

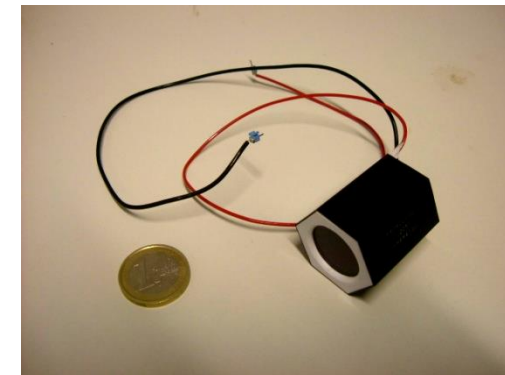
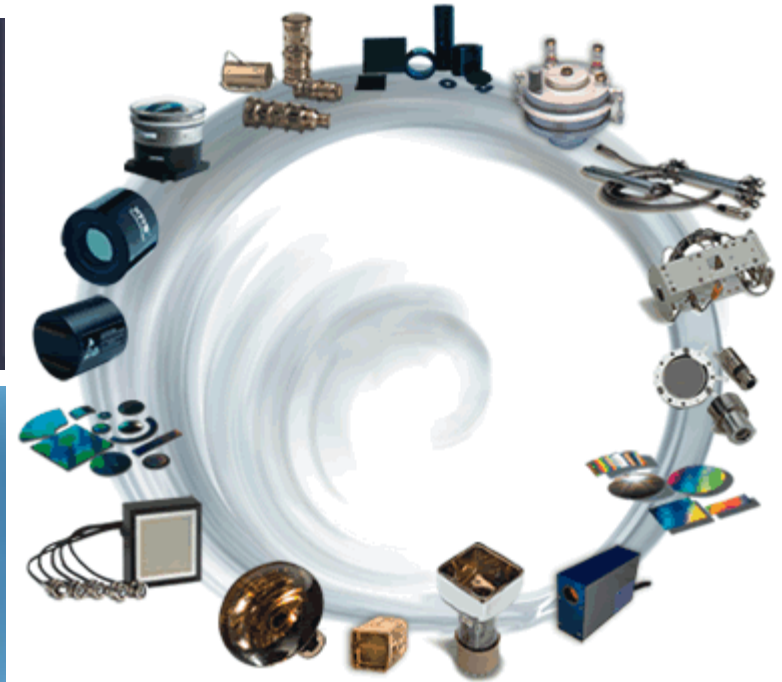
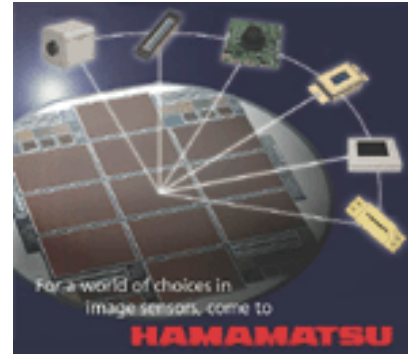
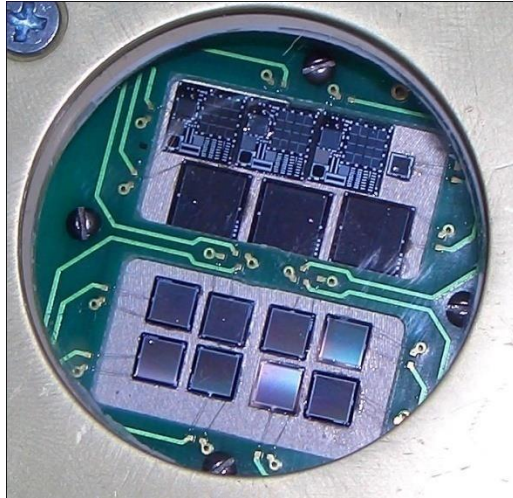
Photo – Multiplier Technologies

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The „zoo“ of LLL sensors



Vacuum Light Sensor Types

- Quantum Efficiency
- Photo-multipliers (including large area ones, e.g. SuperK)
- Multianode (MaPMTs)
- HAPDs
- hybrid HPDs
- Ways to further enhance the QE

Quantum Efficiency

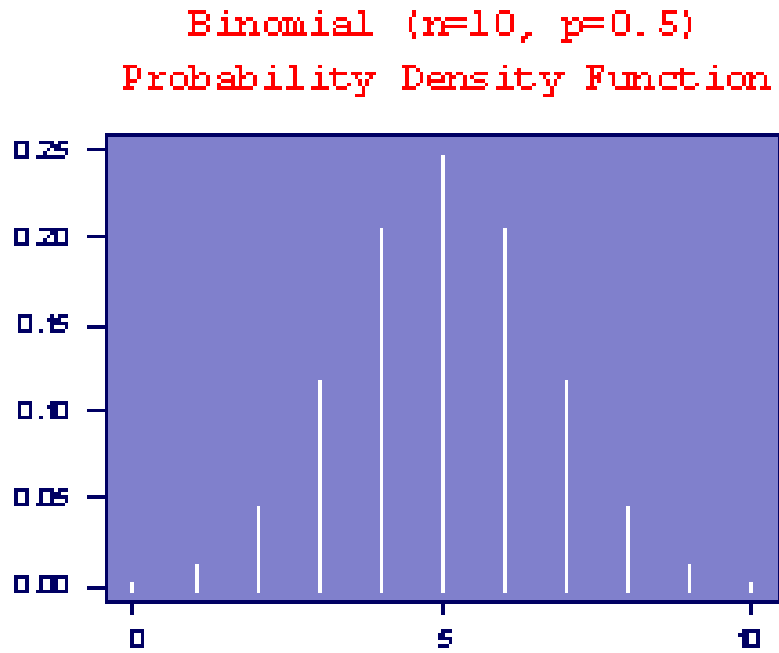
Quantum efficiency (QE) of a sensor is defined as the ratio

$$QE = N(\text{ph.e.}) / N(\text{photons})$$

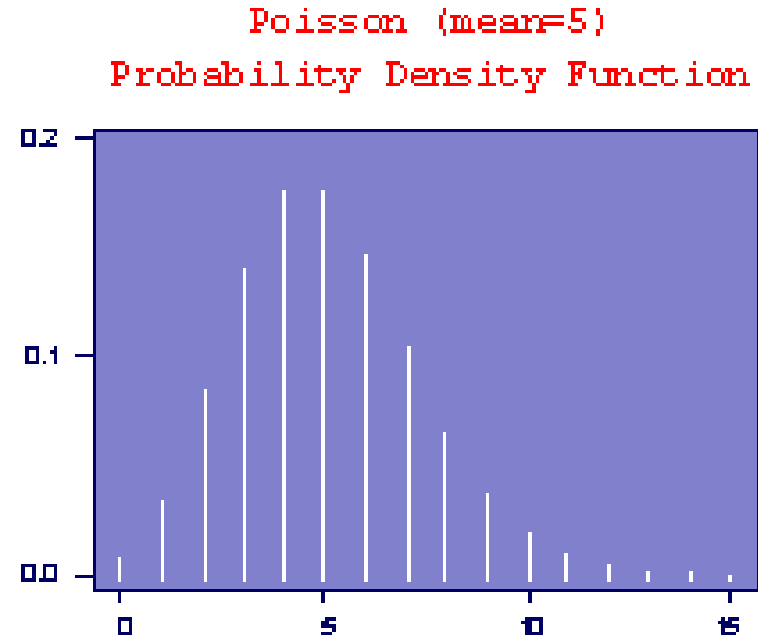
Conversion of a photon into ph.e. is a purely binomial process (and not poisson !)

Light sources of thermal origin can be described by the poisson distribution (including LED)

Differences between binomial and poisson distributions



SNR = 3.16



mean/ σ = 2.24

Signal to noise ratio

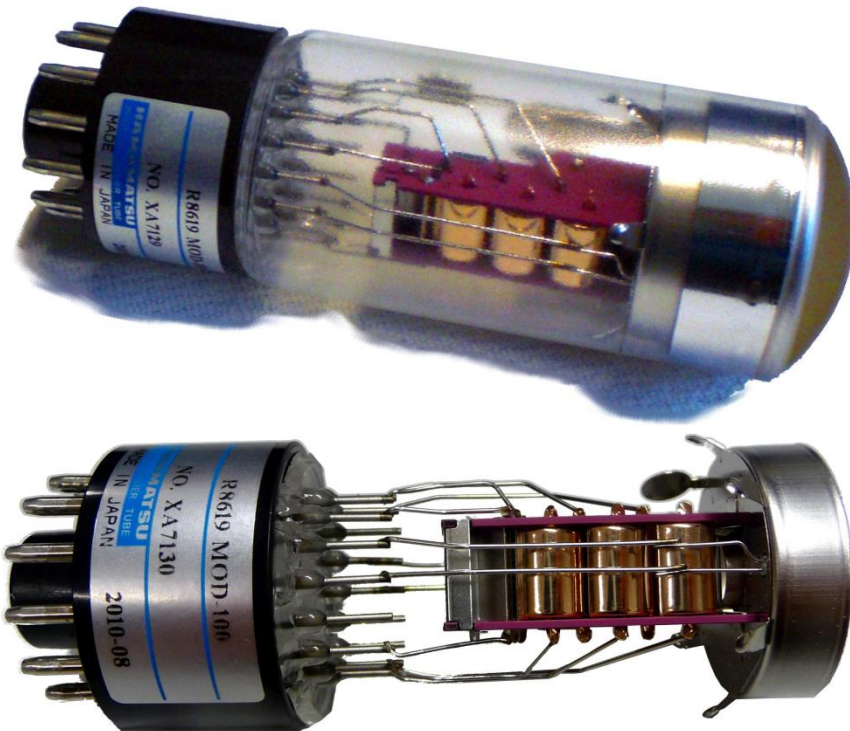
The signal-to noise ratio of a light sensor can be calculated as

$$\text{SNR} = [N \times P / (1 - P)]^{1/2}$$

For example, if $N = 1$ (single impinging photon):

P	0.1	0.3	0.9
SNR	0.33	0.65	3

One of the best known light sensors: the classical PMT



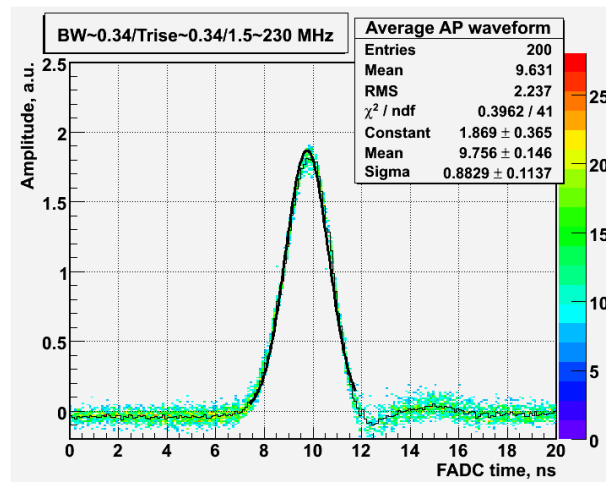
- The impinging photons kick out e⁻ from the thin photo cathode (~25nm)
- e⁻ are accelerated in a static electric field (~100V) and hit dynodes arranged in a sequential topology
- Every dynode enhances the number of e⁻ by a factor 4-5
- The net gain of a PMT could be $3 \times 10^4 - 10^9$
- That allows measuring single photons

PMTs for MAGIC developed by ETE, Hamamatsu, Photonis



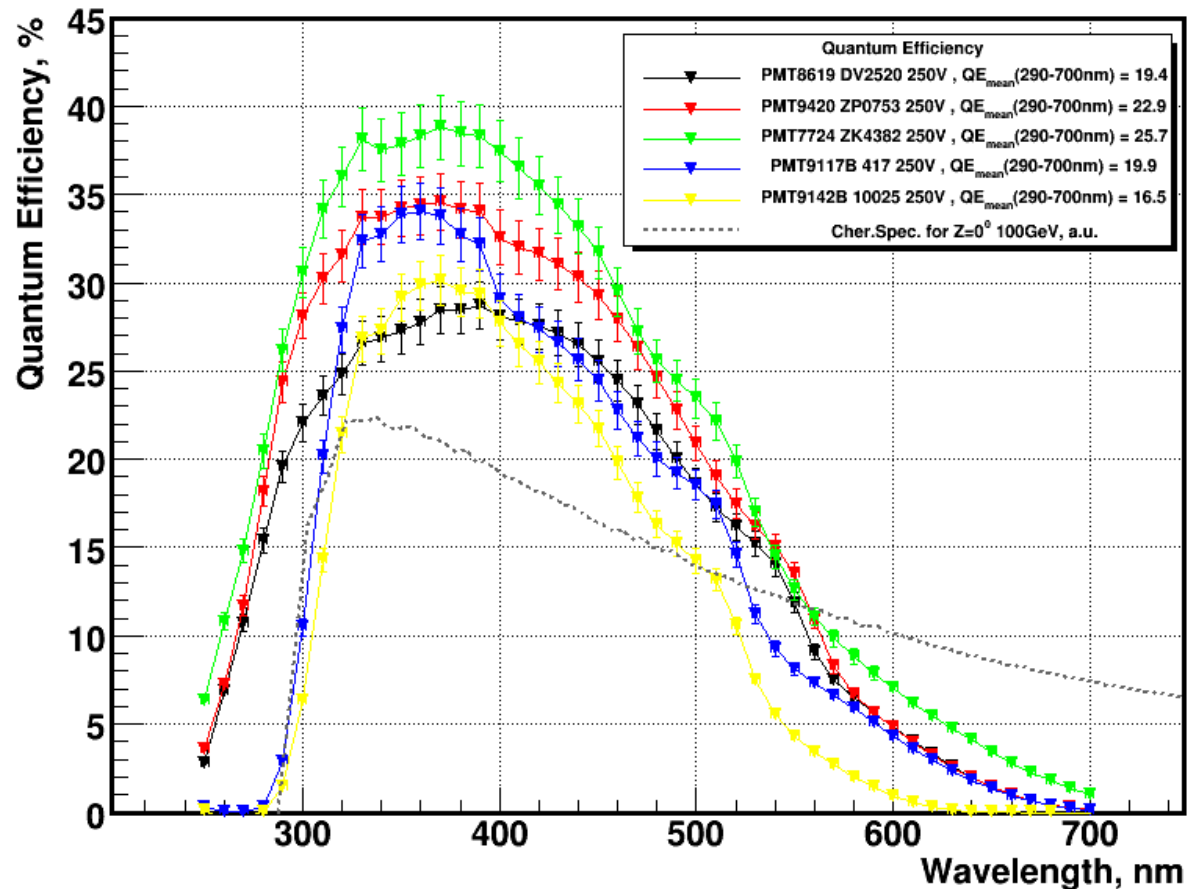
1-inch hemispherical MAGIC-type PMTs:

- ultra-fast resonance; ETE PMT: rise time 600ps, fall time 700ps, FWHM = 1.2ns
- possible due to 6 dynodes
- hemispherical shape photo cathode
- providing double crossing of photons (the highest probability of the semi-transparent photocathode is ~60% @ 400nm) with light guides
- low gain → slow ageing in time



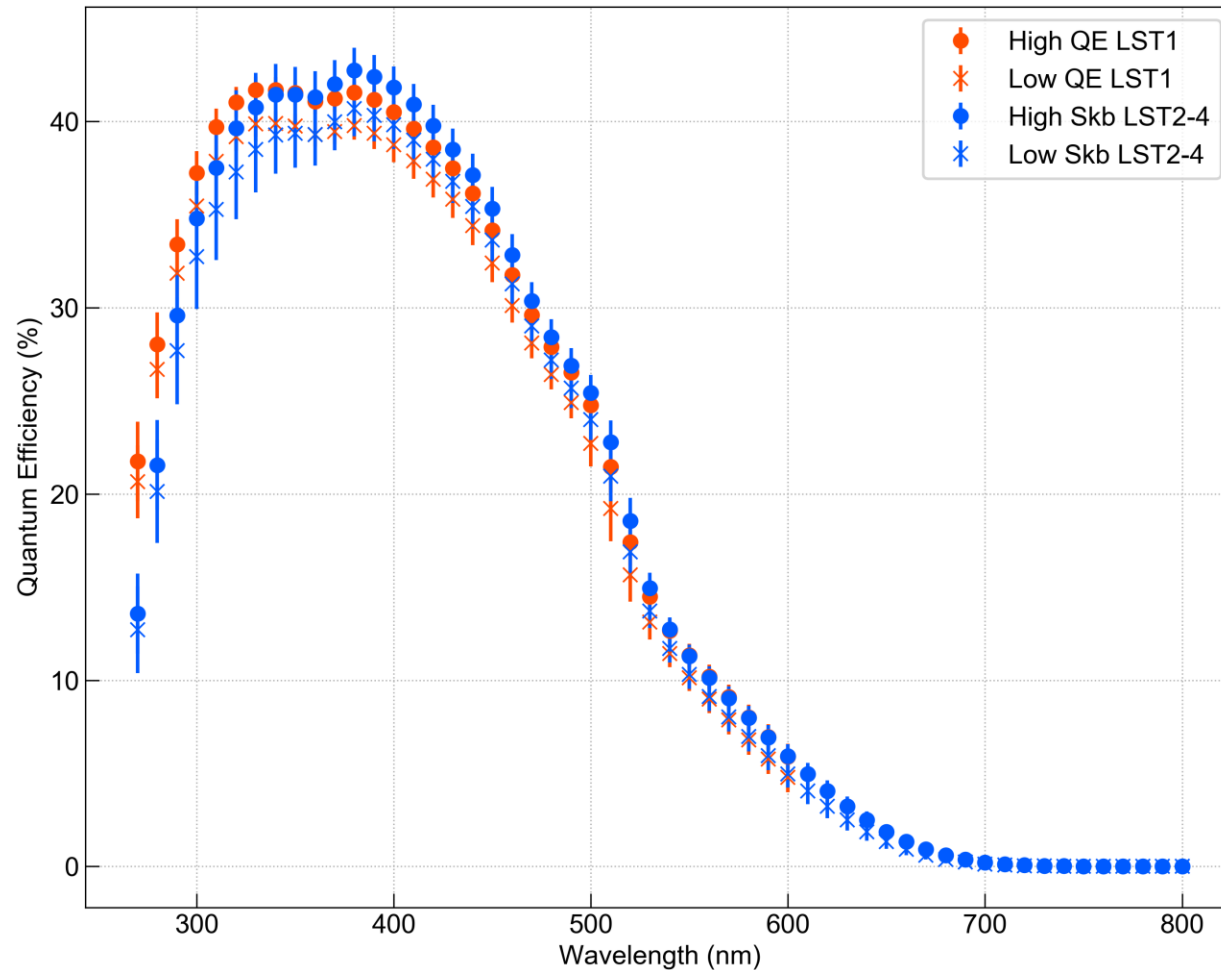
Instrumental/technological improvements

Running target: light sensor improvements. Successfully pushing the PDE higher up. Shown for several types of PMTs



- Some 16 years ago we have launched a QE improvement program with manufacturers Hamamatsu (Japan), Photonis (France) and Electron Tubes Enterprises (England).
- The results were very encouraging, → superbialkali PMTs
- About 11 ago years we launched a new improvement program for CTA PMTs

QE for novel 1.5 inch PMTs from Hamamatsu (~1200 PMTs)



credit: M. Takahashi

Type: R12992-100-05

PMT Light Emission

Ahnen, et al., IEEE TNS, 2015

2.3.1 Set-Up

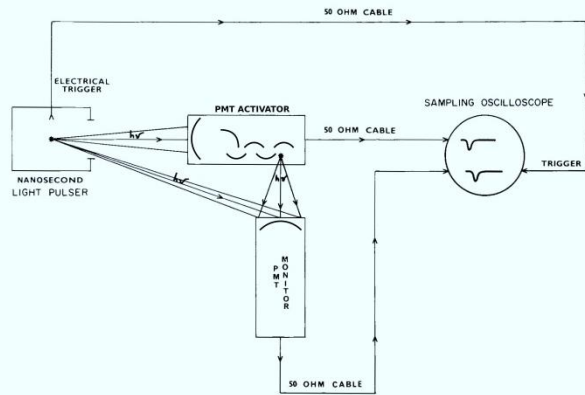
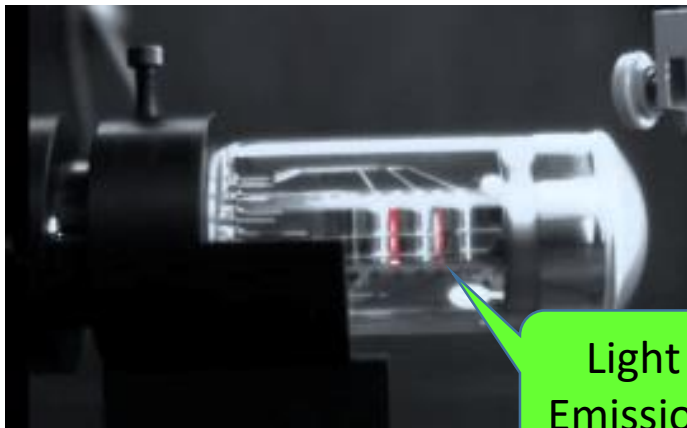


Figure 2.2: The photomultiplier dynode glow test apparatus, sketch adapted from [10]



Light Emission

2.3.3 Results

A screen capture of the oscilloscope with more than 200 million waveforms. Fig.2.3. The individual peaks on the activator photomultiplier are

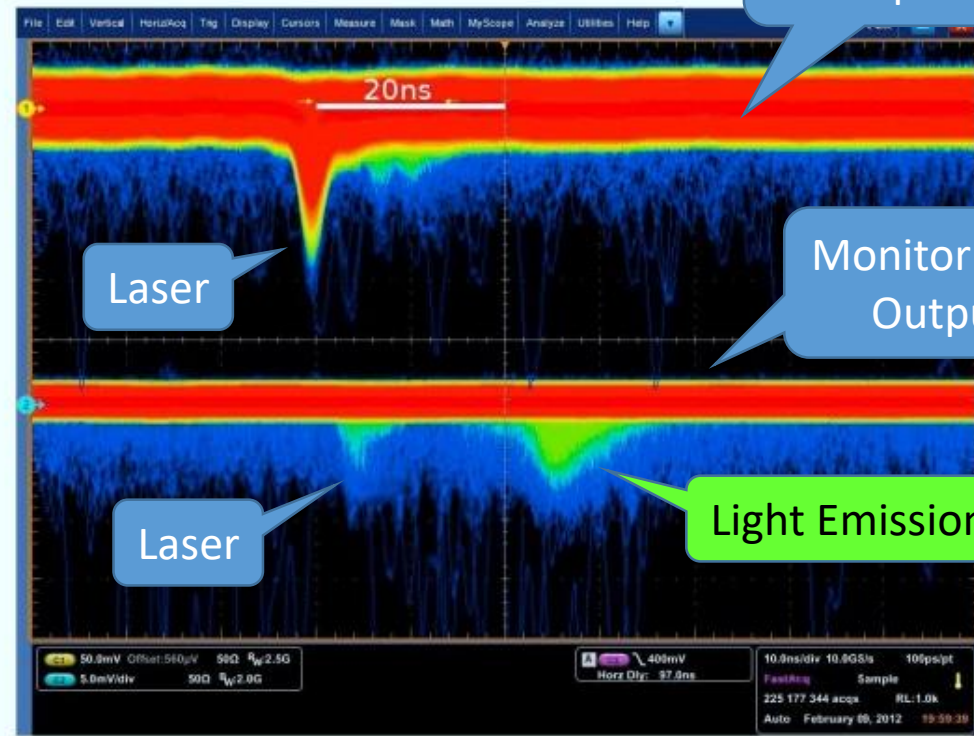


Figure 2.3: Measurement of the activator photomultiplier (top) and the monitor photomultiplier (bottom).

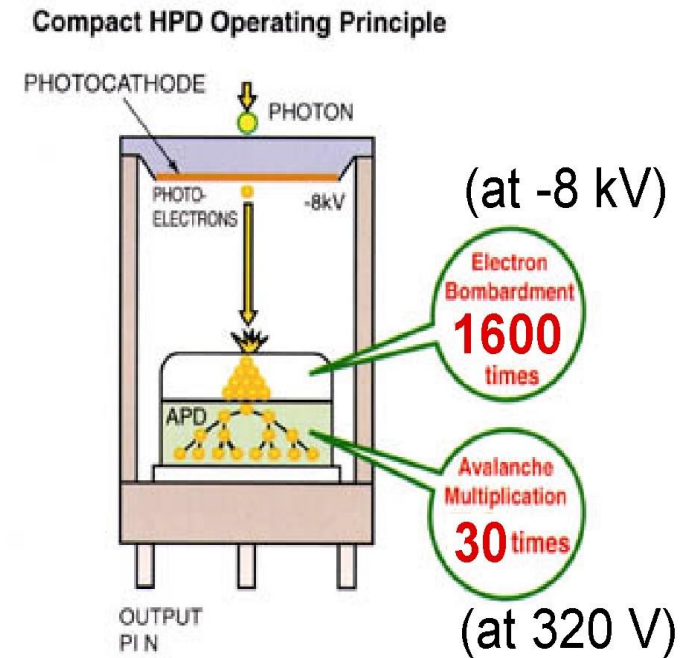
Hamamatsu Solution to Block the Light Emission

Ahnen, et al., IEEE TNS, 2015



HPD Structure

- HPD (Hybrid Photo Diode).
- Structure
 - Photo cathode
 - Avalanche diode as anode.
 - High vacuum tube ($\sim 10^{-7}$ Pa)
- Gain mechanism (2 stages)
 - Electron bombardment $\sim (x 1600)$
 - Avalanche effect $\sim (x 30-50)$

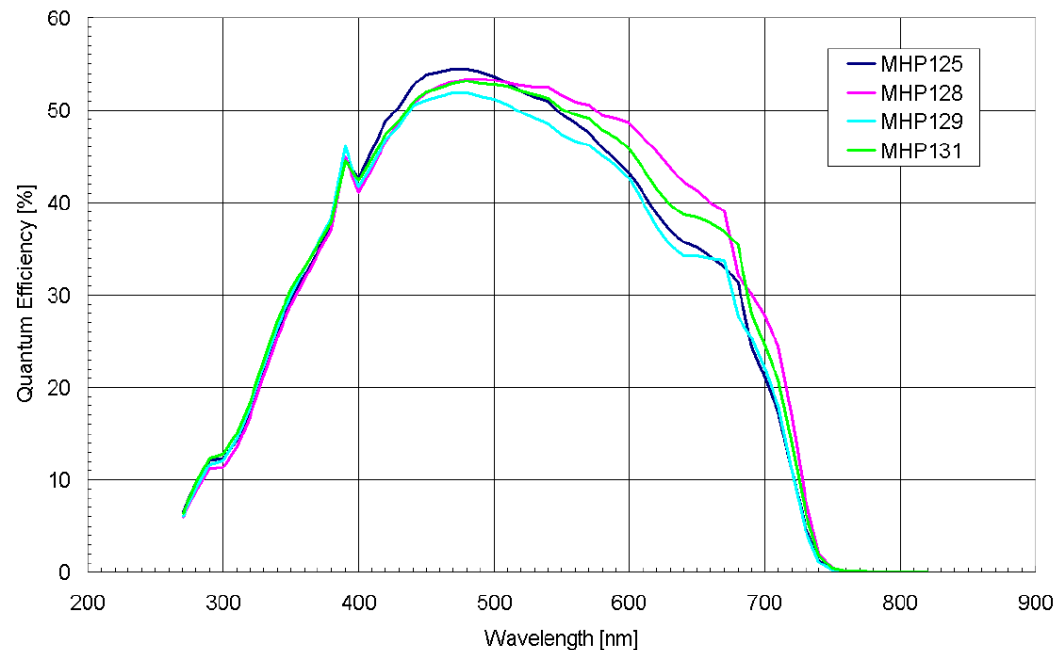


Much better pulse height resolution than PMT.

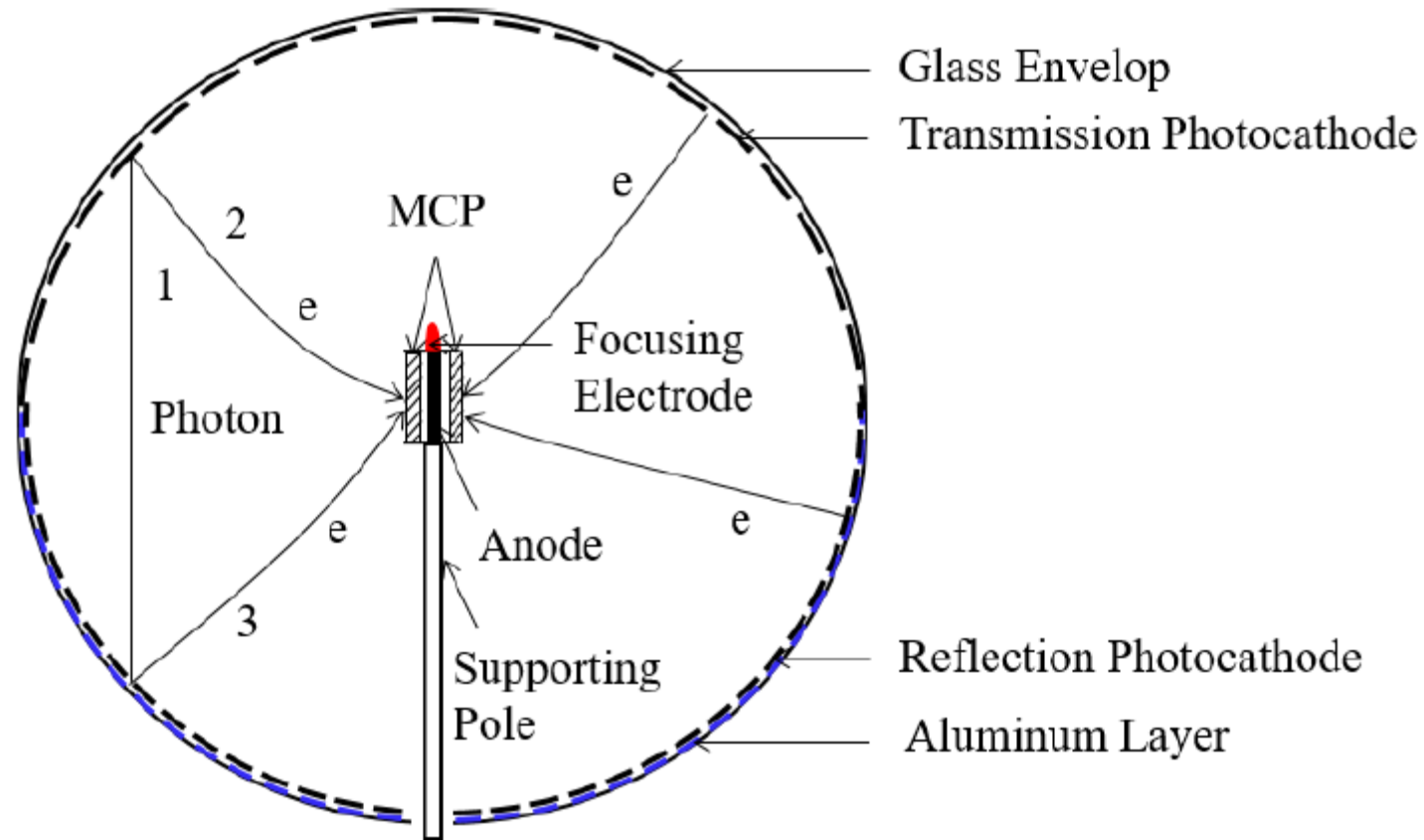
18-mm GaAsP HPD (R9792U-40) (development started ~15 years ago)

Designed for MAGIC-II telescope camera;
(developed with *Hamamatsu Photonics*)

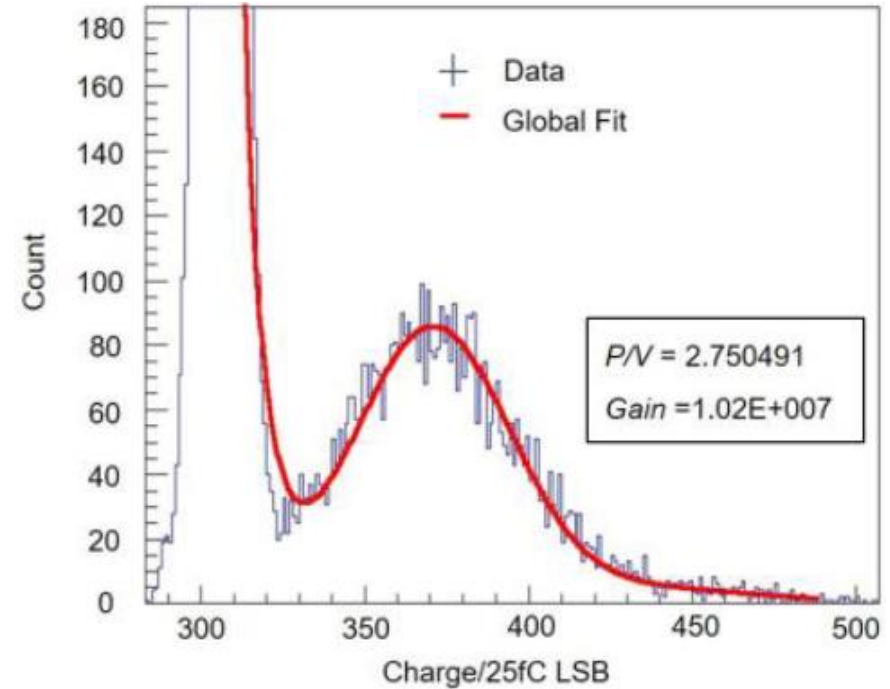
Photocathode(GaAsP) Spectral Response



8-inch spherical MCP-PMT from China

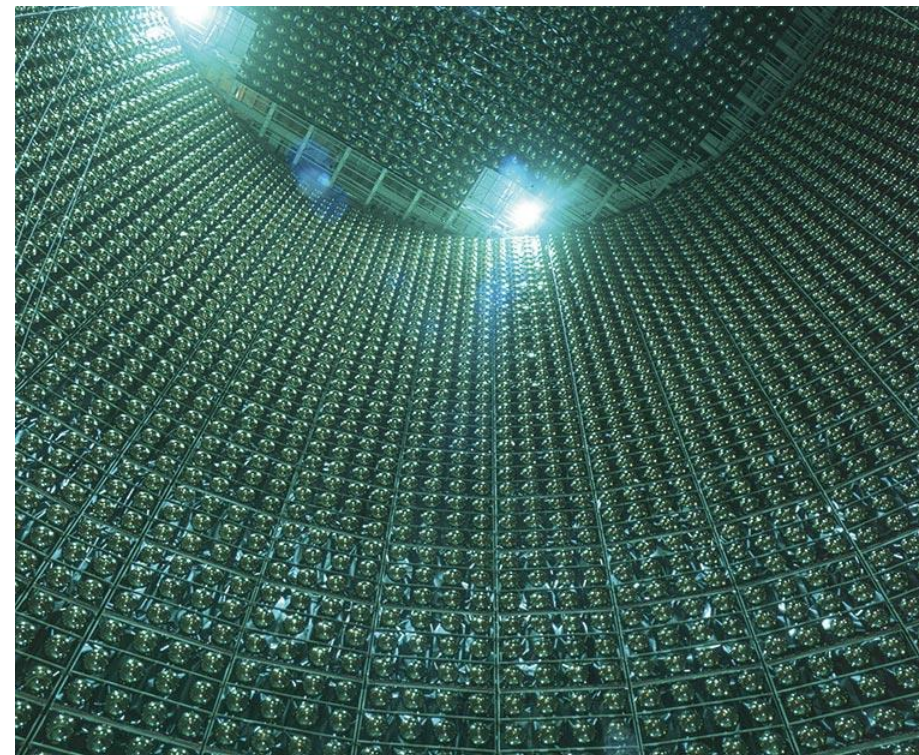
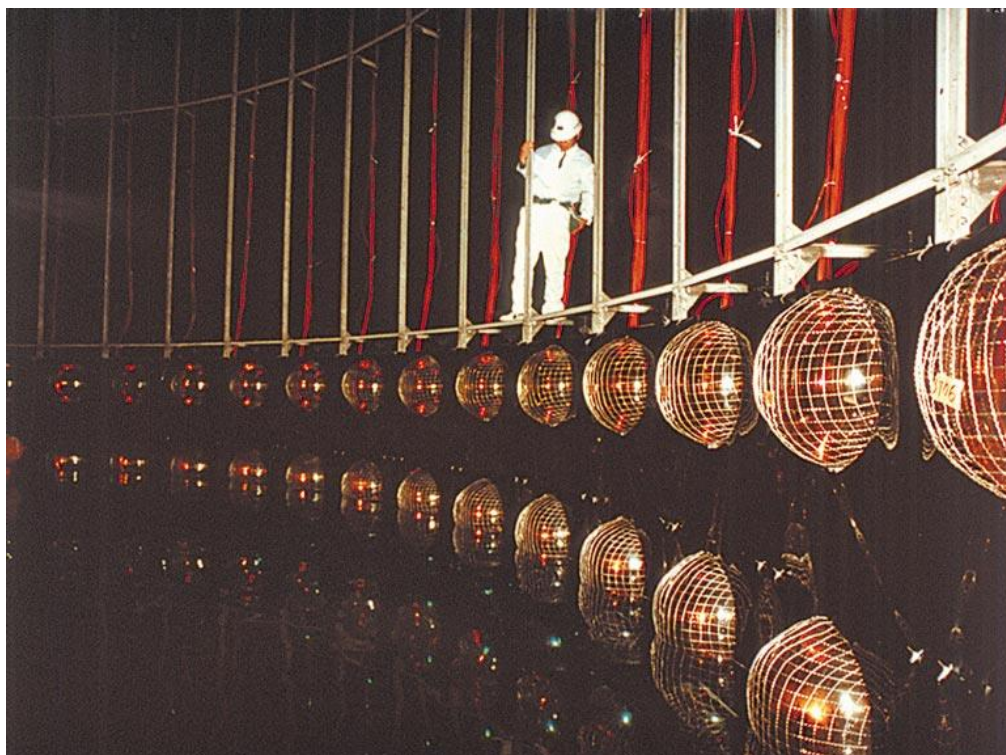


Picture of the 8-inch spherical MCP-PMT prototype



Hamamatsu 20 inch size PMT Hamamatsu R1449 PMT; further improved model under the name R3600-05.

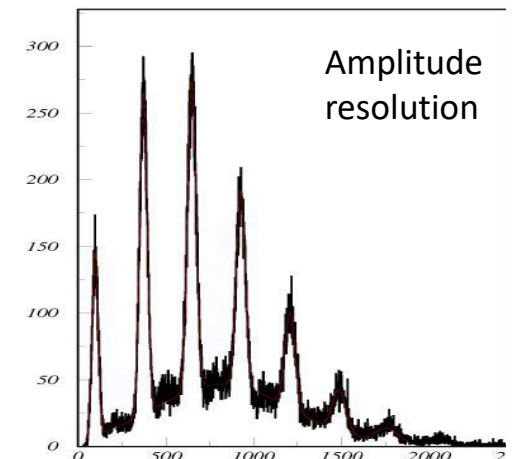
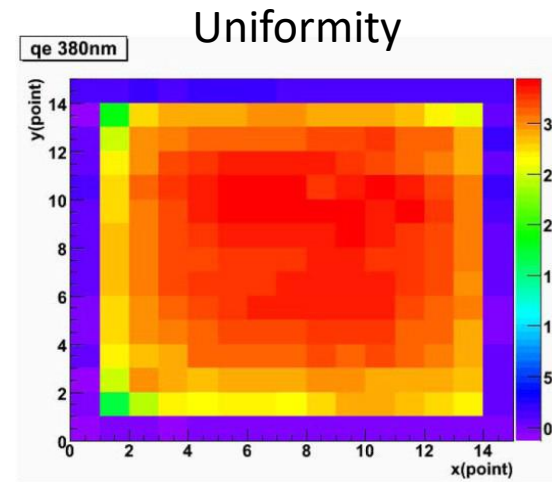
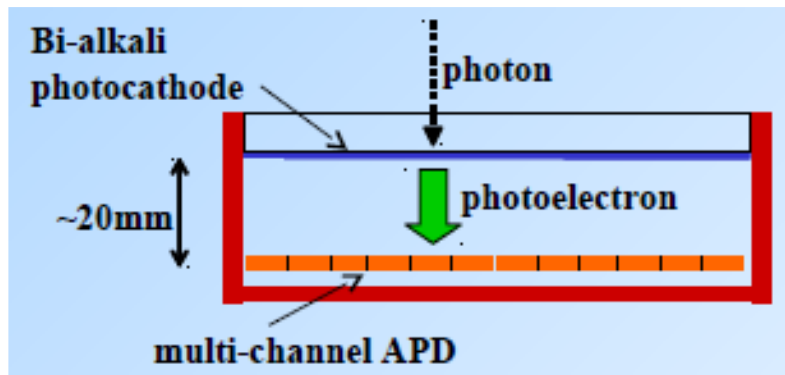
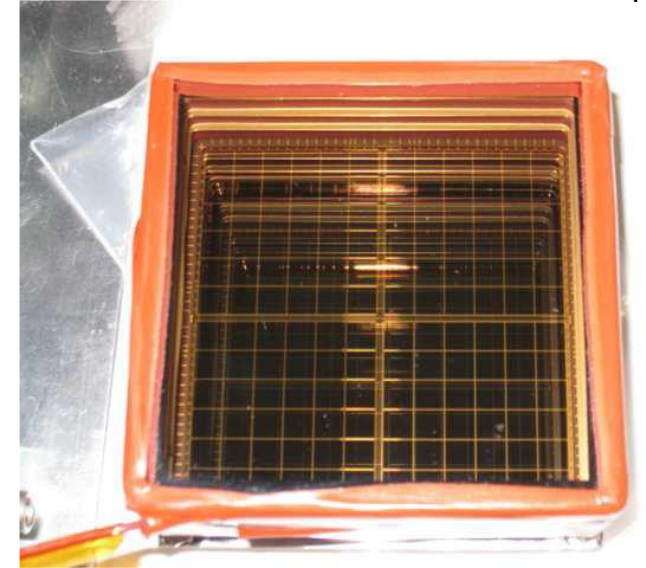
Credit: ICRR, Univ. Tokyo



144-channel HAPD for aerogel RICH at Belle II

Credit: S. Korpar

- Hybrid avalanche photo detector developed in cooperation with Hamamatsu, proximity focusing type
- 12 x 12 channels (5 mm x 5mm); physical size 72mm x 72 mm
- ~65 % effective area
- Net gain: 4.5×10^4 (e- bombardment: x 1500, APD: x 30)
- Superbialkali photocathode, peak QE ~28 %
- Capable to work under the condition of ~perpendicular to entrance window strong magnetic field of 1.5 T
- Output capacitance: ~80 pF/ch.



Flat Panel Multi-anode PMT Hamamatsu H9500



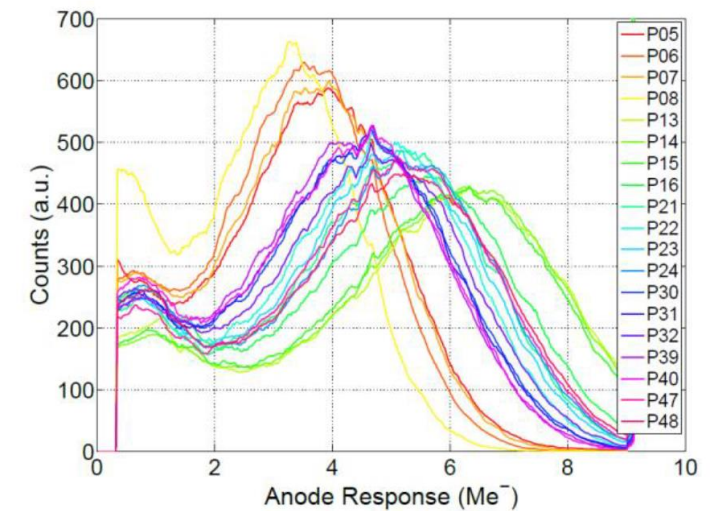
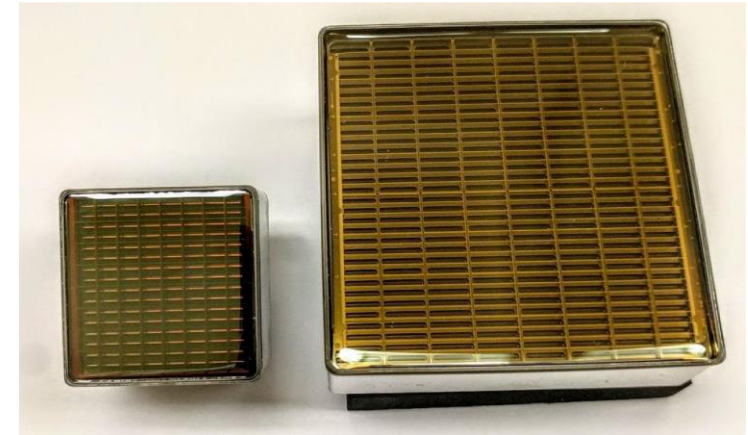
- Large effective area: 49 mm × 49 mm
- 16 × 16 multianode
- Anode pixel size: 2.8 mm × 2.8 mm / anode
- Small dead space
- Fast time response
- Metal channel dynodes
- Number of stages: 12
- Maximum supply voltage: 1.1 kV

MaPMTs

On the example of Hamamatsu R13742 (1'') & R13743 (2'')

Credit: M. Fiorini

- Super-bialkali photo cathode, peak QE $\sim 35\%$
- UV-glass window
- Gain $\geq 10^6$ @ 1 kV
- Channel-to channel gain spread in a given MaPMT: 3:1
- Spread among different MaPMTs: 3:1
- Usual for PMTs low dark count rate
- Single photo-electrons well separated from noise
- Sensitive to magnetic fields, at field strength ~ 30 G edge pixels show $\sim 50\%$ loss of efficiency

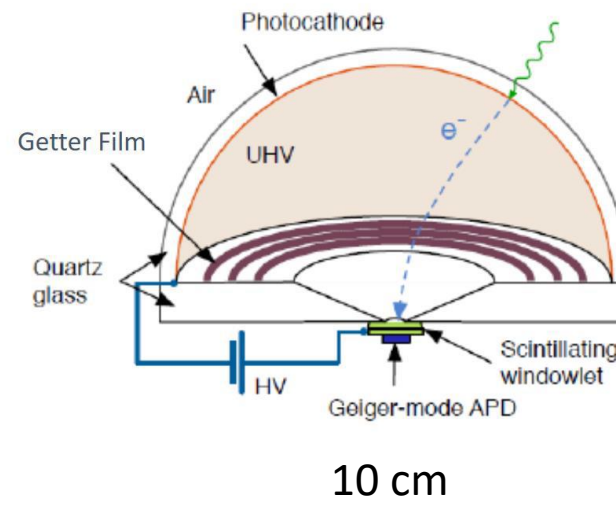


Abalone

D. Ferenc, et al., NIM A 2020

- @ 25 keV a single ph.e. generates ~ 650 photons in LYSO scintillator ($6 \times 6 \times 1.5 \text{ mm}^3$)
- 100 of those detected by SensL J-type G-APD ($6 \times 6 \text{ mm}^2$) with a gain of $\approx 6.1 \times 10^6$
- Combined gain $\approx 6 \times 10^8$
- Cs3Sb photocathode
- After-pulsing rate $\sim 5 \times 10^{-3}$ ions (mostly helium) per photo-electron
- Sub-nanosecond timing resolution
- Single-photon sensitivity
- Very high radio-purity
- UV sensitivity due to fused silica

Schematics

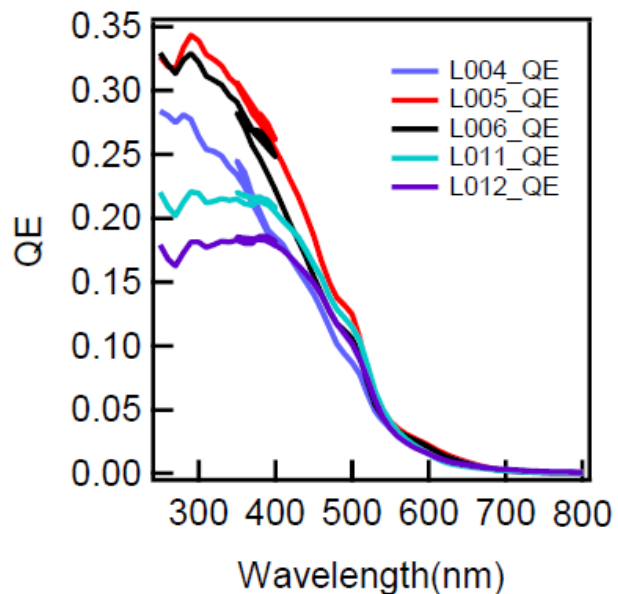


Prototype for IceCube extension



Surface roughness and the QE

J. Smedley, Light-17

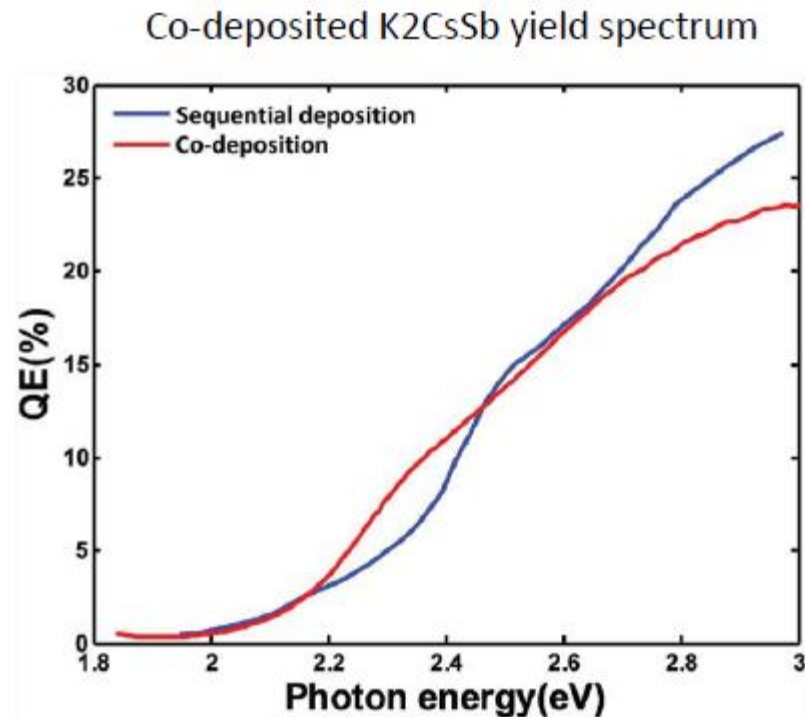


- Simultaneous evaporation of all constituents results in no crystal phase transformation
- Smooth, and high crystal quality
=> provides High QE

	QE@532nm(%)	Roughness(A)	Thickness (A)	Grain size (A)
L004 Si	4.9	3.5	234	155
L005 Si	5.8	11.5	815.3	277
L006 Si	5.4	13.8	757.5	202

Co-deposition of the photo cathode materials leads to very smooth surface quality as well as to high QE

J. Smedley, Light-17



Triple co-evaporation at LBNL

