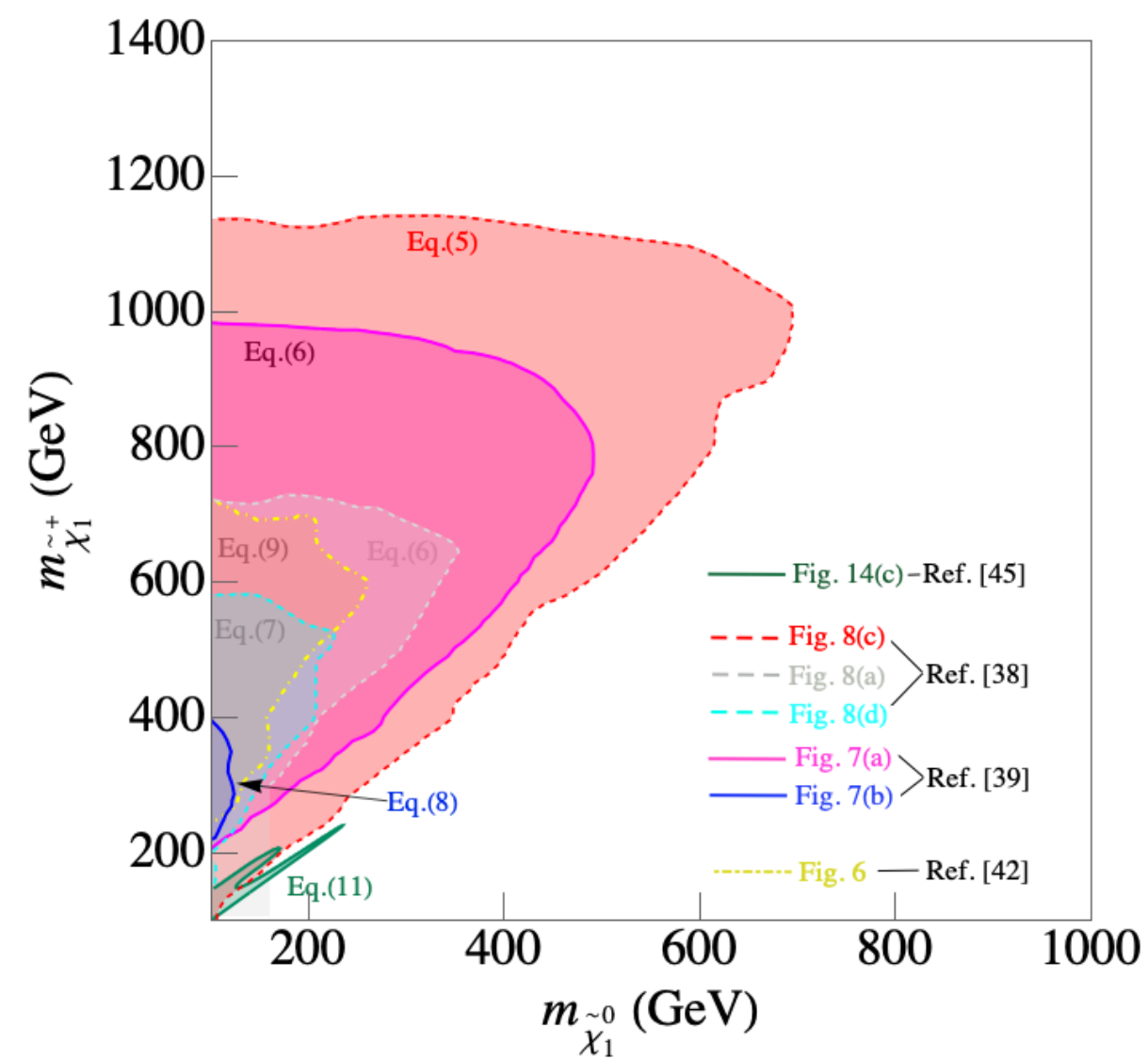
Proper recasting is important → checkMATE

Recasting with CM

Drees, Dreiner, Schmeier, Tattersall, Kim '13

Kim, Schmeier, Tattersall, Rolbiecki '15

Dercks, Desai, Kim, Rolbiecki, Tattersall '16



Show in the color	CheckMATE-implementation
Red dashed	✓
cyan dashed	✓
Gray dashed	✓
Magenta	✓
Blue	✓
Green	✗

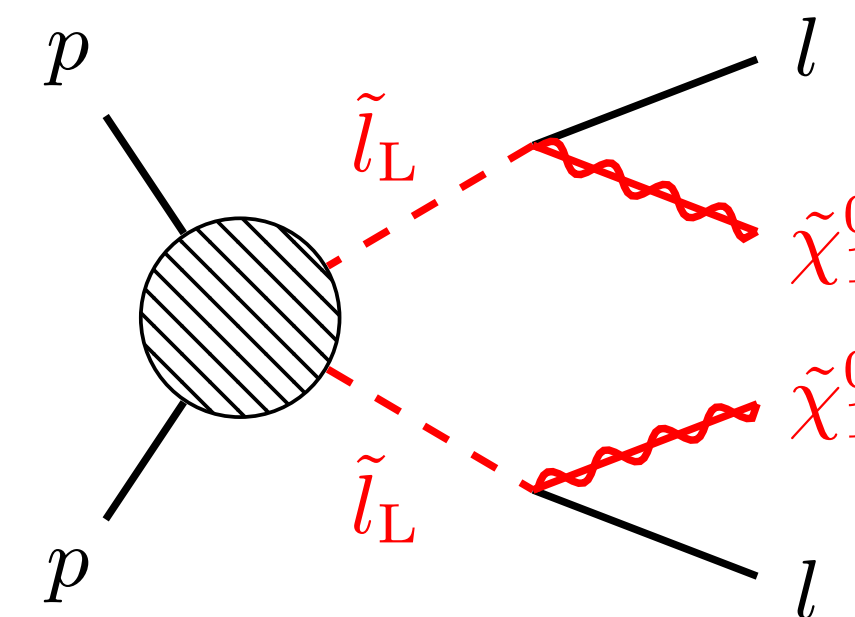
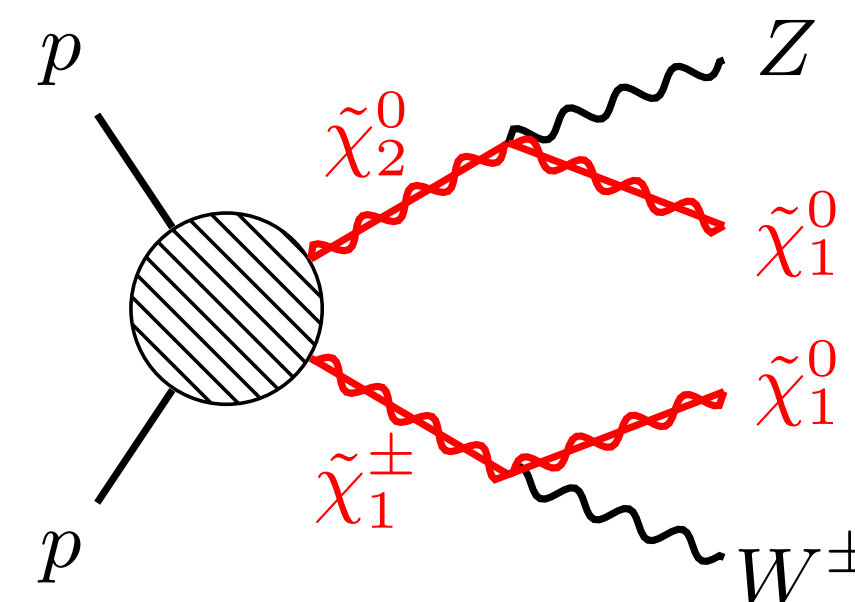
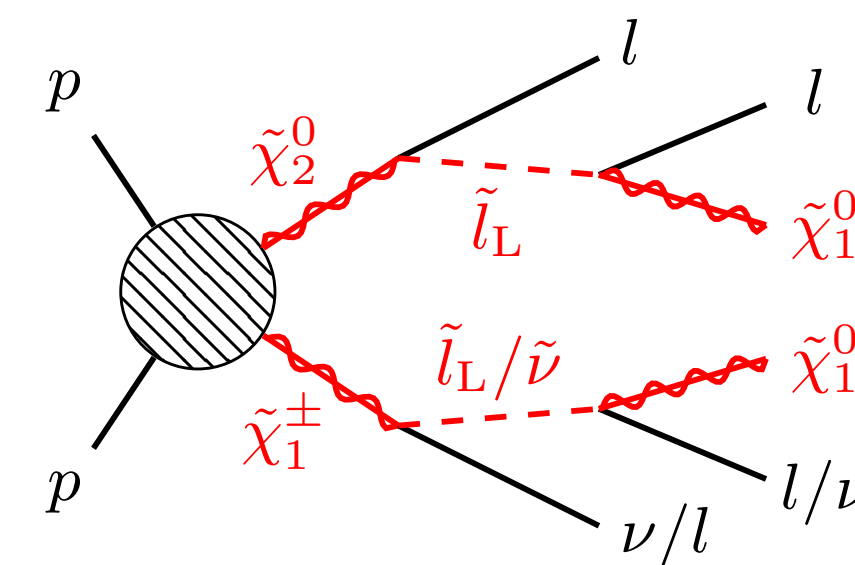
Magenta	✓
Red dashed	✓
Green	✗

Compressed spectra searches applied directly

• ATLAS [1803.02762]

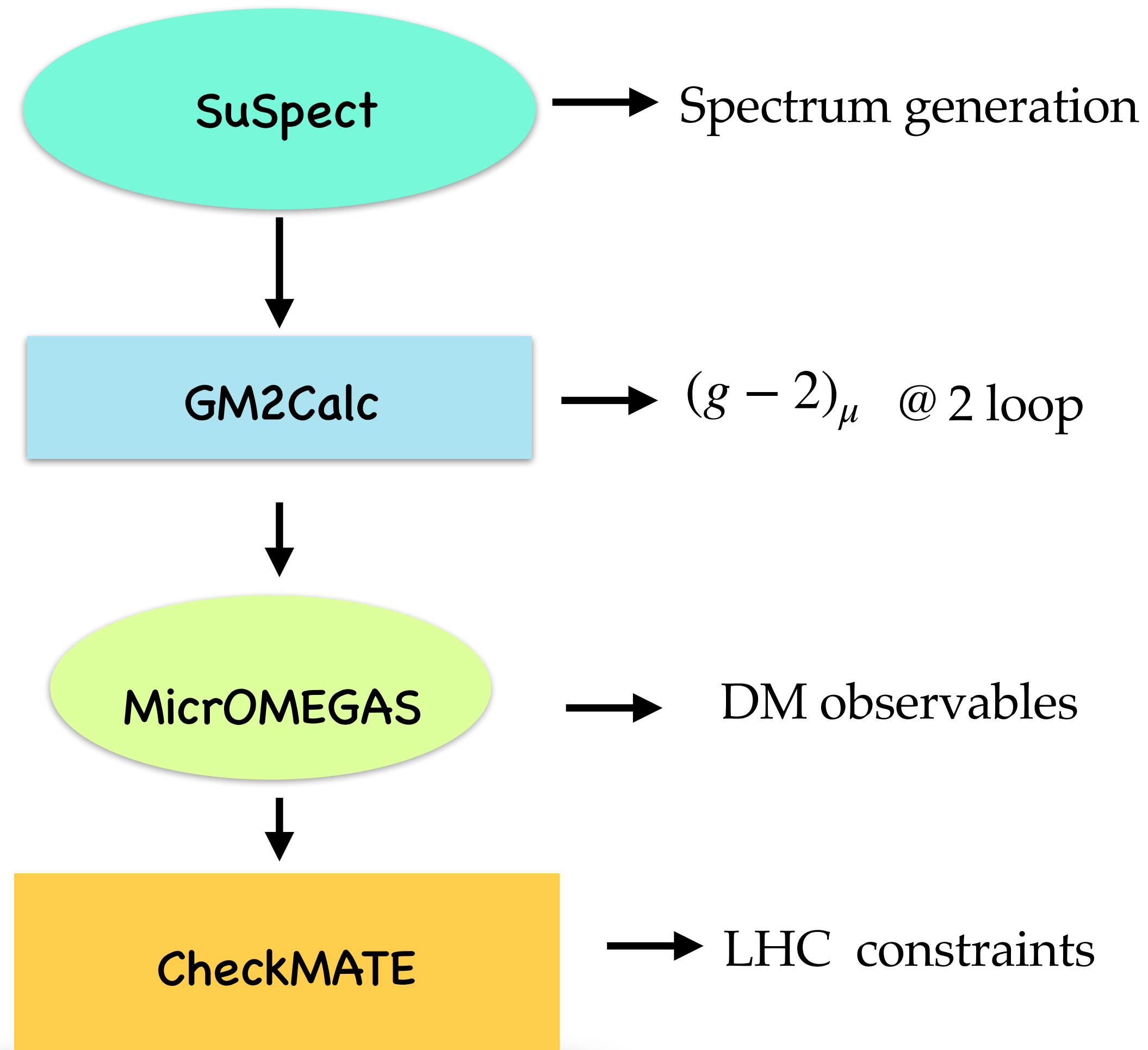
• ATLAS [1803.02762]

• ATLAS [1908.08215]



Most relevant in our case

Analysis flow



- $\Delta a_\mu = (28.02 \pm 7.37) \times 10^{-10}$
- Anticipated future bound
 $\Delta a_\mu^{fut} = (28.02 \pm 5.2) \times 10^{-10}$
- $\Omega_{CDM} h^2 = 0.120 \pm 0.001$
- Direct detection SI bounds from XENON1T

Parameter Scanning

Chargino co-annihilation region:

$$100 \text{ GeV} \leq M_1 \leq 1 \text{ TeV}, \quad M_1 \leq M_2 \leq 1.1M_1, \\ 1.1M_1 \leq \mu \leq 10M_1, \quad 5 \leq \tan \beta \leq 60, \\ 100 \text{ GeV} \leq m_{\tilde{l}_L} \leq 1 \text{ TeV}, \quad m_{\tilde{l}_R} = m_{\tilde{l}_L}.$$

Bino-wino co-annihilation

Slepton co-annihilation region:

Case-L: SU(2) doublet

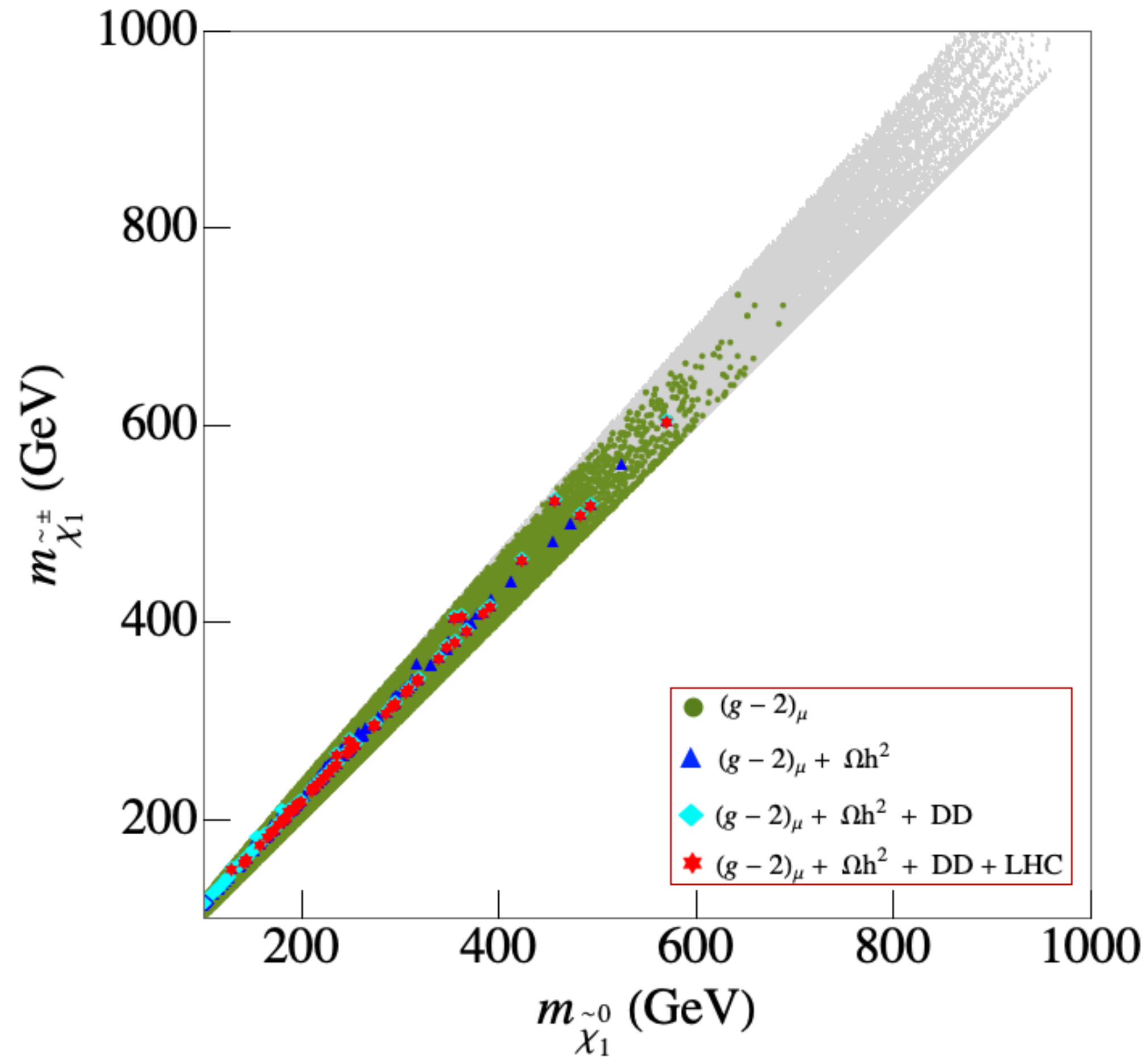
$$100 \text{ GeV} \leq M_1 \leq 1 \text{ TeV}, \quad M_1 \leq M_2 \leq 10M_1, \\ 1.1M_1 \leq \mu \leq 10M_1, \quad 5 \leq \tan \beta \leq 60, \\ M_1 \text{ GeV} \leq m_{\tilde{l}_L} \leq 1.2M_1, \quad M_1 \leq m_{\tilde{l}_R} \leq 10M_1.$$

Case-R: SU(2) singlet

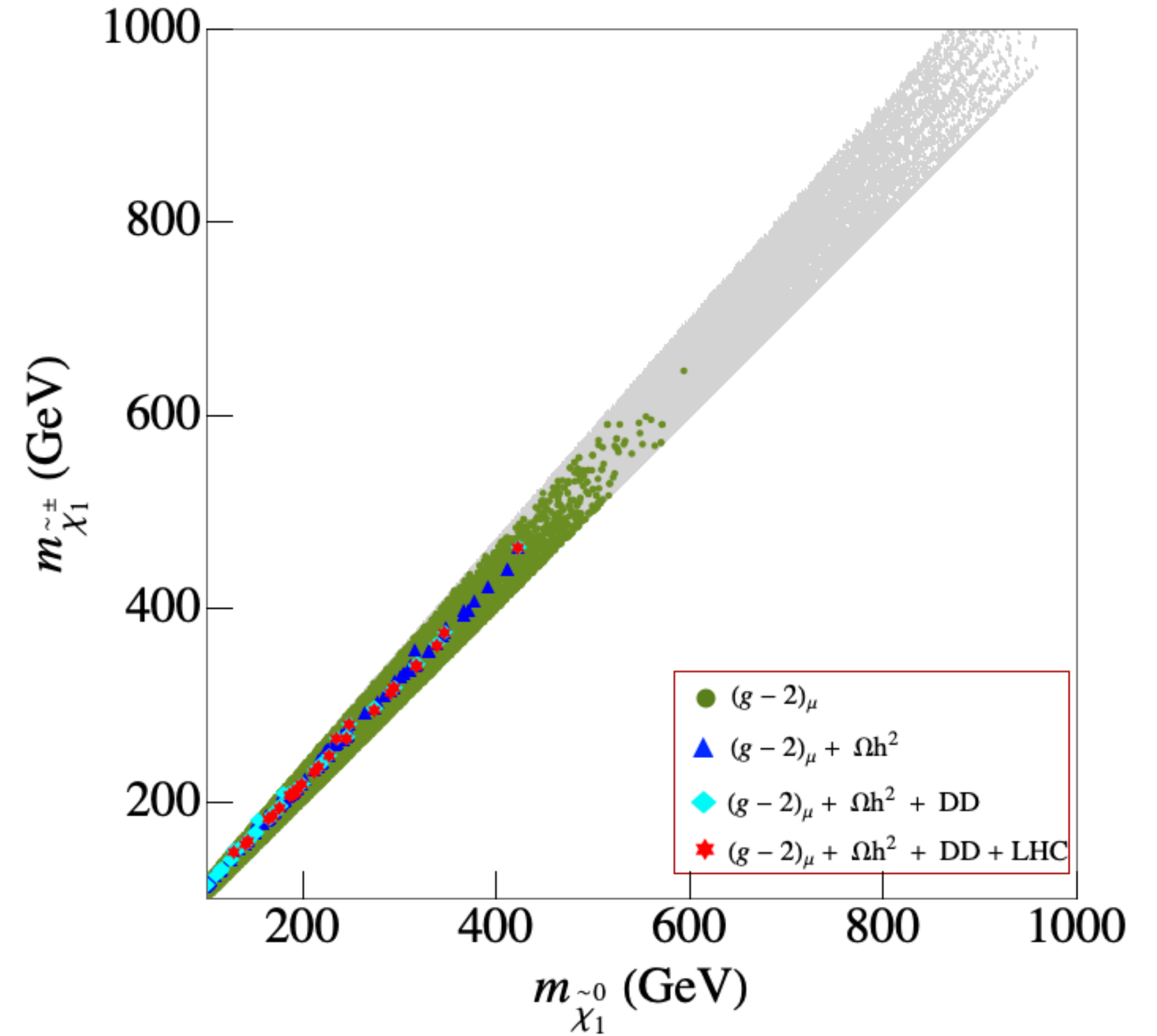
$$100 \text{ GeV} \leq M_1 \leq 1 \text{ TeV}, \quad M_1 \leq M_2 \leq 10M_1, \\ 1.1M_1 \leq \mu \leq 10M_1, \quad 5 \leq \tan \beta \leq 60, \\ M_1 \text{ GeV} \leq m_{\tilde{l}_R} \leq 1.2M_1, \quad M_1 \leq m_{\tilde{l}_L} \leq 10M_1.$$

Chargino Co-annihilation

Current $(g - 2)_\mu$ limit



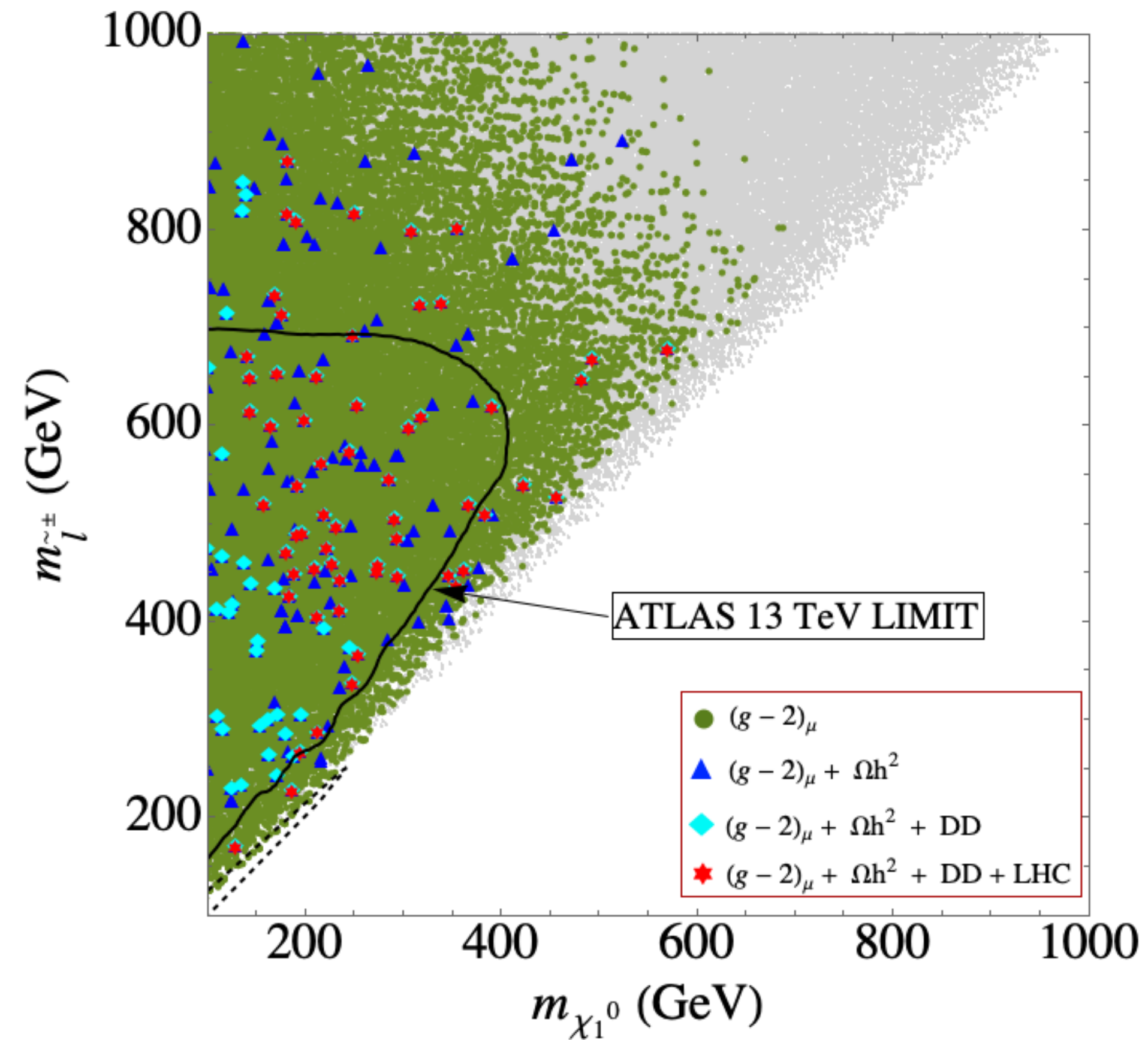
Anticipated future $(g - 2)_\mu$ limit



Upper and lower bounds from $(g - 2)_\mu$ and LHC searches (for compressed spectrum)

Results in the $m_{\tilde{\chi}_1^0} - m_{\tilde{l}_1}$ plane

Current $(g - 2)_\mu$ limit

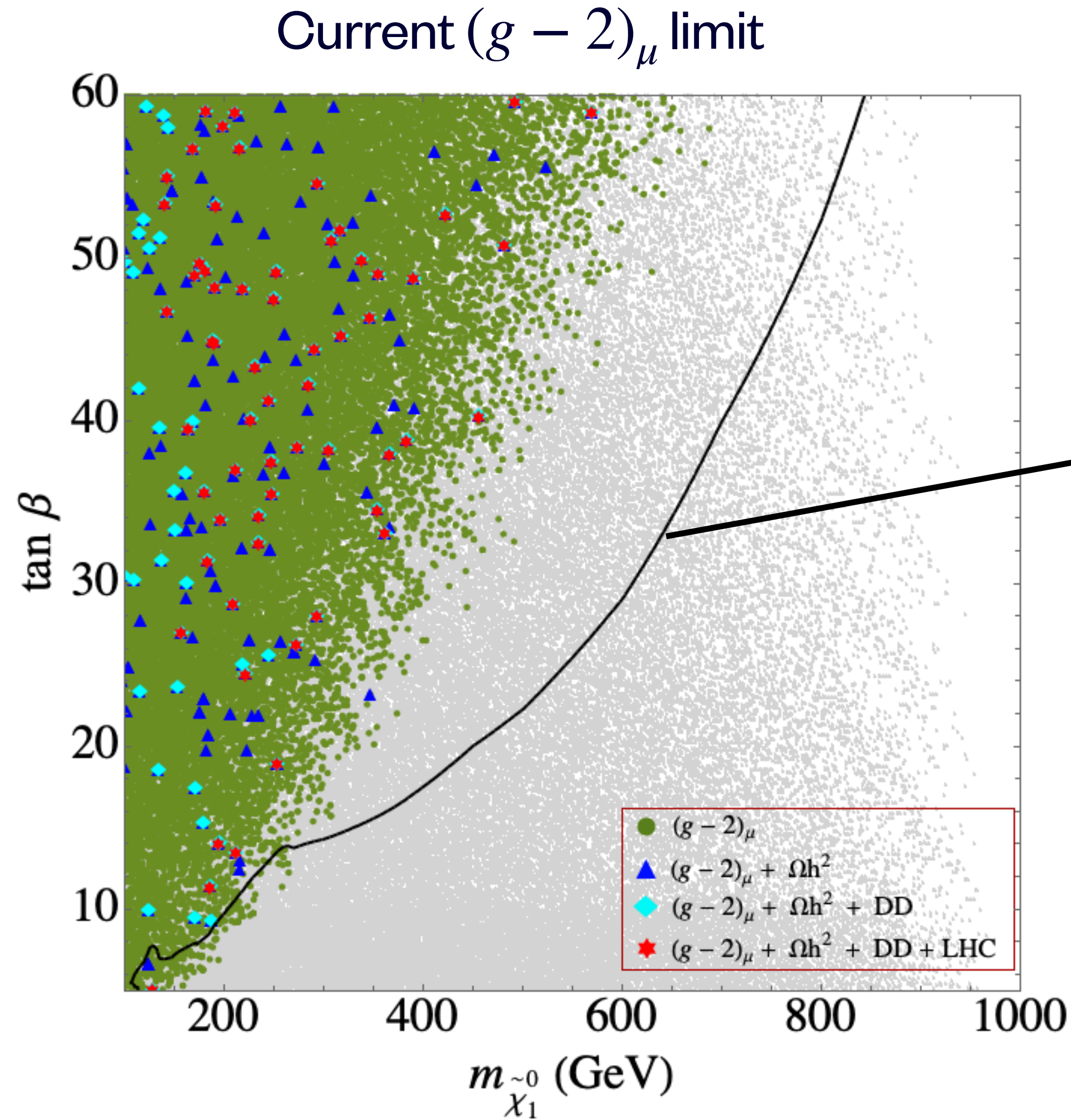


Slepton-pair production $\rightarrow (2l + \text{missing } E_T)$ provides important search channel

R-sleptons heavy, Considerable BR for $\tilde{e}_L(\tilde{\mu}_L) \rightarrow \tilde{\chi}_1^+ \nu_e(\nu_\mu) \rightarrow$ Less no. of signal leptons.

Possibility of A-pole annihilation

$$a_\mu \sim \frac{\tan \beta}{m_{EW}^2}$$



$$m_{\tilde{\chi}_1^0} = \frac{M_A}{2}$$

$M_h^{125}(\tilde{\chi})$ Benchmark scenario

Bagnaschi et al. '18

Black contour : simplified application of $H/A \rightarrow \tau^+ \tau^-$ \longrightarrow A-pole annihilation strongly constrained

DM with Low abundance

$$\Omega_{CDM} h^2 \leq 0.122$$

Wino LSP

$$100 \text{ GeV} \leq M_2 \leq 1.5 \text{ TeV}, \quad 1.1M_2 \leq M_1 \leq 10M_2, \\ 1.1M_2 \leq \mu \leq 10M_2, \quad 5 \leq \tan \beta \leq 60, \\ 100 \text{ GeV} \leq m_{\tilde{l}_L}, m_{\tilde{l}_R} \leq 2 \text{ TeV}.$$

$SU(2)_L$ triplet

Under-abundant upto $\sim 3 \text{ TeV}$

Compressed spectra with $m_{\tilde{\chi}_1^0} \sim m_{\tilde{\chi}_1^\pm}$

Higgsino LSP

$$100 \text{ GeV} \leq \mu \leq 1.2 \text{ TeV}, \quad 1.1\mu \leq M_1 \leq 10\mu, \\ 1.1\mu \leq M_2 \leq 10\mu, \quad 5 \leq \tan \beta \leq 60, \\ 100 \text{ GeV} \leq m_{\tilde{l}_L}, m_{\tilde{l}_R} \leq 2 \text{ TeV}.$$

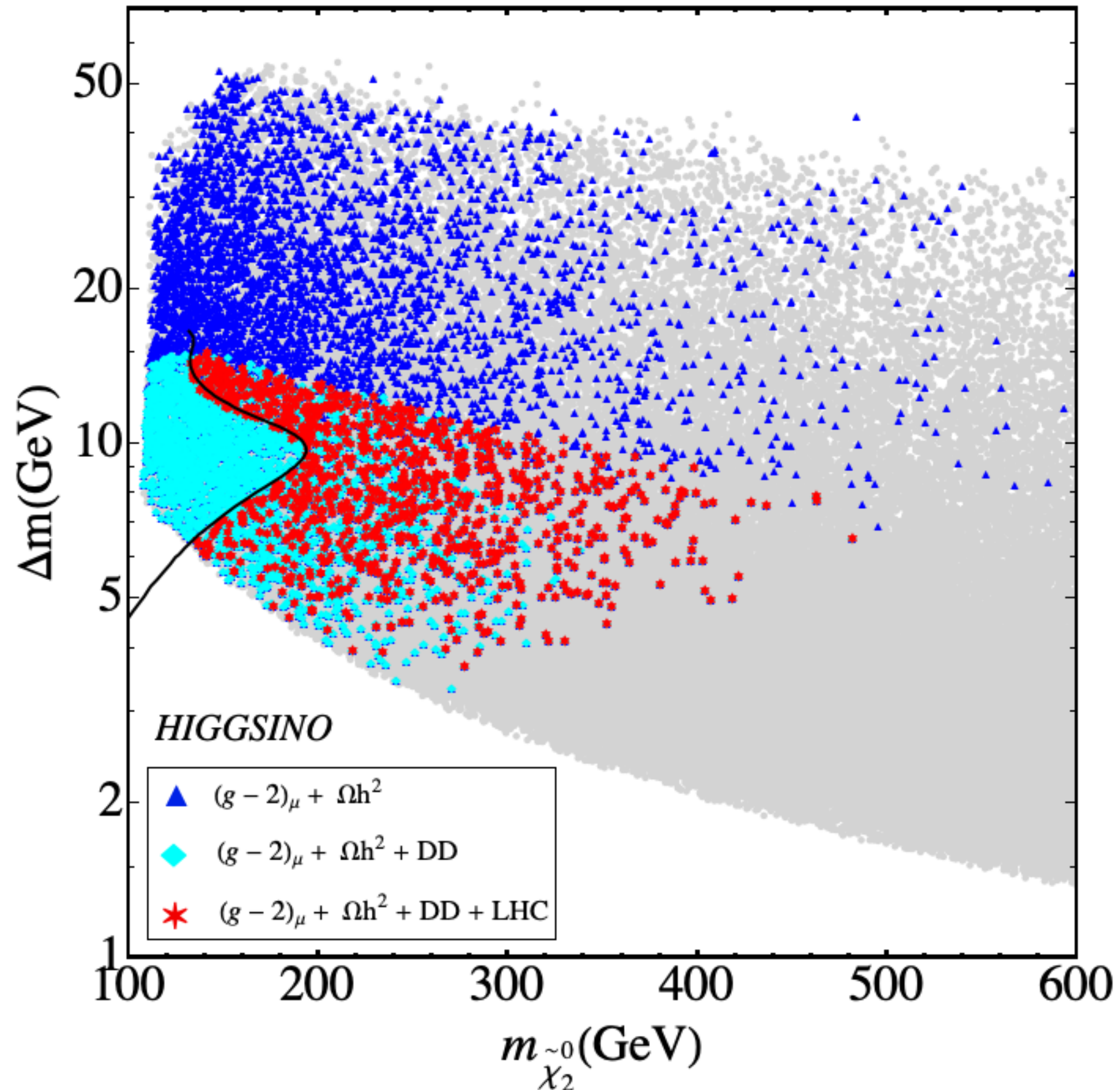
$SU(2)_L$ doublet

Under-abundant upto $\sim 1 \text{ TeV}$

Compressed spectra with $m_{\tilde{\chi}_1^0} \sim m_{\tilde{\chi}_1^\pm} \sim m_{\tilde{\chi}_2^0}$

Higgsino LSP

Current $(g - 2)_\mu$ limit

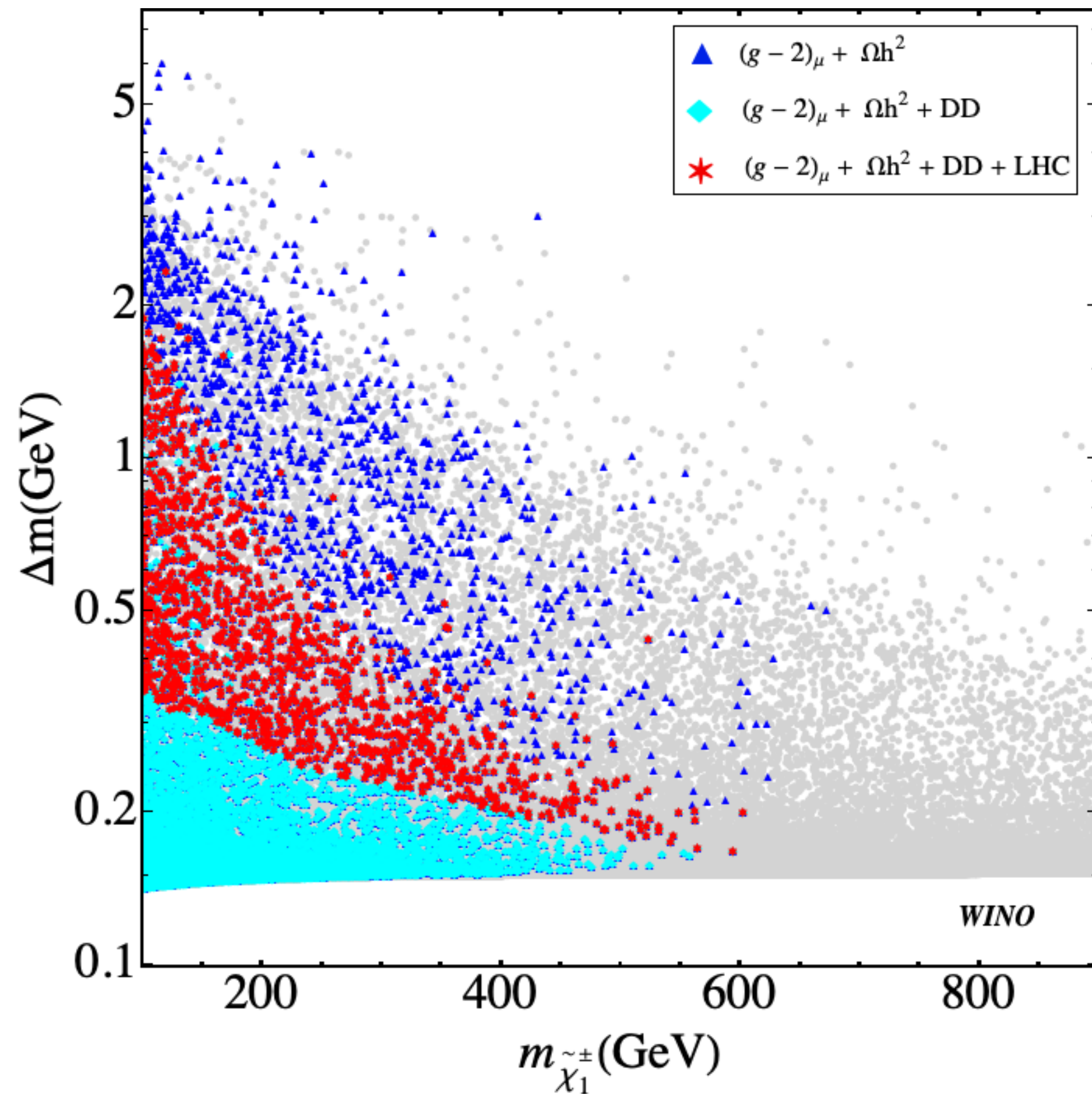


- Compressed spectra searches most important.
- Slepton pair production searches also relevant
- $\Delta m \sim \mathcal{O}(10)$ GeV \rightarrow Disappearing track searches not sensitive

$$c\tau \simeq 0.7 \text{ cm} \times \left[\left(\frac{\Delta m_+}{340 \text{ MeV}} \right)^3 \sqrt{1 - \frac{m_\pi^2}{\Delta m_+^2}} \right]^{-1}$$

Wino LSP

Current $(g - 2)_\mu$ limit



High Δm restricted by DD

Low Δm restricted by LHC

- Tree level splitting from \tilde{h} mixing

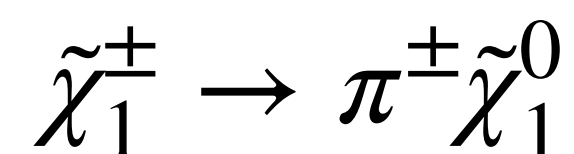
$$\simeq \frac{M_W^4 (\sin 2\beta)^2 \tan^2 \theta_w}{(M_1 - M_2) \mu^2}$$

- Coupling for DD

$$c_{h\tilde{\chi}_1^0\tilde{\chi}_1^0} \simeq \frac{M_W}{M_2^2 - \mu^2} (M_2 + \mu \sin 2\beta),$$

Ibe, Matsumoto, Sato '13

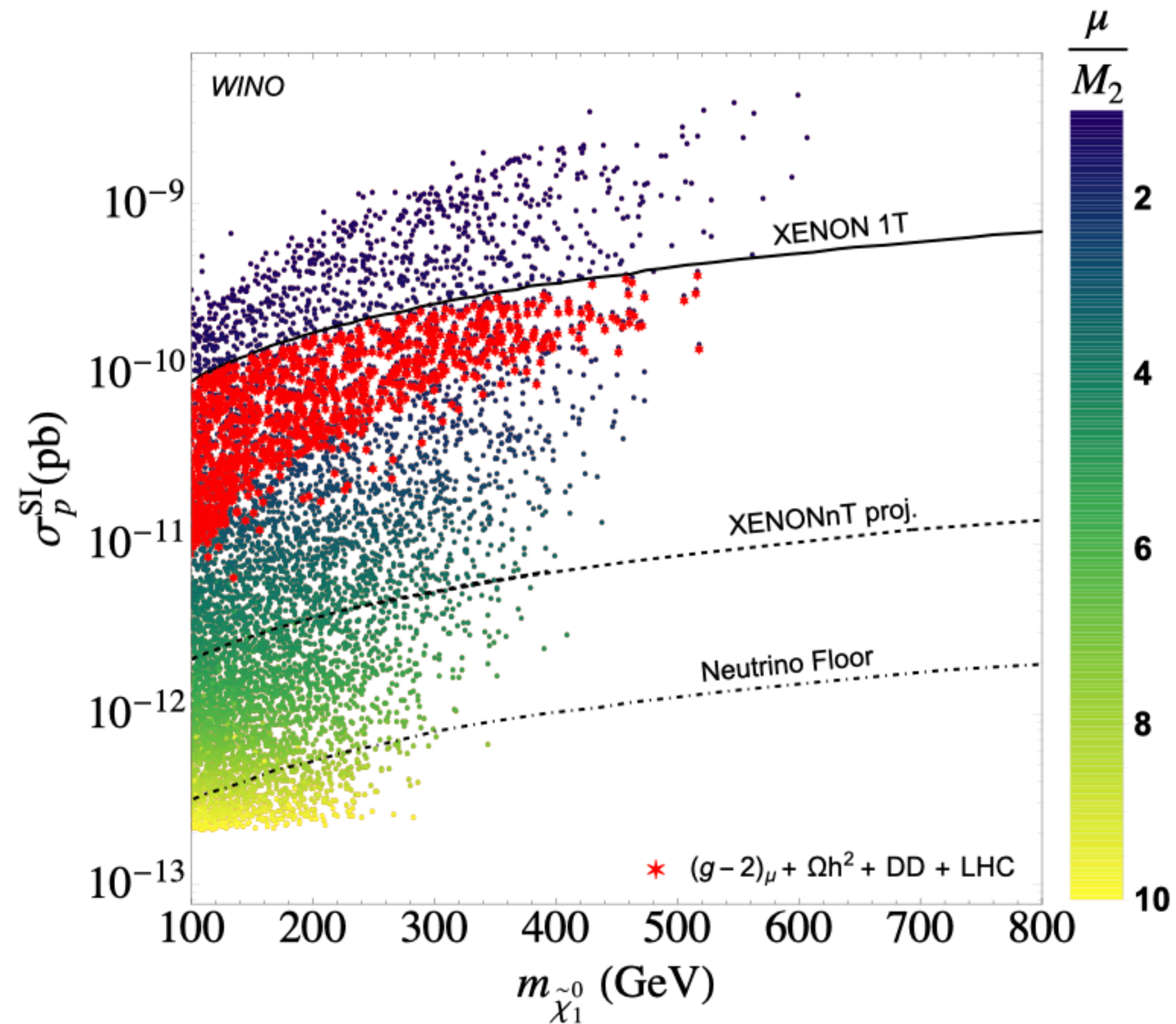
- Disappearing track searches



CMS-EXO-19-010

Wino LSP : Direct detection

Current $(g - 2)_\mu$ limit

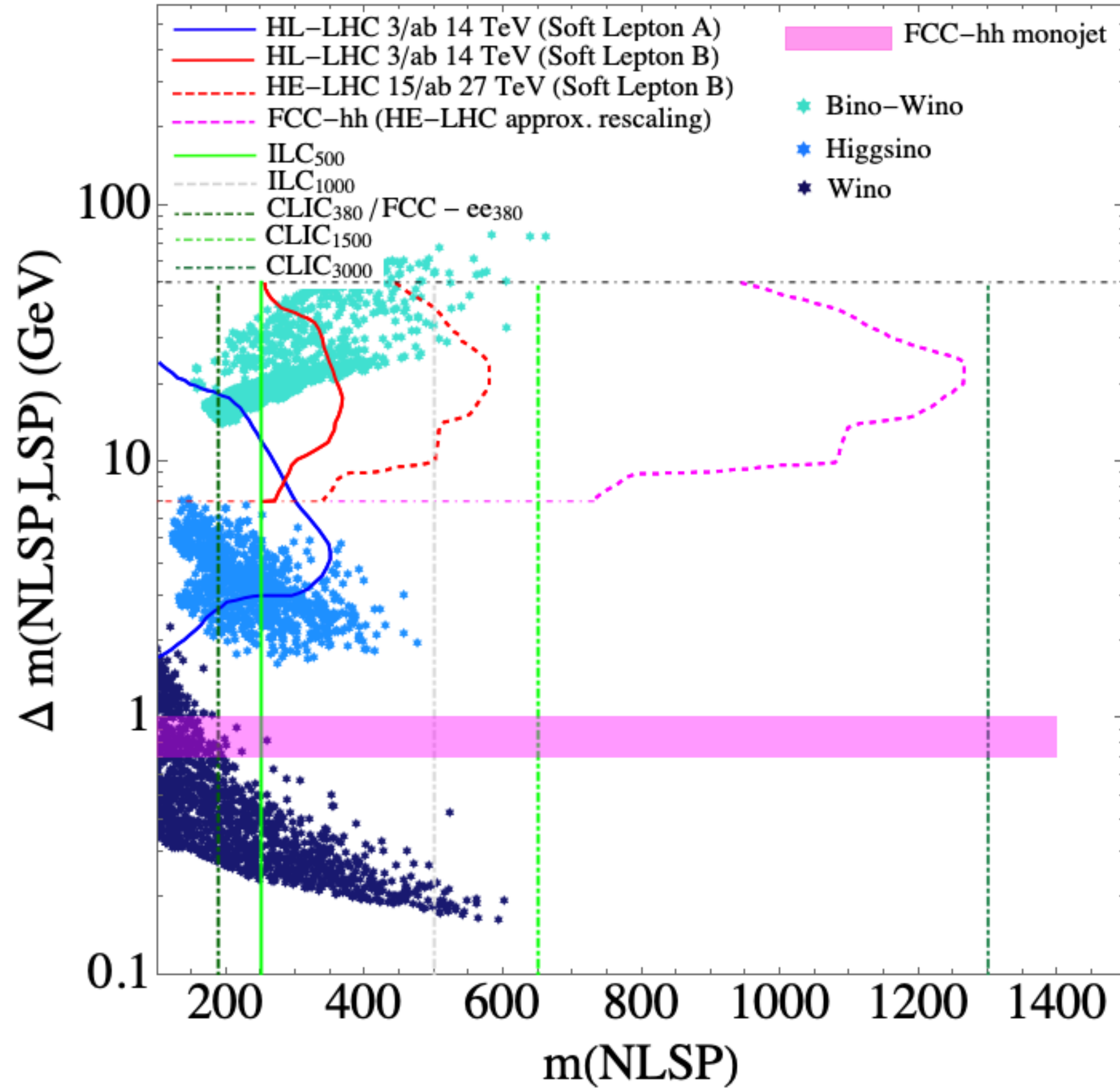


- DD coupling

$$c_{h\tilde{\chi}_1^0\tilde{\chi}_1^0} \simeq \frac{M_W}{M_2^2 - \mu^2} (M_2 + \mu \sin 2\beta),$$

- All allowed points to be checked by XENONnT

Future prospects



Conclusions

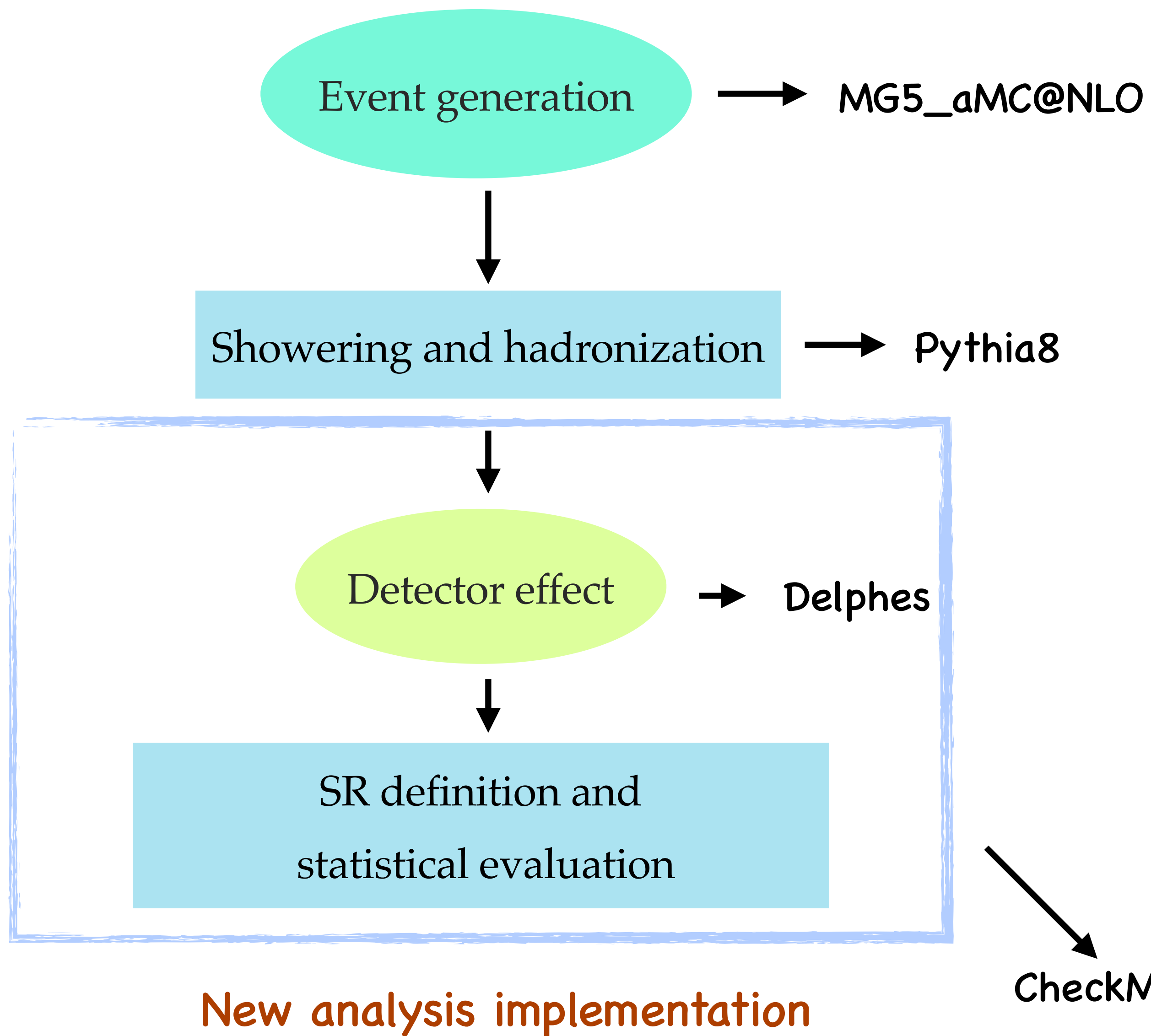
- It is possible to constrain the EW MSSM with the help of indirect constraints along with the direct collider limits.
- DM and muon $(g-2)$ constraints put effective upper limit on EW SUSY masses.
- LHC limits restrict the mass ranges from below.
- Proper recasting of ATLAS/CMS analyses important!
- **Future linear collider** searches will be conclusive.
- New experimental results for $(g - 2)_\mu$ from **Fermilab**, J-PARC **STAY TUNED!!!**

THANK YOU!

BACKUP

Recasting with CM

Drees, Dreiner, Schmeier, Tattersall, Kim '13
Kim, Schmeier, Tattersall, Rolbiecki '15
Dercks, Desai, Kim, Rolbiecki, Tattersall '16



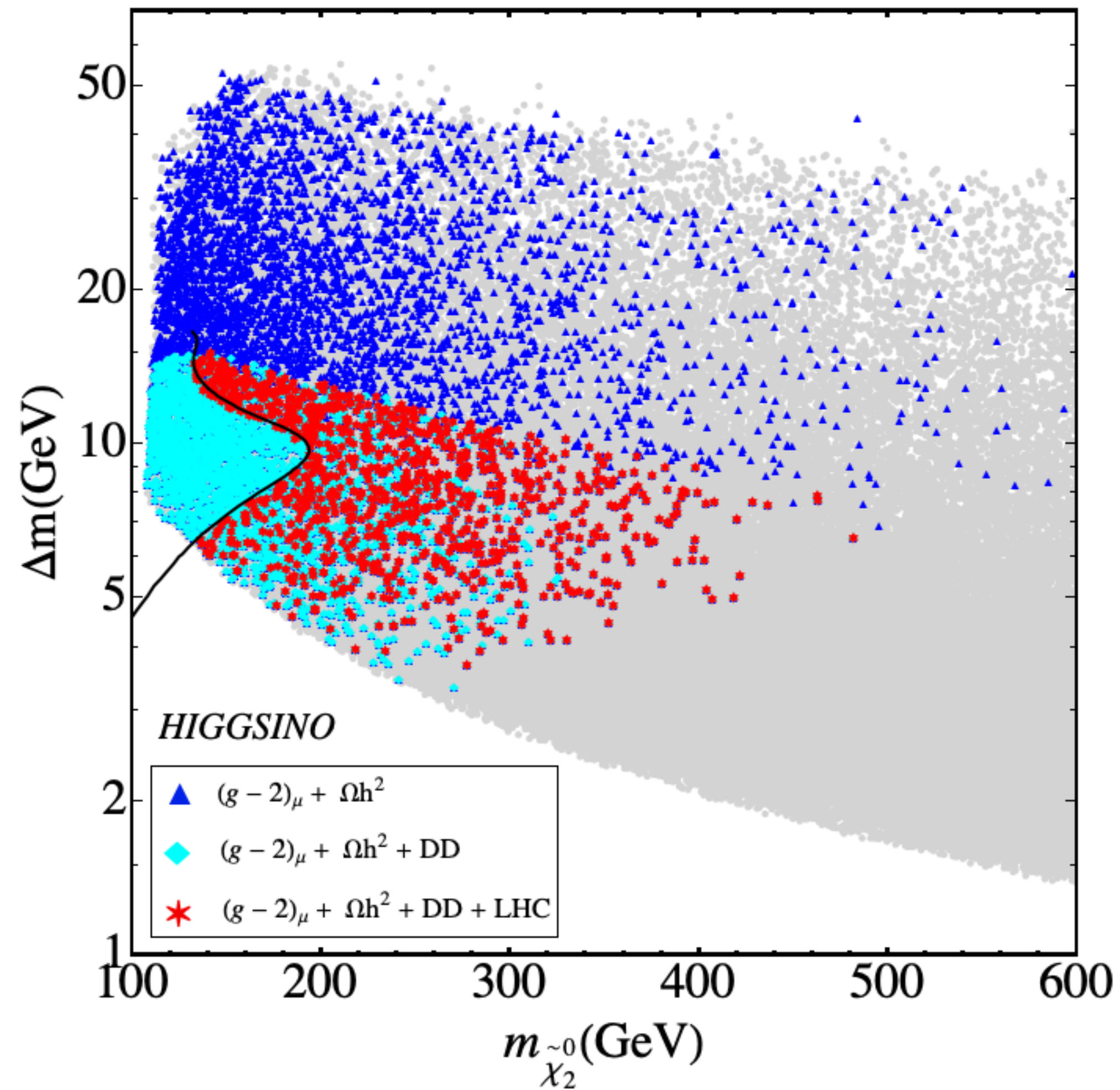
Model testing

- Each parameter point is tested against newly implemented analyses
- Signal events calculated for each SR
- Evaluation of $r = \frac{S - 1.96 \times \Delta S}{S_{exp}^{95}}$
- For the best SR, $r > 1 \rightarrow$ **excluded!**

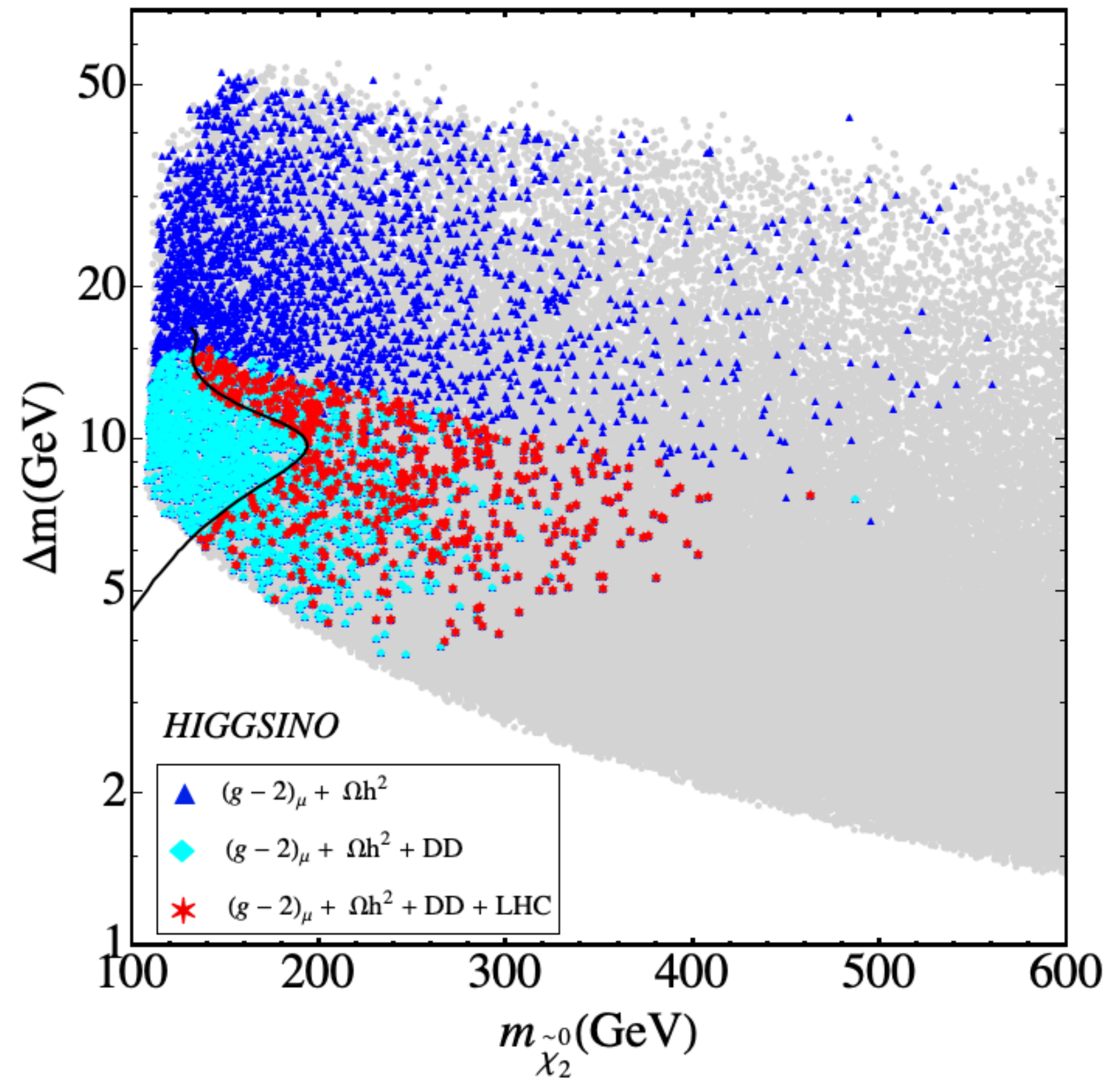
CheckMATE \rightarrow Experimental Cutflow reproduced

Higgsino LSP

Current $(g - 2)_\mu$ limit

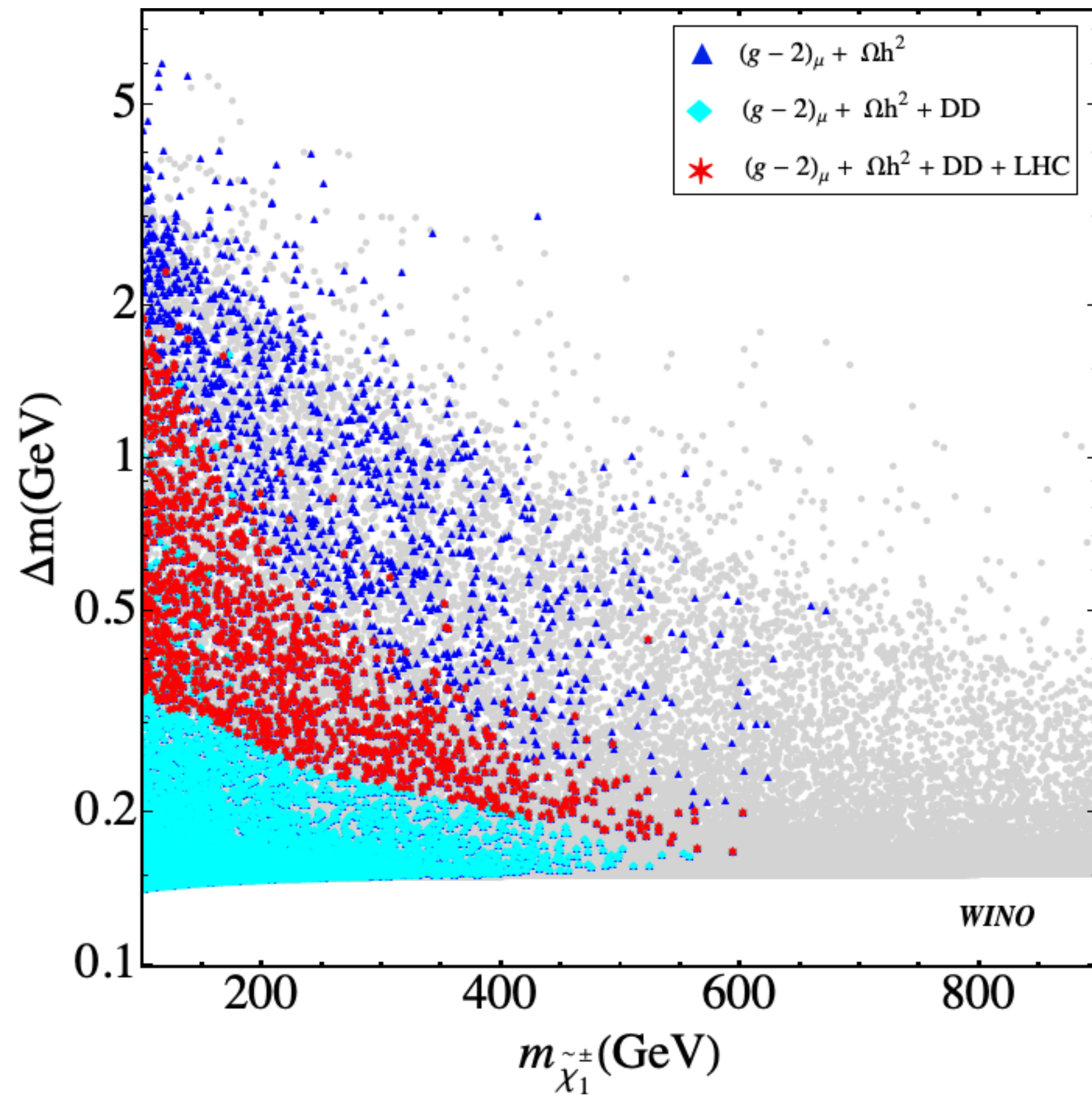


Future $(g - 2)_\mu$ limit

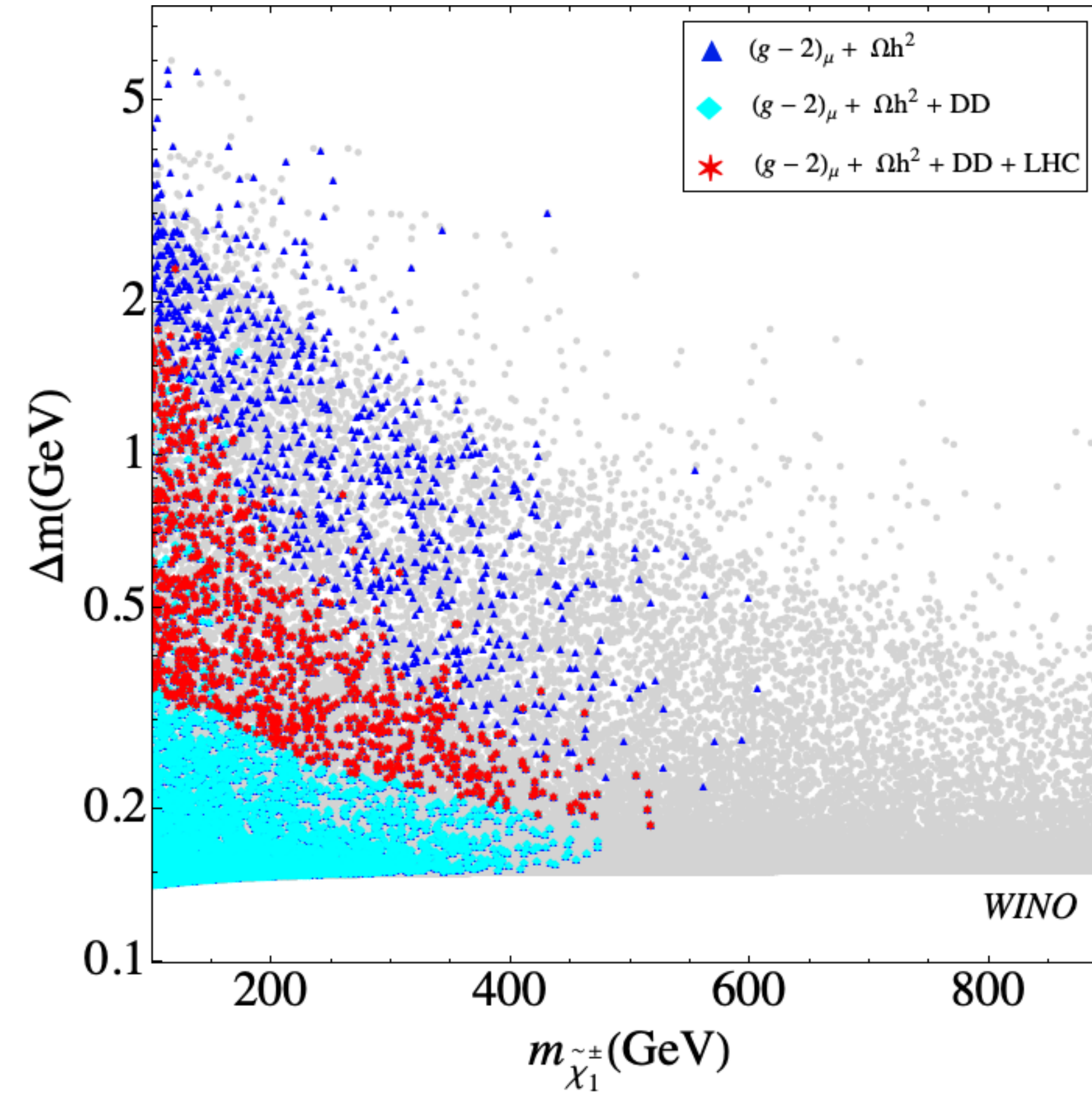


Wino LSP

Current $(g - 2)_\mu$ limit

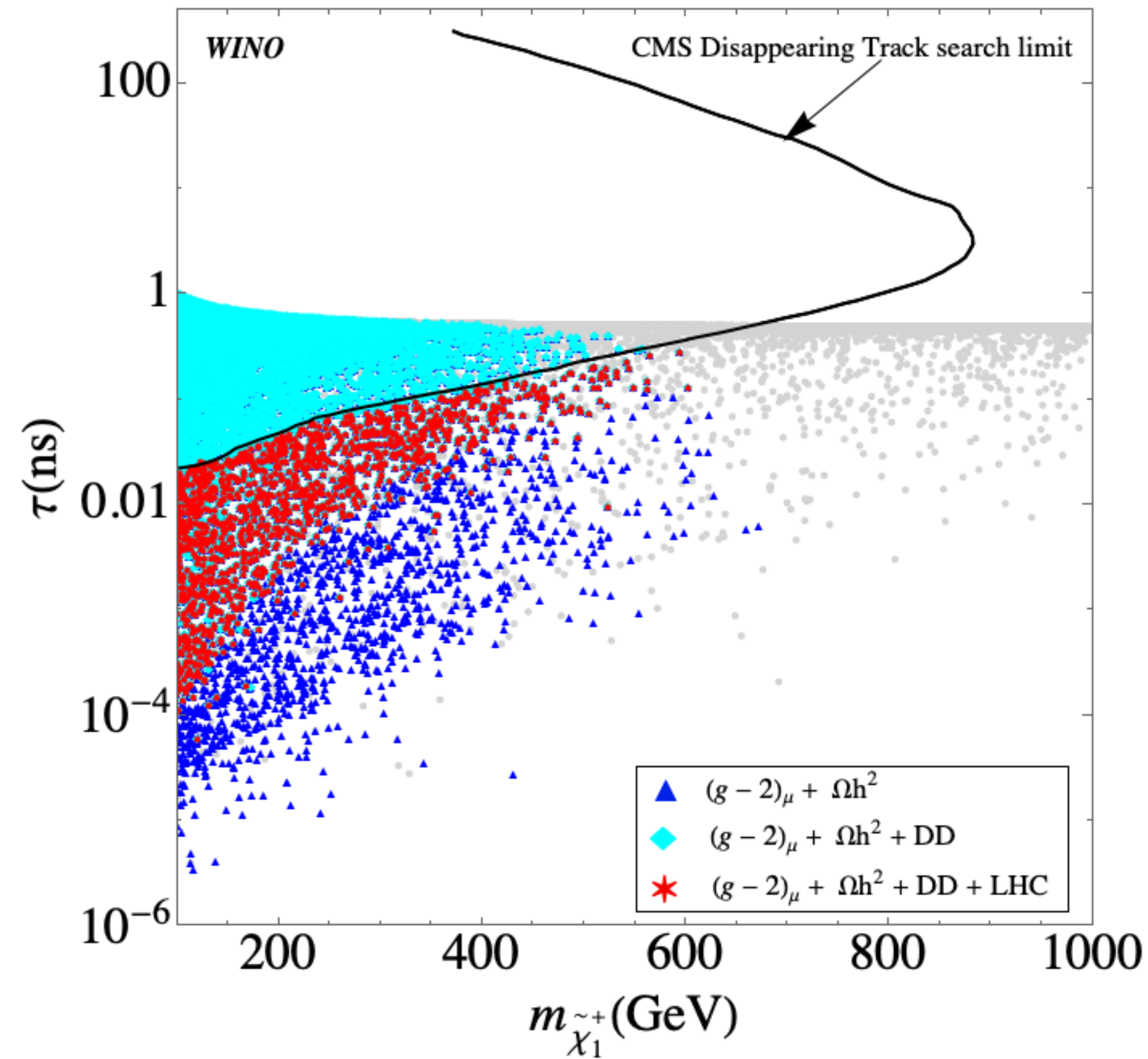


Future $(g - 2)_\mu$ limit



Wino LSP lifetime

Current $(g - 2)_\mu$ limit



$$\Gamma(\tilde{\chi}^\pm \rightarrow \tilde{\chi}^0 \pi^\pm) = \Gamma(\pi^\pm \rightarrow \mu^\pm \nu_\mu) \times \frac{16\delta m^3}{m_\pi m_\mu^2} \left(1 - \frac{m_\pi^2}{\delta m^2}\right)^{1/2} \left(1 - \frac{m_\mu^2}{m_\pi^2}\right)^{-2}$$

$$\Gamma(\tilde{\chi}^\pm \rightarrow \tilde{\chi}^0 e^\pm \nu_e) \simeq \frac{2G_F^2}{15\pi^3} \delta m^5.$$

- Disappearing track searches most important.

MSSM Superpotential

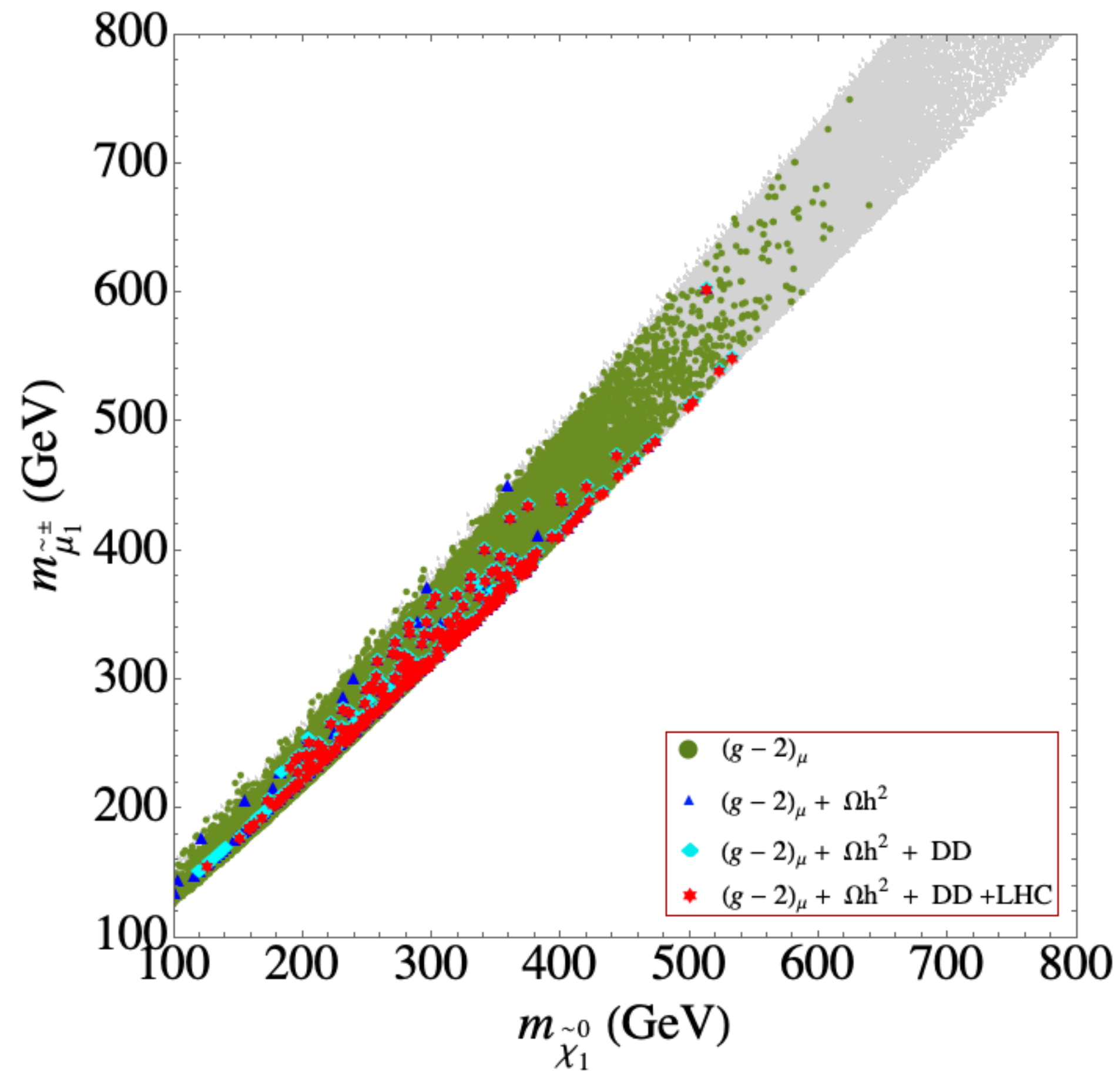
$$W_{\text{MSSM}} = \bar{u}Y_u QH_u - \bar{d}Y_d QH_d - \bar{e}Y_e LH_d + \mu H_u H_d$$

Soft Breaking Terms

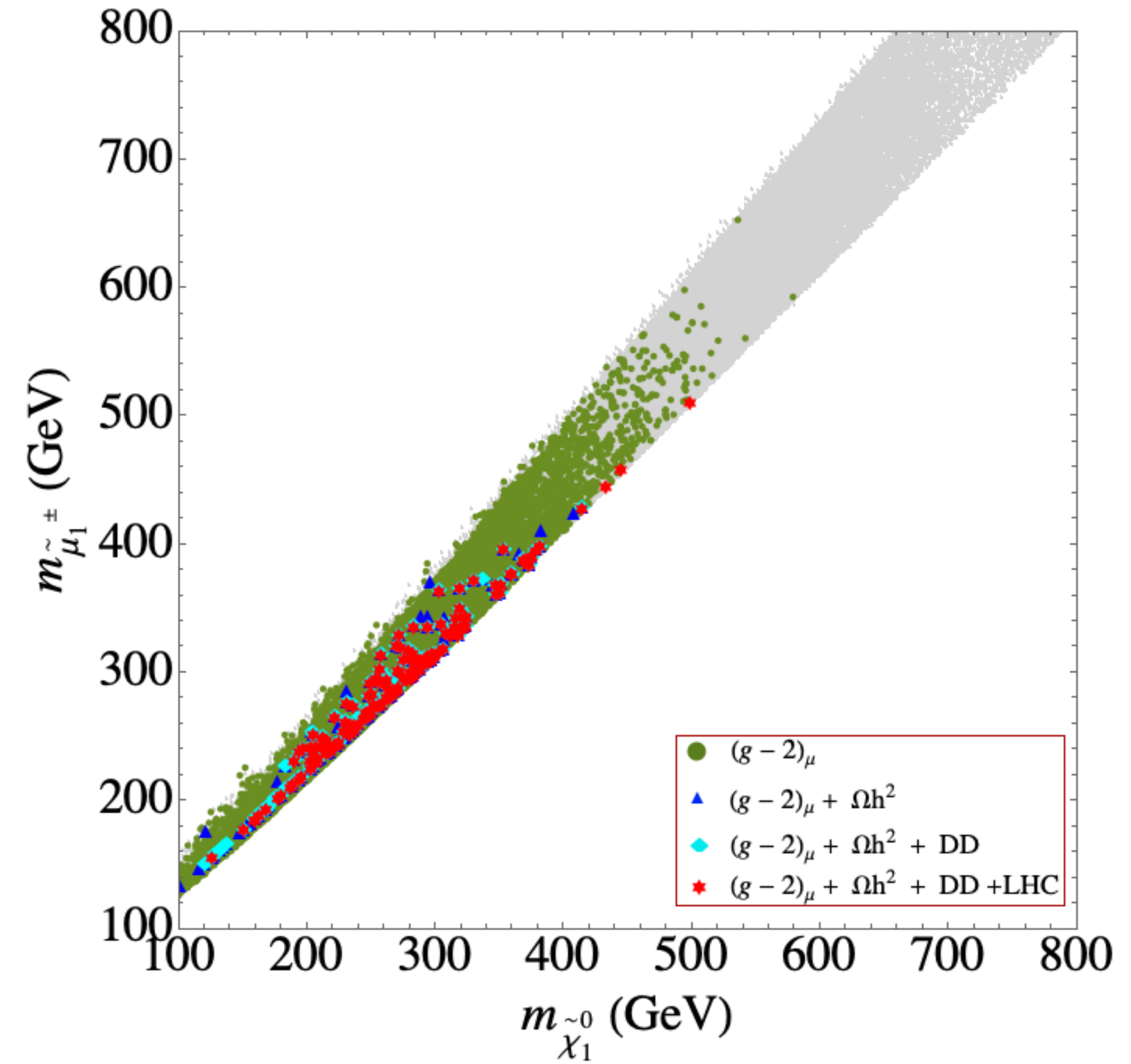
$$\begin{aligned} \mathcal{L}_{\text{soft}}^{\text{MSSM}} = & -\frac{1}{2} (M_3 \tilde{g}\tilde{g} + M_2 \tilde{W}\tilde{W} + M_1 \tilde{B}\tilde{B} + c.c) \\ & - (\tilde{u} \mathbf{a}_u \tilde{Q}H_u - \tilde{d} \mathbf{a}_d \tilde{Q}H_d - \tilde{e} \mathbf{a}_e \tilde{L}H_d + c.c) \\ & - \tilde{Q}^\dagger \mathbf{m}_Q^2 \tilde{Q} - \tilde{L}^\dagger \mathbf{m}_L^2 \tilde{L} - \tilde{u} \mathbf{m}_u^2 \tilde{u}^\dagger - \tilde{d} \mathbf{m}_d^2 \tilde{d}^\dagger - \tilde{e} \mathbf{m}_e^2 \tilde{e}^\dagger \\ & - m_{H_u}^2 H_u^* H_u - m_{H_d}^2 H_d^* H_d - (bH_u H_d + c.c) \end{aligned}$$

Slepton Co-annihilation: Case-L

Current $(g - 2)_\mu$ limit

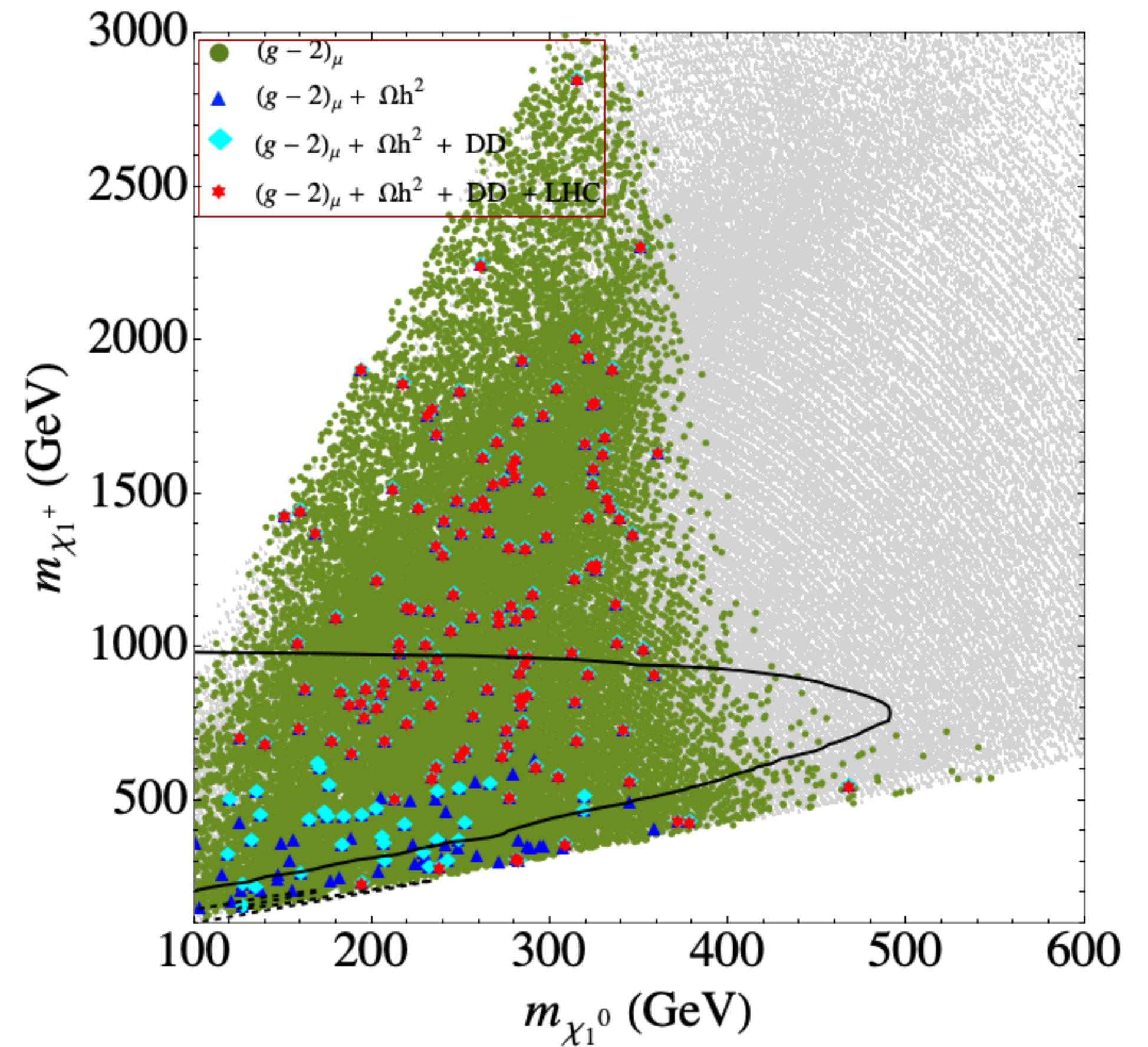
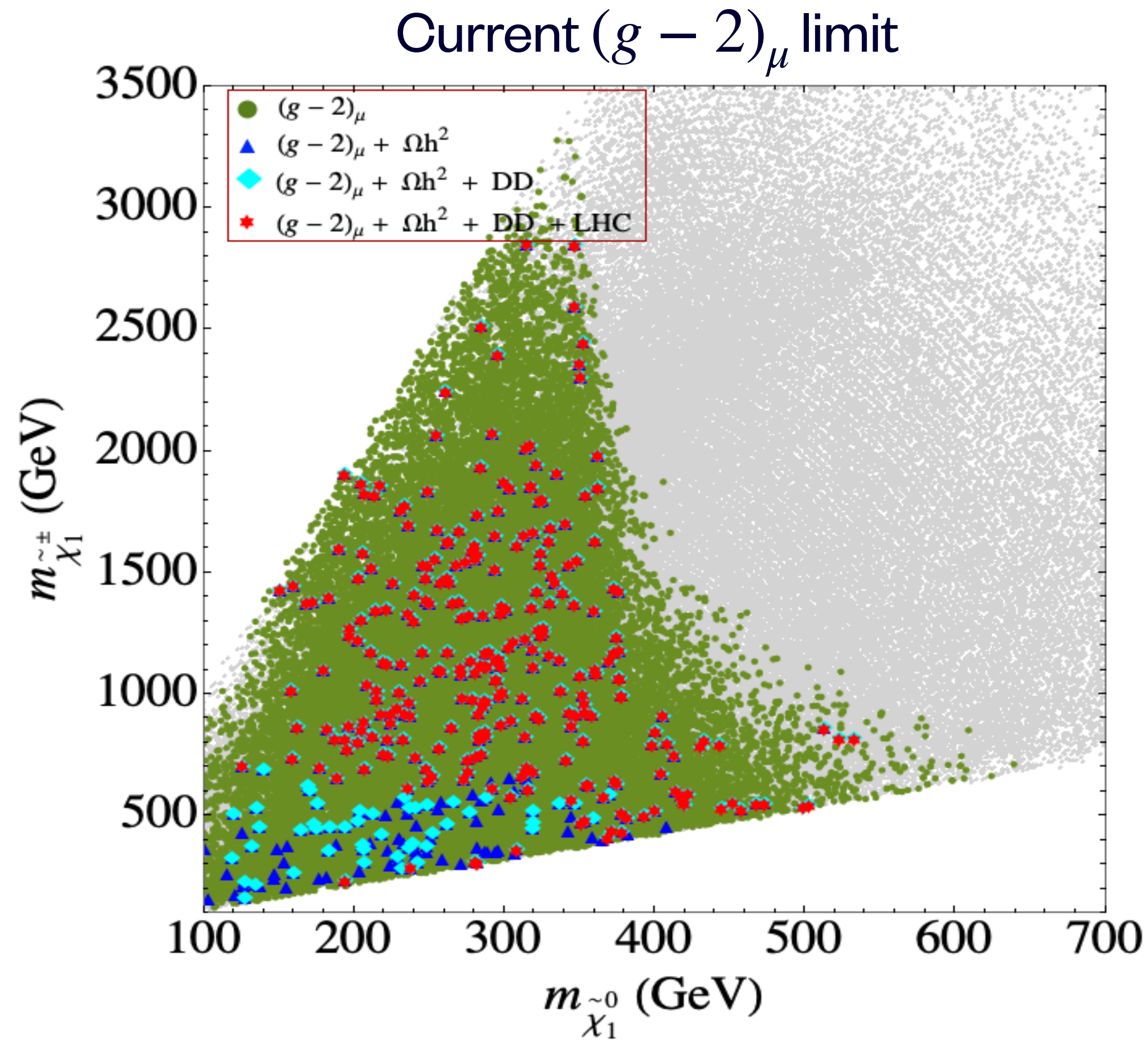


Anticipated future $(g - 2)_\mu$ limit



The left-sleptons and sneutrinos are close in mass to the LSP

Slepton Co-annihilation: Case-L



Large $\text{BR}(\tilde{\chi}_1^\pm \rightarrow \tilde{\tau}_1 \nu_\tau)$ and $\text{BR}(\tilde{\chi}_2^0 \rightarrow \tilde{\tau}_1 \tau)$, $\text{BR}(\tilde{\chi}_2^0 \rightarrow \tilde{\nu} \nu)$

($3l + \text{missing } E_T$) exclusion limit weakens ...

$(g - 2)_\mu$

- Large discrepancy from the SM (more than 3σ):

$$a_\mu^{exp} - a_\mu^{SM} = (28.02 \pm 7.37) \times 10^{-10}.$$

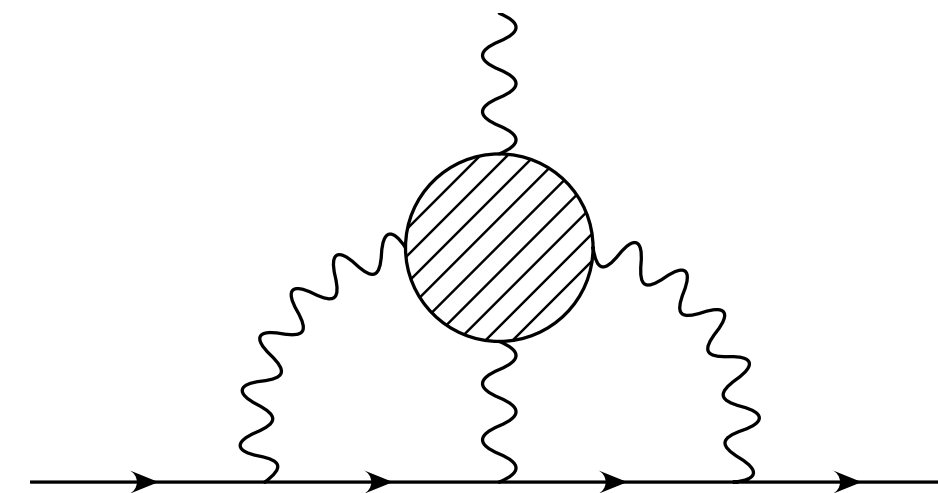
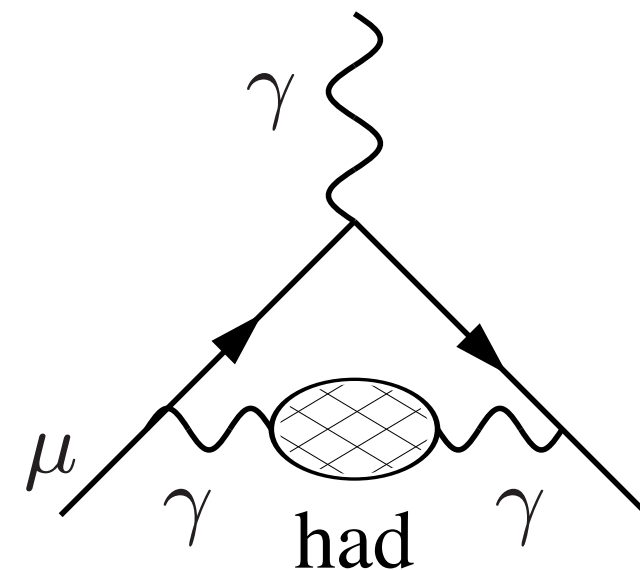
Keshavarzi, Nomura, Teubner '19

- Important probe for new physics. $\frac{\delta a_l}{a_l} \sim \frac{m_l^2}{\Lambda^2}$.
- SM contributions : QED, weak, hadronic vacuum polarization, hadronic light by light scattering.

- QED : complete calculation upto 5 loops. EW : two loops.

Aoyama, Hayakawa, Kinoshita, Nio '17, Ishikawa, Nakazawa, Yasu '18,
Heinemeyer, Stöckinger, Weiglein '04

- Uncertainty dominated by non-perturbative, hadronic sector.



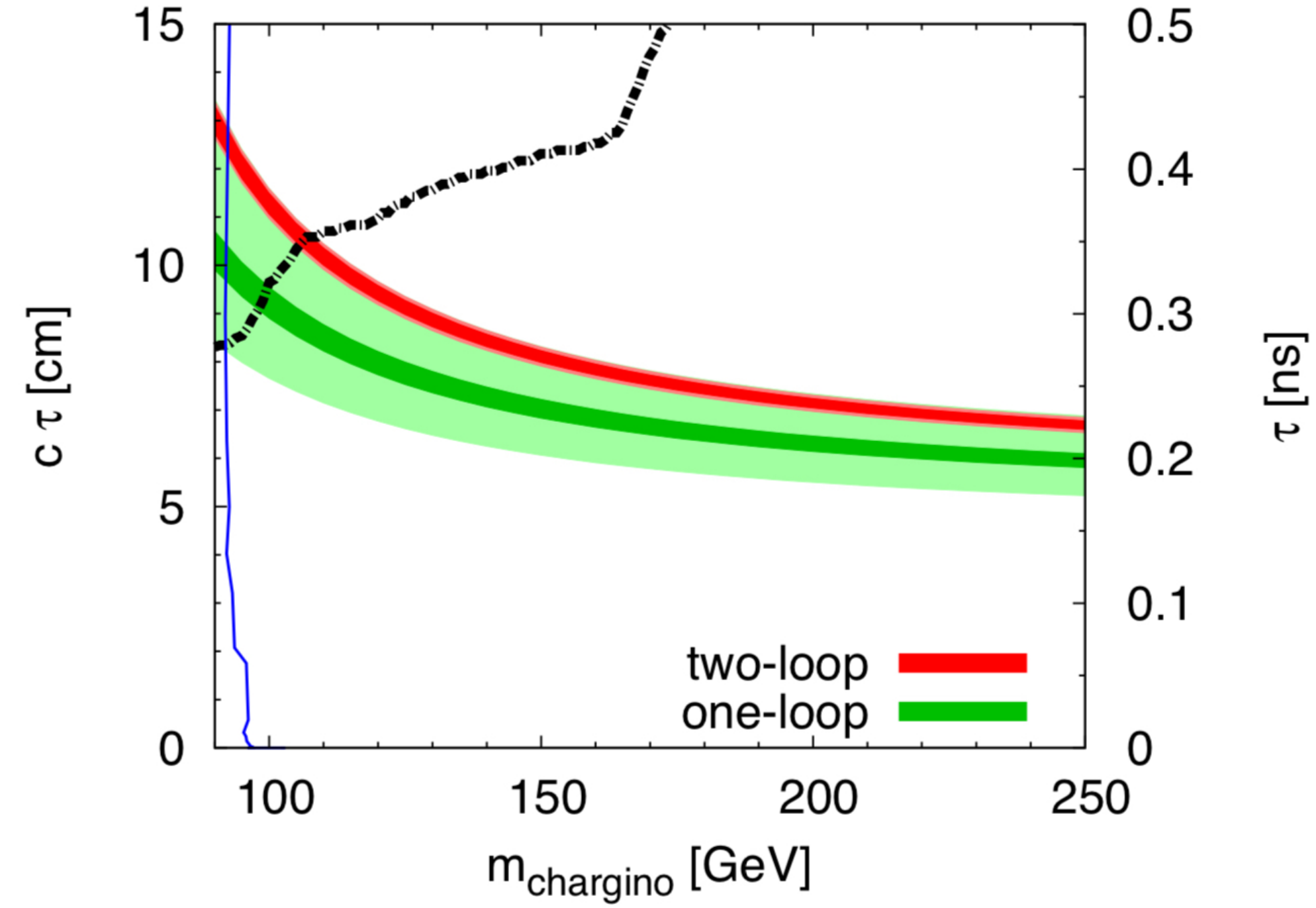


Figure 6: The lifetime of charged wino evaluated by using δm at the one-loop (green band) and two-loop (red band). We neglected the next-to-leading order corrections to the lifetime of the charged wino estimated in terms of the pion decay rate, which is expected to be a few percent correction. The black chain line is the upper limit on the lifetime for a given chargino mass by the ATLAS collaboration at 95% CL ($\sqrt{s} = 7 \text{ TeV}$, $\mathcal{L} = 4.7 \text{ fb}^{-1}$) [28]. The blue line shows the constraints which are given by the LEP2 constraints [30]–[33].